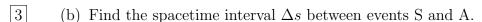
## Physics 370 Exam 1 Solutions October 5, 2016

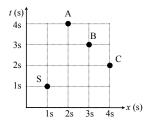
3 1. Explain the difference between inertial and noninertial reference frames.

Newton's First Law is valid in inertial frames

- 3 2. B Which runs more slowly?
  - A) a clock on top of a mountain B) a clock at the base of the mountain
  - C) It depends on where the observer is

- 3. The graph shows four events, in c = 1 units.
- (a) C Which event(s) could be simultaneous with S in some other frame? (That is, for which event(s) does there exist a frame such that S and that event are simultaneous?)
  - A) A B) B C) C D) none
  - E) B and C F) A and B G) A, B, and C

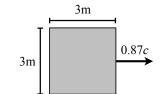




- $\Delta s = \sqrt{\Delta t^2 \Delta x^2} = \sqrt{9 1} = \sqrt{8} \,\mathrm{s} = 2.8 \,\mathrm{s}$
- [2] (c) What is the spacetime interval  $\Delta s$  between events S and A, in a frame that is moving at speed 0.8c to the right?

 $\Delta s = 2.8\,\mathrm{s}$ , the same in all frames

4. A  $3 \text{ m} \times 3 \text{ m}$  square (in its frame) is moving at 0.87c to the right in the "rest" frame (i.e. the frame of this paper).



- (a) B What does the square look like in the rest frame?
- (b) B What is the (horizontal) length of the square as seen 3 in the rest frame?
  - **A)** 1.35 m

3

- **B)** 1.5 m **C)** 1.7 m **D)** 3 m

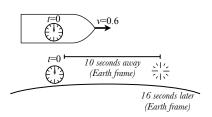
$$\gamma = \frac{1}{\sqrt{1-(0.87)^2}} = 2$$
 , so the width is  $3\,\mathrm{m}/2 = 1.5\,\mathrm{m}$ 

- 5. Passengers in a rocket travel to a distant star. It takes them 18 years to get there from our perspective, travelling at a velocity such that  $\gamma = 1.5$ .
- 3 (a) If event A is the moment the rocket leaves Earth, and event B is the moment the rocket arrives at the star, what is the distance  $\Delta x'$  between those events, in the passenger's frame?

Zero: both events occur in the same place in the passenger's frame

- (b) B How long does it take to reach the star, from a passenger's perspective? 3
  - A) 9 years
- B) 12 years
- C) 15 years
- **D)** 18 years
  - E) 27 years
- **F**) 36 years

3 6. When a rocket moving at 0.6c in the +x direction flies by an observer on Earth, clocks on the rocket and on Earth both read t=0. According to the Earthly observer, a firecracker goes off 16s later, at a distance of 10s away (using c=1 units). When the firecracker goes off, what time does the rocket observer see on her clock?



$$t' = \gamma(t - vx) = 1.25(16 - 0.6(10)) = \boxed{12.5 \text{ s}}$$

[3] 7. A rocket is travelling by the Earth at 0.7c to the right. Someone inside the rocket throws a ball that moves at 0.5c to the left, from their perspective. How fast is the ball moving in the Earth's frame, and in which direction?





$$\frac{0.7-0.5}{1-0.7(0.5)} = +0.3$$
 which is to the right

3 8. A True or false: Massless particles can be deflected by a gravitational field.

4 9. Fill in the blanks in this paragraph: The ultraviolet catastrophe was the paradox which predicted that a

blackbody emits <u>infinite</u> energy, due to there being too

may standing waves with **high** frequency. Planck solved the paradox by assuming that the energy of a given standing wave with frequency f must be

## an integer multiple of hf

3 10. Explain the difference between the photoelectric effect and bremsstrahlung. For the photoelectric effect,

photons hit metal and electrons are released

while for bremsstrahlung,

electrons hit metal and photons are created

311. A gamma ray hits a stationary electron, and is reflected directly backwards with a wavelength of 8 picometers. What was the wavelength of the original gamma ray? (Note:  $h/m_e c = 2.42 \,\mathrm{pm.}$ )



**A)** 3.16 pm **B)** 5.58 pm **C)** 8 pm **D)** 10.42 pm

**E)** 12.84 pm

$$8 = \lambda + \frac{h}{mc}(1 - \cos\theta) = \lambda + (2.42 \,\mathrm{pm})(2)$$

- 12. Green light (500 nm) shines on metal, and a stream of (nonrelativistic) electrons are detected from the metal moving at  $5 \times 10^5 \,\mathrm{m/s}$ .
- 3 (a) How much energy does each photon of green light possess?

