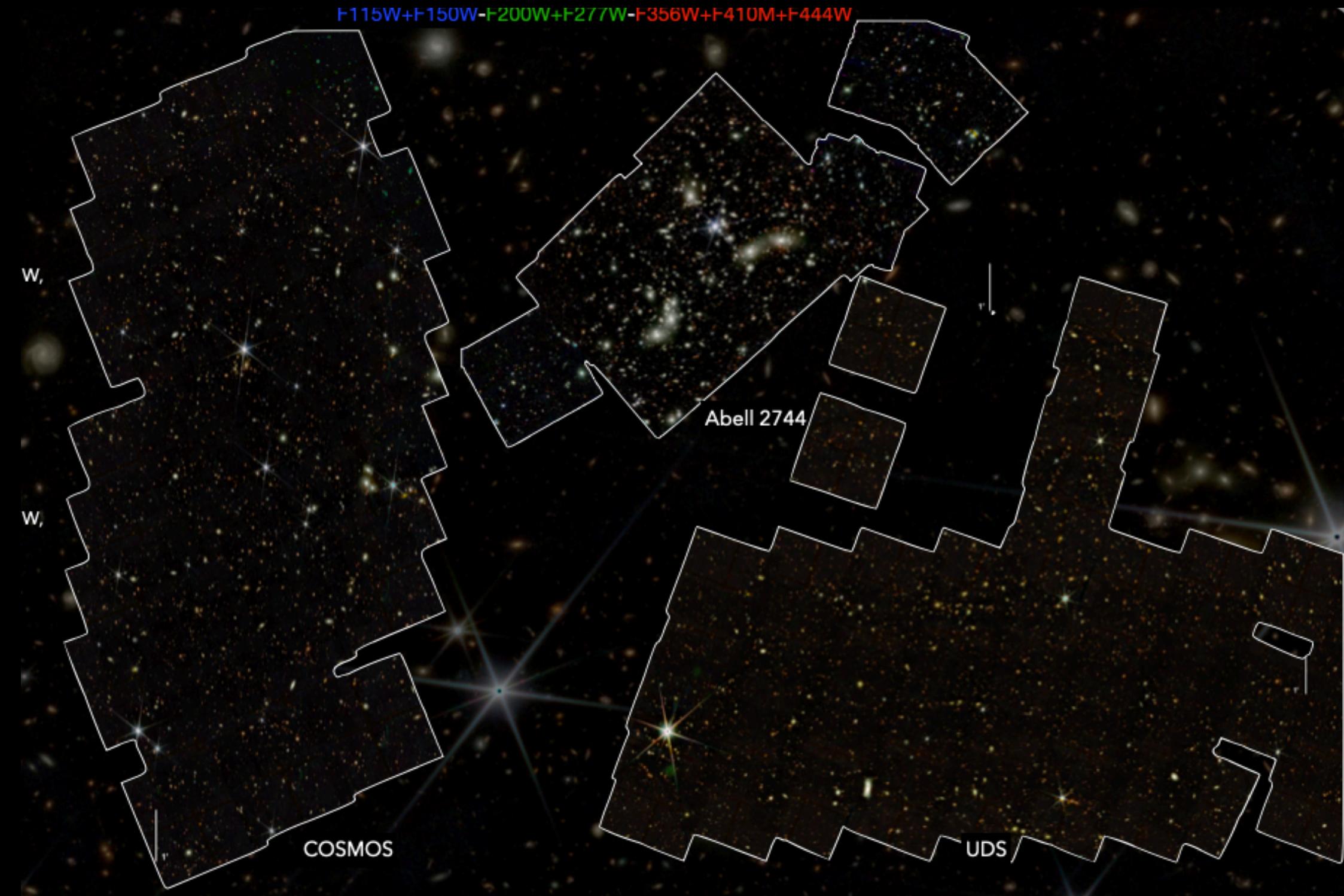


Low-Mass Quiescent Galaxy Sizes in JWST PRIMER and UNCOVER

Revealing Two Distinct Quiescent Galaxy Populations at Cosmic Noon

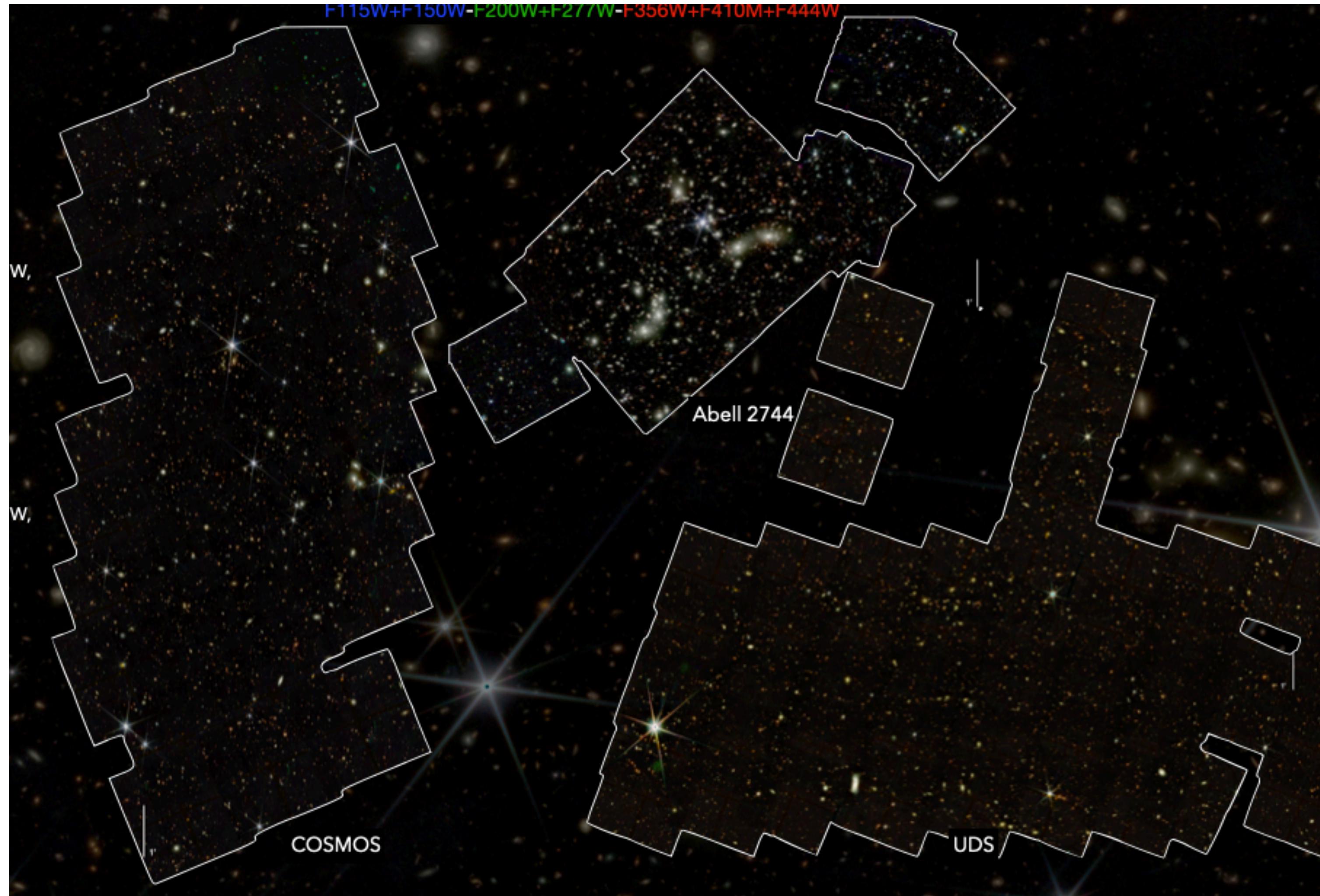


Sam Cutler, Kate Whitaker (UMass Amherst)
The PRIMER and UNCOVER Teams
@secutler



Low-Mass Quiescent Galaxy Sizes with JWST

The UNCOVER and PRIMER Treasury Surveys

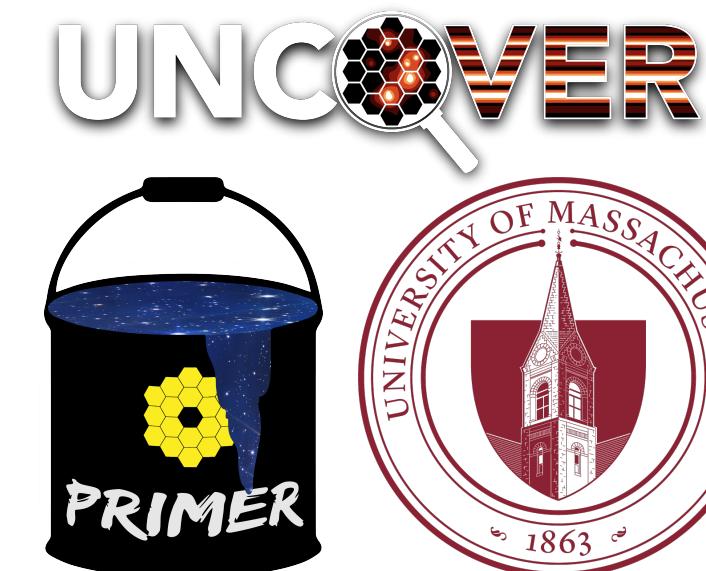


PRIMER (JWST-GO-1837):

- Covers two legacy fields (COSMOS and UDS)
- Homogeneous depth (~ 28 ABmag in F200W)
- 378 sq. arcmin. total
- Observations in F090W, F115W, F150W, F200W, F277W, F356W, F410M, and F444W
- Archival HST data

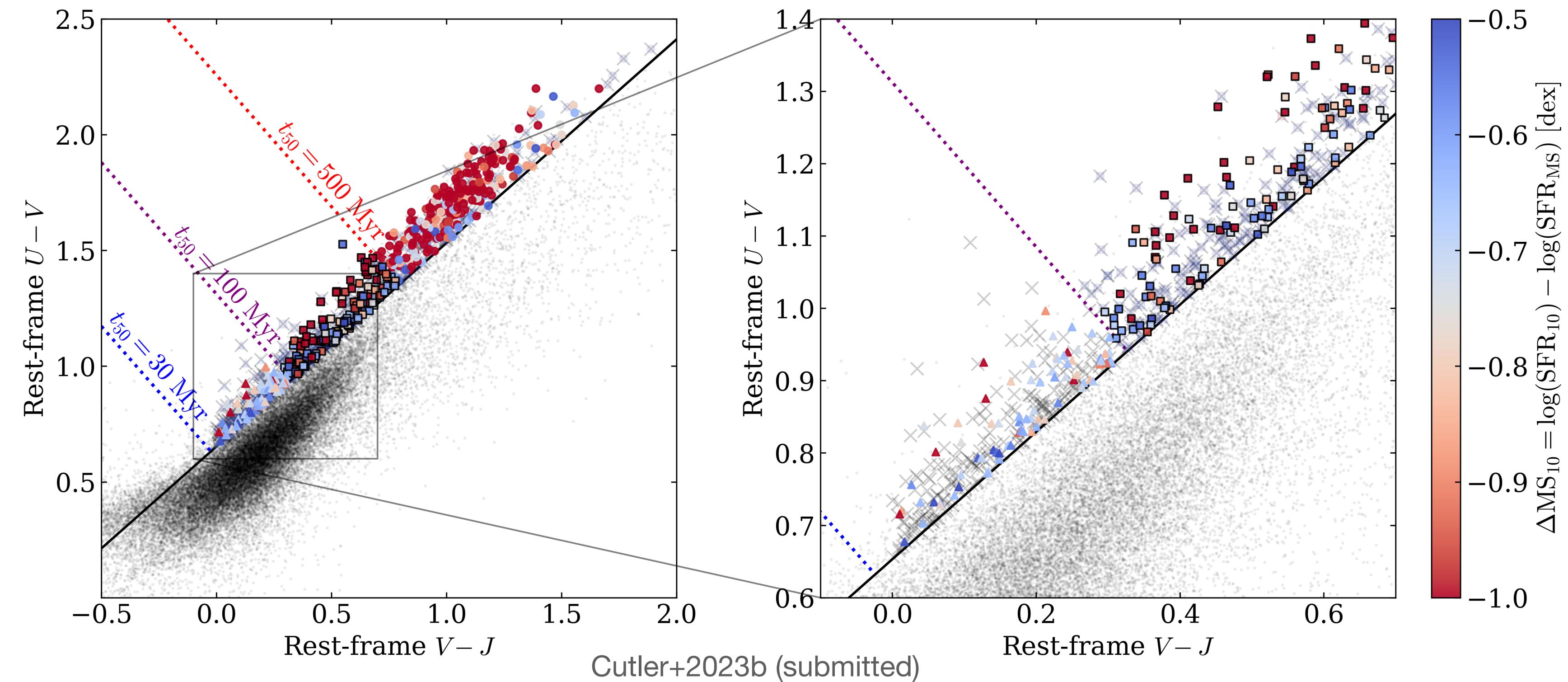
UNCOVER (JWST-GO-2561):

- Targets the Abell-2744 lensing cluster
- Deepest-to-date publicly available survey (> 29 ABmag in F200W without corrections for lensing)
- 45 sq. arcmin. total
- Observations in F090W, F115W, F150W, F200W, F277W, F356W, F410M, and F444W
- Archival HST data



