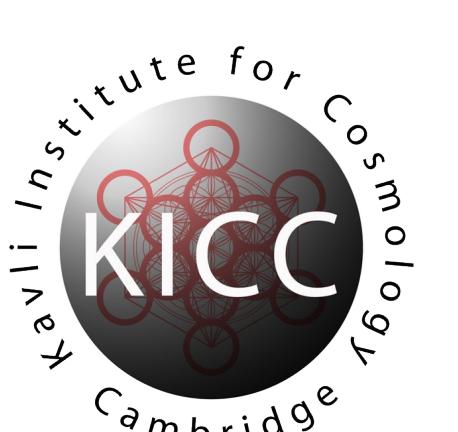


# Bayesian Anomaly Detection for Ia Supernovae using JAX-bandflux

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**Robust parameter estimation for Type Ia supernovae is crucial for precision cosmology.** Anomalous data points from instrumental artifacts or astrophysical contamination can bias SALT3 fits and propagate into cosmological parameters. We present a differentiable Bayesian framework that simultaneously fits supernova light curves while identifying anomalous observations, implemented in JAX for GPU acceleration.



## Bayesian Anomaly Detection

### Problems with current methods:

- Traditional methods are not model aware
- Anomalies sought before/after fitting, not simultaneously
- Binary classification (no encoding of belief)
- Standard likelihoods cannot incorporate anomalous data

### Our solution

We model each data point as either expected or anomalous:

1. Define anomaly mask:  $\varepsilon_i \in \{0, 1\}$  for each data point
2. Bernoulli prior:  $P(\varepsilon_i) = p^{\varepsilon_i}(1-p)^{1-\varepsilon_i}$

### 3. Piecewise likelihood:

$$P(\vec{D}, \vec{\varepsilon} | \theta) = \prod_{i=1}^N (L_i(\theta)(1-p))^{(1-\varepsilon_i)} \left(\frac{p}{\Delta}\right)^{\varepsilon_i}$$

4. Marginalize over  $\varepsilon$ :  $P(\mathcal{D} | \theta) = \sum_{\varepsilon \in \{0,1\}^N} P(\mathcal{D}, \varepsilon | \theta)$

### 5. Dominant mask approximation:

$$P(\mathcal{D} | \theta, \varepsilon_{\max}) \gg \max_j P(\mathcal{D} | \theta, \varepsilon^{(j)})$$

### 6. Final log-likelihood:

$$\log P(\mathcal{D} | \theta) = \begin{cases} \log \mathcal{L}_i + \log(1-p), & \text{if } \log \mathcal{L}_i + \log(1-p) > \log p - \log \Delta \\ \log p - \log \Delta, & \text{otherwise} \end{cases}$$

