

October 28, 2022

```
[311]: import numpy as np
import camb
from matplotlib import pyplot as plt
import time
from tqdm.auto import tqdm
```

1 1. Run Jon's sample script to plot the power spectrum of the CMB

The χ^2 value for the default parameters is 15267, which is fairly large, so probably not a very good fit.

```
[98]: def get_spectrum(pars,lmax=3000):
    #print('pars are ',pars)
    H0=pars[0]
    ombh2=pars[1]
    omch2=pars[2]
    tau=pars[3]
    As=pars[4]
    ns=pars[5]
    pars=camb.CAMBparams()
    pars.set_cosmology(H0=H0,ombh2=ombh2,omch2=omch2,mnu=0.06,omk=0,tau=tau)
    pars.InitPower.set_params(As=As,ns=ns,r=0)
    pars.set_for_lmax(lmax,lens_potential_accuracy=0)
    results=camb.get_results(pars)
    powers=results.get_cmb_power_spectra(pars,CMB_unit='muK')
    cmb=powers['total']
    tt=cmb[:,0]    #you could return the full power spectrum here if you wanted_
    ↪to do say EE
    return tt[2:]

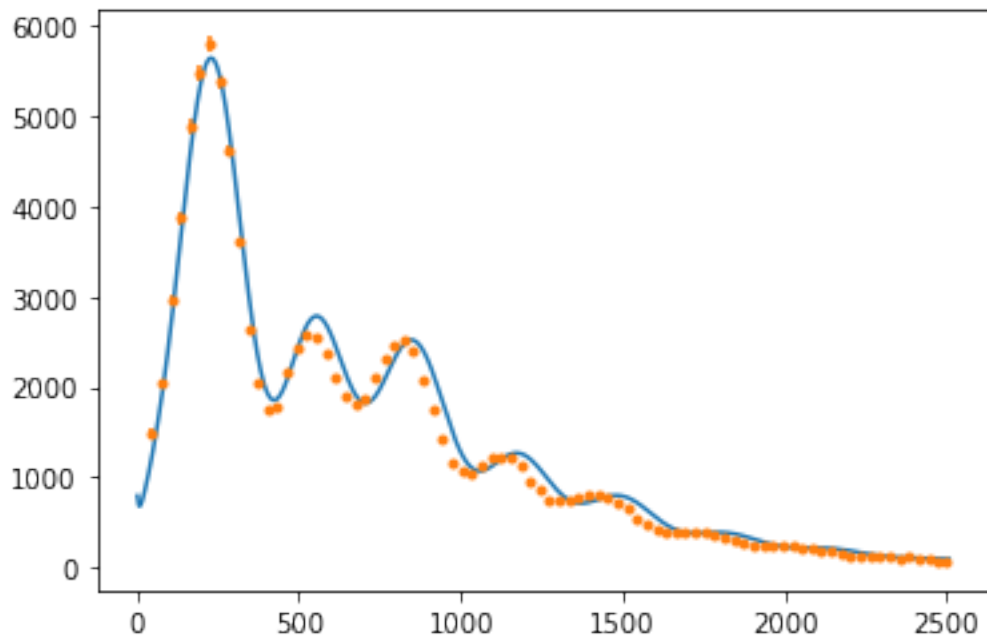
pars=np.asarray([60,0.02,0.1,0.05,2.00e-9,1.0])
planck=np.loadtxt('mcmc/COM_PowerSpect_CMB-TT-full_R3.01.txt',skiprows=1)
ell=planck[:,0]
spec=planck[:,1]
errs=0.5*(planck[:,2]+planck[:,3]);
model=get_spectrum(pars)
```

```

model=model[:len(spec)]
resid=spec-model
chisq=np.sum( (resid/errs)**2)
print("chisq is ",chisq," for ",len(resid)-len(pars)," degrees of freedom.")
#read in a binned version of the Planck PS for plotting purposes
planck_binned=np.loadtxt('mcmc/COM_PowerSpect_CMB-TT-binned_R3.01.
↳txt',skiprows=1)
errs_binned=0.5*(planck_binned[:,2]+planck_binned[:,3]);
plt.clf()
plt.plot(ell,model)
plt.errorbar(planck_binned[:,0],planck_binned[:,1],errs_binned,fmt='.')
plt.show()

```

chisq is 15267.937968194292 for 2501 degrees of freedom.



1.0.1 Try fitting again with different initial guesses for parameters

We can see that χ^2 improves by an order of magnitude (now 3272.2). I still wouldn't consider these values an acceptable fit since we can see that the variance is 6255001, which is still much too large an error bar for what we want for the fit.

```

[99]: pars=np.asarray([69,0.022,0.12,0.06,2.10e-9,0.95])
planck=np.loadtxt('mcmc/COM_PowerSpect_CMB-TT-full_R3.01.txt',skiprows=1)
ell=planck[:,0]
spec=planck[:,1]
errs=0.5*(planck[:,2]+planck[:,3]);

```

```

model=get_spectrum(pars)
model=model[:len(spec)]
resid=spec-model
chisq=np.sum( (resid/errs)**2)
print("chisq is ",chisq," for ",len(resid)-len(pars)," degrees of freedom.")
#read in a binned version of the Planck PS for plotting purposes
planck_binned=np.loadtxt('mcmc/COM_PowerSpect_CMB-TT-binned_R3.01.
    ↳txt',skiprows=1)
errs_binned=0.5*(planck_binned[:,2]+planck_binned[:,3]);
plt.clf()
plt.plot(ell,model)
plt.errorbar(planck_binned[:,0],planck_binned[:,1],errs_binned,fmt='.')
plt.show()

```

chisq is 3272.203604462886 for 2501 degrees of freedom.



```

[100]: dof = len(resid)-len(pars)
var = dof**2

print("Mean of chi^2 is:", dof)
print("Variance of chi^2 is", var)

```

Mean of chi² is: 2501
Variance of chi² is 6255001

```

[202]: print(pars)

```

```
[6.76128759e+01 2.22013439e-02 1.18295357e-01 2.46443537e+04
 3.86602109e-09 9.59422321e-01]
```

2. Use Newton's method to find the best fit parameters, using numerical derivatives

2.0.1 Function to take numerical derivatives:

This is the 2 point numerical derivative I wrote in PS1 but adapted in `partial_derivative` so that it is able to take arguments in order to compute partial derivatives.

To do the partial derivative, I use a wrapper (from my PS4 code) that evaluates the function we want to differentiate at each parameter as well as the array of parameters along the axis we're differentiating along. The optimal δ is $1/\sqrt{\varepsilon}$, but in this case, we need step sizes small enough so that τ converges, so we use a non-optimal $\delta = 1e-13$.

χ^2 from Newton's method is slightly larger but comparable to the LS fit using improved parameters above. I take into account the noise matrix here, which is just $1/\text{variance}$ of the errors given in the problem placed along the diagonal of a matrix. The errors are just the average of the upper (column 2 of the dataset) and lower (column 3) bounds of the error bars

```
[213]: def derivative(f,x,delta):
        return (1/(12*delta))*(f(x-2*delta)-8*f(x-delta)+8*f(x+delta)-f(x+2*delta))

def partial_derivative(func, var, point=[]):
    args = point[:]
    def wraps(x):
        args[var] = x
        return func(args)
    return derivative(wraps, point[var], delta = 1e-13)
```

```
[370]: def fit(func,p):
        y = func(p)
        grad=np.zeros([len(y),len(p)])
        for i in range(len(p)):
            grad[:,i] = partial_derivative(func,i,p)
        return y,grad
```

```
[656]: p0 = np.asarray([69.,0.022,0.12,0.06,2.10e-9,0.95])
        p = p0.copy()
        spec=planck[:,1]
        err_avg = (planck[:,2]+planck[:,3])/2
        Ninv= np.diag(1/err_avg**2)

        for i in tqdm(range(7)):
            pred,grad = fit(get_spectrum,p)
            pred,grad = pred[:len(spec)],grad[:len(spec)]
            r = spec - pred
```

```

err = np.sum(r**2)
r=r.T
grad = grad[:len(spec)]
lhs=grad.T@Ninv@grad
rhs=grad.T@Ninv@r
dp=np.linalg.inv(lhs)@(rhs)
for i in range(len(p)):
    p[i]=p[i]+dp[i]

```

0%| | 0/7 [00:00<?, ?it/s]

2.0.2 Note that the covariance (curvature) matrix is calculated below - we'll use this for the MCMC in the next question!

```

[657]: cov = np.linalg.inv(lhs)
param_errs = np.sqrt(np.diag(cov))

print(f"Best-fit parameters are:\n\
H0 = {p[0]} +/- {param_errs[0]}\n\
 $\Omega_{bh}^2$  = {p[1]} +/- {param_errs[1]}\n\
 $\Omega_{ch}^2$  = {p[2]} +/- {param_errs[2]}\n\
    = {p[3]} +/- {param_errs[3]}\n\
As = {p[4]} +/- {param_errs[4]}\n\
ns = {p[5]} +/- {param_errs[5]}")

```

Best-fit parameters are:

```

H0 = 69.00259551783105 +/- 0.0003826488765396756
 $\Omega_{bh}^2$  = 0.02199951565154579 +/- 6.357161383414019e-08
 $\Omega_{ch}^2$  = 0.11999955916739918 +/- 4.018629892439298e-08
    = 0.08247408643159534 +/- 0.030853467319198837
As = 2.232716888071639e-09 +/- 1.4025755984955074e-10
ns = 0.9725744241242387 +/- 0.005219030675878392