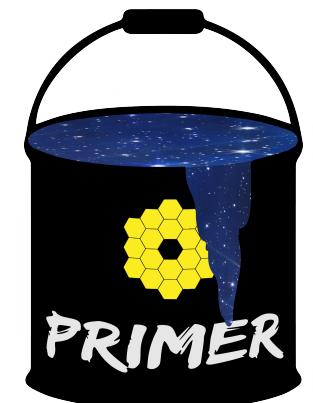
Low-Mass Quiescent Galaxy Sizes with JWST

Sample selection: extended UVJ + sSFR criteria



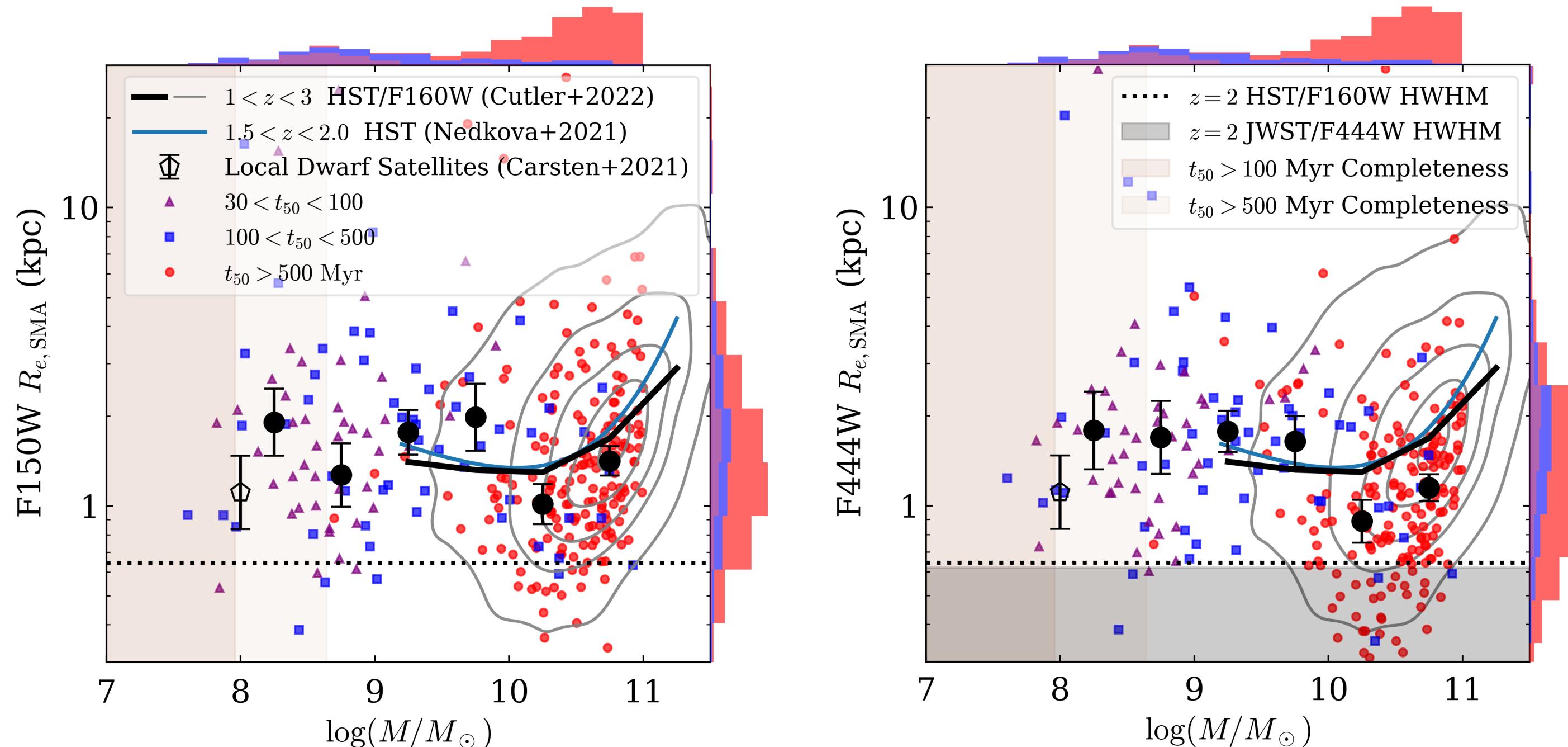
Cutler+2023b (submitted)

UNCOVER



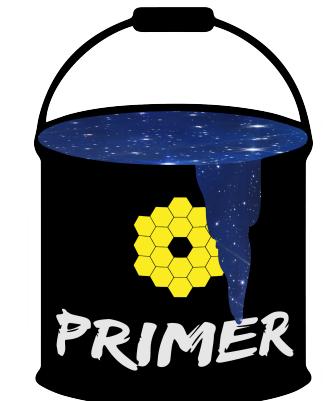
Low-Mass Quiescent Galaxy Sizes with JWST

The size-mass relation at cosmic noon



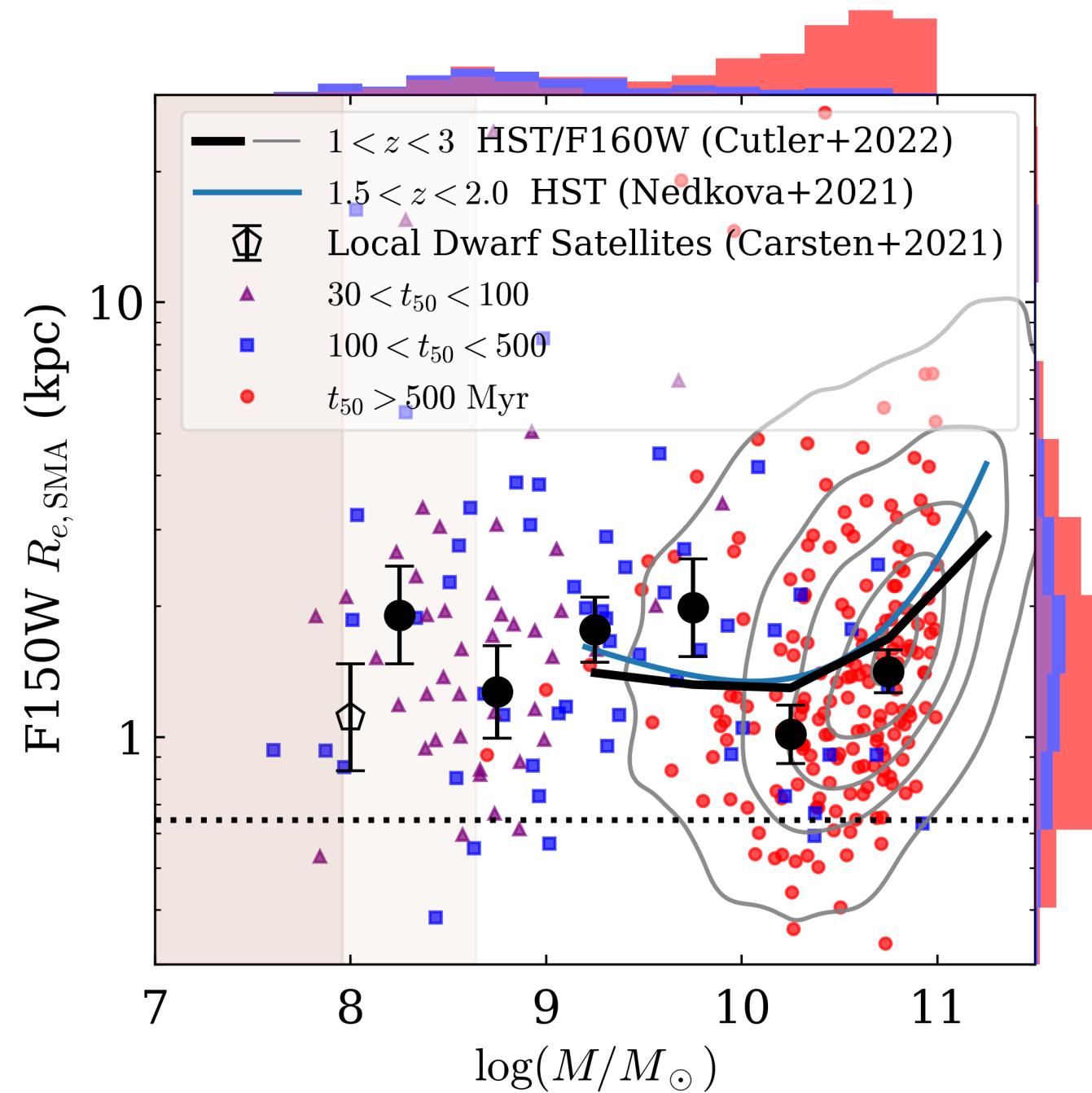
Cutler+2023b (submitted)

UNCOVER

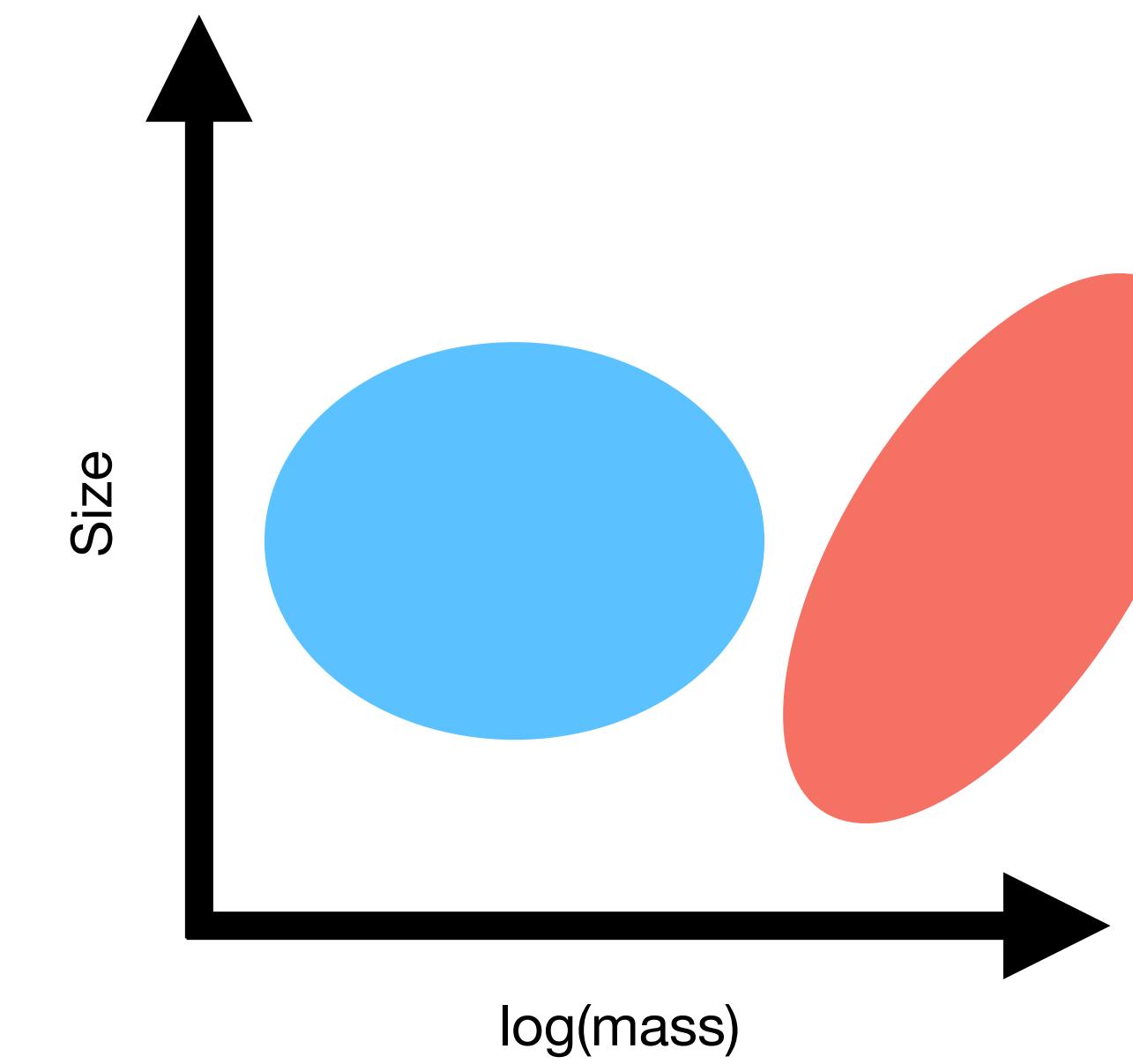
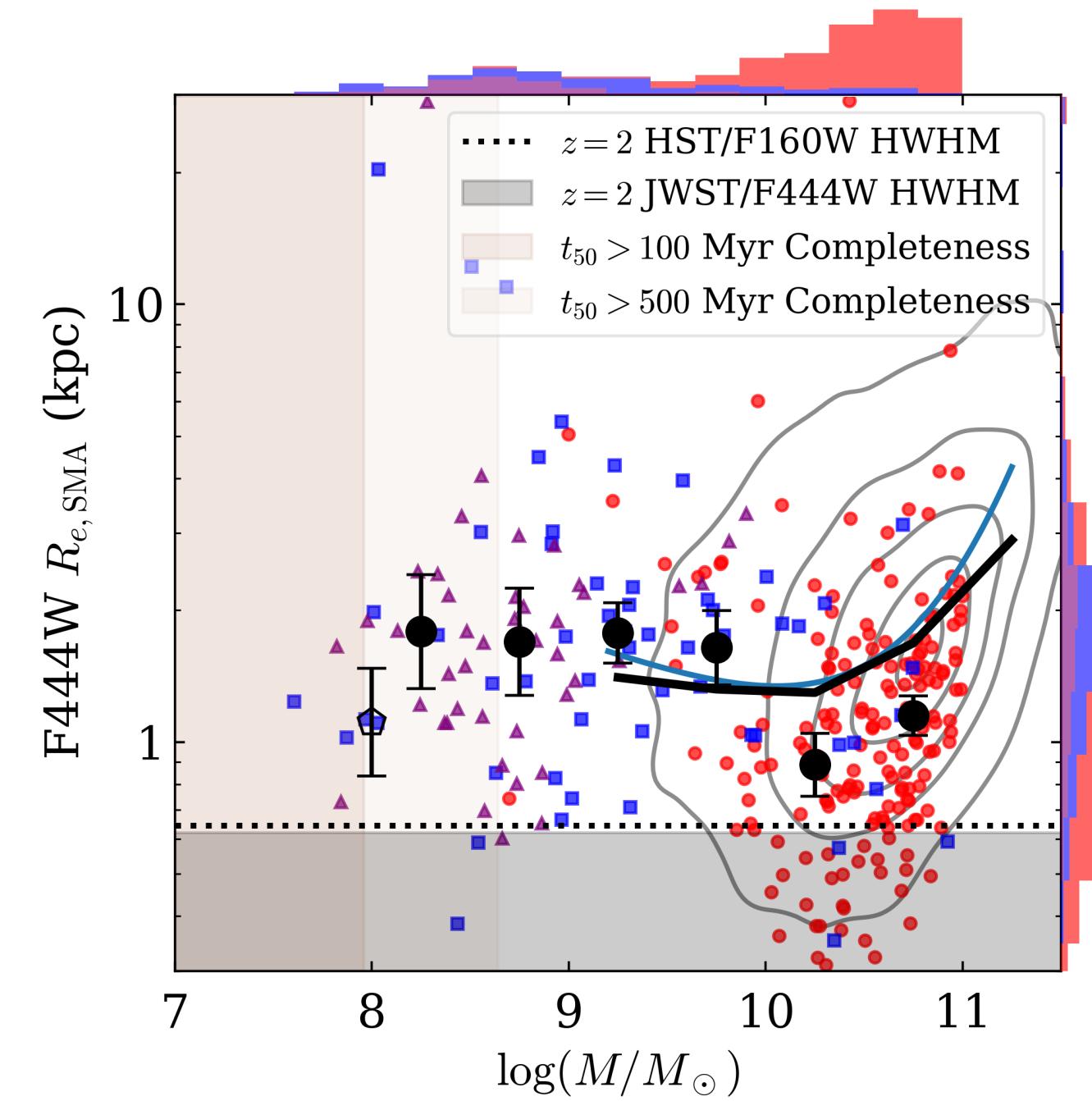


