

# CLASSICAL MECHANICS PROBLEMS (MAY2025)

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## Problem 1

A charged particle with mass  $m$ , charge  $e$ , and velocity  $\mathbf{v}$  moves in an electromagnetic field described by the electric scalar potential  $\phi$  and magnetic vector potential  $\mathbf{A}$ . The speed of light is denoted by  $c$ . The Lagrangian of the system is given by:

$$L = T - U = \frac{mv^2}{2} - e\phi + \frac{e}{c} \mathbf{v} \cdot \mathbf{A}$$

where  $T$  is the kinetic energy,  $U$  is the potential energy, and  $\mathbf{v} \cdot \mathbf{A}$  represents the dot product between the velocity and the magnetic vector potential. prove this statement.

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## Problem 2

Huygens' cycloidal pendulum consists of a particle which oscillates under gravity in the vertical plane along a frictionless cycloidal track with parametric equations where the vertical y-axis points downward. Calculate that a Lagrangian for this system.

$$x = R(\theta - \sin \theta), \quad y = R(1 - \cos \theta),$$

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## Problem 3

 Compute the EOM for this Lagrangian

$$L = e^{\lambda t} \left( \frac{m}{2} \dot{x}^2 - \frac{m\omega^2}{2} x^2 \right)$$

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## Problem 4

Show that the Lagrangian

$$L = \frac{m}{2} (\dot{r}^2 + r^2 \dot{\theta}^2 + r^2 \dot{\phi}^2 \sin^2 \theta) - \frac{eg}{c} \dot{\phi} \cos \theta$$

describes a charged particle in the magnetic field  $\mathbf{B} = g\mathbf{r}/r^3$  of a magnetic monopole and find Lagrange's equations. Hint: check that the vector potential for a magnetic monopole in spherical coordinates has the components  $A_r = 0$ ,  $A_\theta = 0$ ,  $A_\phi = g(1 - \cos \theta)/r \sin \theta$ .

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## Problem 5

 For this Lagrangian  $\mathcal{L}$  find the conserved momenta.

$$\mathcal{L} = \frac{1}{2} \left[ -f(r)\dot{t}^2 - 4nf(r)(\cos \theta - 1)\dot{t}\dot{\phi} + \frac{\dot{r}^2}{f(r)} + (r^2 + n^2)\dot{\theta}^2 + ((r^2 + n^2)\sin^2 \theta - 4n^2 f(r)(\cos \theta - 1)^2) \dot{\phi}^2 \right]. \quad (1)$$

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**Problem 6** A particle  $P$  of mass  $m$  moves in a central force field described by the force:

$$\mathbf{F} = -\frac{\gamma m}{r^2} \hat{r},$$

where  $\gamma$  is a positive constant, and  $\hat{r}$  is the radial unit vector. The force is attractive and follows an inverse-square law, similar to gravity. We need to determine the conditions for bound and unbound orbits based on the total mechanical energy  $E$ .

An asteroid approaches the Sun from a great distance with constant speed  $u$  in a straight line, where the perpendicular distance from the Sun is  $p$ . We need to find the equation satisfied by the subsequent orbit. For the special case where  $u^2 = \frac{4M_\odot G}{3p}$  (where  $M_\odot$  is the mass of the Sun, and  $G$  is the gravitational constant), find:

- (i) the distance of closest approach of the asteroid to the Sun, and
- (ii) the speed of the asteroid at the time of closest approach.

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**Problem 7** Find the force law for a central-force field that allows a particle to move in a logarithmic spiral orbit given by  $r = ke^{\alpha\theta}$ , where  $k$  and  $\alpha$  are constants.

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**Problem 8** Determine  $r(t)$  and  $\theta(t)$  for the problem in previous problem.

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**Problem 9** What is the total energy of the orbit of the previous two examples? \_\_\_\_\_

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**Problem 10**

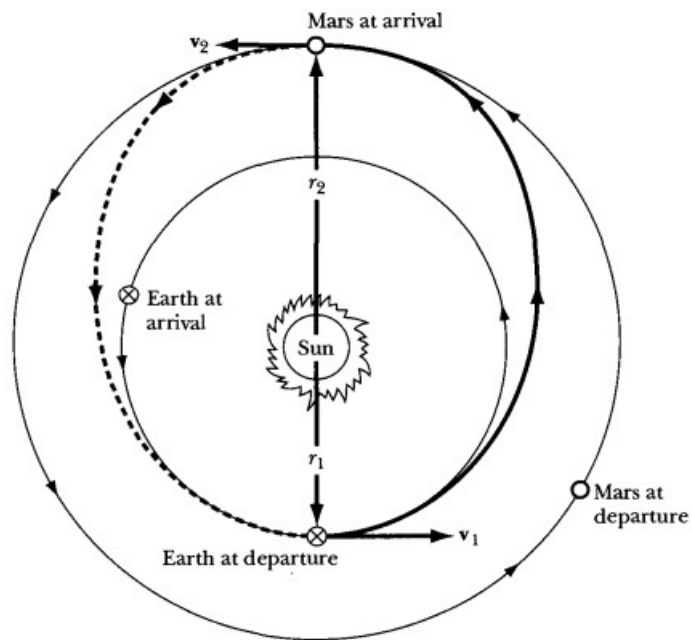
Halley's comet, which passed around the sun early in 1986, moves in a highly-elliptical orbit with an eccentricity of 0.967 and a period of 76 years. Calculate its minimum and maximum distance from the Sun.

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**Problem 11**

Halley's comet, which passed around the sun early in 1986, moves in a highly-elliptical orbit with an eccentricity of 0.967 and a period of 76 years. Calculate its minimum and maximum distance from the Sun.

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### Problem 12

Calculate the time it takes for a spacecraft to make a Hohmann transfer from Earth to Mars and the heliocentric transfer speed required assuming that both planets are in coplanar orbits.