

# GRE Physics Study Notes - Lessons from Practice Exams

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Last Edited 6th November 2012

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# Introduction

By no means comprehensive, this list is meant to serve as additional study material to re-reading textbooks, practicing GRE tests, and nagging your physics friends for study help.

## 1 Some Stuff

### 1.1 Cherenkov Radiation

A charged particle which passes through a media with a speed greater than the speed of light in that media will emit E&M radiation (light)

$$n = \frac{c}{v}, v = \frac{c}{n} \leftarrow \text{minimum velocity of the particle} \quad (1)$$

This does not violate relativity because  $n > 1$  so  $v < c$ .

### 1.2 Bremsstrahlung Radiation:

A continuous spectra of radiation emitted when a charged particle is decelerated in a metal target.

## 2 Problems

### 2.1 Hoops Hung on a Nail, Undergoing Small Oscillations

1. Approximate using small  $\theta$  pendulum, with  $\omega = \sqrt{\frac{g}{l}}$
2. Works because it's a hoop, with extended object use a physical pendulum.

### 2.2 Physical Pendulum

$$\begin{aligned} \tau &= I\dot{\omega} = I\alpha & \tau &= mgL_{\text{cm}} \sin \theta \\ \ddot{\theta} &= \frac{mgL_{\text{cm}} \theta}{I} & \omega &= \sqrt{\frac{mgL_{\text{cm}}}{I}} \end{aligned}$$

For hoop,  $I = MR^2 = ML_{\text{cm}}^2 \rightarrow \omega = \sqrt{\frac{mg}{L}} = \sqrt{\frac{g}{\text{radius}}}$   
Often dealing with ratios so reduces out mass + constant factor in I.

$$\frac{\omega_1}{\omega_2} = \sqrt{\frac{L_1 \cdot R_2^2}{L_2 \cdot R_1^2}} \quad (2)$$

### 2.3 Intrinsic Magnetic Moment

$$\bar{\mu}_s = \frac{gq}{2m} \bar{s} \quad (3)$$

Mass is the dominant factor because  $g \sim 10$  s,  $q \sim 100$  s but  $m$  is many orders of magnitude.

### 2.4 Equations of Motion

Look for boundary values. A given  $x(t)$ ,  $y(t)$ , and  $v_0$ , differentiate and evaluate at  $t = 0$  to see which one yields  $v_0$ .

## 2.5 Moments of Inertia

The moment of a plate is  $I + \frac{1}{3}Md^2$  where  $d$  is the width from the center rotation axis of the plate to the edge of the plate.

### 2.5.1 Stretch Axis Theorem

If you stretch an object along the dimension it rotates about, then the moment is unchanged.

e.g., Disk and cylinder both have  $I = \frac{1}{2}MR^2$

## 2.6 Moments

1. Hoop:  $I = MR^2$
2. Disk:  $I = \frac{1}{2}MR^2$
3. Hollow Sphere:  $I = \frac{2}{3}MR^2$
4. Solid Sphere:  $I = \frac{2}{5}MR^2$
5. Rod Center:  $I = \frac{1}{12}ML^2$
6. Rod End:  $I = \frac{1}{3}ML^2$
7. Thick Cylinder:  $I = \frac{1}{2}m(r_1^2 + r_2^2)$
8. Cuboid:  $I_{\hat{y}} = \frac{1}{12}M(x_l^2 + y_l^2)$

## 2.7 Hermitian Matrix

1. Described observables: The eigenvalues are real
2. Square matrix

$$A = A^T \rightarrow \text{the entries are equal to their conjugate transpose} \quad (4)$$

$$A = \begin{bmatrix} 3 & 2+i \\ 2-i & 1 \end{bmatrix}$$

$$A^T = \begin{bmatrix} 3 & (2-i)^* \\ (2+i)^* & 1 \end{bmatrix} = \begin{bmatrix} 3 & 2+i \\ 2-i & 1 \end{bmatrix} = A$$

Transpose:

1. Row A  $\rightarrow$  column A
2. Row B  $\rightarrow$  column B
3. The entries on the diagonal are real
4. The sum of any 2 Hermitian matrices is Hermitian
5. The product is Hermitian if and only if  $\hat{A}\hat{B} = \hat{B}\hat{A} \Rightarrow [\hat{A}\hat{B}] = 0$  (Commutes)