Low-Mass Quiescent Galaxy Sizes with JWST

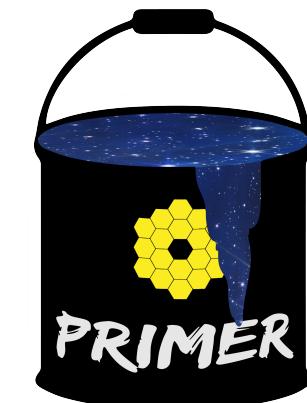
The size-mass relation at cosmic noon



Cutler+2023b (submitted)

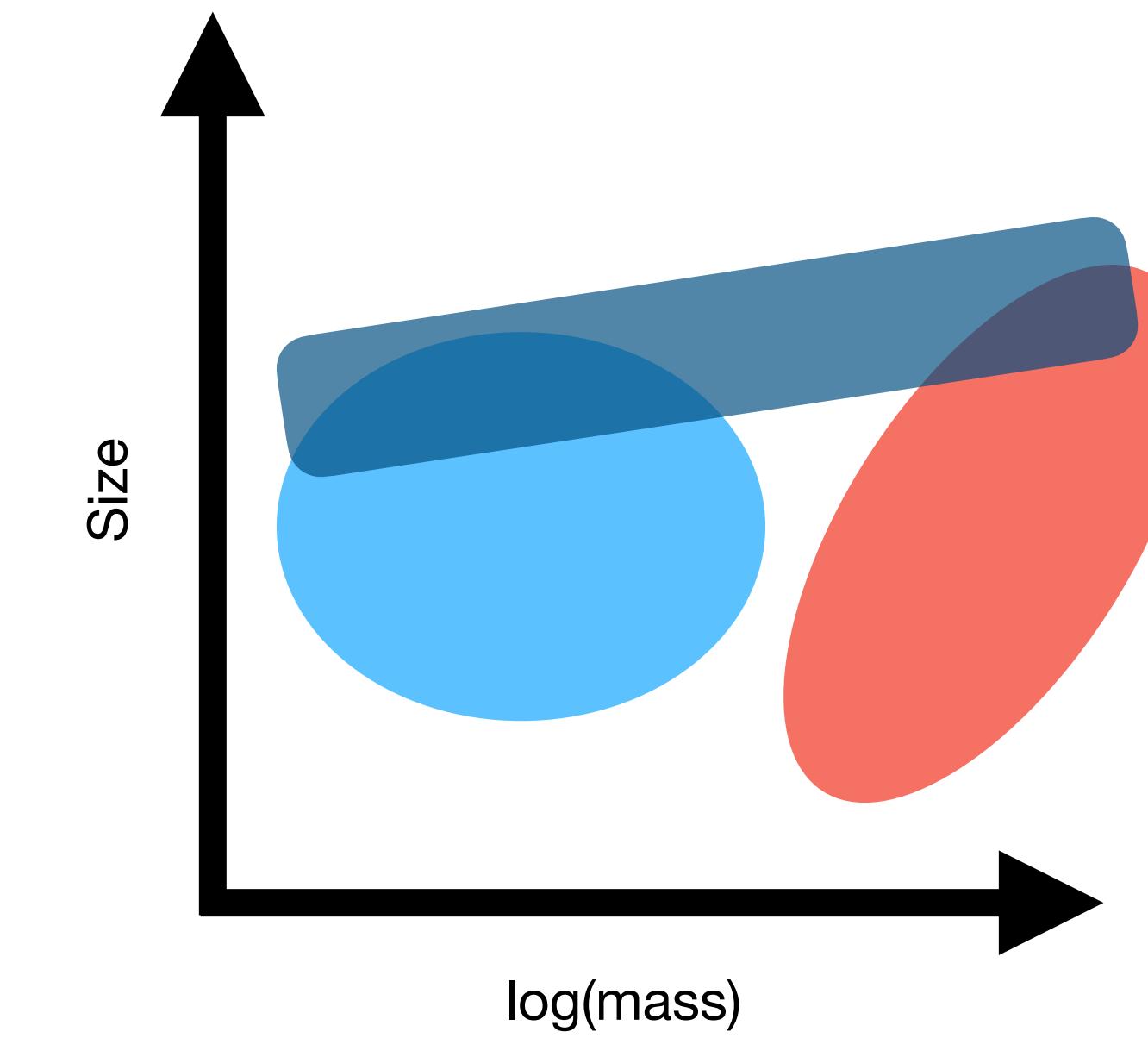
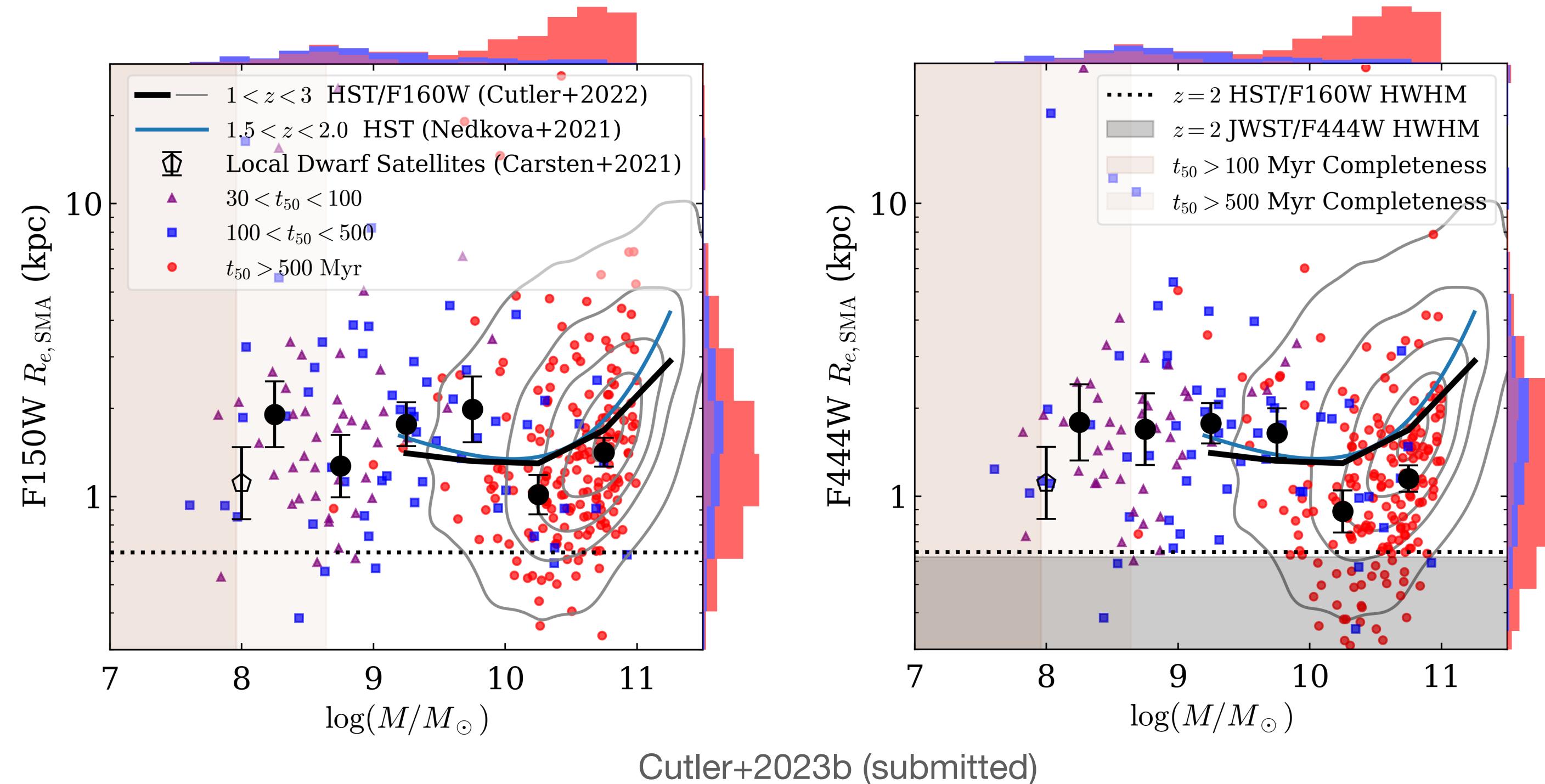


