

$$p = \gamma_u m u = \frac{m u}{\sqrt{1-u^2}} \quad c=1$$

$$\text{total energy } E = \gamma_u m = \overset{\text{rest energy}}{m} + \overset{\text{kinetic energy}}{(\gamma_u - 1)m}$$

for massless particles,

$$E = p$$

— — — —

Collision



hit & stick together

Find mass & velocity of final object.

Classical: $21m_0$

0

$$\text{Momentum conserved: } \gamma_{u_1} m_1 u_1 + \gamma_{u_2} m_2 u_2 = \gamma_{u_f} m_f u_f$$

$$\gamma_{0.8} = \frac{5}{3}$$

$$\gamma_{0.6} = \frac{5}{4}$$

$$\frac{5}{3}(9m_0)(0.8) - \frac{5}{4}(12m_0)(0.6) = \gamma_{u_f} m_f u_f$$

$$12m_0 - 9m_0$$

$$3m_0 = \gamma_{u_f} m_f u_f$$

Total

$$\text{Energy conserved (!): } \gamma_{u_1} m_1 + \gamma_{u_2} m_2 = \gamma_{u_f} m_f$$

$$\frac{5}{3}(9m_0) + \frac{5}{4}(12m_0)$$

$$15m_0 + 15m_0$$

$$30m_0 = \gamma_{u_f} m_f$$

$$3m_0 = (30m_0)u_f \rightarrow u_f = 0.1c$$

$$\gamma_{u_f} = \frac{1}{\sqrt{1-u_f^2}} = \frac{1}{\sqrt{1-0.1^2}} = 1.005$$

$$\therefore m_f = \frac{30m_0}{1.005} = 29.9m_0 > 21m_0$$

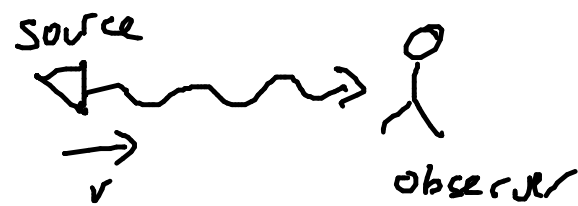
$$\Delta KE = -8.9m_0$$

$$KE = (\gamma - 1)m$$

When particles collide inelastically,
KE decreases & mass increases

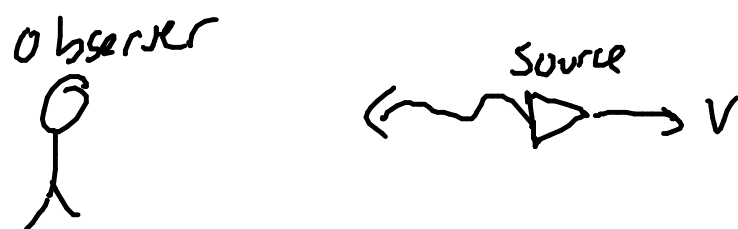
this is what particle accelerators do
to create massive particles. (e.g. Higgs)

Doppler Effect for Light



$$f' = f \frac{v}{v - v_s}$$

for light, math is a little different



Let Δt_p be time between wavefronts produced by source, as seen in observer's frame

Let Δt_a be time between wavefronts arriving to the observer, in her frame.

$$\Delta t_a = \Delta t_p + \frac{v \Delta t_p}{c}$$

Second wavefront has to travel additional distance $v \Delta t_p$

period of produced light is $T = \Delta t_p'$
in source's frame.

$$\Delta t_p = \gamma \Delta t_p' = \gamma T$$

$$c=1 \quad \Delta t_a = \Delta t_p (1+v) = \Delta t_p' \gamma (1+v)$$

$$\Delta t_a = \frac{1+v}{\sqrt{1-v^2}} T$$