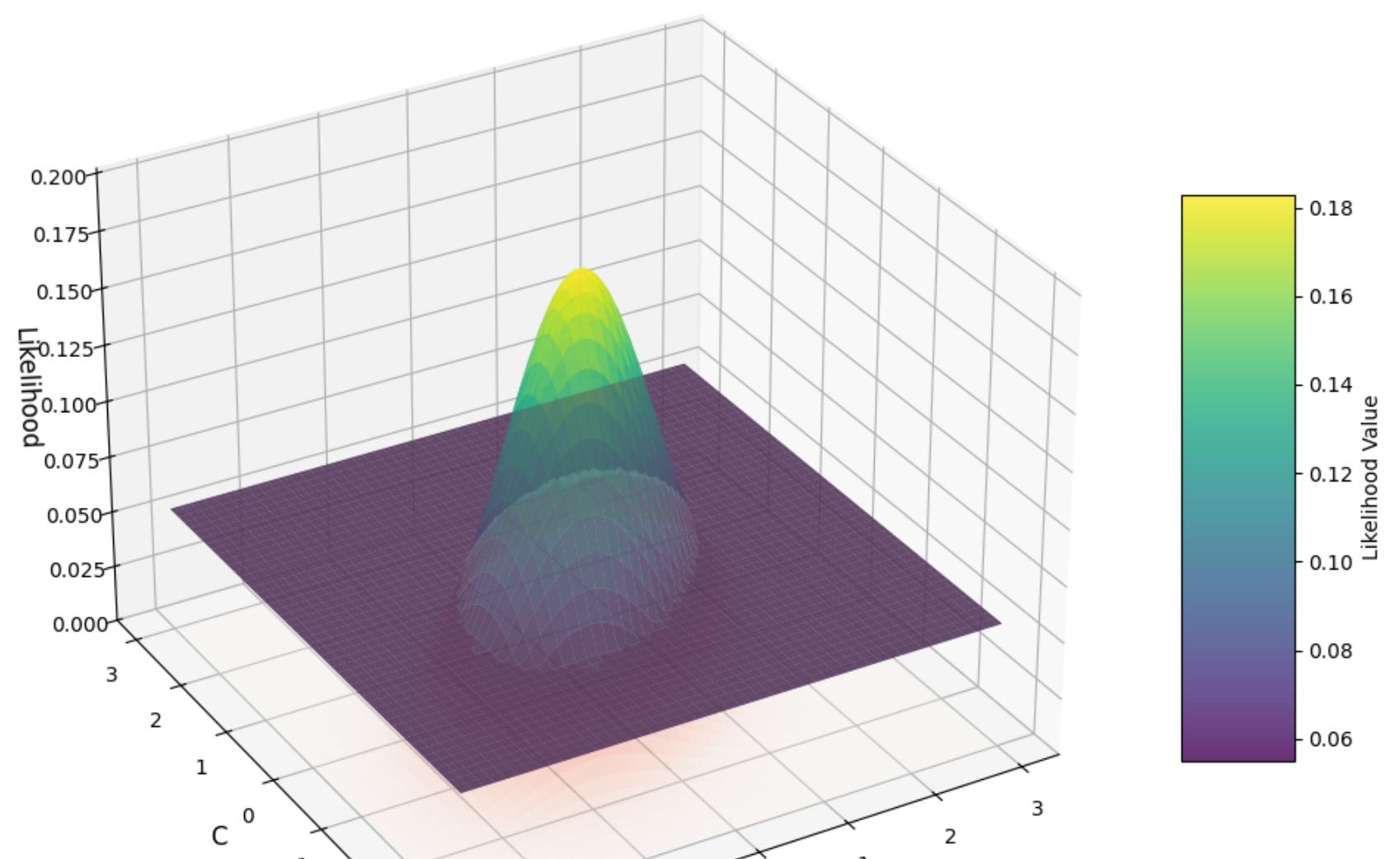
### Key Insight

We fit for this 'floor' as a free parameter, fully automating the anomaly detection process.