Low-mass quiescent galaxies are a distinct population

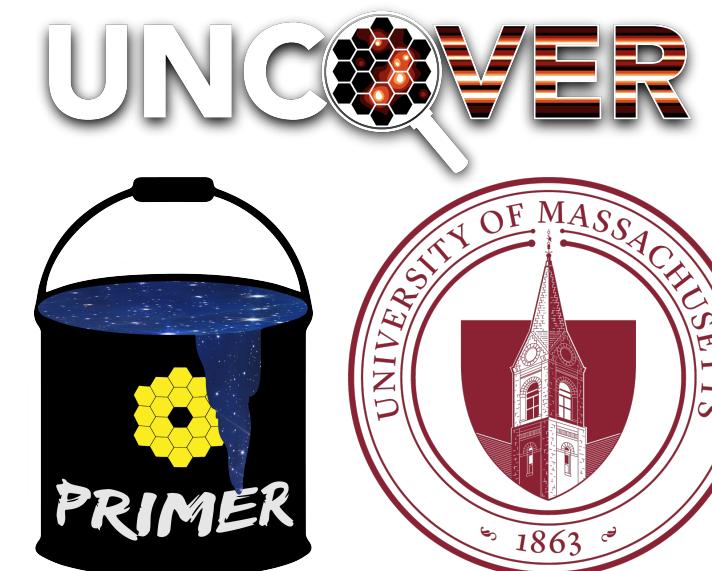


Low-Mass Quiescent Galaxy Sizes with JWST

The size-mass relation at cosmic noon

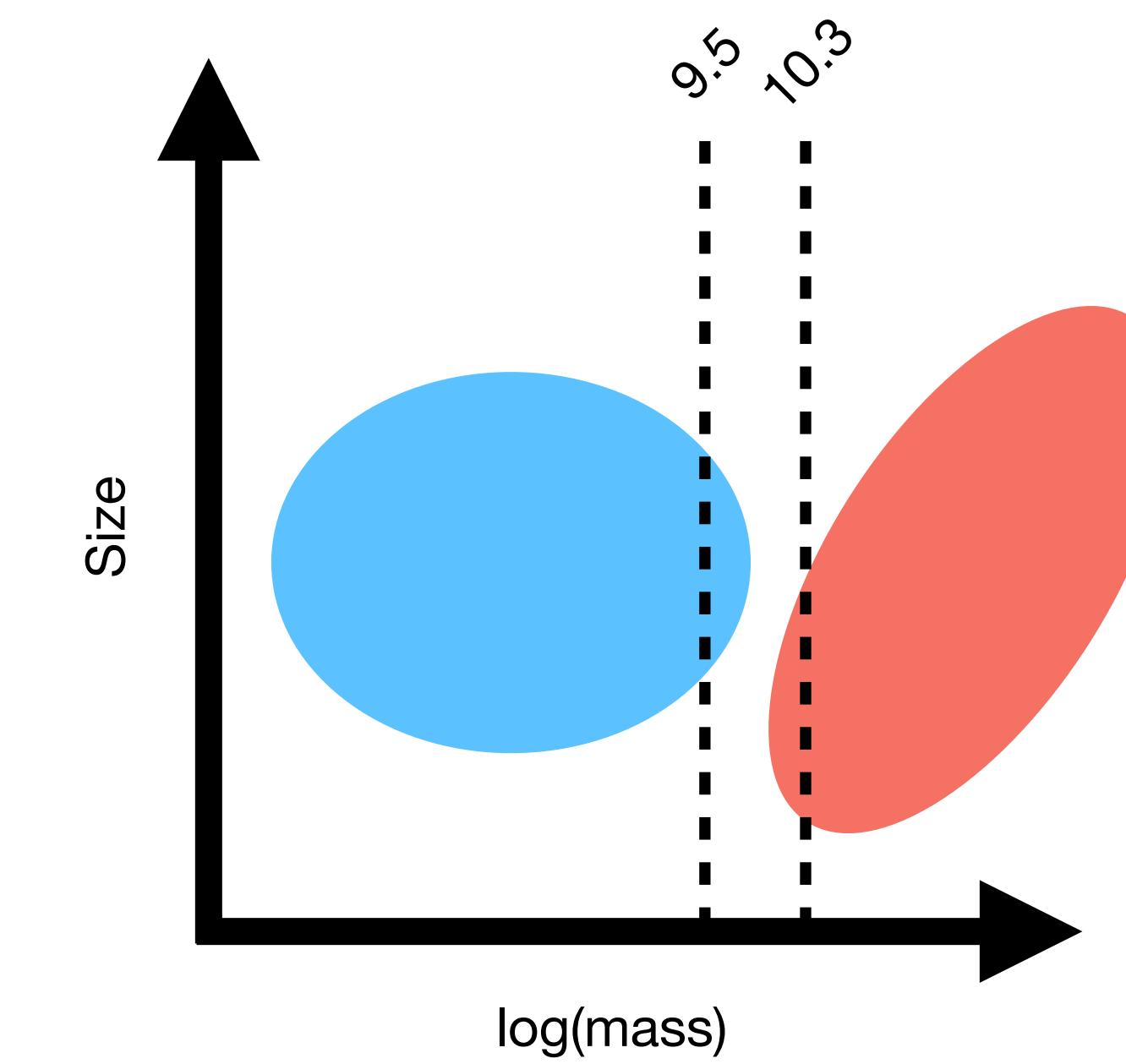
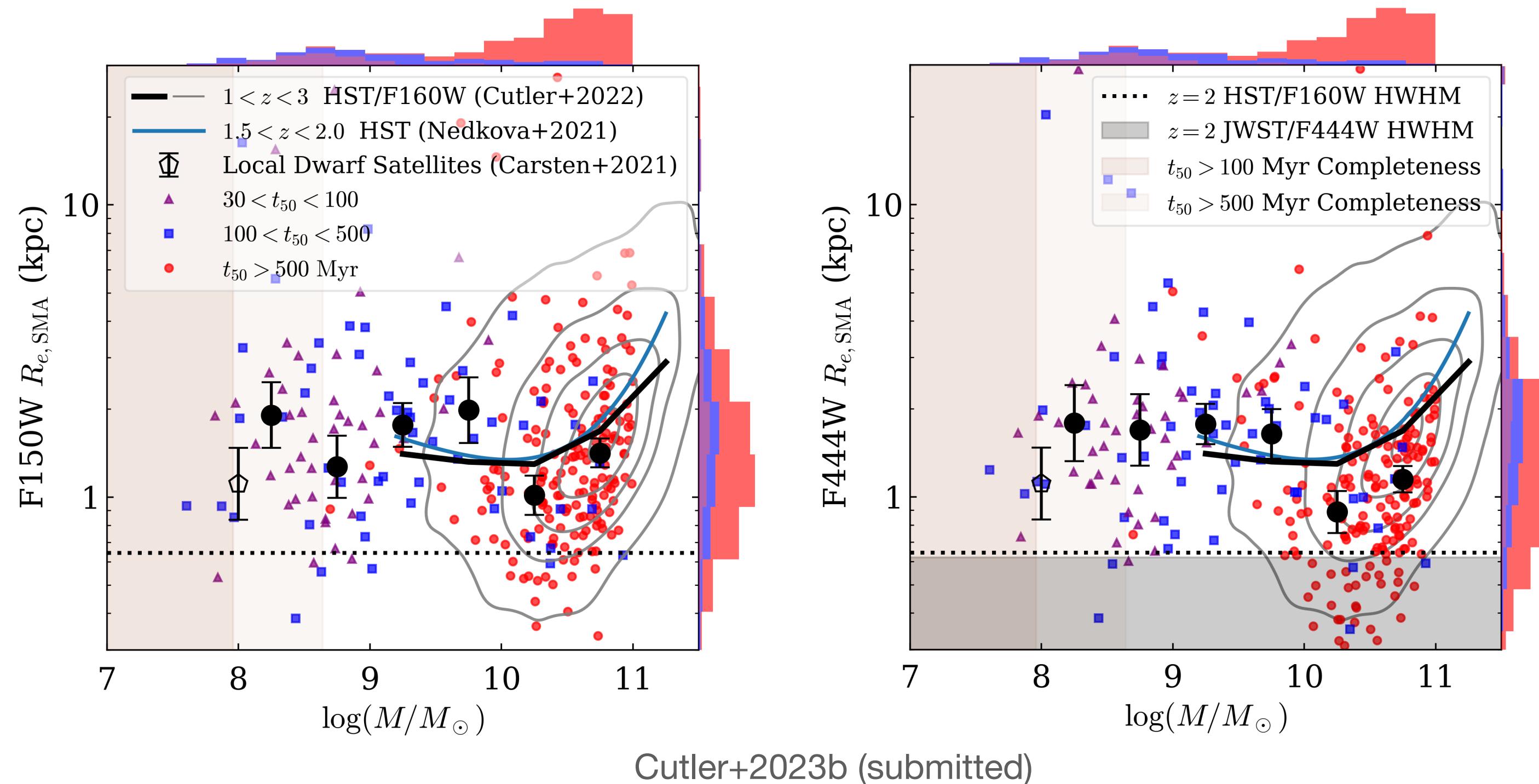


Low-mass quiescent galaxies are a distinct population
similar to star forming galaxies

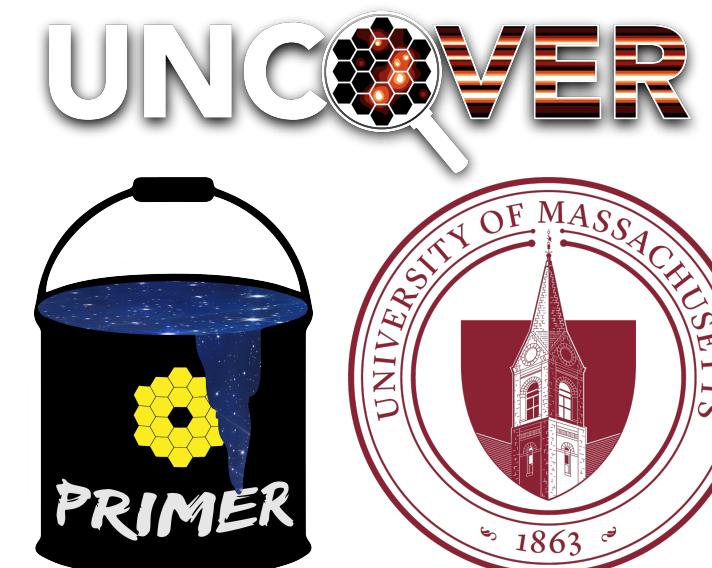


Low-Mass Quiescent Galaxy Sizes with JWST

The size-mass relation at cosmic noon

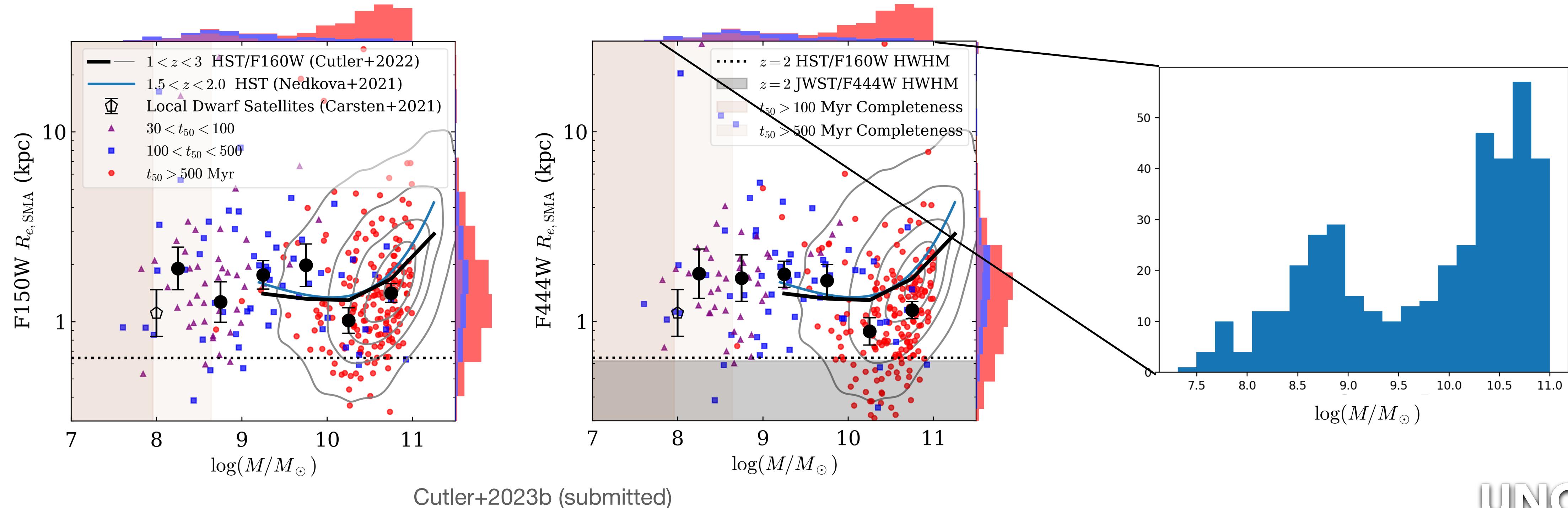


Low-mass quiescent galaxies are a distinct population
similar to star forming galaxies

