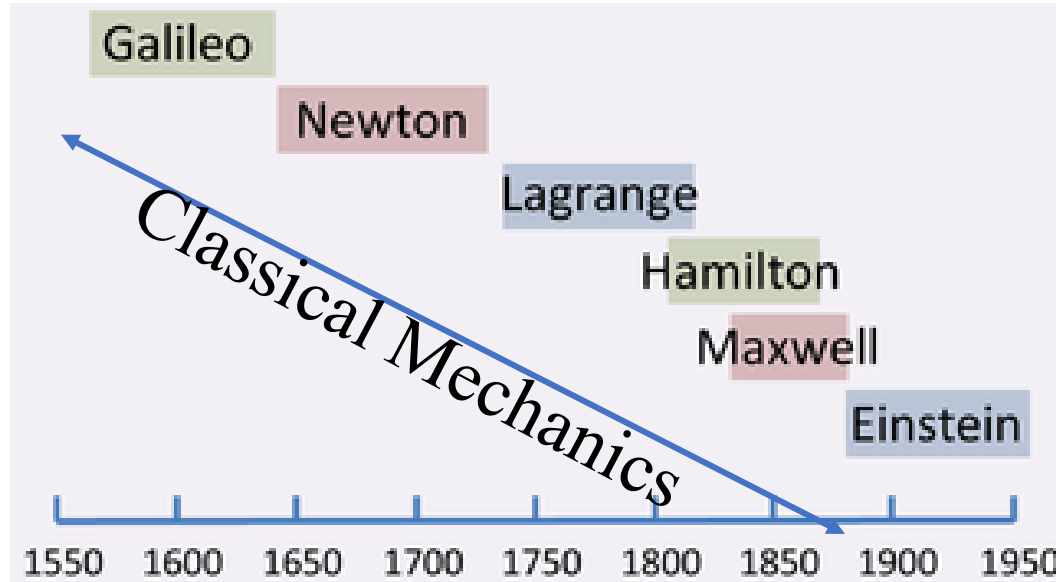


Evolution in Mechanics



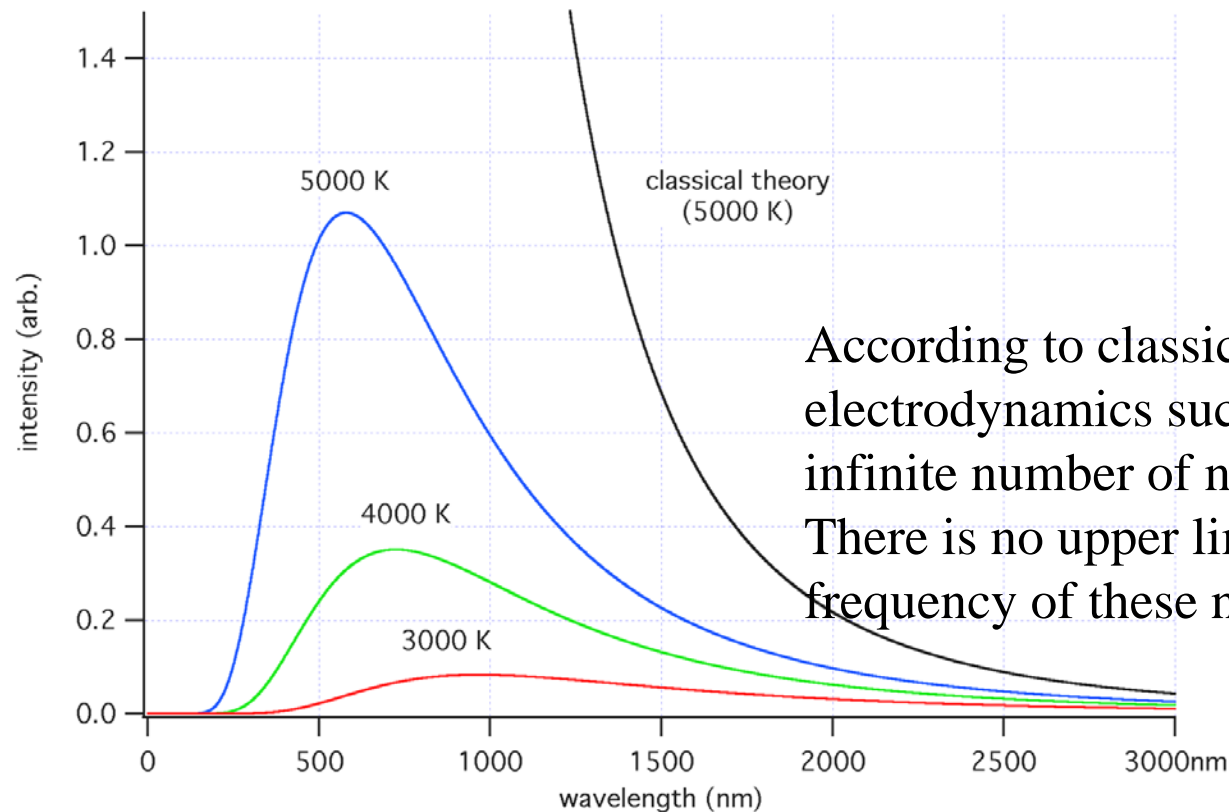
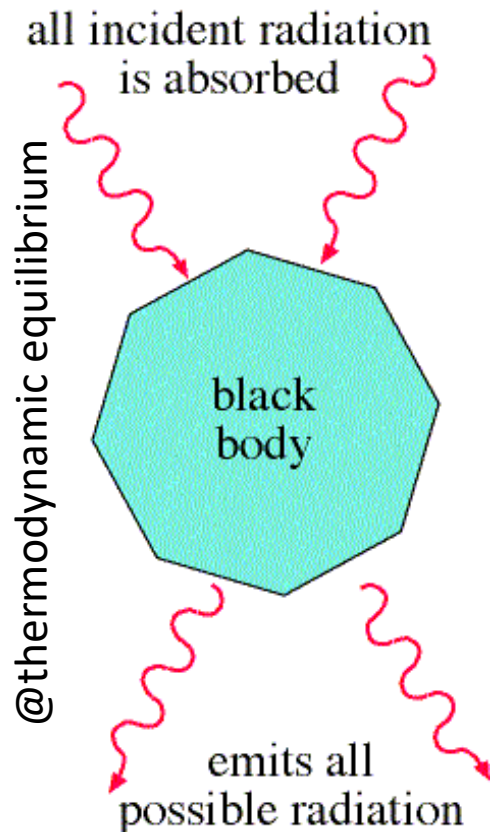
	Newtonian Mechanics	Lagrangian Mechanics	Hamiltonian Mechanics
Principle	Force = $m \times$ <i>accln</i>	Principle of Least Action	Reformulation of Lagrangian Mechanics (in Phase space coordinates (x,p))
Driving Equation	$m\ddot{x} = F$	Lagrange's Eq. $\frac{d}{dt} \left\{ \frac{\partial L}{\partial \dot{x}} \right\} = \frac{\partial L}{\partial x}$ $L = K.E. - P.E.$	$\dot{p} = -\frac{\partial H}{\partial x}, \dot{x} = \frac{\partial H}{\partial p}$ $H = K.E. + P.E.$