```

Save output to planck_fit_params.txt

```

[658]: write = np.column_stack([p,param_errs])
file = 'planck_fit_params.txt'
np.savetxt(file , write, fmt=['%f','%f'])

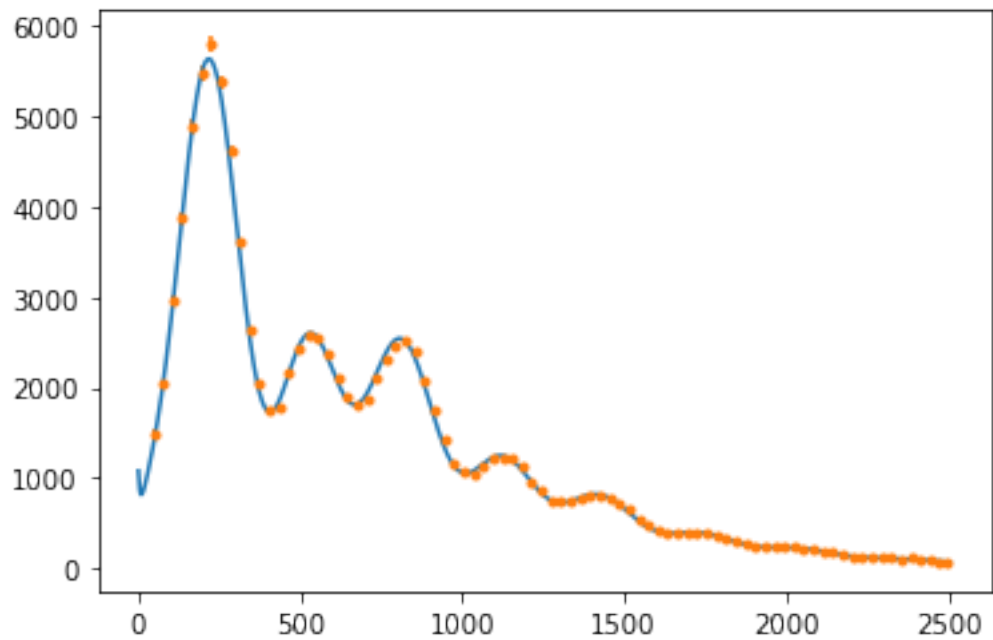
```

```

[659]: plt.plot(pred)
plt.errorbar(planck_binned[:,0],planck_binned[:,1],errs_binned,fmt='. ')

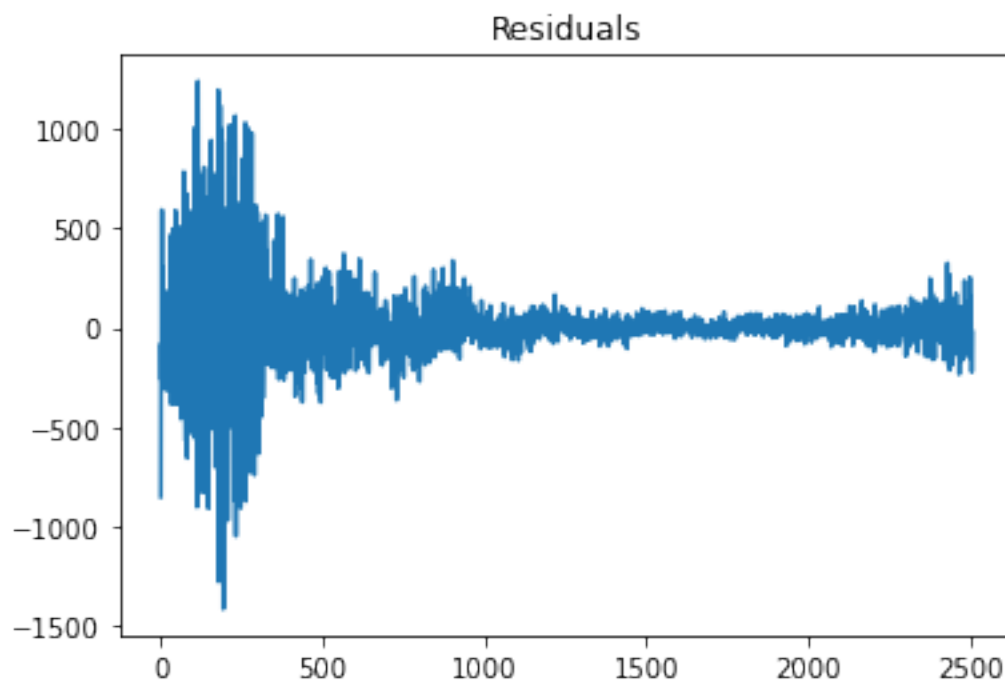
```

[659]: <ErrorbarContainer object of 3 artists>



```
[660]: plt.plot(spec-pred)
plt.title("Residuals")
```

```
[660]: Text(0.5, 1.0, 'Residuals')
```



```
[661]: errs=0.5*(planck[:,2]+planck[:,3])
resid = spec-pred
chisq=np.sum((resid/errs)**2)
print("Chi^2 is:", chisq, "for", len(resid), "degrees of freedom")
```

Chi^2 is: 2757.862625093855 for 2507 degrees of freedom

3 3. Estimate parameters and error using an MCMC

The `mcmc` function is adapted from Jon's `mcmc_class.py` (from 2021 github). I added a constraint that τ must be > 0.1 in order for the CAMB package to stop throwing errors if τ becomes too large and therefore unphysical, τ should also be positive, so I added in a constraint to reject if it is ever < 0 .

I only used 5000 steps (ideally I'd want 10000+) because I was struggling all week trying to debug my τ and n_s parameters running away from their expected magnitudes from Newton's method. In the interest of time, I decided to just run 5000 steps and finish the assignment with an unconverged chain rather than having nothing :)

If the chains are converged, the first plot should look like white noise and the second plot (power spectrum) should be flat at low k values - the left of the plot.

Neither of the plots look like this, and n_s is still moving far away from where it should be, so, since the parameters are correlated, more iterations don't really produce any better results.

```
[666]: def calc_chisq(pars,data,err):
    pred=get_spectrum(pars)
    pred=pred[:len(err)]
    residuals = data-pred
    return np.sum((residuals/err)**2)

def mcmc(pars,err,y,cov,fun,scale,nstep):
    accept_num = 0
    chi_cur=fun(pars,y,err)
    npar=len(pars)
    chain=np.zeros([nstep,npar])
    chivec=np.zeros(nstep)
    for i in tqdm(range(nstep)):
        trial_pars=pars+scale*np.random.multivariate_normal(np.
→zeros(len(pars)),cov)
        trial_chisq=fun(trial_pars,y,err)
        delta_chisq=trial_chisq-chi_cur
        accept_prob=np.exp(-0.5*delta_chisq)
        accept=np.random.rand(1)<accept_prob

        if accept:
            accept_num = accept_num + 1
```

```

        pars=trial_pars
        chi_cur=trial_chisq
        chain[i,:]=pars
        chivec[i]=chi_cur
    return chain,chivec,accept_num