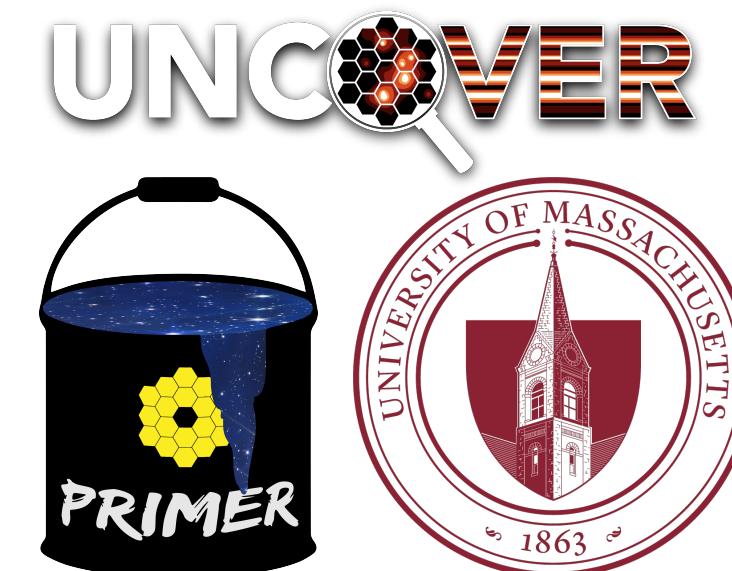


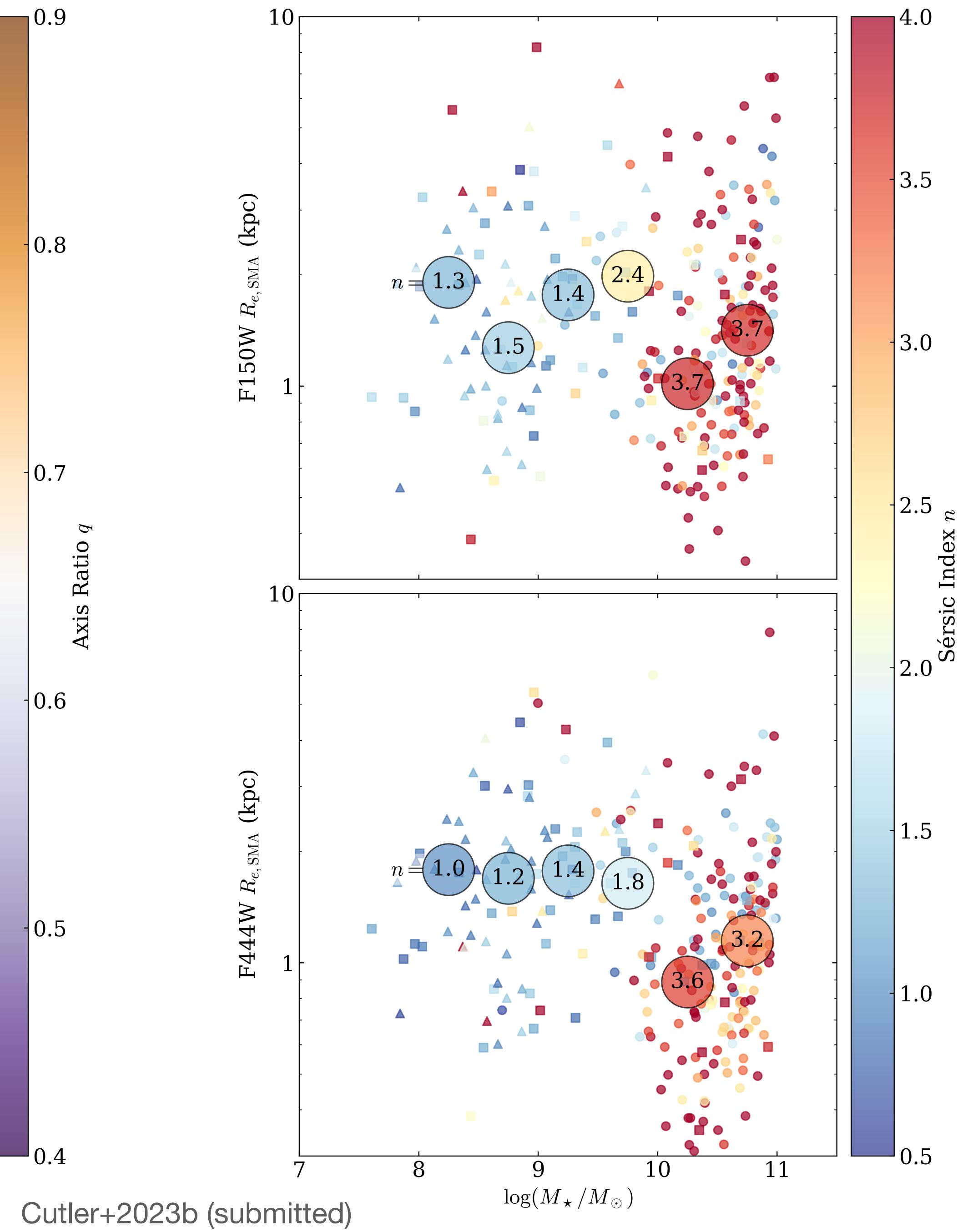
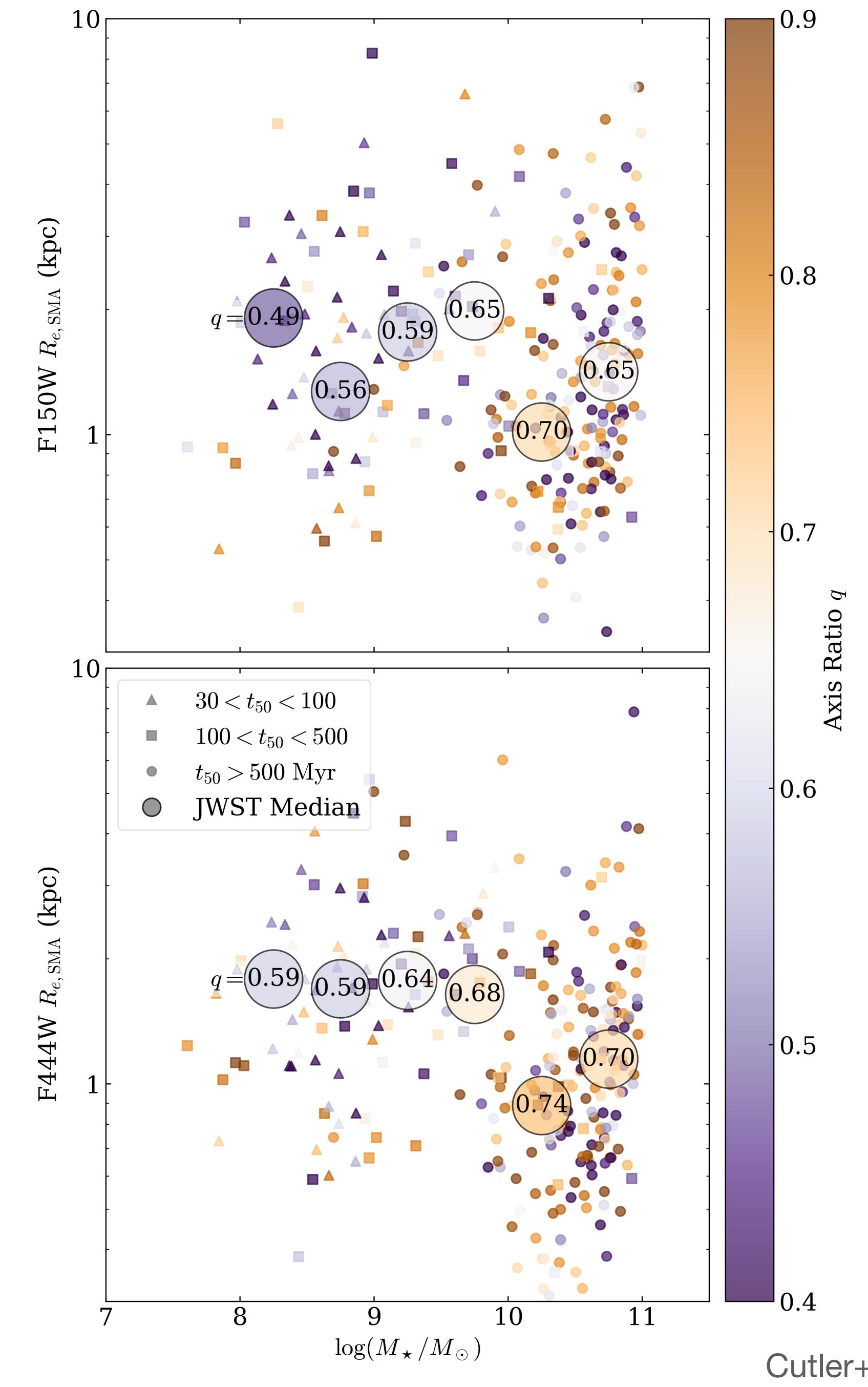
Low-Mass Quiescent Galaxy Sizes with JWST

The size-mass relation at cosmic noon



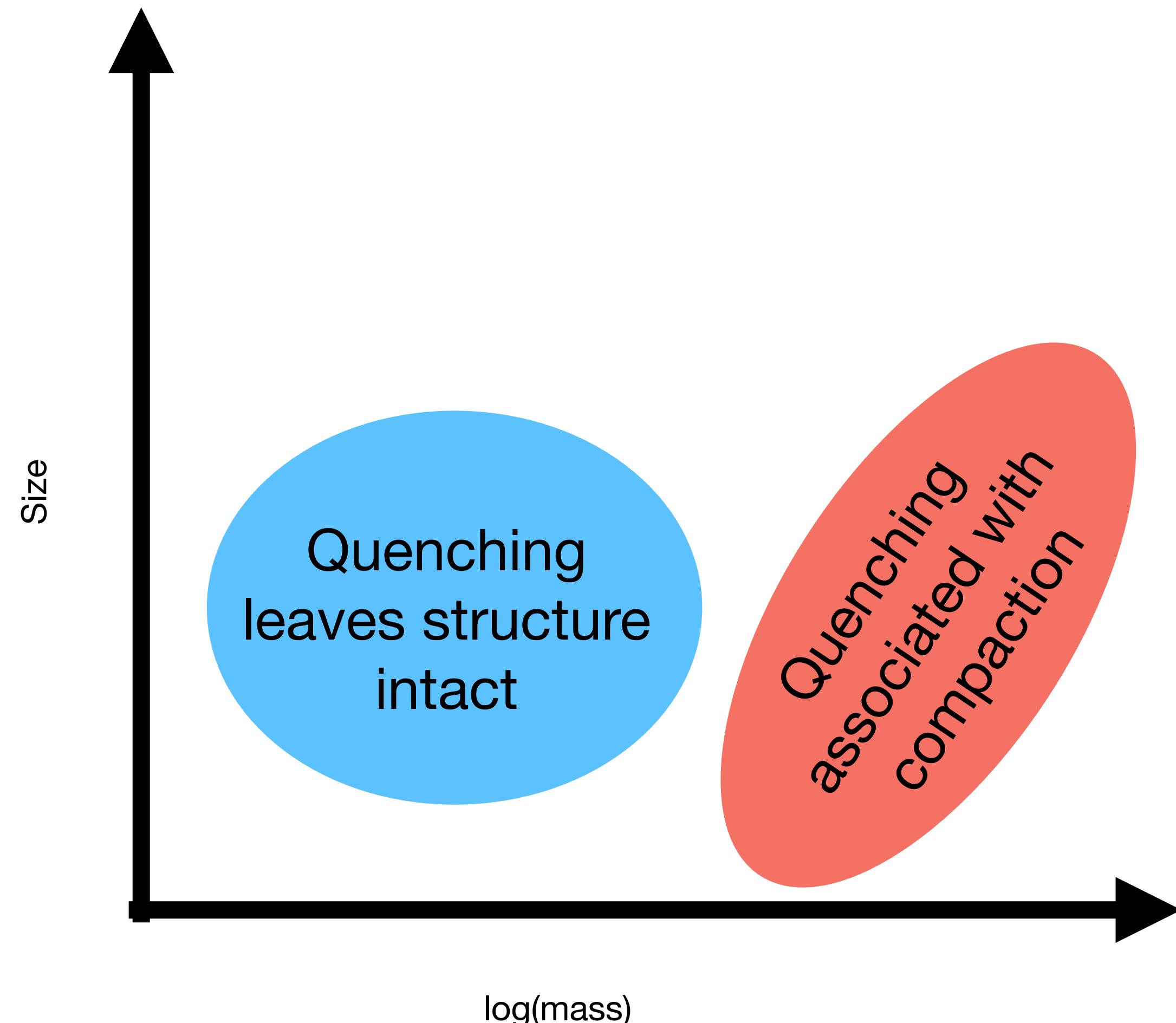
Also apparent in stellar mass distribution



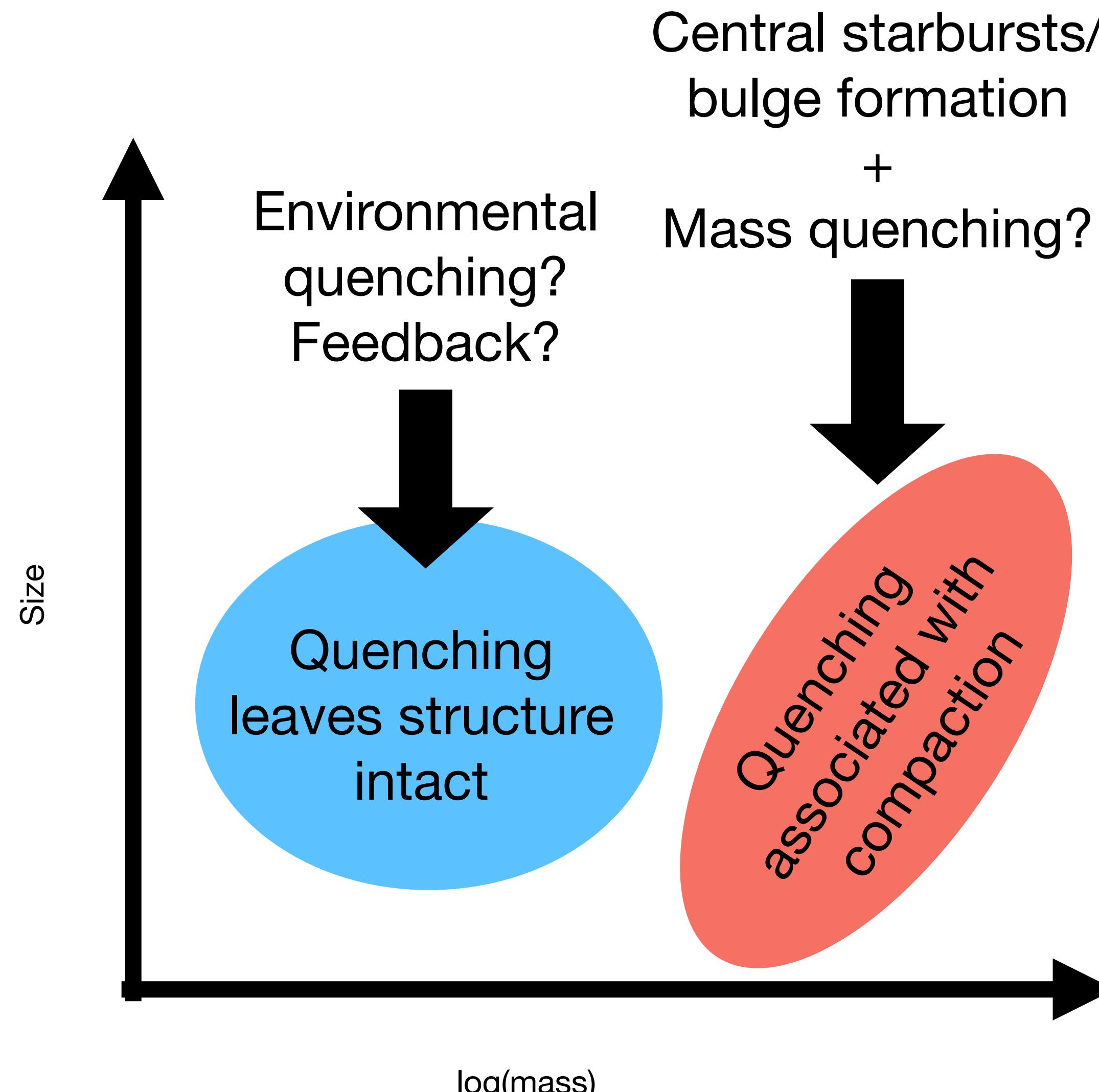


Cutler+2023b (submitted)