## 2.8 Hermitian Operators

Condition:

$$\langle f | \hat{A} f \rangle = \langle \hat{A} f | f \rangle \quad (5)$$

is equivalent to saying  $A = A^*$

## 2.9 Balancing Problem

Torque is the long and hard way to find where the pivot should be. Classic problem, where to put the fulcrum. Want to set up  $\sum \tau = 0$  problem. Instead, put the origin somewhere convenient (center of rod if it has mass) and calculate the center of mass.

$$CM = \frac{\sum_i m_i r_i}{\sum_i m_i} \quad (6)$$

## 2.10 Decay Rates

$$\frac{dA}{dt} = -kA \rightarrow A = A_0 e^{-kt} \rightarrow \frac{A}{A_0} = \frac{1}{2} e^{-kt} \rightarrow \frac{\ln(2)}{k} = t_{1/2} \quad (7)$$

for a substance with multiple decay modes,

$$k_{\text{tot}} A = (k_1 + k_2) A \quad (8)$$

## 2.11 Conservation of Energy

Total energy must be conserved.

$$U + KE_{\text{roll}} + KE_{\text{trans}} = E_{\text{tot}} \quad (9)$$

Example: Roll down a hill at a given  $v_{\text{transl}}$

$$U = KE_{\text{roll}} + KE_{\text{transl}} \quad (10)$$

$$mgh = \frac{1}{2} I \omega^2 + \frac{1}{2} m v_{\text{transl}}^2 \quad (11)$$

## 2.12 Interferometer

Fringe shifts occur for changing distances of mirrors or changing wavelengths. A fringe shift occurs for  $2d = \lambda$ . When using a gas to change wavelength, it changes with  $n$ .

$$2d = m(\lambda_{\text{gas}} - \lambda_{\text{vac}}) = m\lambda_{\text{vac}} \left( \frac{1}{n} - 1 \right), \lambda_{\text{gas}} = \frac{\lambda_{\text{vac}}}{n} \quad (12)$$

## 2.13 Vector Calculus

1. The div of a curl is 0  $\rightarrow \nabla \cdot (\nabla \times F) = 0$
2. The curl of a gradient is 0  $\rightarrow \nabla \times (\nabla \cdot F) = 0$

## 2.14 Thermodynamic Work

The area under the  $PV$  graph, or in a closed cycle. If clockwise,  $+W$ , if counter clockwise,  $-W$ .

## 2.15 Special Relativity, Momentum

$$E_{\text{rel}} \neq \frac{p_{\text{rel}}^2}{2m} \quad (13)$$

Instead use

$$E_{\text{rel}}^2 = (pc)^2 + E_0^2 \quad (14)$$

Only photons move at  $c$ , duh. However, particles can move faster than the speed of light in a medium,  $v_\phi = \frac{c}{n}$  and it will emit Cherenkov radiation.

## 2.16 Decay

1. Write out coefficients!
2.  $e^-/e^+$  always accompanied by  $\bar{\nu}/\nu$  by conservation of lepton #. Any combo of capture and emission.

$${}_Z^AX = {}_{p^+}^{p^++n^0} X \quad (15)$$

## 2.17 Springs

Add like capacitors.

For series:

$$\frac{1}{k_{\text{tot}}} = \frac{1}{k_1} + \frac{1}{k_2} \quad (16)$$

For parallel:

$$k_{\text{tot}} = k_1 + k_2 \quad (17)$$

## 2.18 Speed of Sound in an Ideal Gas

$$v \propto T^{1/2} \quad (18)$$

## 2.19 Conservative Field

$$\nabla \times F = 0 \rightarrow F = -\nabla V \quad (19)$$

## 2.20 Orbit Problems

First, think Kepler ( $T^2 \propto R^3$ ). Minimum  $E$  is a circular orbit.

## 2.21 Intensity / Radiation Problems

Radiation spreads like a spherical wavefront.

$$\# \text{ particles detected} / \text{ counts} = \frac{\text{Area detector}}{\text{Area sphere @ detector}} \quad (20)$$

## 2.22 Partition Function

$$\frac{f}{2}NkT, f = \# \text{ squared terms in Hamiltonian} \quad (21)$$

$f$  = Degrees of freedom; is a corollary.