```

```

[668]: scale_vals = 1
p_new = p.copy()
chain2,chisq2,accept=mcmc(p_new,spec,errs,cov,calc_chisq,scale=scale_vals,nstep=5000)

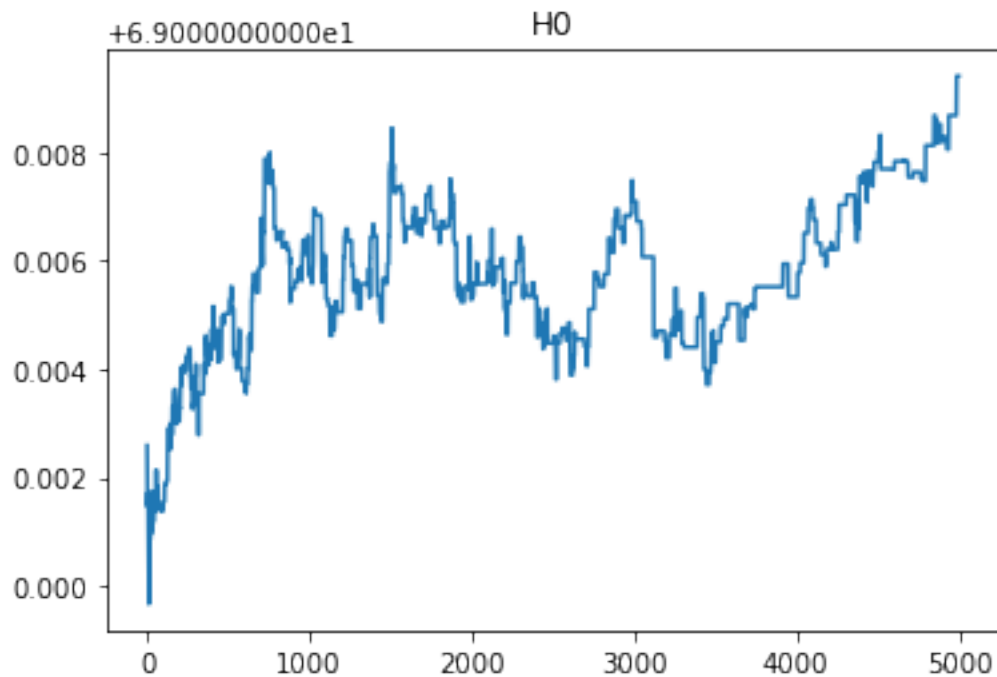
0%|          | 0/5000 [00:00<?, ?it/s]

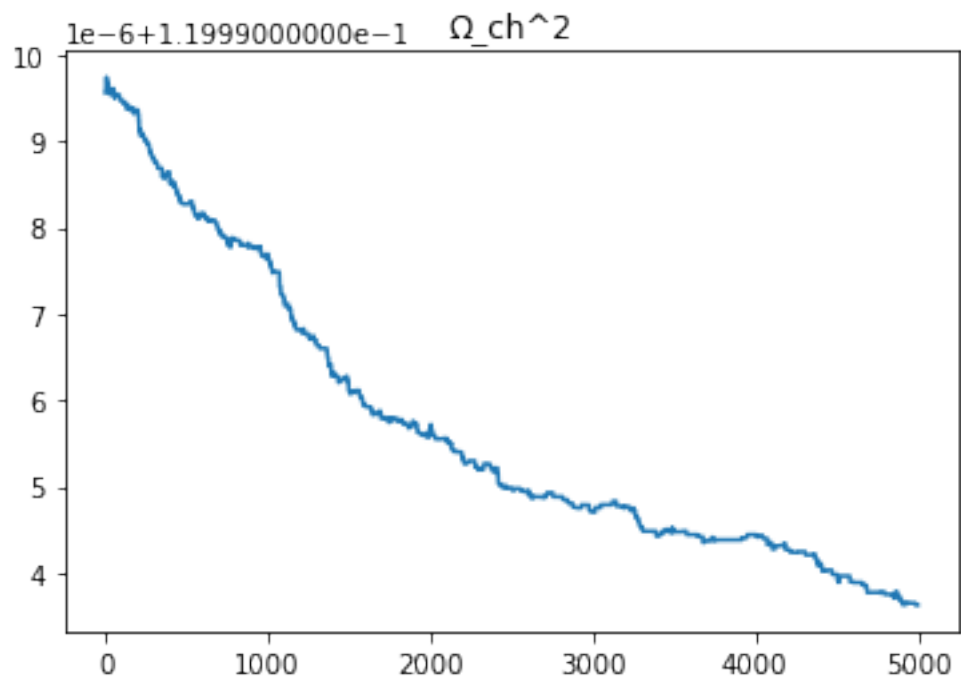
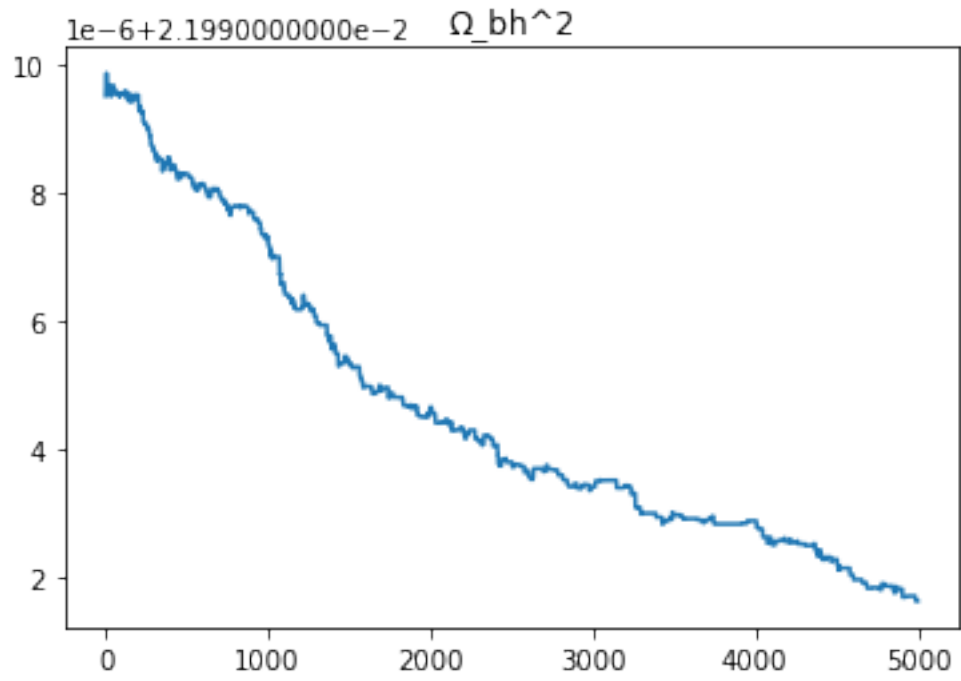
```

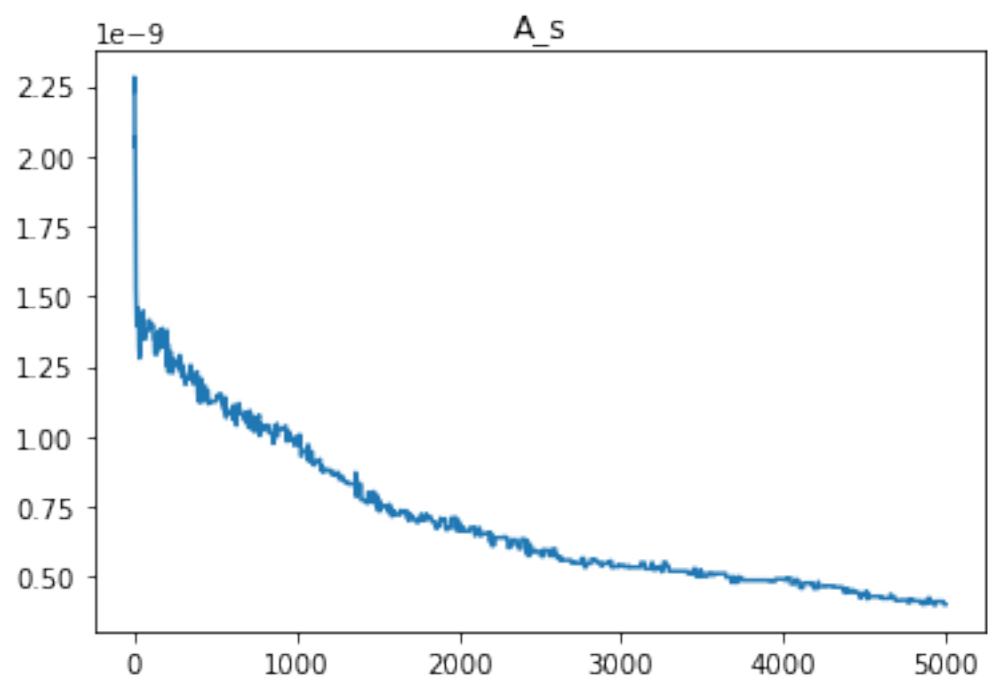
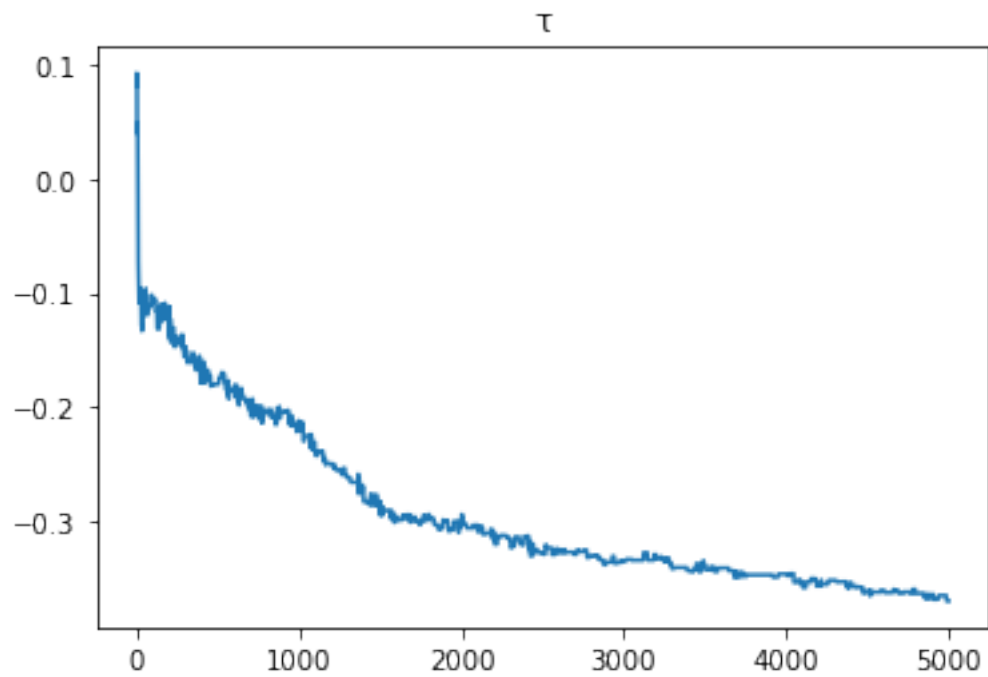
```

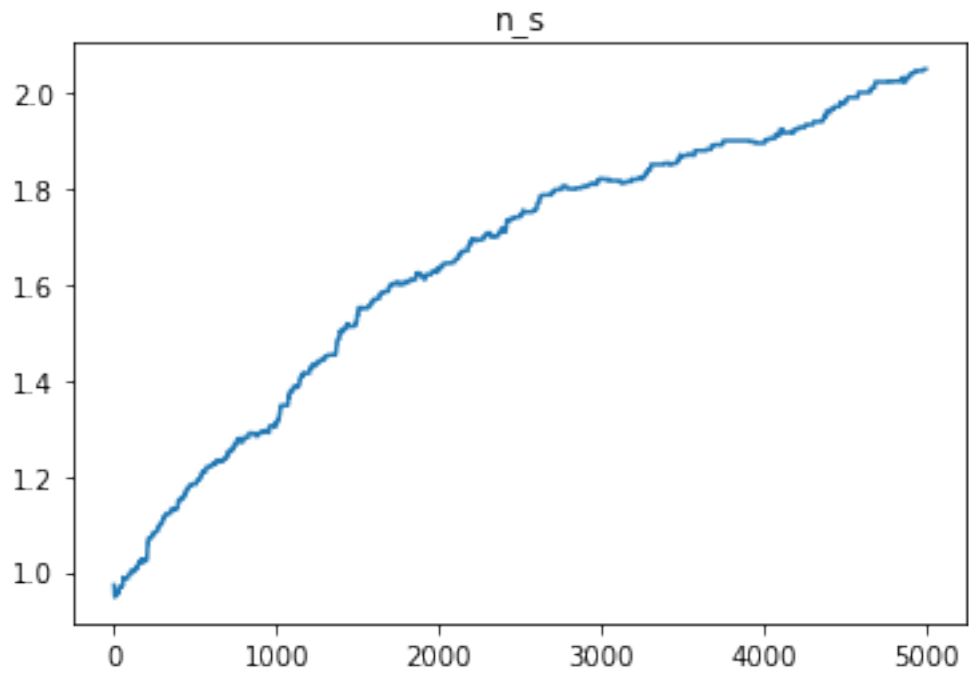
[672]: titles = np.asarray(['H0', ' $\Omega_{bh}^2$ ', ' $\Omega_{ch}^2$ ', ' ', 'A_s', 'n_s'])
for i in range(len(p_new)):
    plt.plot(chain2[:,i])
    plt.title(titles[i])
    plt.show()
for i in range(6):
    ft_chain = np.abs(np.fft.rfft(chain2[:,i]-np.mean(chain2[:,i]))**2)
    plt.loglog(ft_chain)
plt.legend(['H0', ' $\Omega_{bh}^2$ ', ' $\Omega_{ch}^2$ ', ' ', 'A_s', 'n_s'])