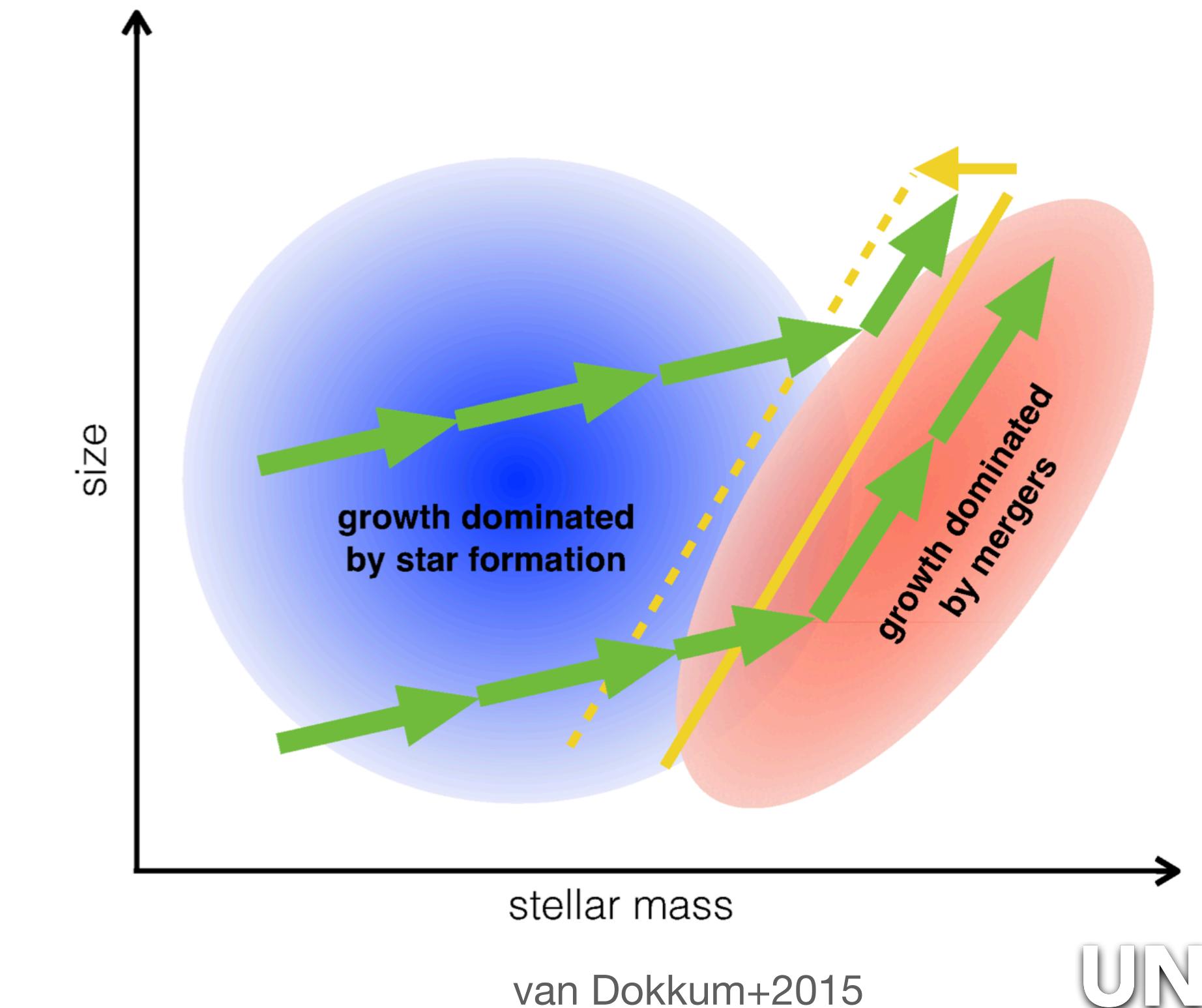
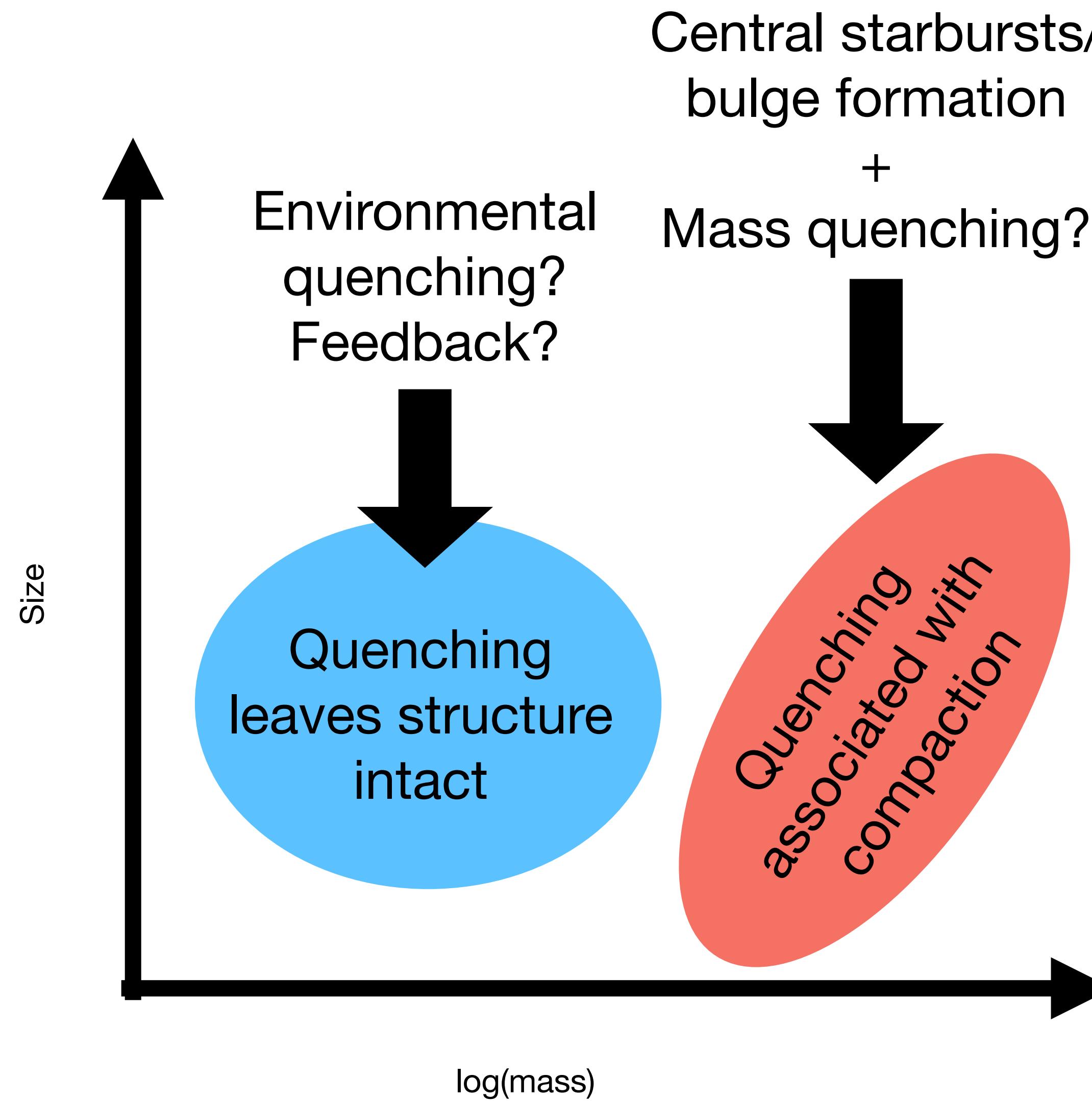
Galaxy Evolution at Low Mass



Galaxy Evolution at Low Mass



Galaxy Evolution at Low Mass

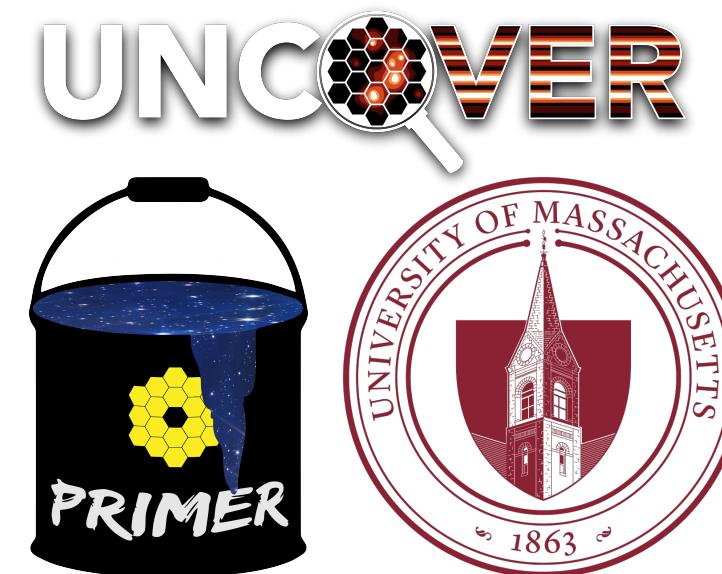
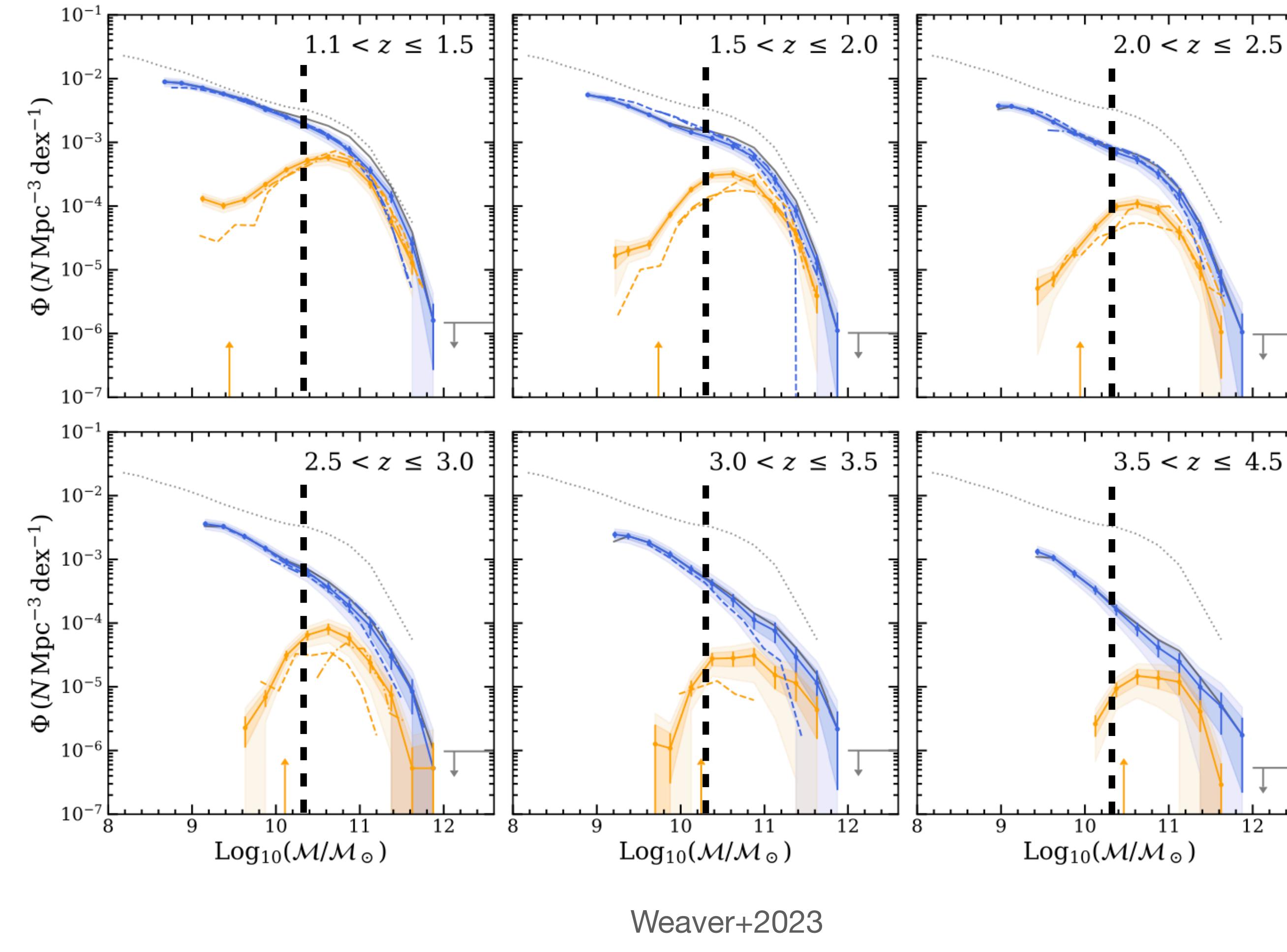
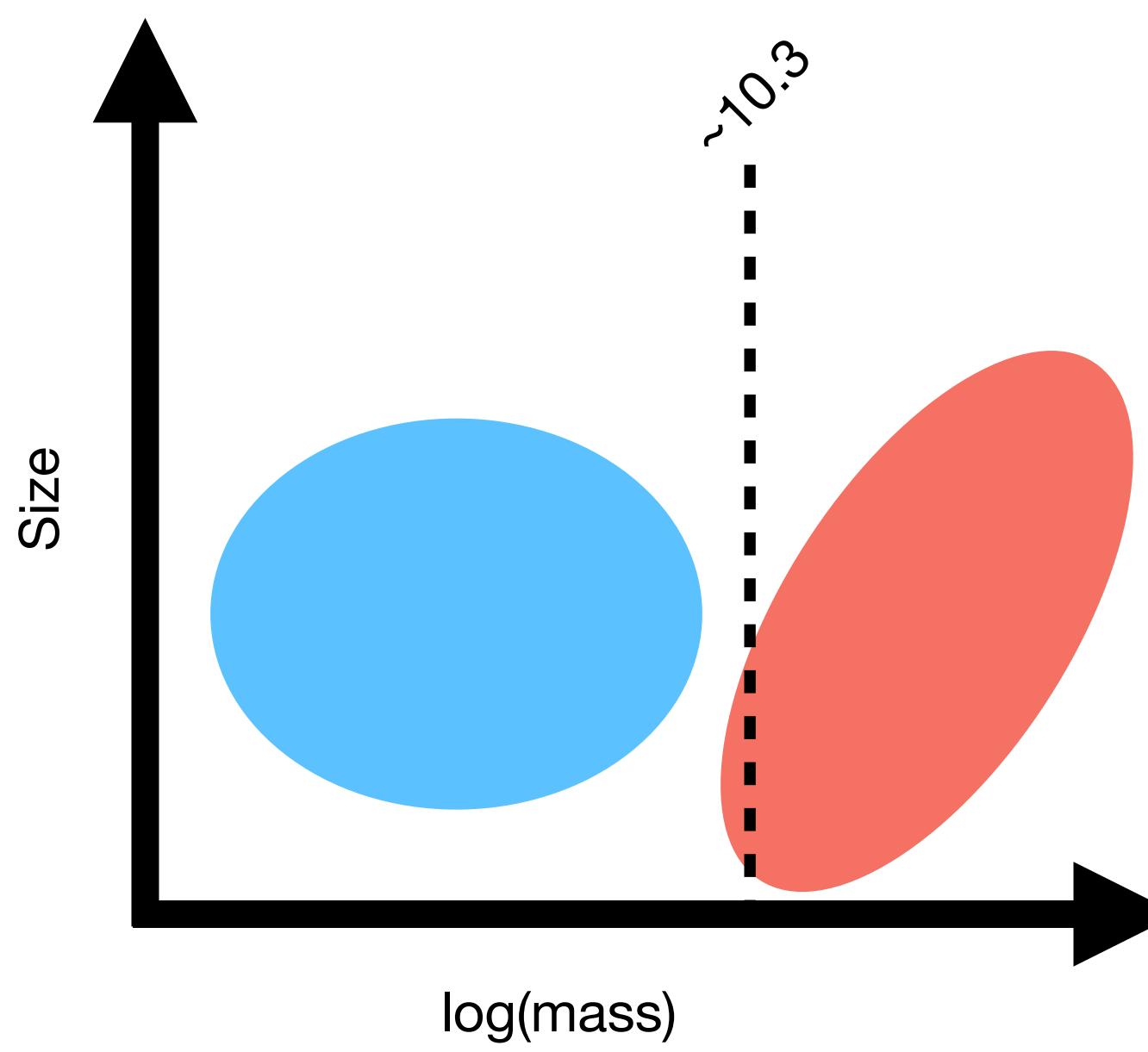