## 2.23 Expectation Value

$$\langle \hat{Q} \rangle = \langle \Psi | \hat{Q} | \Psi \rangle = q \langle \Psi | \Psi \rangle \text{ or } q_1 \langle \Psi_1 | \Psi_1 \rangle + q_2 \langle \Psi_2 | \Psi_2 \rangle + \dots \quad (22)$$

## 2.24 Commutator Identities

$$[A, B] = -[B, A] \quad (23)$$

$$[A, BC] = B[A, C] + [A, B]C \quad (24)$$

$$[AB, C] = A[B, C] + [A, C]AB \quad (25)$$

## 2.25 Motion in a Circle

Always  $a_{\text{radial}}$  component, only  $a_{\parallel}$  if  $v_{\text{tan}}$  is changing.

$$F = \frac{mv^2}{r} = ma_r \rightarrow a_r = \frac{v^2}{r} \quad (26)$$

Compare to:

$$v = r \times \omega \quad (27)$$

$$a_r = r \times \alpha \quad (28)$$

$$a = a_{\parallel}^2 + a_r^2 \quad (29)$$

## 2.26 Particle Decay

$$\frac{dN}{dt} = -kN \rightarrow N = N_0 e^{-kt}, \text{ decay is exponential} \quad (30)$$

$k$  = decay constant,  $\tau$  = average, lifetime =  $\frac{1}{k}$

### 2.26.1 Half Life

$$\frac{N}{N_0} = \frac{1}{2} = e^{-kt} \rightarrow t_{hl} = \frac{\ln(2)}{k} \quad (31)$$

### 2.26.2 Multiple Decay Channels

$$(K_{\text{tot}})N = K_1N + K_2N + \dots \quad (32)$$

so,  $k = \frac{\ln(2)}{t_{hl}}, \frac{1}{t_{hl}} = \frac{1}{t_1} + \frac{1}{t_2} + \dots$



## 2.27 Specific Heat in a Solid

Best fit for when accounting for  $e^-$  specific heat with FD.

1. Einstein Model: Treat atoms as 3N Harmonic Oscillators. They all have same E, (frequency).
2. Debye: Also 3N Harmonic Oscillators. Assigns a range of energies and treats lattice vibrations as phonons in box.  
Correctly predicts low temperature  $C_V \propto T^3 Nk$
3. Dulong-Petit: High temps, uses equipartition with harmonic oscillator  $f = 6, c = 3Nk$   
Debye and Einstein reduce to this in high  $T$  limit.

## 2.28 Relativistic Doppler Shift

$$\frac{f_0}{f} = \frac{\lambda}{\lambda_0} = \sqrt{\frac{1+\beta}{1-\beta}}, \beta = \frac{v}{c} \quad (33)$$

The “redshift”:  $z = \frac{\lambda_0 - \lambda}{\lambda} = \frac{f - f_0}{f}$

## 2.29 Fission

1. Conservation of Energy
2. Binding Energies of nucleus is always  $(-)$ , like BE  $e^-$ 's

$$-BE_i + KE_i = -BE_f + KE_f \quad (34)$$

## 2.30 Wire Resistance

$$R = \frac{\rho L}{A} \quad (35)$$

## 2.31 Inside a Non-conducting Sphere of Uniform Charge Density

With constant surface potential, like a conductor.  $\nabla V = 0 = E$

## 2.32 Spin Matrices

$$S_i \Psi = \frac{\hbar}{2} \sigma_i \Psi \quad (36)$$

$$S_z |\uparrow\rangle = \frac{\hbar}{2} |\uparrow\rangle = \frac{\hbar}{2} \begin{pmatrix} 1 \\ 0 \end{pmatrix} \quad (37)$$

$$S_z |\downarrow\rangle = -\frac{\hbar}{2} |\downarrow\rangle = -\frac{\hbar}{2} \begin{pmatrix} 0 \\ 1 \end{pmatrix} \quad (38)$$

For example, eigenstate of  $S_x$  with  $-\frac{\hbar}{2}$ ,

$$\sigma_x = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$$

$$\frac{1}{\sqrt{2}}(|\uparrow\rangle - |\downarrow\rangle) = \frac{1}{\sqrt{2}} \left[ \begin{pmatrix} 1 \\ 0 \end{pmatrix} - \begin{pmatrix} 0 \\ 1 \end{pmatrix} \right] = \frac{1}{\sqrt{2}} \begin{pmatrix} \Psi \\ -1 \end{pmatrix} \quad (39)$$

$$\begin{pmatrix} 1 \\ -1 \end{pmatrix} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} = \begin{pmatrix} -1 \\ 1 \end{pmatrix} \rightarrow S_x \Psi = \frac{\hbar}{2} \begin{pmatrix} 1 \\ \sqrt{2} \end{pmatrix} \begin{pmatrix} -1 \\ 1 \end{pmatrix} = -\frac{\hbar}{2} \Psi \quad (40)$$