```

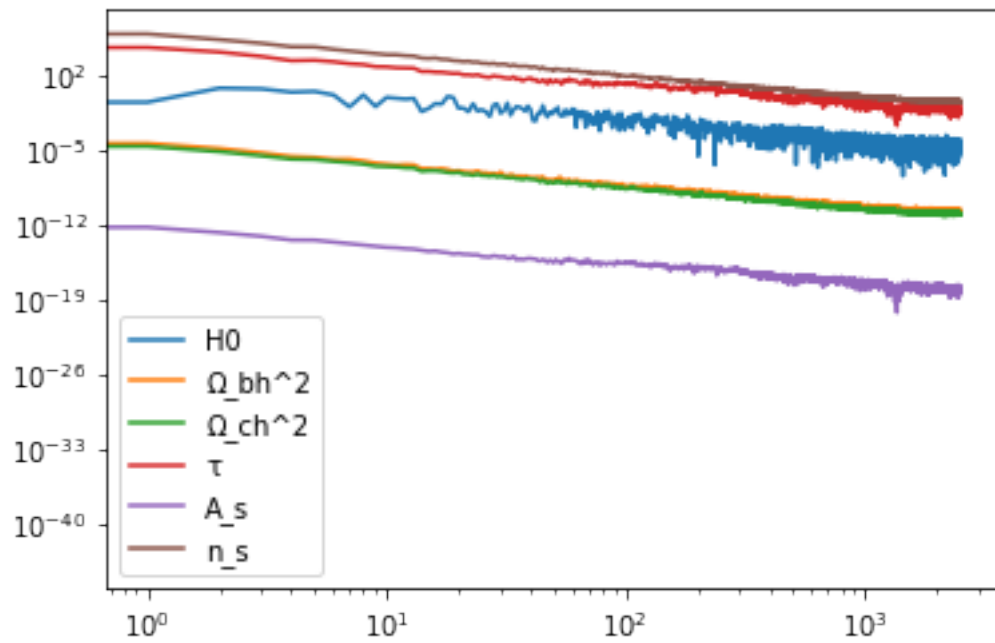








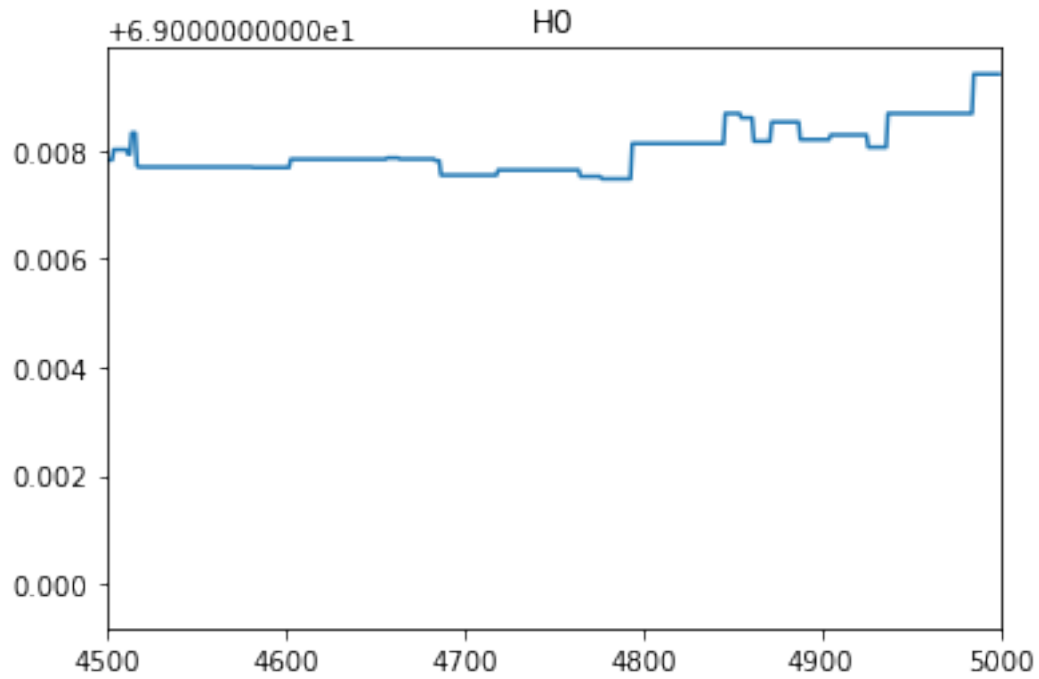
[672]: <matplotlib.legend.Legend at 0x7f7bc0e29790>

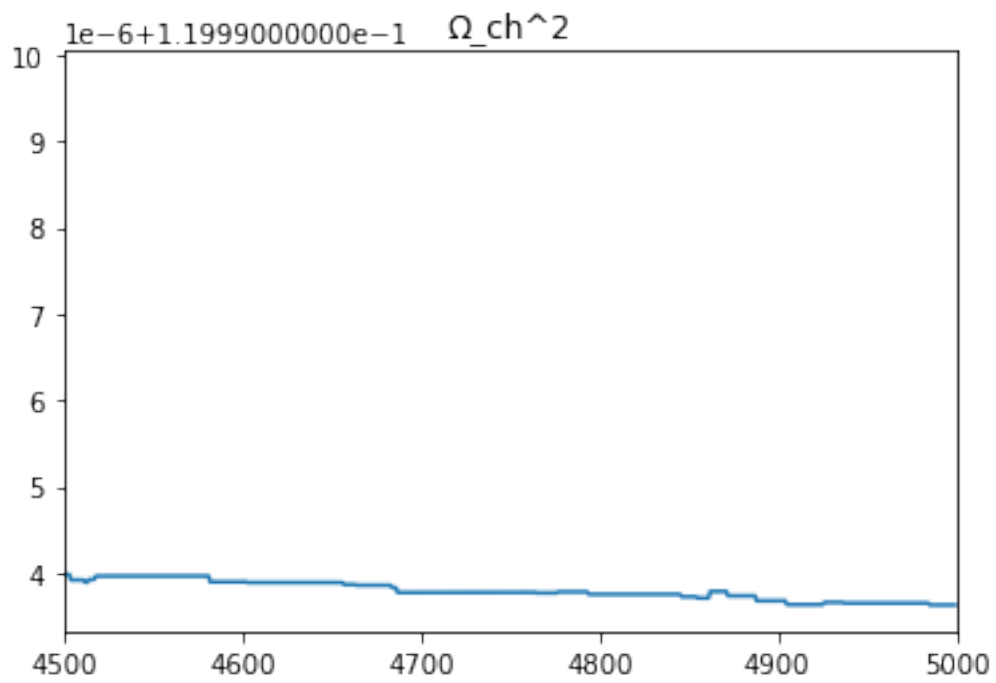
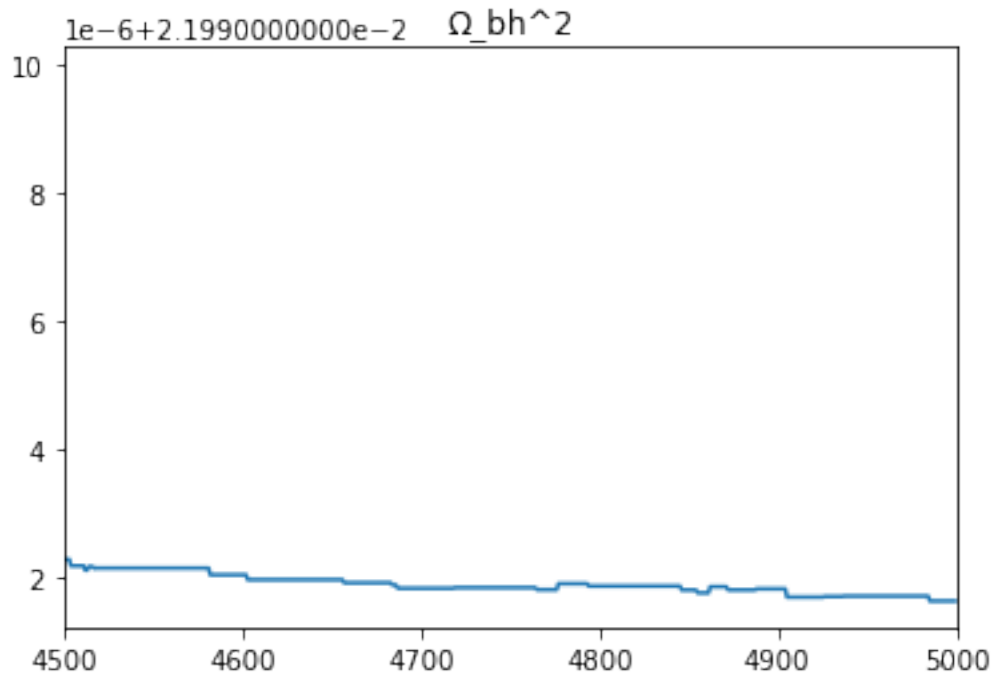