UNCOVER



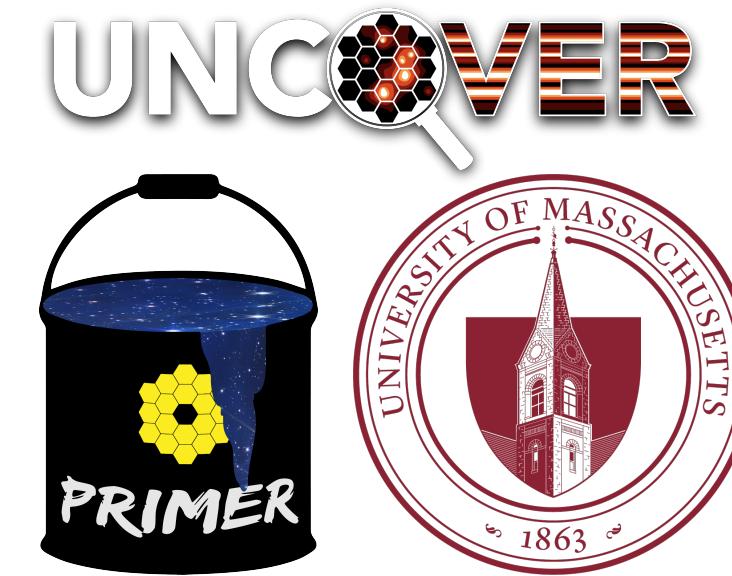
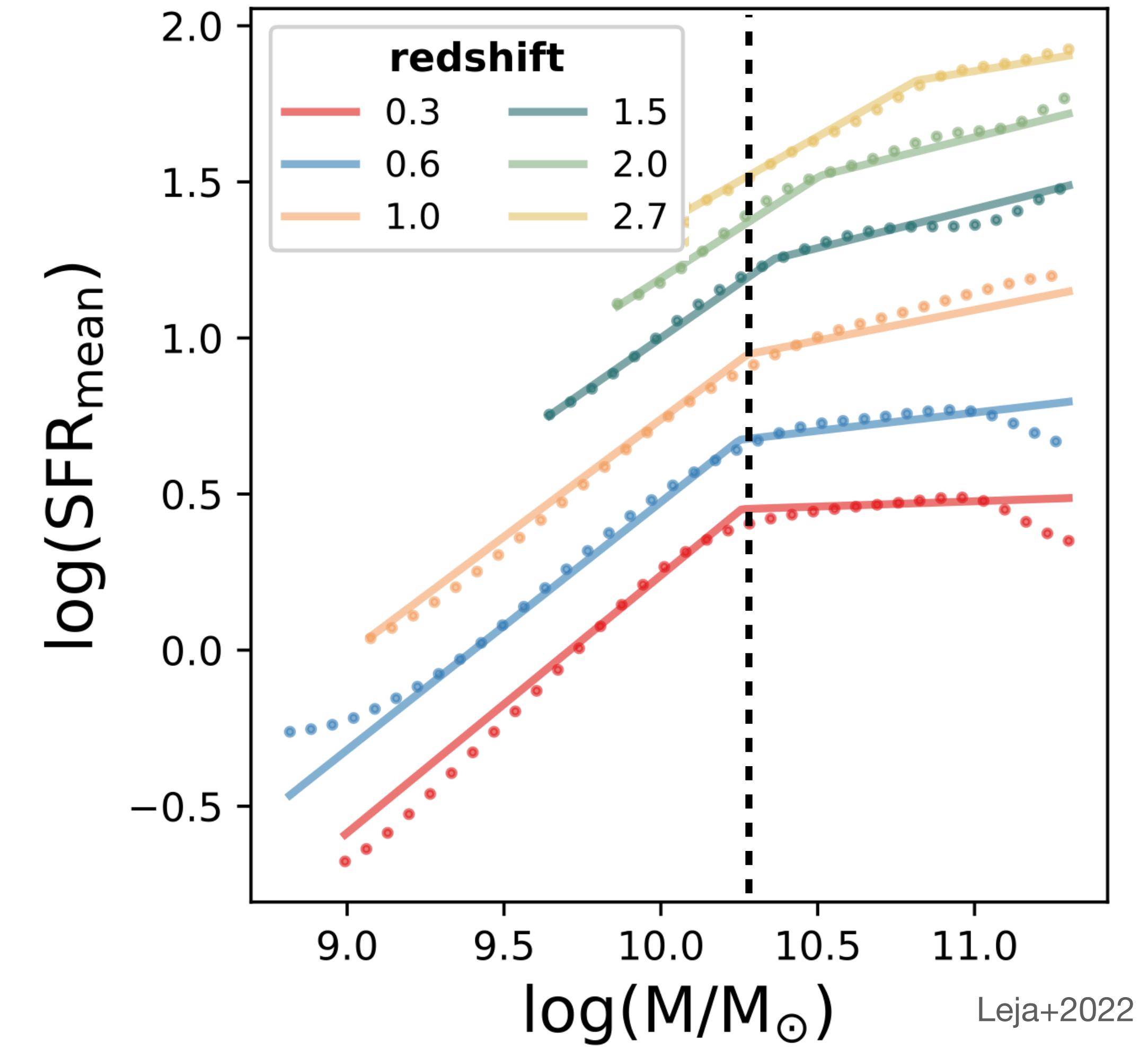
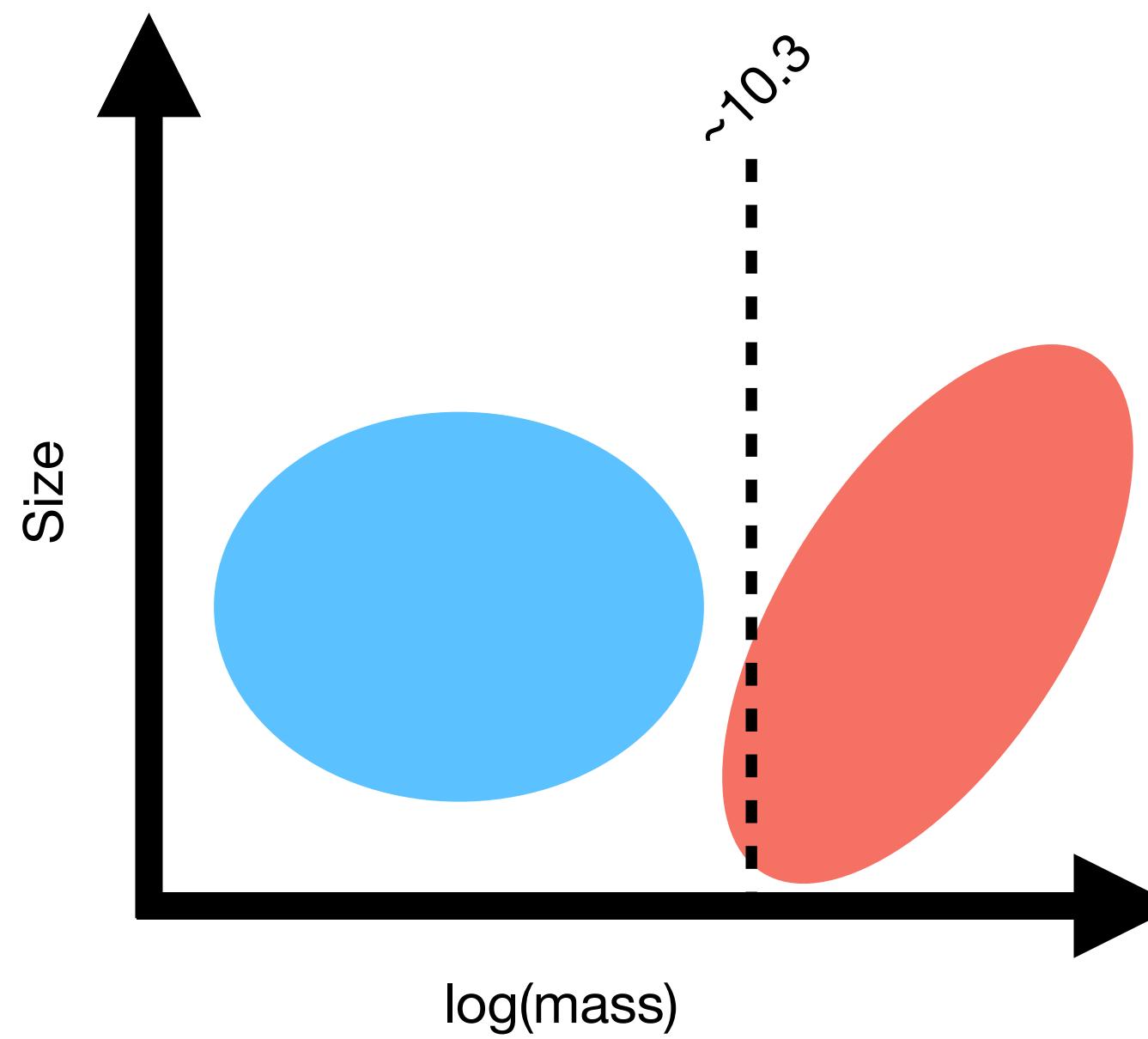
Galaxy Evolution at Low Mass

