

## Lecture 3

### Learning Goals:

By the end of class today you should be able to ...

- Use scientific notation to write very large and very small numbers
- Define the celestial sphere

Reading for Today: Units 3, 5.1, 5.2

Reading for Next Time: 5, 6

### Scientific Notation

- As we have seen in astronomy we typically deal with numbers that are very large (or small).

– Distance of earth from sun:

93,000,000 miles

– Velocity of light: 186,000 miles/second

– Density of interstellar space:

0.000,000,000,000,000,000,0017 that of water

- Need a way of dealing with this  
→ scientific notation

- Way of writing numbers as a decimal between 1 and 10 multiplied by a power of 10

- Benefits of using scientific notation:

– Easier to write large and small numbers

– Easy to see how many significant figures a number has

- I.e. how well a number is known

Makes arithmetic easier

## Lecture 3

### Scientific Notation

• **In scientific notation, 2000 would be written as  $2 \times 10^3 =$**

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• To convert, count the number of spaces to the left, between the decimal point and the leftmost non-zero digit.

### Scientific Notation

• For positive numbers less than one, exponent is negative

**E.g. 0.002 would be written as  $2 \times 10^{-3} =$**

- To convert, count the number of spaces to the right, between the decimal point and the leftmost, non-zero number.

### Units

• In this course we will primarily use the metric system.

- Length (meter, m)

- Mass (kilogram, kg)

- Time (second, sec)

• Units for other measured quantities can be written in terms of these fundamental units.

• Speed (distance/time): [m/s]

• Density (mass/volume)

= mass/(length x length x length):

[kg/m<sup>3</sup>]

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### Special Astronomy Units

- Even in scientific notation hard to think about what a number like

$1.99 \times 10^{30}$  kg means

- So, often use special units

- Astronomical Unit (Earth – Sun distance)

  - $1 \text{ AU} = 1.50 \times 10^{11} \text{ m}$

- Solar Radii

  - $1 R_{\text{Sun}} = 6.97 \times 10^8 \text{ m}$

- Solar Masses

  - $1 M_{\text{Sun}} = 1.99 \times 10^{30} \text{ kg}$

- Earth Radii

  - $1 R_{\text{Earth}} = 6.37 \times 10^6 \text{ m}$

- Earth Masses

  - $1 M_{\text{Earth}} = 5.97 \times 10^{24} \text{ kg}$

- For more practice with scientific notation and units see the Basic

Mathematics Review Guide

– Under ‘Course Documents’ on Blackboard

- Excellent illustration of the scale of the Universe:

– Video: *Powers of 10* by Charles & Ray Eames

– Watch the video online for homework – ‘External Links’ on BB

### Celestial Sphere

Treat stars as if they are all at the same distance from us with the Earth at the center

→ Useful fiction

### Finding Your Way Around The Celestial Sphere

- Zenith – \_\_\_\_\_

- Nadir – \_\_\_\_\_

### Lecture 3

- Extend poles and equator of Earth out to celestial sphere
  - **Celestial poles**
  - **Celestial equator**

- Meridian –
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#### Celestial Sphere

- Stars (and the Sun and Moon) rise and set as Earth rotates on it's axis
  - celestial sphere appears to rotate around celestial poles

#### Daily Motion

- As Earth rotates, stars rise in the East, move across the sky, and set in the West.
- We see stars rise in the East and Set in the West because the Earth rotates counterclockwise (as looking down on the North Pole)