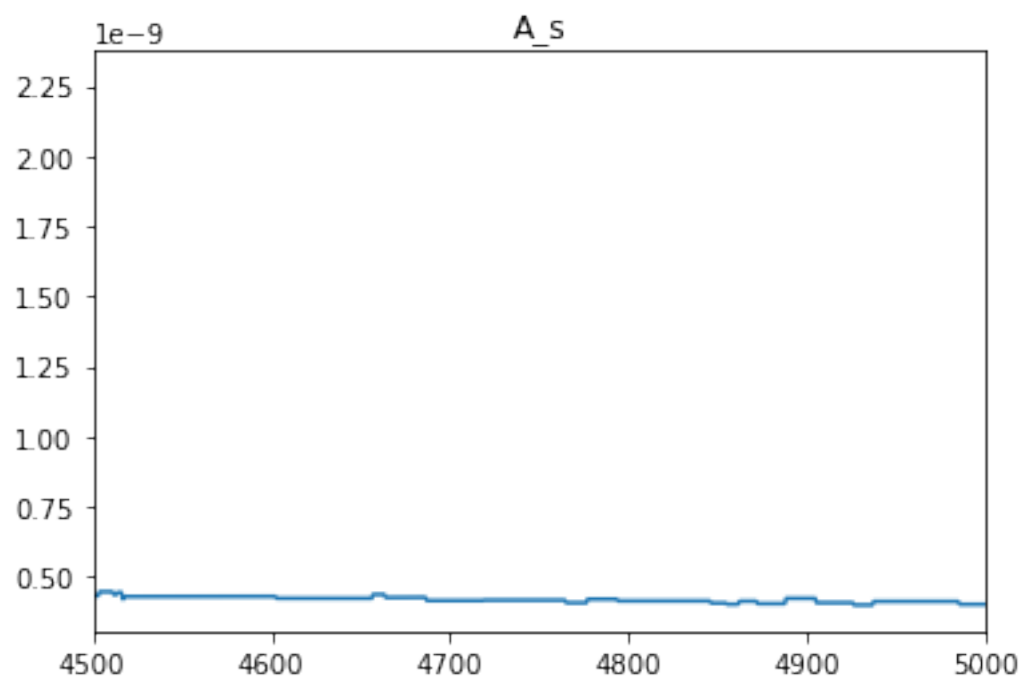
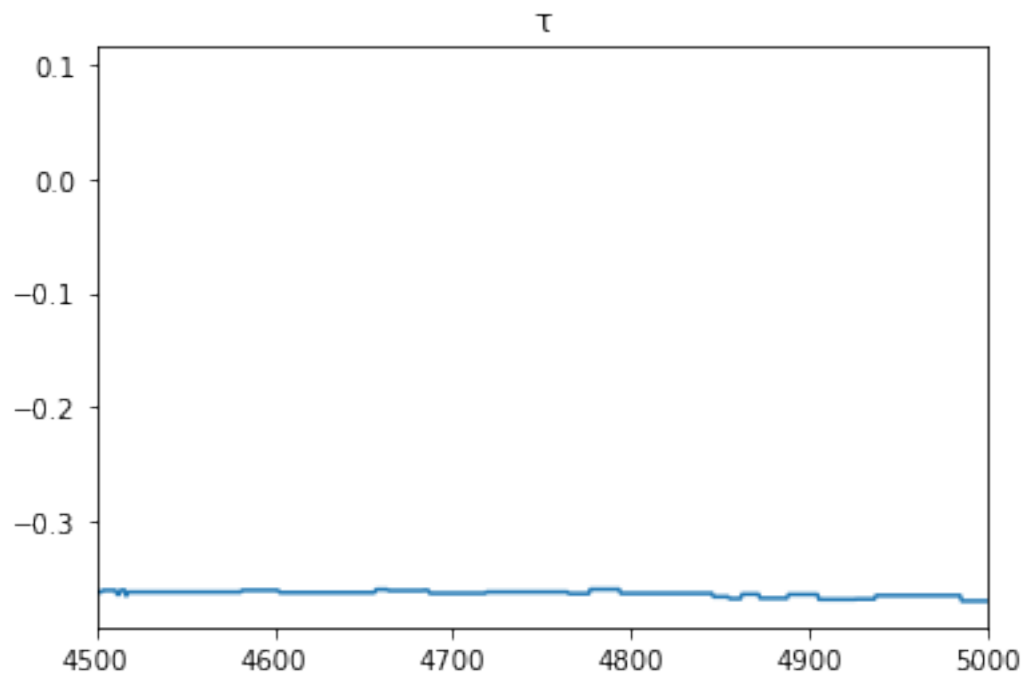
3.0.1 Is this any better if we zoom in to the ends of the chains?

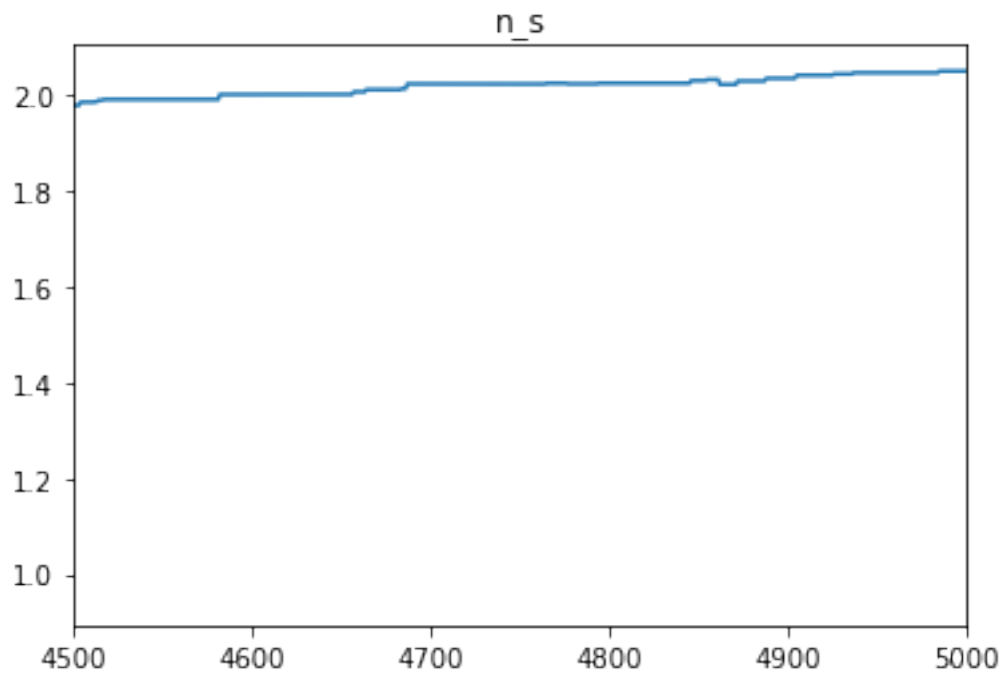
No, but it's not bouncing around like crazy or anything, so I think we can continue

```
[673]: titles = np.asarray(['H0', ' $\Omega_{bh}^2$ ', ' $\Omega_{ch}^2$ ', ' ', 'A_s', 'n_s'])
for i in range(len(p_new)):
    plt.plot(chain2[:,i])
    plt.title(titles[i])
    plt.xlim(4500,5000)
    plt.show()
```



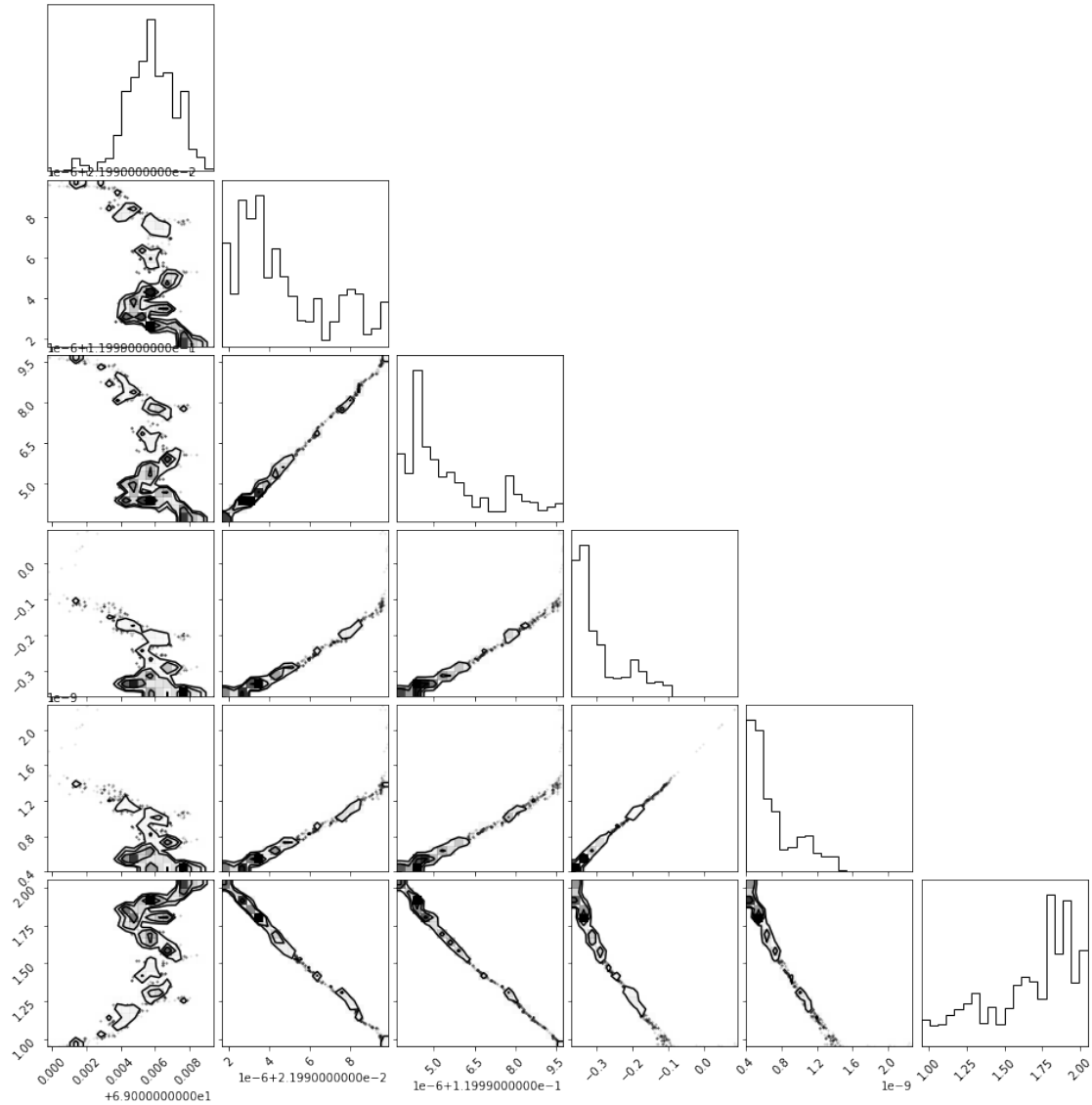






```
[674]: import corner  
figure = corner.corner(chain2)
```