### 2.33 Collisions

1. For  $i \rightarrow f$  conditions, use conservation of momentum only!
2. For converting between  $U$  and  $KE$ , use  $KE$  only.
3.  $\epsilon = 1$  if elastic,  $\epsilon = 0$  if inelastic

$$\frac{\overbrace{|V_2| + |V_1|}^{\text{final}}}{\underbrace{|U_2| + |U_1|}_{\text{before}}} \quad (41)$$

4.  $KE$  is conserved in elastic collisions only

### 2.34 For what $v$ will car stay on hill?

$$F_c = F_g = mg \quad (42)$$

$$\frac{mv^2}{r} = mg \quad (43)$$

### 2.35 Böhr Model

1.  $e^-$  have classical motions
2.  $\Delta E = hf$
3. Quantization of angular momentum,  $L = n\hbar$
4.  $E_n = -\frac{Z^2 E_0}{n^2}$ ,  $E_n \propto \mu$
5.  $\Delta E = E_0 \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right) \rightarrow \frac{1}{\lambda} = R_y \left( \frac{1}{n_f^2} - \frac{1}{n_i^2} \right)$ ,  $R_y = 1 \times 10^7 \text{ m}^{-1}$
6. Positronium:  $\mu = \frac{m_e}{2} \rightarrow E_p = \frac{E_0}{2n^2}$

### 2.36 Hydrogen Spectral Series

1. Lyman:  $n_f = 1 \rightarrow \text{UV}$
2. Balmer:  $n_f = 2 \rightarrow \text{Visible}$
3. Paschen:  $n_f = 3 \rightarrow \text{IR}$

### 2.37 Fluids

Equilibrium when  $F_a = F_b$

## 2.38 Gauss with non-uniform densities

Must integrate. For example:

$$\rho = Ar^2, \rho \propto r^2, dV = 4\pi r^2 dr \quad (44)$$

$$\int E \cdot dA = \int \frac{\rho dV}{\epsilon_0} = \int_0^R \frac{\rho(4\pi r^2 dr)}{\epsilon_0} = 4\pi \int \frac{r^4 dr}{\epsilon_0} \quad (45)$$

## 2.39 Capacitors

If capacitors in series,  $Q_1 = Q_2$ . If parallel,  $V_1 = V_2$ .

## 2.40 Diffraction Limit

Airy Disk: circular aperture diffraction

$$\theta = \frac{1.22\lambda}{d}, \Delta l = \frac{1.22f\lambda}{d} \quad (46)$$

## 2.41 Time Dilation and Mass Contraction

$$t = \gamma t_0, X = \frac{X}{\gamma} \quad (47)$$

Used to relate a moving frame  $t, x$  to a rest frame's  $x_0, t_0$ . Cannot use these equations to relate two moving frames.

## 2.42 Expectation Value Problems

Look for even/odd functions

$$\int_0^T \sin x \cos x = 0 \text{ because orthogonal, } T = \text{period} \quad (48)$$

If  $\frac{\partial}{\partial x}$  for  $\langle p \rangle$  doesn't bring out an  $i$ , then  $\langle p \rangle = 0$

## 2.43 Normal Modes

1. Highest normal mode frequency when out of phase
2. Use limits if possible, if  $M \rightarrow \infty$ , etc.
3. # Frequencies = # masses
4. If odd # masses, one  $\omega$  will be  $\omega_0$ , others will be above/below.

For 2 masses hung by strings and connected by a spring,

1. In phase:  $\omega = \sqrt{\frac{g}{l}}$
2. Out of phase:  $\omega = \sqrt{\frac{2k}{m} + \frac{g}{l}} \rightarrow F = ma = K_{\text{eff}}x - mg \cos \theta = -2Kx - mg \cos \theta$

For 3 masses held together by strings in an m-M-m configuration:

1. 2 moving in opposite directions,  $M$  @ rest,  $\omega = \sqrt{\frac{k}{m}}$ , like they're attached to a wall.
2. Side masses are in phase, mid mass is out,  $\omega_2 = \sqrt{\frac{2k}{m}}$