Corresponds with peak in quiescent galaxy mass function



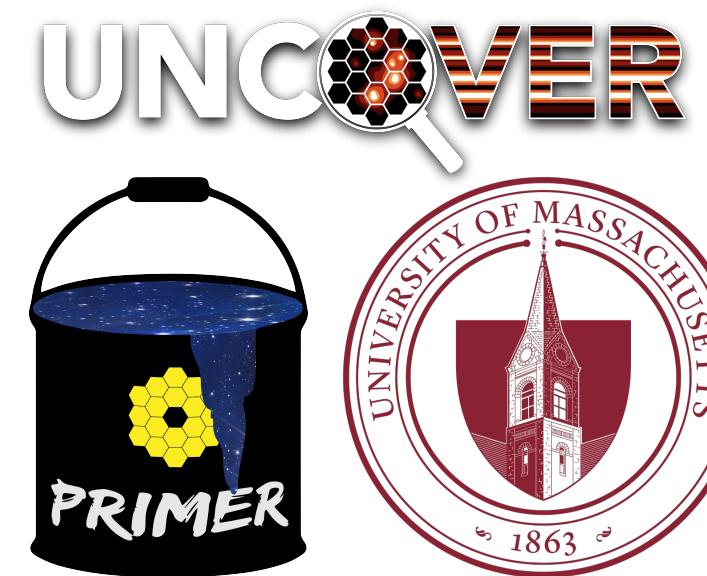
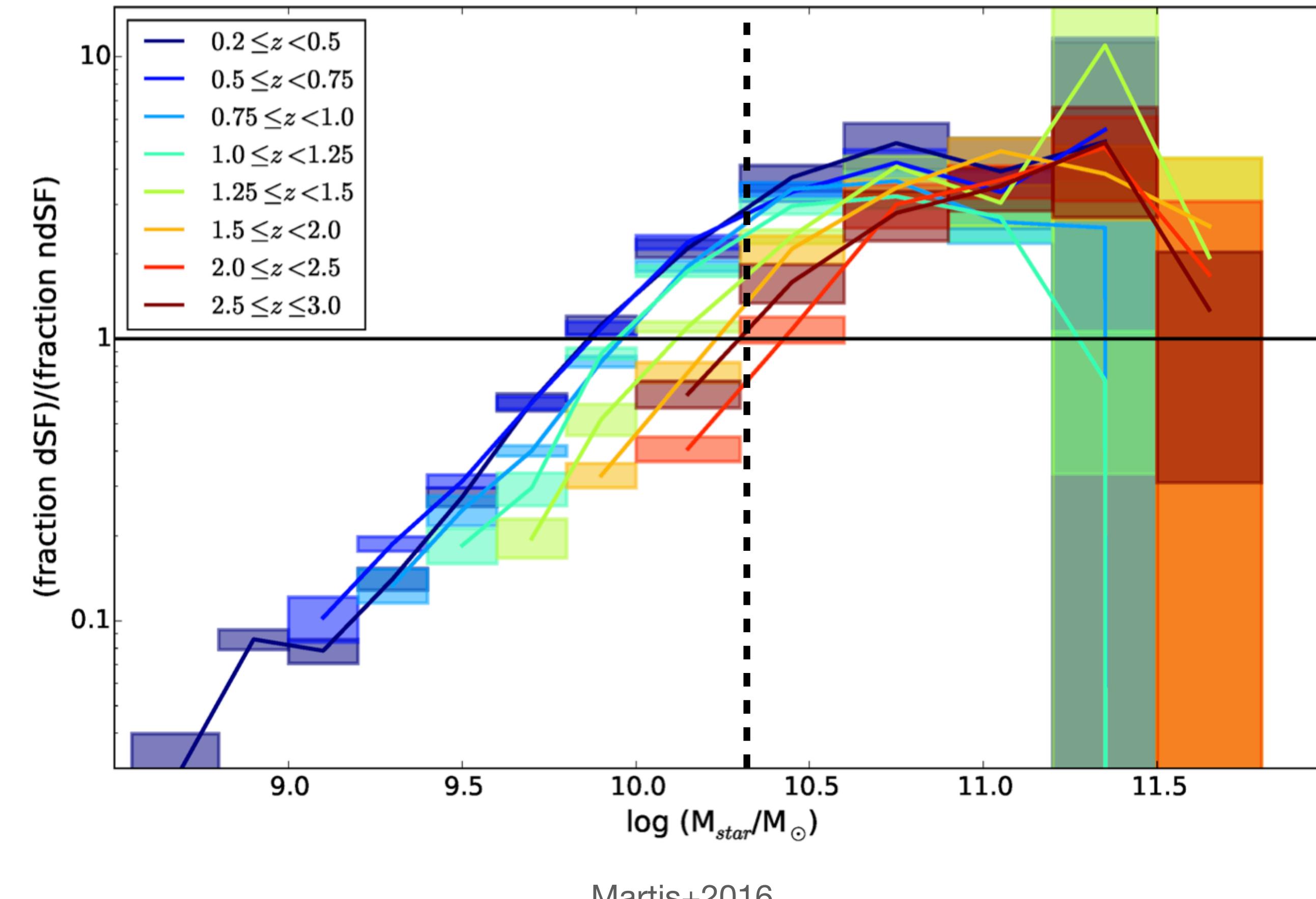
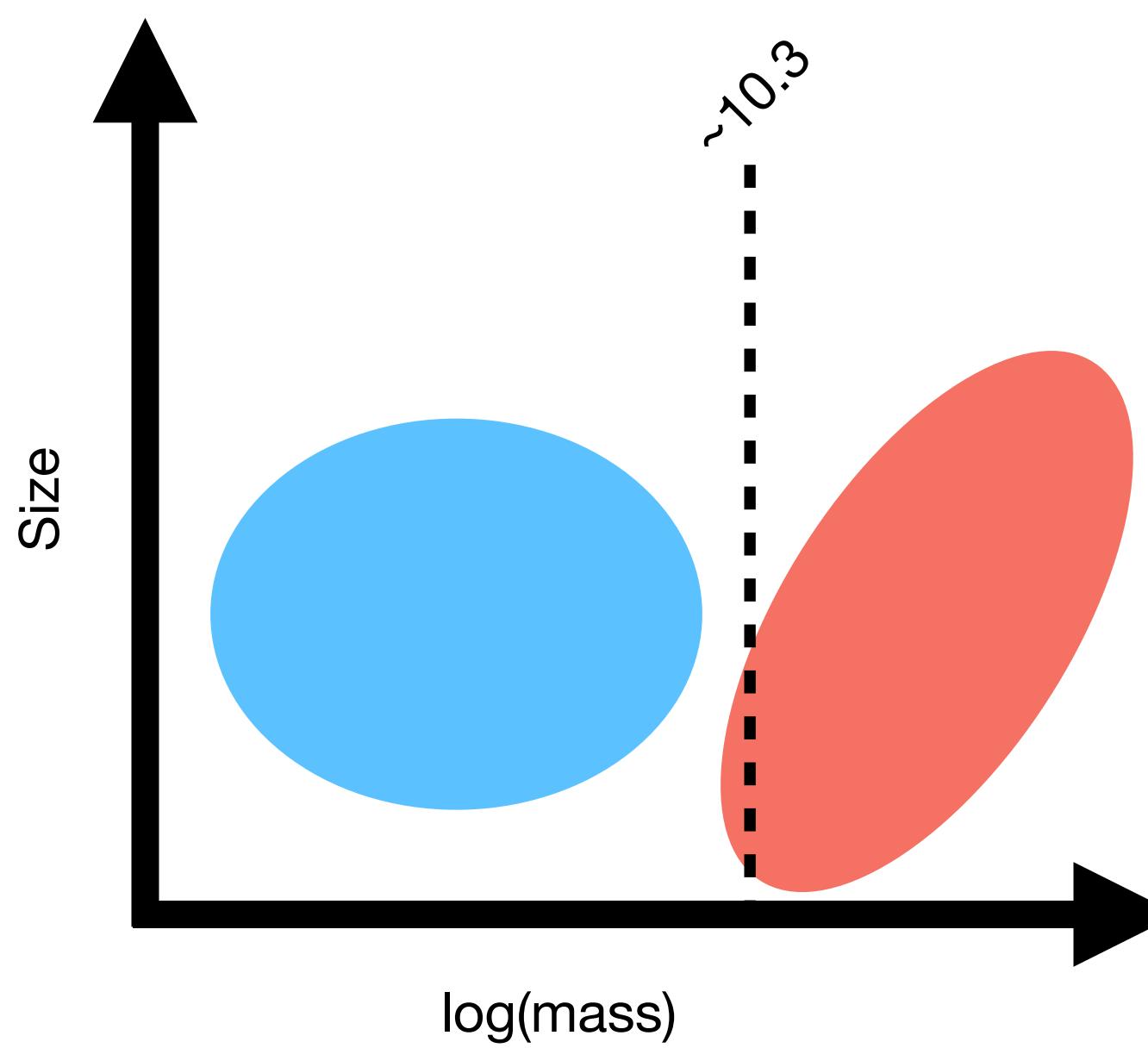
Galaxy Evolution at Low Mass

Corresponds with change in SFMS slope



Galaxy Evolution at Low Mass

Corresponds with transition to predominantly dusty SFGs



Quiescent galaxies at cosmic noon fall into two classes:

1. Young*, low-mass, and disk-like
2. Old, massive, and spheroidal

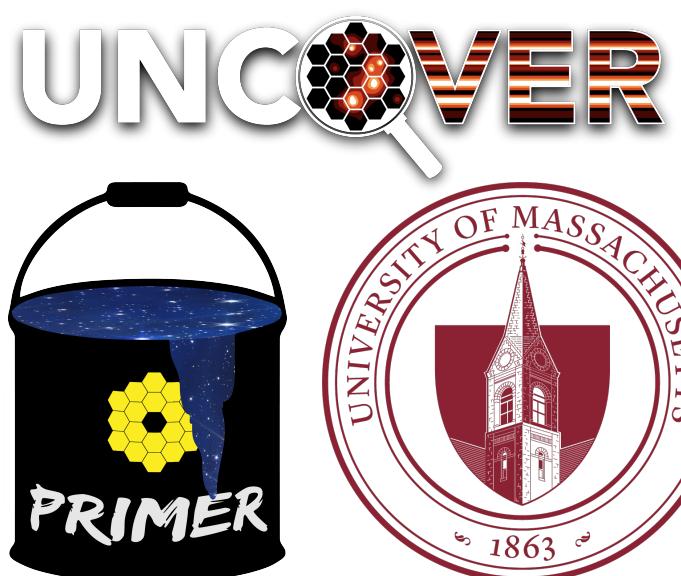
The separation between these populations occurs at $\log M \sim 10.3$

- This coincides with several significant transitions in galaxy evolution
- Galaxy evolution is dramatically different at low masses

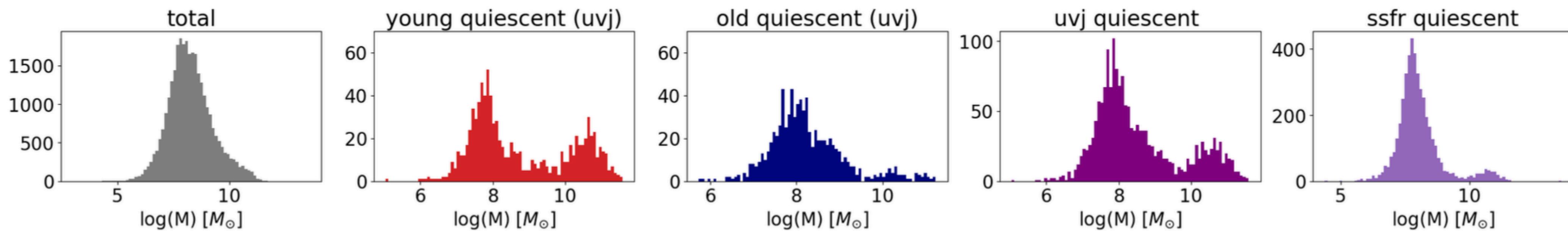
Cutler et al. 2023b
(Submitted,
[arXiv:2312.15012](https://arxiv.org/abs/2312.15012))



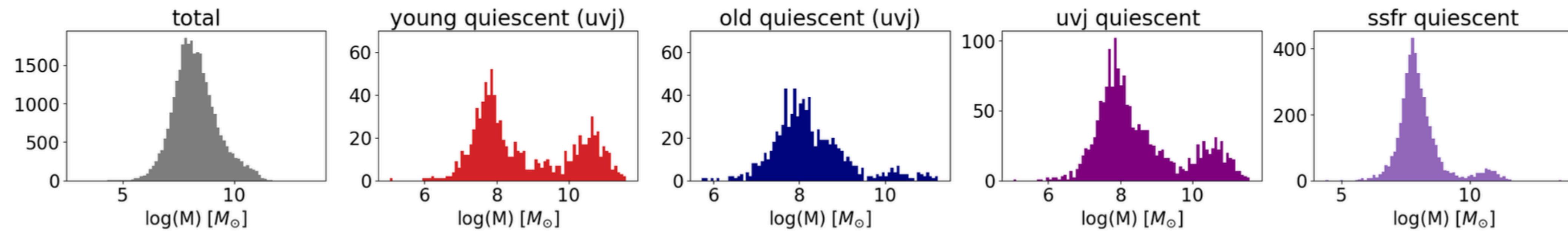
Contact me with any thoughts or suggestions!
secutler@umass.edu; [@secutler](https://twitter.com/secutler); [samecutler.github.io](https://github.com/samecutler)



PRIMER-COSMOS



PRIMER-UDS



UNCOVER