WARNING:root:Too few points to create valid contours



3.0.2 Save chain to text file

```
[676]: mcmc_write = np.column_stack([chain2, chisq2])
mcmc_file = 'planck_chain.txt'
np.savetxt(mcmc_file, mcmc_write, fmt=['%f', '%f', '%f', '%f', '%f', '%f', '%f'])
```

3.0.3 Calculate the parameters and errors from the average/std of the converged part of the MCMC chain

```
[703]: mcmcp = np.mean(chain2[4500:5000], axis=0)
mcmce = np.std(chain2[4500:5000], axis=0)
```



```
print(f"MCMC best-fit parameters are:\n\
H0 = {mcmcp[0]} +/- {mcmce[0]}\n\
Ωbh2 = {mcmcp[1]} +/- {mcmce[1]}\n\
Ωch2 = {mcmcp[2]} +/- {mcmce[2]}\n\
    = {mcmcp[3]} +/- {mcmce[3]}\n\
As = {mcmcp[4]} +/- {mcmce[4]}\n\
ns = {mcmcp[5]} +/- {mcmce[5]}")
```

MCMC best-fit parameters are:

```
H0 = 69.0080142863614 +/- 0.0004417930181979411
Ωbh2 = 0.021991879515911097 +/- 1.496922294700448e-07
Ωch2 = 0.11999379063187567 +/- 1.0657011555891349e-07
    = -0.36429926421939707 +/- 0.002493067559703605
As = 4.1873911042186683e-10 +/- 9.808733821763588e-12
ns = 2.020364732315966 +/- 0.01889960727008938
```

3.0.4 Estimate on dark energy and its uncertainty

We know that $\Omega_b + \Omega_c + \Omega_\lambda = 1$, where $h = H_0/100$ and Ω_b and Ω_c are parameters in our chain.

$$\Rightarrow \Omega_\lambda = 1 - \Omega_b - \Omega_c$$

Error can be calculated by error propagation using partial derivatives (I used this resource:

[https://chem.libretexts.org/Bookshelves/Analytical_Chemistry/Supplemental_Modules_\(Analytical_Chemistry](https://chem.libretexts.org/Bookshelves/Analytical_Chemistry/Supplemental_Modules_(Analytical_Chemistry)

This shouldn't be negative, so there's clearly an error with the parameters from our MCMC

```
[704]: omega_b = mcmcp[1]
omega_b_err = mcmce[1]

h = mcmcp[0]
h_err = mcmce[0]

omega_c = mcmcp[2]
omega_c_err = mcmce[2]

omega_lambda = omega_b/((h/100)**2) - omega_c/((h/100)**2)
err_omega_lambda = (omega_b_err/omega_b + (2*h_err/h))*(omega_b/((h/100)**2)) +
    ↪ (omega_c_err/omega_c + (2*h_err/h))*(omega_c/((h/100)**2))

print("Energy density of dark energy is", omega_lambda, "+/-", err_omega_lambda)
```

Energy density of dark energy is -0.20579531488865443 +/- 4.355766377088003e-06

4. Run a new chain including polarization data constraint

Total likelihood = old likelihood (previous chain) * new likelihood

We can weight each chain sample by the new likelihood \Rightarrow new distributions = weighted averages over chain

The likelihood ratio is given by $e^{-\frac{1}{2}(\frac{\delta\tau}{\tau_{\text{uncertainty}}})^2}$

The `process_chain` function is adapted from Jon's `mcmc_high_T.py`

```
[705]: def process_chain(chain,tau_prior,tau_prior_err):
    dchi=chain[:,3] - tau_prior

    wt=np.exp(-0.5*(dchi/tau_prior_err)**2) #the magic line that importance_
    ↪ samples

    #calculate the weighted sum of the chain and the chain squared
    npar=chain.shape[1]
    tot=np.zeros(npar)
    totsqr=np.zeros(npar)
    for i in range(npar):
        tot[i]=np.sum(wt*chain[:,i])
        totsqr[i]=np.sum(wt*chain[:,i]**2)
    #divide by sum of weights
    mean=tot/np.sum(wt)
    meansqr=totsqr/np.sum(wt)

    #variance is <x^2>-<x>^2
    var=meansqr-mean**2
    return mean,np.sqrt(var),wt
```

```
[706]: mcmc_mean,mcmc_sig,weights = process_chain(chain2.T,0.0540,0.0074)
```

4.0.1 Use these weights to compute a new covariance matrix before running our next chain and run a new chain starting at the end values of our chain from problem 3

We need to change the χ^2 function slightly, so that we factor in χ^2 associated with the difference between our parameter τ and the prior 0.054 ± 0.0074 - this is encompassed in the function `calc_chisq_tau` and we just run `mcmc`, the function from problem 3.

```
[717]: new_cov = np.cov(m=chain2,aweights=weights)

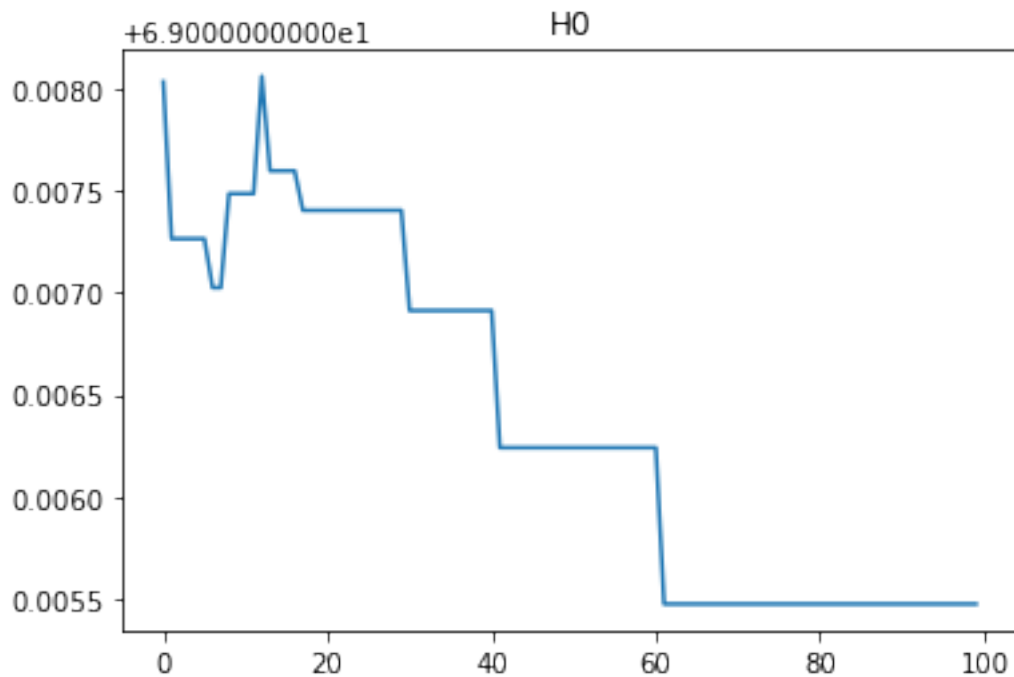
old_chain = chain2[:,-1]
p_tau = mcmcp.copy()
p_tau[3] = 0.054
print(p_tau)

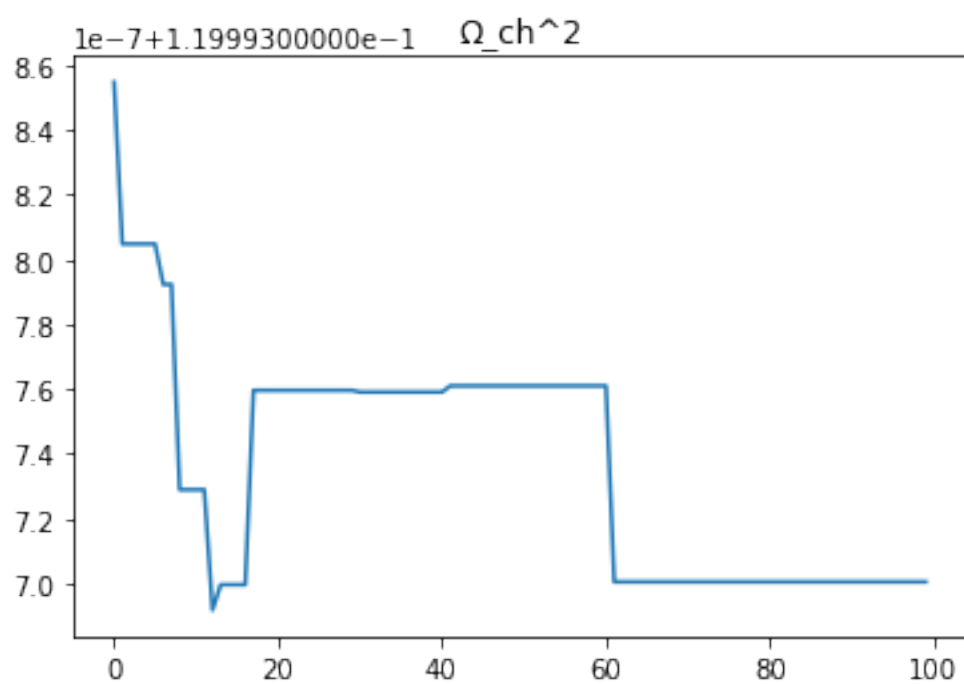
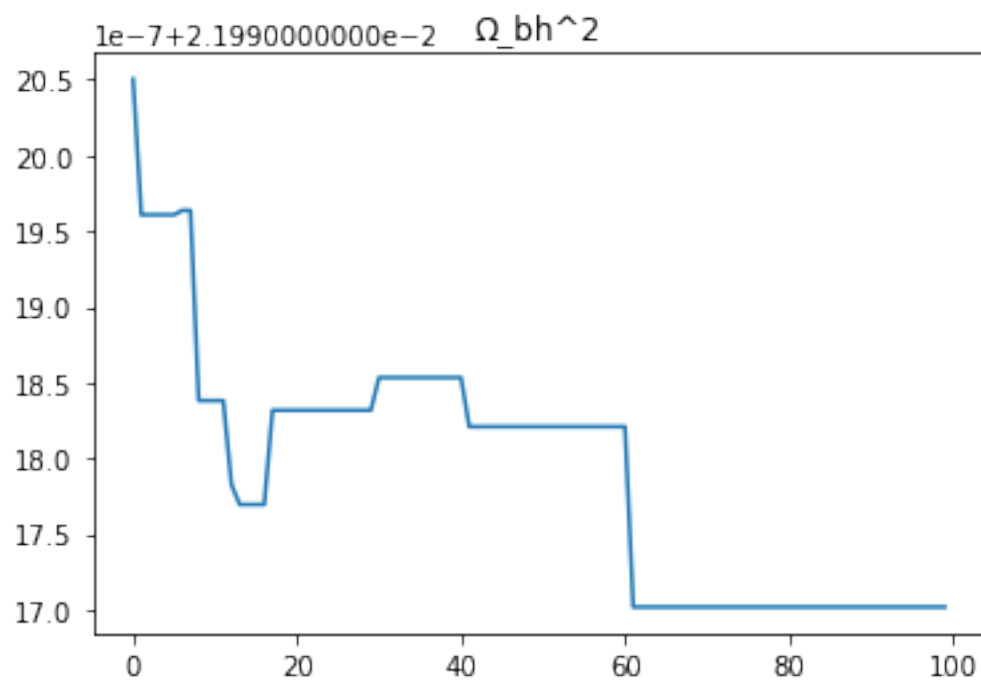
def calc_chisq_tau(pars,data,err):
    pred=get_spectrum(pars)
    pred=pred[:len(err)]
    residuals = data-pred
    chisq = np.sum((residuals/err)**2)
    return chisq + ((pars[3]-0.054)/0.0074)**2
```