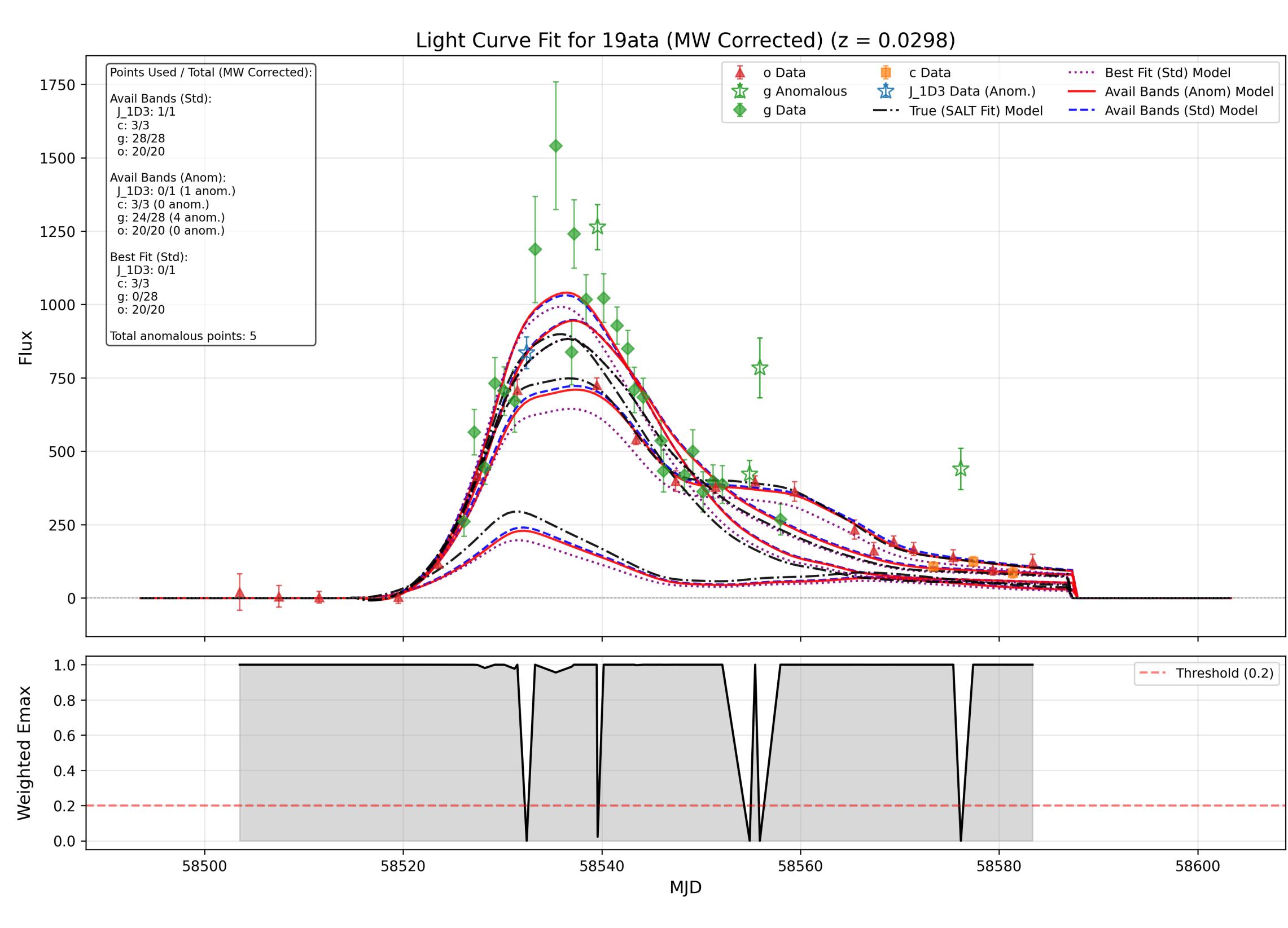
### Demo:

[github.com/samleene/Bayesian-Anomaly-Detection/blob/main/demo\\_anomaly\\_detection.ipynb](https://github.com/samleene/Bayesian-Anomaly-Detection/blob/main/demo_anomaly_detection.ipynb)

2D Gaussian Likelihood with a Flat Floor at  $p$



## SN 19amo: Light Curve



## JAX-bandflux

### A differentiable framework for supernova analysis:

#### Key Features:

- GPU-accelerated SALT3 model fitting
- Built on JAX for automatic differentiation
- Integrates with BlackJAX for nested sampling
- Handles large supernova datasets efficiently

