

21/5/25

Lec 9 & 10

Problem 6

A particle  $P$  of mass  $m$  moves in a central force field described by the force:

$$F = -\frac{\gamma m}{r^2} \hat{r},$$

where  $\gamma$  is a positive constant, &  $\hat{r}$  is the radial unit vector. The force is attractive & follows an inverse square law, similar to gravity. We need to determine the conditions for bound & unbound orbits based on the total mechanical energy  $E$ .

An asteroid approaches the Sun from a great distance with const. speed  $u$  in a straight line, where the perpendicular distance from the Sun is  $p$ . We need to find the equation satisfied by the subsequent orbit. For the special case where

$$u^2 = \frac{4M_{\odot}G}{3p} \quad \left( \text{where } M_{\odot} \text{ is the mass of the Sun, \& } G \text{ is the Gravitational const.} \right),$$

find:

- (i) the distance of closest approach of the asteroid to the Sun, &
- (ii) the speed of the asteroid at the time of closest approach.

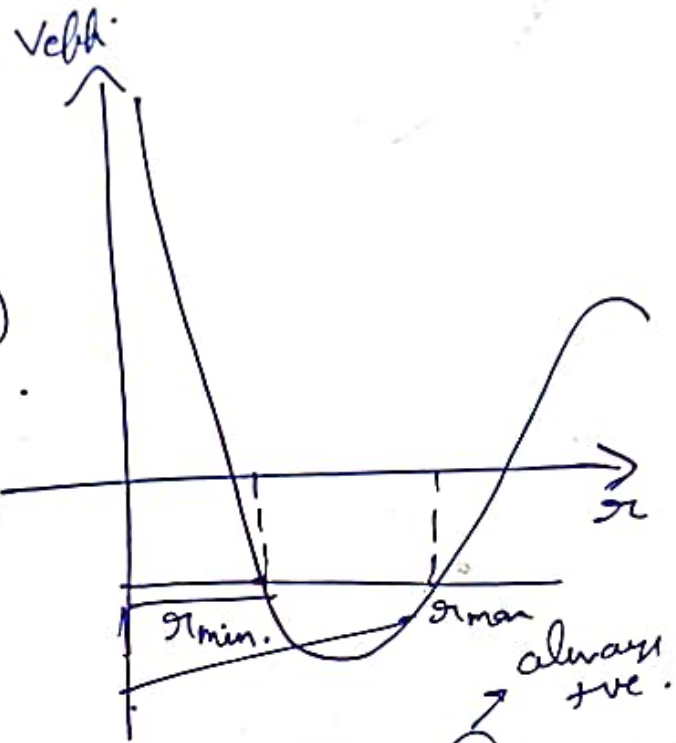
$$L = T - V$$

$$= \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\theta}^2) - V(r)$$

$$E = \frac{1}{2} m \dot{r}^2 + V(r) + \frac{L^2}{2mr^2}$$

$$\frac{-k}{r}$$

$$V_{\text{eff.}}$$



$$E = \bar{T} + V_{\text{eff.}}$$

$$V_{\text{eff.}} = E$$

$$\frac{-k}{r} + \frac{L^2}{2mr^2} = E$$

(acc to que.)