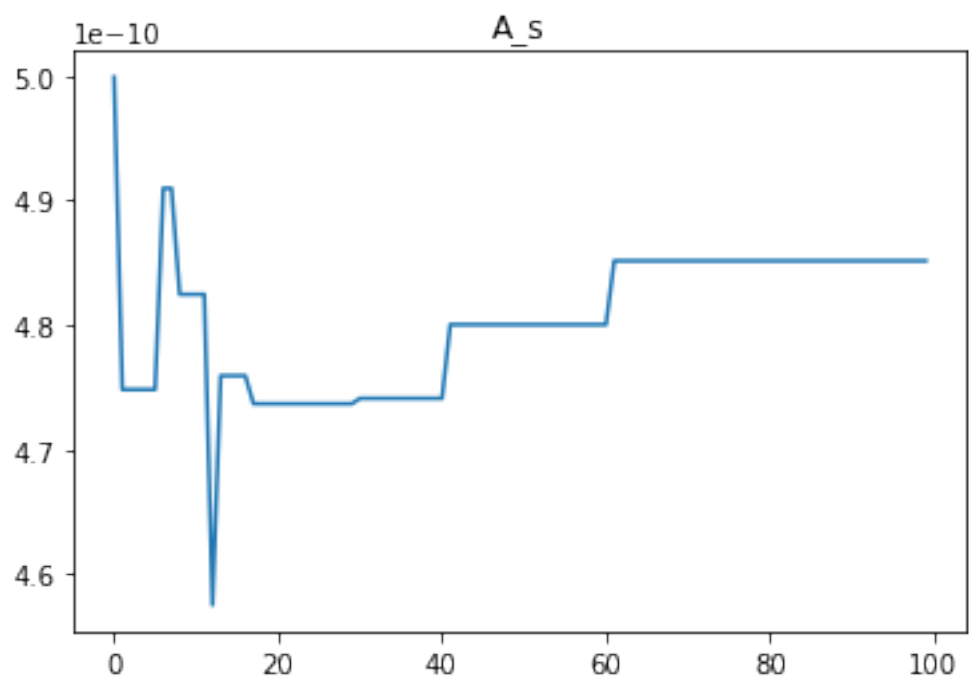
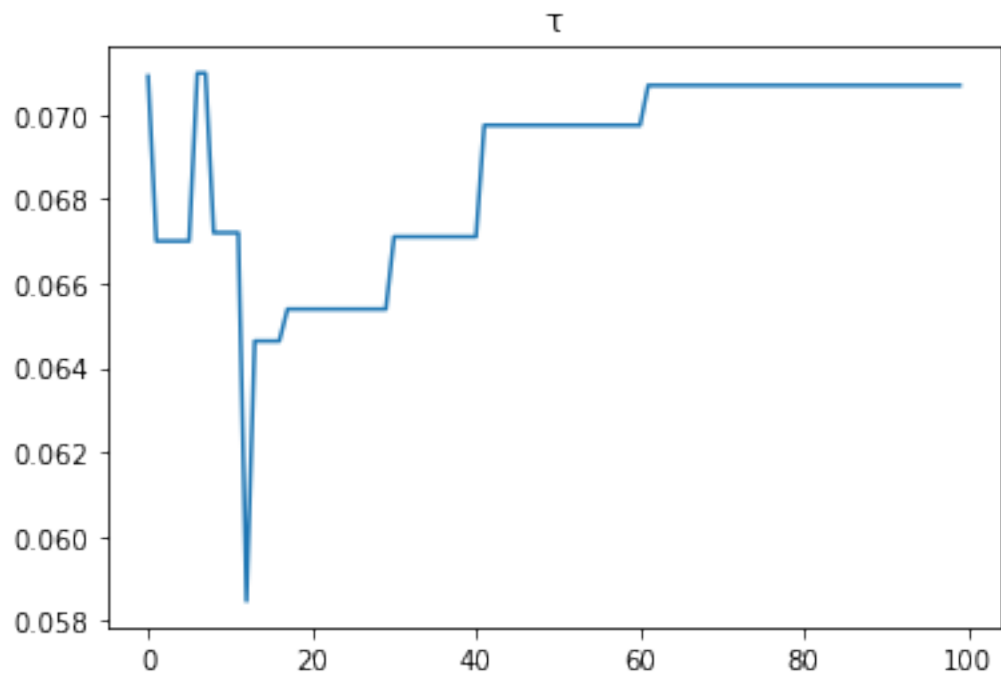
```
[6.90080143e+01 2.19918795e-02 1.19993791e-01 5.40000000e-02
 4.18739110e-10 2.02036473e+00]
```