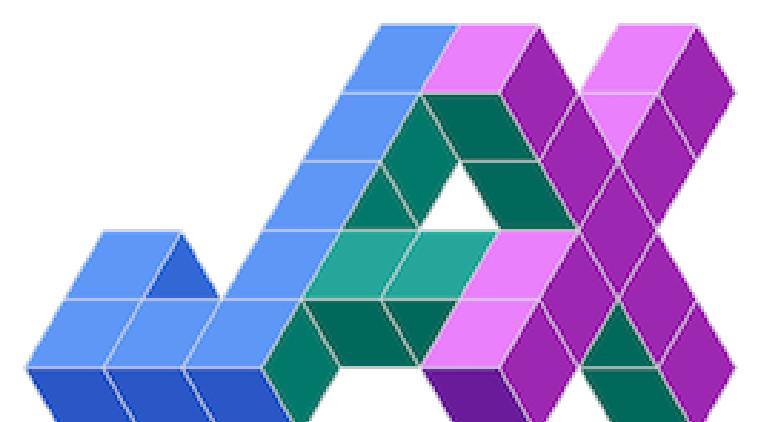
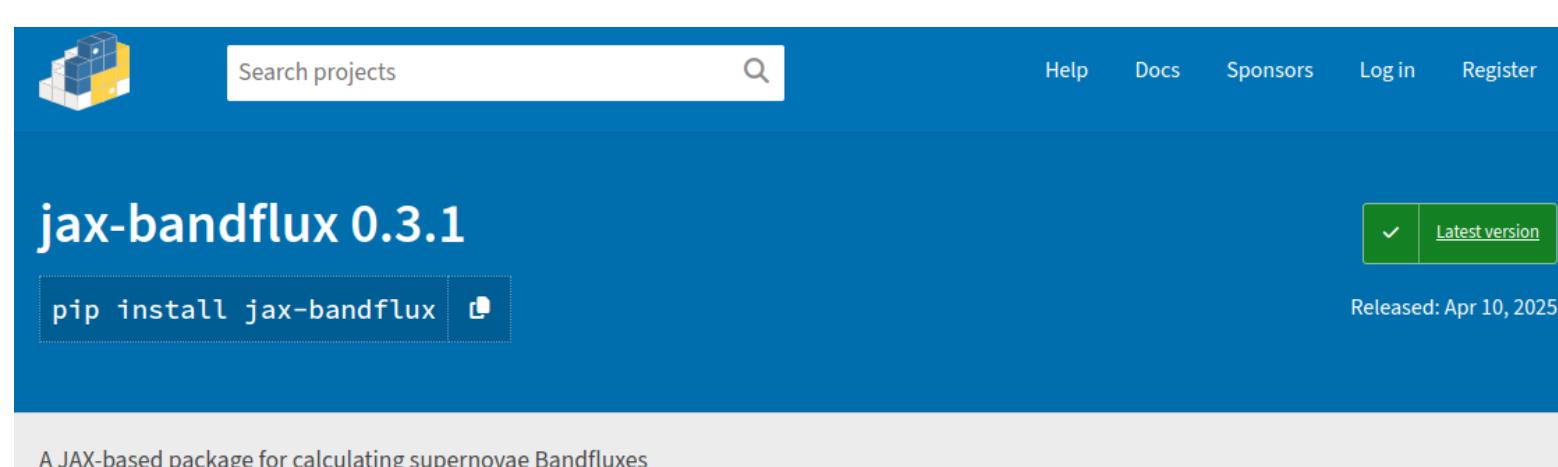
**Example:** [handley-lab.co.uk/nested-sampling-book/physics/supernovae.html](http://handley-lab.co.uk/nested-sampling-book/physics/supernovae.html)

#### SALT3 Model:

$$F(p, \lambda) = x_0 [M_0(p, \lambda) + x_1 M_1(p, \lambda)] \times e^{c \times CL(\lambda)}$$

#### Bandflux Integration:

$$F_{\text{band}} = \int_{\lambda_{\min}}^{\lambda_{\max}} F(p, \lambda) \cdot T(\lambda) \cdot \frac{\lambda}{hc} d\lambda$$