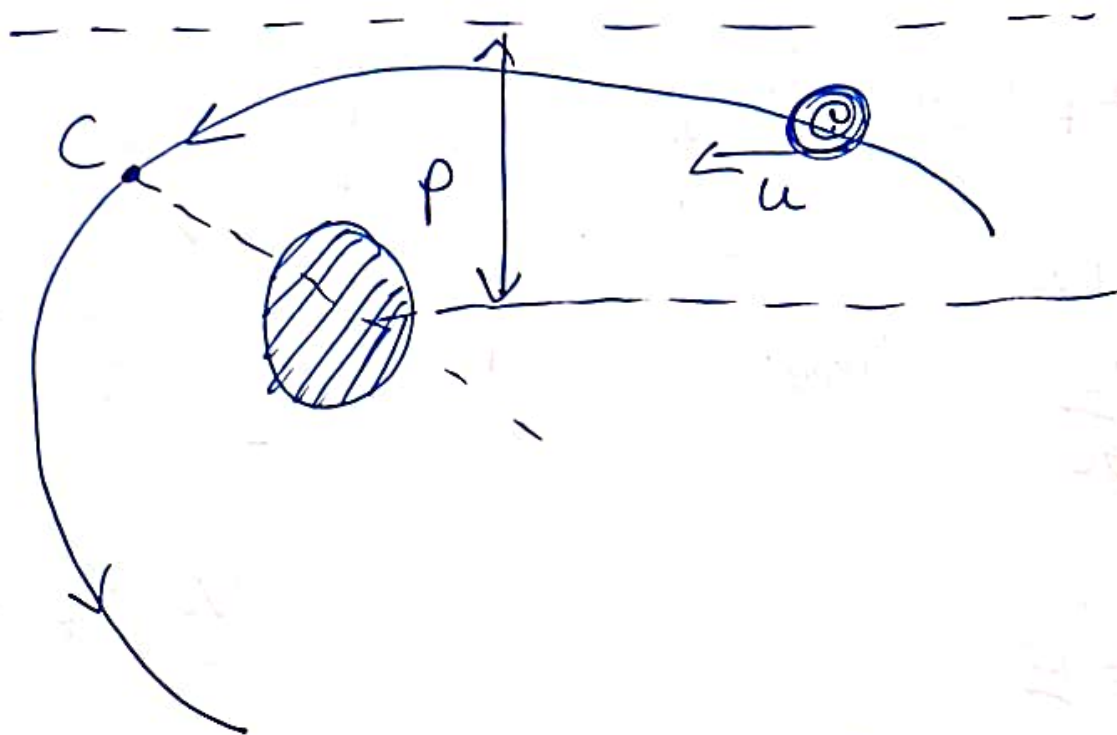
$$\frac{-k}{r}$$

$$(V_{\text{eff.}} - E) = -T$$

$$(V_{\text{eff.}} - E) \leq 0$$

$$V_{\text{eff.}} \leq E$$

Bounded means  $\rightarrow E < 0$   
 $F > 0$



Sun in a asteroid system.

$$E = \frac{1}{2} m v^2 \quad V_{eff} = E$$

$$-\frac{\gamma}{r} + \frac{L^2}{2mr^2} = E$$

$$-\frac{GM_\odot m_a}{r} + \frac{L^2}{2m_a r^2} = E$$

$$\tilde{L} = \frac{L}{m} = r^2 \dot{\theta}$$

$L \Rightarrow$  conserved quantity

$$\frac{-GM_{\odot}}{r} + \frac{L^2}{2r^2} = E$$

moment of

$$L = p u$$

$$\frac{-GM_{\odot}}{r} + \frac{p^2 u^2}{2r^2} = \frac{1}{2} u^2$$

$$\left( \Rightarrow R_h \text{ Gravity} \right) \quad E = \frac{1}{2} u^2$$

$$\frac{-GM_{\odot}}{r} + \frac{p^2 u^2}{2r^2} = \frac{1}{2} u^2$$

$$\frac{-GM_{\odot}}{r} + p^2$$

$$\frac{-GM_{\odot}}{r}$$

$$\frac{-2GM_{\odot} r + p^2 u^2}{2r^2} = \frac{1}{2} u^2$$

$$-2GM_{\odot} r + p^2 u^2 = \frac{2r^2 u^2}{2}$$



$$-\frac{GM_{\odot}}{r} + \frac{p^2 u^2}{2r^2} = \frac{1}{2} u^2$$

$$\cancel{-\frac{GM_{\odot}}{r}} + \cancel{p^2} \quad \cancel{-\frac{GM_{\odot}}{r}}$$

$$\frac{-2GM_{\odot}r + p^2 u^2}{2r^2} = \frac{1}{2} u^2$$

$$-2GM_{\odot}r + p^2 u^2 = \frac{2r^2 u^2}{2}$$

$$-2GM_{\odot}r + p^2 u^2 = r^2 u^2$$

$$r^2(u^2) + (2GM_{\odot})r + (-p^2 u^2) = 0$$

$$r = \frac{-(2GM_{\odot}) \pm \sqrt{(2GM_{\odot})^2 - 4(u^2)(-p^2 u^2)}}{2(u^2)}$$

$$= \frac{-2GM_{\odot}}{2u^2} \pm \frac{\left((2GM_{\odot})^2 + 4u^2 \cdot u^2 \cdot p^2\right)^{1/2}}{2u^2}$$

$$r = \frac{GM_0 \cdot 3p}{4M_0G} \pm \frac{\left[ 4G^2M_0^2 + 4p^2 \cdot \left( \frac{4M_0G}{3p} \right)^2 \right]}{2 \cdot \frac{4M_0G}{3p}}$$

$$= \cancel{GM_0} \cdot \frac{3p}{\cancel{4M_0G}} \pm \frac{\frac{10}{3} \cancel{GM_0}}{\frac{8}{3p} \cancel{M_0G}}$$

$$= \frac{3p}{4} \pm \frac{\frac{5}{10} \times \frac{3p}{84}}{\frac{1}{84}} \quad (\cancel{2GM_0})^2 \pm$$

$$= \frac{3p}{4} \pm \frac{5p}{4}$$

$$= \frac{2p}{4_1} \text{ or } -\frac{2p}{4_2}$$

$$= 2p \text{ or } \frac{p}{2}$$

$$\boxed{r=2p}$$

So

$$\boxed{\boxed{r=\frac{p}{2}}}$$

closest pt.

$$4G^2M_0^2 + 4p^2 \cdot \frac{4^2M_0^2G^2}{3^2p^2}$$

$$\sqrt{4G^2M_0^2 \left( 1 + \frac{4^2}{3^2} \right)}$$

$$2GM_0 \sqrt{1 + \left( \frac{4}{3} \right)^2}$$

$$\frac{2 \times 5GM_0}{3}$$

$$\frac{10}{3} GM_0$$

$$1 + \frac{16}{9} = \frac{9+16}{9} = \frac{25}{9}$$

$$= \frac{5}{3}$$

Hence the closest point is  $\frac{P}{2}$ .

So the velocity at that point would be, ?  
angular momentum is conserved.

$$P u = \frac{P}{2} v$$

$$v = 2u$$