ASTROMETRIC PLATE SOLVING

PHYS 605 Term Project Presentation

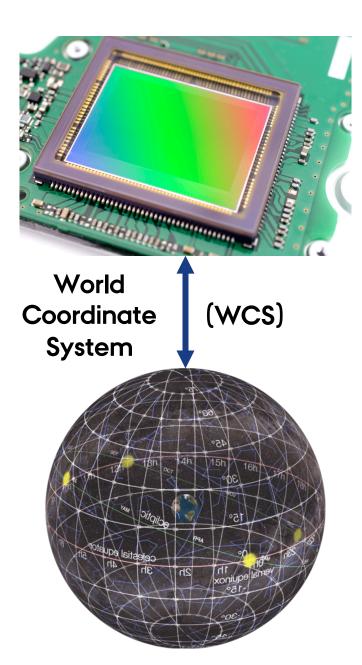




INTRODUCTION

Astrometric plate solving

- The process of obtaining WCS information from an image. [1]
- Extremely important technique in <u>astronomy</u> and Attitude Determination and Control Systems in <u>satellites</u> [2]







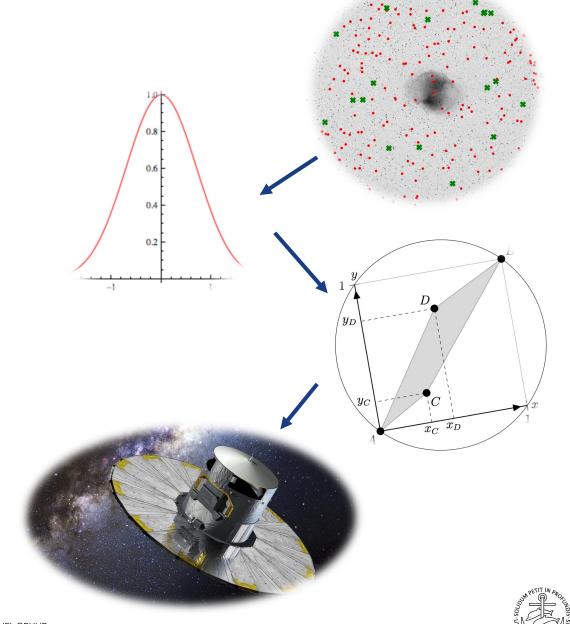




SAMUEL GRUND STUDENT

METHODS

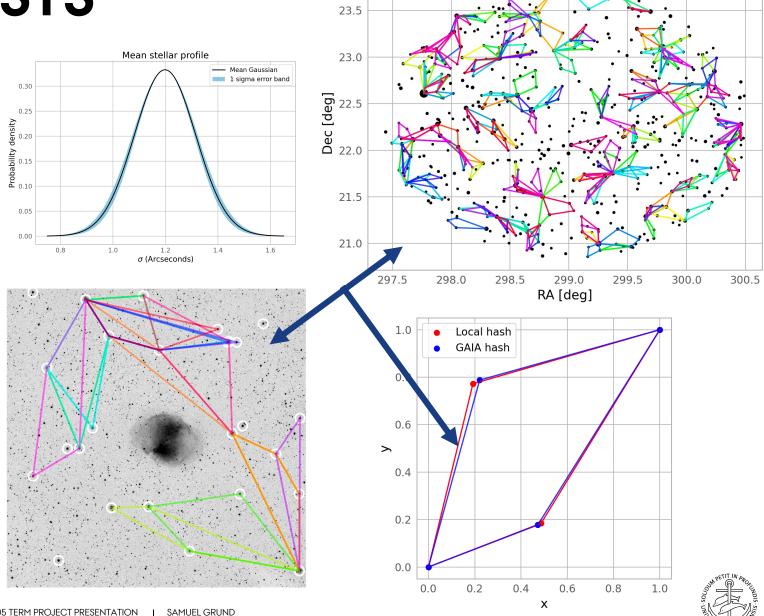
- 2D Convolution techniques for star detection
- Fitting of stars to gaussian profiles
- Examination of statistical properties of stars and using this to reject outliers
- Dimensionality reduction (by Hashing algorithm [1], not PCA)
- Data fetching and handling from remote databases (GAIA DR2[3])
- TODO: Bayesian inference for comparing the quality of proposed solutions





RESULTS AND TESTS

- Stars detected and asterisms built
- Matched to GAIA catalogue index asterisms
- Input image stellar shape parameters determined
- Original idea: Hash from both geometry and relative flux
 - Does not work because of saturation effects and band variance.





REFERENCES

[1] Lang, D., Hogg, D. W., Mierle, K., Blanton, M., & Roweis, S. (2010). Astrometry. net: Blind astrometric calibration of arbitrary astronomical images. The astronomical journal, 139(5), 1782. (https://arxiv.org/abs/0910.2233)

[2] Starin, Scott R. and John S. Eterno. "Attitude Determination and Control Systems." (2010) (https://ntrs.nasa.gov/api/citations/20110007876/downloads/20110007876.pdf)

[3] Katz, D., Antoja, T., Romero-Gómez, M., Drimmel, R., Reylé, C., Seabroke, G. M., ... & Gaia Collaboration. (2018). Gaia data release 2: mapping the Milky Way disc kinematics. arXiv preprint arXiv:1804.09380 (https://arxiv.org/abs/1804.09380).