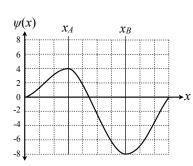
$$hf = h\frac{c}{\lambda} = \frac{(6.626 \times 10^{-34})(3 \times 10^8)}{500 \times 10^{-9}} = 4.0 \times 10^{-19}$$

3 (b) How much energy is required to liberate an electron from the surface of this metal?

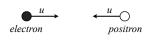
> Each electron emerges with kinetic energy  $KE=\frac{1}{2}mv^2=\frac{1}{2}(9.11\times 10^{-31})(5\times 10^5)^2=1.1\times 10^{-19}\,\mathrm{J}$ . The remaining energy from each photon is needed to liberate the electron; that is,  $\phi = 4.0 \times 10^{-19} - 1.1 \times 10^{-19} =$  $2.9 \times 10^{-19} \,\mathrm{J}$

(c) B We could turn off the flow of electrons by changing the light's 2 A) brightness B) color C) polarization

- $\boxed{3}$  13.  $\underline{\mathbf{A}}$  This graph shows a wavefunction  $\psi(x)$ . If P(x) is the probability that the object will be found at position x, what is  $P(x_A)/P(x_B)$ ?
  - **A)** 0.25 **B)** 0.5 **C)** 2
- D) 4
- **E**) 8



14. An electron and a positron are moving towards each other at u=0.6 (in c=1 units) and  $\gamma=1.25$ . The rest mass of an electron to 0.5 MeV. Answer the following questions in c=1 units.



(a) What is the total energy of the electron, in MeV?

$$E = \gamma m = (1.25)(0.5 \,\text{MeV}) = 0.625 \,\text{MeV}$$

(b) What is the total kinetic energy of the electron, in MeV?

$$KE = (\gamma - 1)m = (1.25 - 1)(0.5 \,\text{MeV}) = 0.125 \,\text{MeV}$$

(c) What is the momentum of the electron, in MeV?

$$p = (\gamma m)u = (0.625 \,\text{MeV})(0.6) = 0.375 \,\text{MeV}$$

(d) When the two particles collide, two identical photons are emitted in opposite directions. What is the frequency f of each photon, in Hz? (Be careful with the units here.)

$$f = \frac{E}{h} = \frac{0.625 \times 10^6 \,\text{eV}}{4.14 \times 10^{-15} \,\text{eV} \cdot \text{s}} = 1.5 \times 10^{20} \,\text{Hz}$$

215. B The figures show two wavefunctions,  $\psi_A(x)$  and  $\psi_B(x)$ , and their two Fourier transforms  $A_1(k)$  and  $A_2(k)$ , not necessarily in that order.  $A_1(k)$  is the Fourier transform of which of the wavefunctions?



- **A)**  $\psi_A(x)$  **B)**  $\psi_B(x)$
- 16. Consider a plane wave, i.e. a solution to the forceless Schrodinger equation

$$-\frac{\hbar^2}{2m}\frac{\partial^2\Psi}{\partial x^2} = i\hbar\frac{\partial\Psi}{\partial t}$$

with k = 0.1 / m and mass  $9 \times 10^{-31} \text{ kg}$ .

(a) Find the plane wave's angular frequency  $\omega$ .

$$\hbar\omega = \frac{\hbar^2 k^2}{2m} \implies \omega = \frac{\hbar}{2m} k^2 = \frac{1.055 \times 10^{-34} \,\mathrm{J \cdot s}}{2(9 \times 10^{-31} \,\mathrm{kg})} (0.1)^2 = 5.9 \times 10^{-7} \,\mathrm{rad/s}$$

(b) Write the plane wave  $\Psi(x,t)$  in terms of x and t, and no other variables. Ignore the normalization constant.

$$\Psi(x,t) = e^{i(kx - \omega t)} = e^{i(0.1x - 5.9 \times 10^{-7}t)}$$