COMP IV Fall 2020 Dr. Rykalova

PS2: N-Body Simulation

Part A: loading universe files; body class; graphics

In 1687 Sir Isaac Newton formulated the principles governing the motion of two particles under the influence of their mutual gravitational attraction in his famous *Principia*. However, Newton was unable to solve the problem for three particles. Indeed, in general, solutions to systems of three or more particles must be approximated via numerical simulations. Your challenge is to write a program to simulate the motion of *N* particles in the plane, mutually affected by gravitational forces, and animate the results. Such methods are widely used in cosmology, semiconductors, and fluid dynamics to study complex physical systems. Scientists also apply the same techniques to other pairwise interactions including Coulombic, Biot-Savart, and van der Waals.

Program specification

For this part of the assignment, Part A, we will create a program that loads and displays a static universe. In Part B, we will add the physics simulation, and animate the display!

Here are the particular assignment requirements for us:

- Make sure to download the universe specification files and image files from nbody.zip
- You should build a command-line app which reads the universe file
 (e.g., pl anets. txt) from stdi n. Name your executable NBody, so you would run it
 with e.g:

```
./NBody < planets.txt
```

Reading in the universe

The input format is a text file that contains the information for a particular universe (in SI units). The first value is an integer N which represents the number of particles. The second value is a real number R which represents the radius of the universe, used to determine the scaling of the drawing window. Finally, there are N rows, and each row contains 6 values. The first two values are the x- and y-coordinates of the initial position; the next pair of values are the x- and y-components of the initial velocity; the fifth value is the mass; the last value is a String that is the name of an image file used to display the particle. As an example, planets.txt contains data for our own solar system (up to Mars):

```
. . . xvel . . .
                                       . . . yvel . . .
                                                                    filename
. . . xpos. . .
             . . . ypos. . .
                                                    ...mass...
             0.0000e+00
1. 4960e+11
                          0.0000e+00
                                       2. 9800e+04
                                                     5.9740e+24
                                                                    earth. gif
                                                                    mars. gi f
2. 2790e+11
             0.0000e+00
                          0.0000e+00
                                       2. 4100e+04
                                                    6. 4190e+23
5. 7900e+10
             0.0000e+00
                          0.0000e+00
                                       4. 7900e+04
                                                     3. 3020e+23
                                                                    mercury. gif
             0.0000e+00
                          0.0000e+00
                                       0.0000e+00 1.9890e+30
0.0000e+00
                                                                    sun. gi f
1. 0820e+11
             0.0000e+00
                          0.0000e+00
                                       3. 5000e+04
                                                     4.8690e+24
                                                                    venus. gi f
```

You should read in exactly as many rows of body information as are indicated by N, the first value in the file.

The < pl anets. txt construct is known as an "input redirect".

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The planets. txt universe file contains the Sun and the first four planets, with the Sun at the center of universe (x=0, y=0) and the four planets in order toward the right (per below). When this is working, you should be rewarded with:



- For this project you'll implement a **Uni verse** class which will contain all celestial bodies, and a **Cel esti al Body** class representing the celestial bodies.
- The class Universe should create celestial bodies.
- The class Cel esti al Body should have the following features:
 - o It must be sf:: Drawable with a private virtual void method named draw.
 - Each instance of the class should contain all properties needed for the simulation;
 e.g.: x and y position, x and y velocity, mass, and image data.
 - It probably should contain an sf:: Spri te object (as well as the sf:: Texture object needed to hold the sprite's image).

For full credit:

- you must override the input stream operator >>, and use it to load parameter data into an object
- vou must to use the smart pointers
- Please see the grading rubric for all the details and pieces of the project.
- Please submit all files needed to build your project: . cpp, any header files, and a Makefile.
- Please submit the pl anet. txt file, and the specific GIF images associated with it.
- Please submit a screenshot of your running code, named screenshot. png.
- Fill out and include this ps2a-readme. txt file with your work.

Development process

There are a lot of parts to this assignment. We'd suggest the following incremental development process:

- 1. Create a bare-bones implementation of your Celestial Body class that has a constructor where you specify all the initial parameters (x,y position and velocity; mass; image filename).
 - Have the constructor load the image into a new Texture object; create a new Sprite with that Texture.
 - o Given the initial x,y position in the universe, figure out the corresponding pixel-position for display in an SFML window.

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 Hint 1: your class will need to know and store the universe radius, and display window dimensions.

- Hint 2: the universe's center is (0,0), and SFML's (0,0) point is in the upper-left.
- 2. Implement the draw method in your Cel esti al Body class.
- 3. Write a main file that manually creates a Body object, by copying initialization parameters from the pl anets. txt file into your source code.
- 4. Have the main file draw that object in the SFML display loop.

At this point, you should be able to display one planet (or the Sun). You can add one or two more manually, to know you're on the right track.

- 5. Create class **Uni verse** that contains a vector of pointers to **Cel esti al Body** objects and instantiate them with **new**.
- 6. When you have the vector working, you can write the code to overload the stream input operator >>, and read in the universe file to set up your bodies. (See http://www.tutorialspoint.com/cplusplus/input_output_operators_overloading.htm for example code.)

How to turn it in

Submit your work on Blackboard.

The executable file that your Makefile builds should be called NBody (the grading script checks that this executable builds successfully.)

Grading rubric

Feature	Value	e Comment
core implementation	8	full & correct implementation
		1 pt celestial body object is Drawabl e and SFML while loop uses wi ndow. draw(obj)
		1 pt implementation loads universe from stdi n
		1 pt body class has >> overloaded to read in a row from universe file
		1 pt supports arbitrary number of body objects (per universe file)
		1 pt scaling works for arbitrary universe size and given SFML window size declared in main. cpp
Makefile	2	Makefile included
		targets all and clean must exist
		al 1 should build NBody
		must have dependencies correct
screenshot.png	1	included
ps3a- readme.txt	3	must explain how each of the features noted above is implemented to receive credit for those features!
Total	14	
extra credit	+1	Use background image