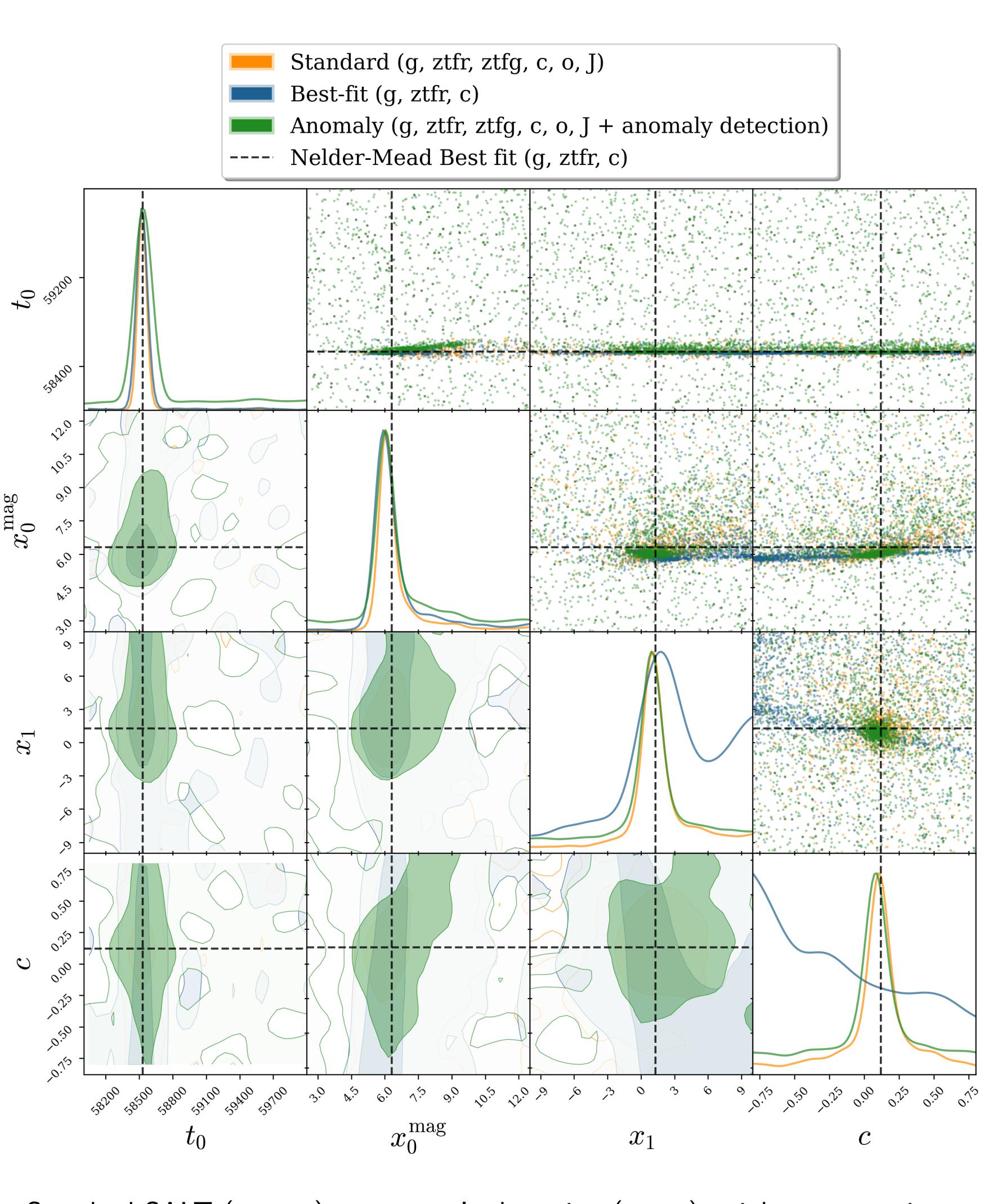


## Main Advantages

### Utility of Bayesian anomaly detection for Ia analysis:

1. Fully automated anomaly detection
2. Fully automated filter selection
3. Preserves previously 'bad' filters by mitigating anomalies

## Corner Plot Comparison



## Next Steps

### Future Applications:

- Extend framework for distance measurements
- Estimate  $H_0$  using robust supernova samples
- Apply to other astronomical datasets

Code: [github.com/samleene](https://github.com/samleene)

Paper: [arXiv:2211.15448](https://arxiv.org/abs/2211.15448)