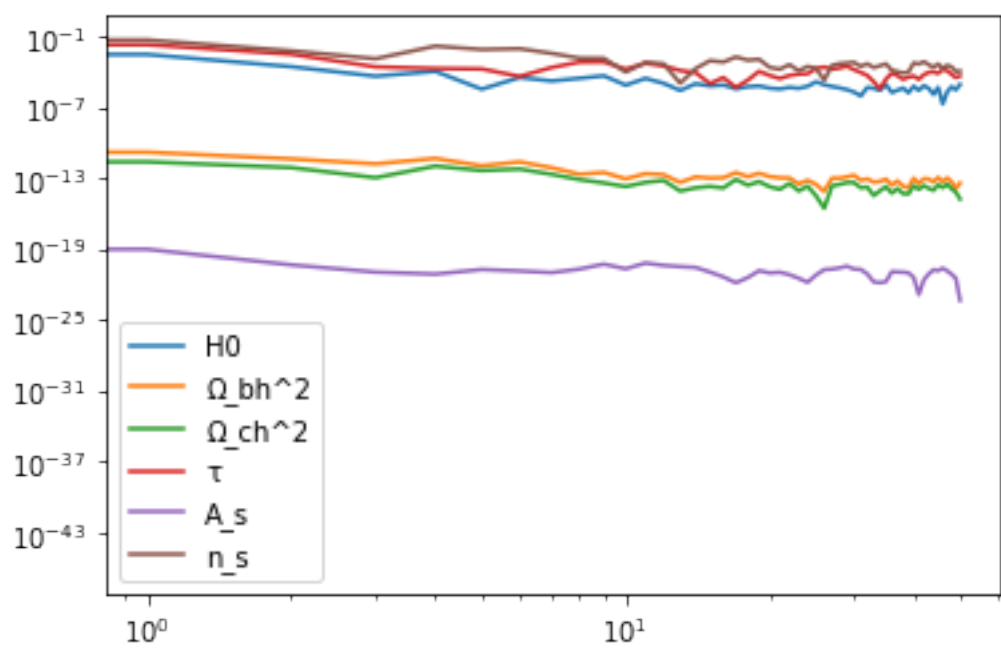
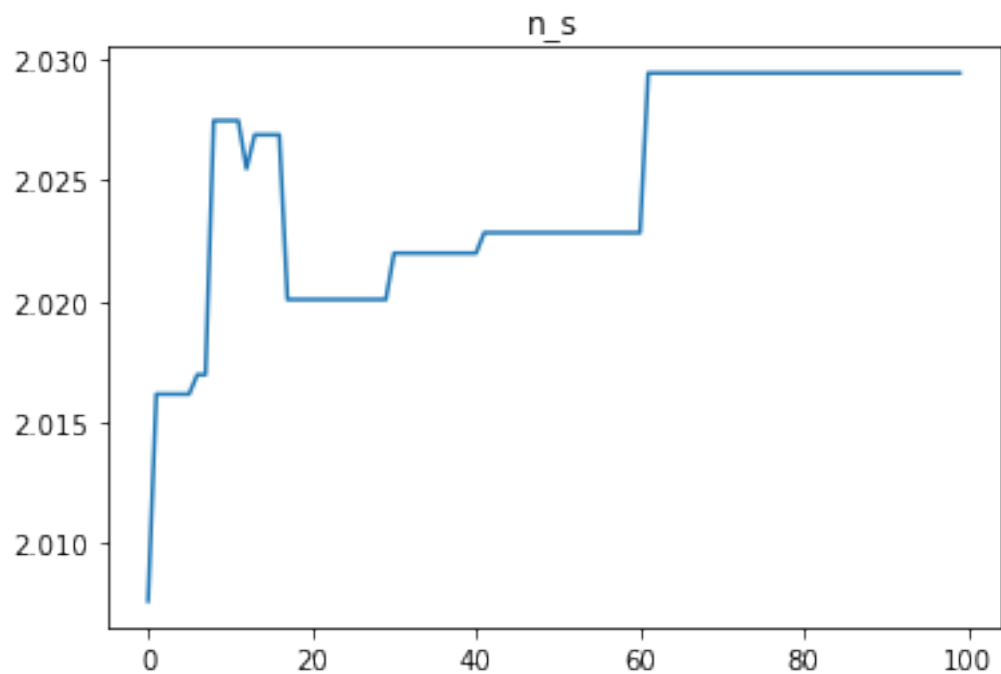
```
[709]: chain_tau,chisq_tau,accept_tau=mcmc(p_tau,spec,errs,cov,calc_chisq_tau,scale=scale_vals,nstep=
0%|          | 0/100 [00:00<?, ?it/s]
```

```
[711]: for i in range(len(p_new)):
    plt.plot(chain_tau[:,i])
    plt.title(titles[i])
    plt.show()
    for i in range(6):
        ft_chain = np.abs(np.fft.rfft(chain_tau[:,i]-np.mean(chain_tau[:,i]))**2)
        plt.loglog(ft_chain)
    plt.legend(['H0', ' $\Omega_{bh}^2$ ', ' $\Omega_{ch}^2$ ', ' ', 'A_s', 'n_s'])
    plt.show()
    figure = corner.corner(chain2)
```

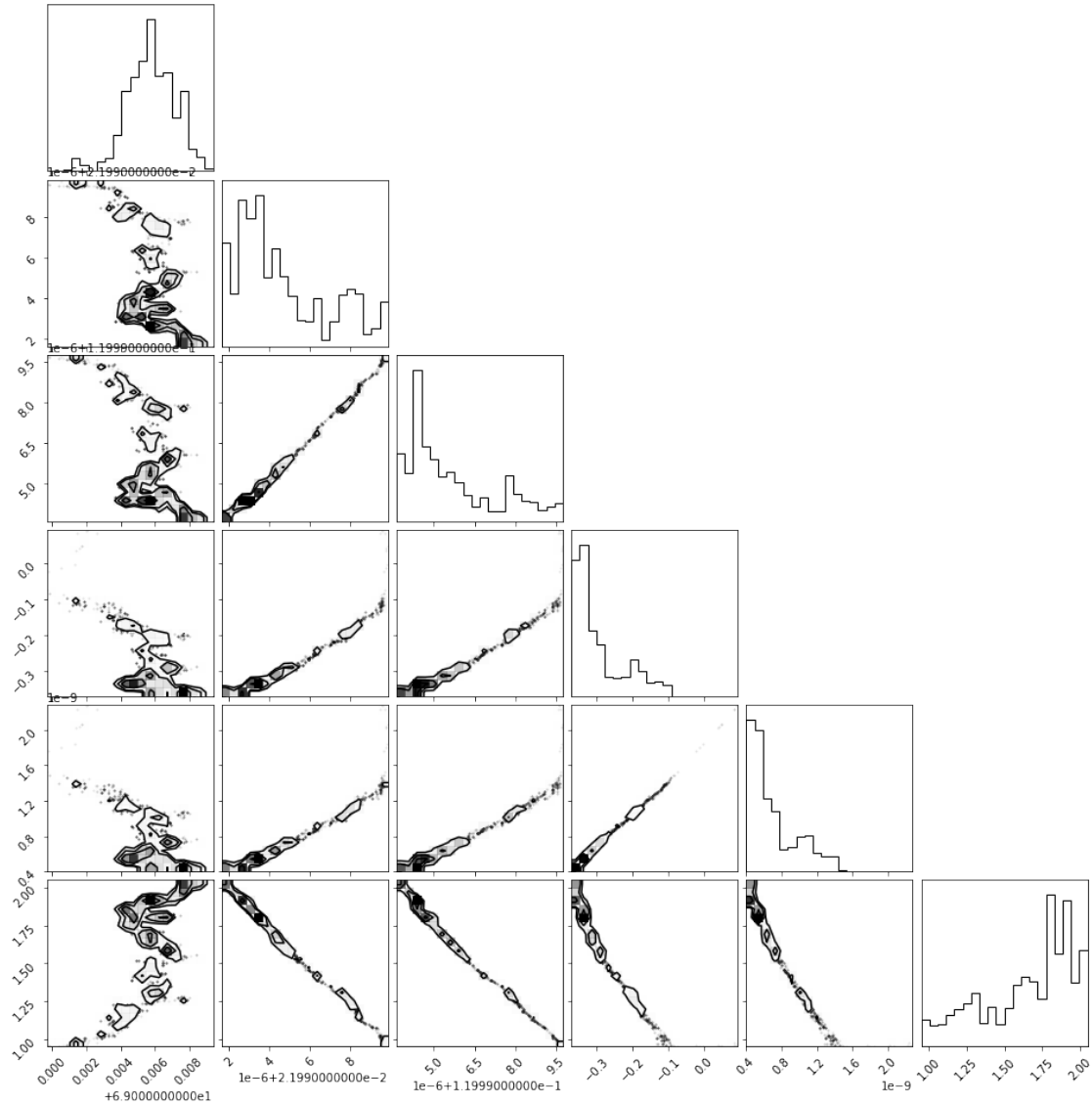








WARNING:root:Too few points to create valid contours



```
[722]: mcmc_tau = np.mean(chain_tau,axis=0)
mcmc_tau_err = np.std(chain_tau,axis=0)

print(f"MCMC best-fit parameters are (with prior):\n\
H0 = {mcmc_tau[0]} +/- {mcmc_tau_err[0]}\n\
 $\Omega_b h^2$  = {mcmc_tau[1]} +/- {mcmc_tau_err[1]}\n\
 $\Omega_c h^2$  = {mcmc_tau[2]} +/- {mcmc_tau_err[2]}\n\
 $\sigma_8$  = {mcmc_tau[3]} +/- {mcmc_tau_err[3]}\n\
 $A_s$  = {mcmc_tau[4]} +/- {mcmc_tau_err[4]}\n\
 $n_s$  = {mcmc_tau[5]} +/- {mcmc_tau_err[5]}")
```

MCMC best-fit parameters are (with prior):
 $H_0 = 69.00637449580441 \pm 0.0008379169743202922$

```

 $\Omega_{bh}^2 = 0.02199178988600887 \pm 8.207588479127709e-08$ 
 $\Omega_{ch}^2 = 0.11999373636198926 \pm 3.5532107683995074e-08$ 
      =  $0.06873358618246474 \pm 0.0023589815275466683$ 
As =  $4.804164673916872e-10 \pm 5.737589464281155e-12$ 
ns =  $2.024734559166282 \pm 0.004610957982233009$ 

```

```

[720]: tau_write = np.column_stack([chain2, chisq2])
tau_file = 'planck_chain_tauprior.txt'
np.savetxt(tau_file, tau_write, fmt=['%f', '%f', '%f', '%f', '%f', '%f', '%f'])

```

4.0.2 Compare these with the values from problem 3:

H0 isn't statistically different, but the other values are. They shouldn't be, since the importance sampling should get us closer to the actual values by adding a new constraint. This is likely because my parameters from problem 3 aren't correct.

```

[725]: print(f"MCMC best-fit parameters are (without prior):\n\
H0 = {mcmcp[0]+mcmc_mean[0]} +/- {mcmc_sig[0]}\n\
 $\Omega_{bh}^2$  = {mcmcp[1]+mcmc_mean[1]} +/- {mcmc_sig[1]}\n\
 $\Omega_{ch}^2$  = {mcmcp[2]+mcmc_mean[2]} +/- {mcmc_sig[2]}\n\
      = {mcmcp[3]+mcmc_mean[3]} +/- {mcmc_sig[3]}\n\
As = {mcmcp[4]+mcmc_mean[4]} +/- {mcmc_sig[4]}\n\
ns = {mcmcp[5]+mcmc_mean[5]} +/- {mcmc_sig[5]}")

```

```

MCMC best-fit parameters are (without prior):
H0 = 69.07684470000497 +/- 0.025277370540507448
 $\Omega_{bh}^2$  = 0.09928366999793603 +/- 0.029844444964762996
 $\Omega_{ch}^2$  = 0.195503184223957 +/- 0.028882346598117527
      = -0.2944301969877605 +/- 0.02583788798217511
As = 0.03639623060197444 +/- 0.0077706031787391616
ns = 2.0632277695894143 +/- 0.011261069429483987

```

```
[ ]:
```