For 2 masses connected to each other and then connected to 2 walls, all on springs with  $k$ - $k'$ - $k$  spring coefficient configuration:

1. In phase  $\Delta x_1 = \Delta x_2$  and  $k'$  isn't expanded  $\omega = \sqrt{\frac{k}{m}}$
2. Out of phase  $\Delta x_1 = -\Delta x_2$  CM  $k'$  stays in place so  $k'$  is split between  $m_1/m_2 \rightarrow$  force on each mass  $2k'$  since only  $1/2$  moves  $\rightarrow F = k + 2k'$ ,  $\omega = \sqrt{\frac{k+2k'}{m}}$

## 2.44 Radiation in Atoms

series	K	L	M	N	Specify what the final states are when coming from infinity.
$n_f =$	1	2	3	4	

## 2.45 Ionization Energy

1.  $E$  required to liberate outermost  $e^-$

## 2.46 Binding Energy

1. How tightly bound nucleons are
2. Reaches peak at Iron-56  
Elements below iron release  $E$  by fusion  
Elements above iron release  $E$  by fission
3. The mass of a nucleus is always less than  $\Sigma$  particle's mass  
The  $\Delta$  energy is the binding energy  
 $c^2(M_{\text{nucleus}} - \Sigma_{\text{nucleons}}) = BE$
4. More tightly bound means less mass/nucleon, more BE/nucleon
5. Created by the strong force
6. The energy given off during fusion/fission is the  $\Delta E$  between binding energies of fuel and products.

## 2.47 Hierarchy of Forces

1. Strong (100x E&M,  $10^5$ x weak,  $10^{39}$ x gravity)
2. E&M
3. Weak
4. Gravity

## 2.48 Pair Production

1. Creation of an elementary particle and it's anti-particle usually from a photon.

2. Cannot occur in free-space since the original momentum of the photon must be absorbed by something  
Usually near a nucleus or other photon
3. For  $e^-$  production, the  $E_{\text{photon}}$  must exceed 2x the rest energy of  $e^- = 1 \text{ MeV}$  or if 2 photons involved,  $E/\text{photon} = 500 \text{ keV}$
4. Dominates at high energy ( $> \text{MeV}$ )
5. Strangeness, momentum, electric charge, must be conserved

## 2.49 Spectral Lines

1. Less Dense  $\rightarrow$  more sharp/precise lines - don't lose  $E$  due to collisions
2. Sodium, famous yellow doublet  
created by spin/orbit coupling  
coupling becomes more pronounced in an external  $B$

## 2.50 Photon Interactions with Matter

1. Low  $E$ , elastically scatter  $\rightarrow$  Compton  $< 10^6 \text{ MeV}$
2. Med/Low  $E \rightarrow$  photoelectric  $< 10^7 \text{ MeV}$
3. High  $E \rightarrow$  pair production  $> 10^6 \text{ MeV}$

## 2.51 Neutron

1. Fermion,  $\text{spin} = \frac{1}{2}$
2.  ${}^1_0n$
3. Decay:  ${}^1_0n \rightarrow {}^1_1p^+ + {}^0_{-1}e + {}^0_0\bar{\nu}$
4. Capture:  ${}^1_1p^+ + {}^0_{-1}e \rightarrow {}^1_0n + {}^0_0\bar{\nu}$

## 2.52 Deuteron

1. Deuterium nucleus:  ${}^2_1H \leftrightarrow$  heavy hydrogen
2. Boson

## 2.53 Protium

1. Hydrogen nucleus  ${}^1_1H$
2. Fermion,  $s = \frac{1}{2}$

## 2.54 Davisson-Germer

1. Found diffraction pattern of  $e^-$  scattering off  $N_i$
2. Confirmed wave nature of matter
3. Plane spacing  $d = D \sin \theta$ ,  $D =$  interatomic spacing
4. Bragg:  $2d \sin \theta = n\lambda$

## 2.55 Kepler

1.  $T^2 \propto R^3$
2.  $\frac{dA}{dt} \propto L \rightarrow \frac{\text{area swept}}{\text{time}} \propto \text{angular momentum}$

## 2.56 Beats

1. Beats occur when 2 frequencies are similar
2. The # of beats  $\Rightarrow f_b = f_1 - f_2$
3. The harmonic is the index of  $n$  or whatever...

$$f_n = \underbrace{n}_{\text{this is the harmonic}} \overbrace{f_0}^{\text{fundamental}} \quad (49)$$