	Newtonian Mechanics	Lagrangian Mechanics	Hamiltonian Mechanics
Principle	Force = $m \times accln$	Principle of Least Action: minimum energy path	Reformulation of Lagrangian Mechanics (in Phase space coordinates (x,p))
Driving Equation	$m\ddot{x} = F$	Lagrange's Eq. $\frac{d}{dt} \left\{ \frac{\partial L}{\partial \dot{x}} \right\} = \frac{\partial L}{\partial x}$ $L = K.E. - P.E.$	$\dot{p} = -\frac{\partial H}{\partial x}, \dot{x} = \frac{\partial H}{\partial p}$ $H = K.E. + P.E.$
Co-ordinates	Cartesian Coordinates	Cartesian Coordinates	Canonical Coordinates
Requirements & Basis	Cause = Effect; Vectorial in Nature; Does not have systematic method for deriving conservation law	Start from 1st Principle; Based on scalar function: Energy in terms of Lagrangian. Not ideal for Non-conservative system, Like friction	Converting 2 nd order differential Lagrange's equation to 1 st order Hamilton's Equations.

Classical → Quantum Mechanics

Despite the huge success of Classical Mechanics

- There are certain experimental results bothered physicists during end part of 19th Century
- Most of the properties of light ← Explained with Classical Mech.
- Means light as wave following wave equation
- **Ultraviolet catastrophe of Black Body Radiation could not**

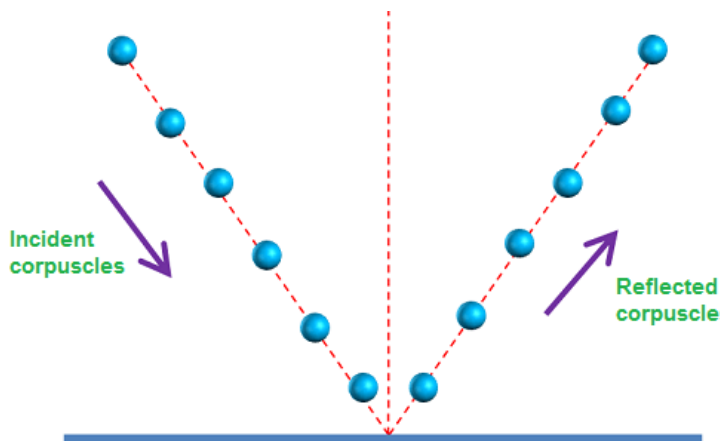


Classical → Quantum Mechanics

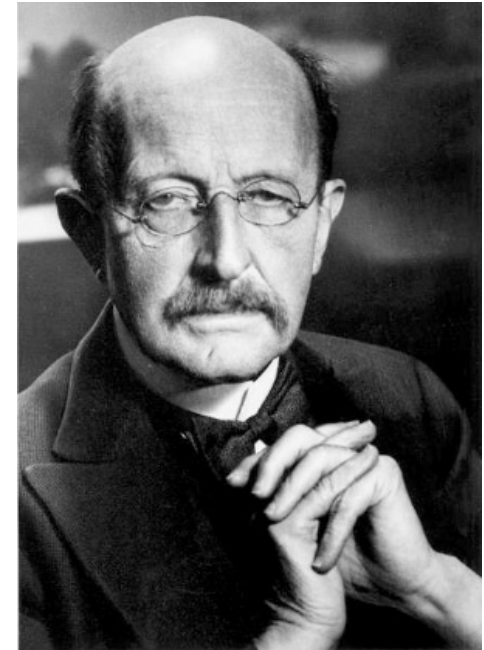
Despite the huge success of Classical Mechanics

- There are certain experimental results bothered physicists during end part of 19th Century
- Most of the properties of light ← Explained with Classical Mech.
- Means: light as wave following wave equation
- Ultraviolet catastrophe of Black Body Radiation could not
- Eventually understood discrete quanta !!!
- Particle (Photon) picture of light wave

Many years back Newton proposed → Corpuscular Theory



- However,
- Particle (Photon) picture of light wave
 - Max Planck's hypothesis of light (@1900)
 - laid foundation of particle-like behavior of light



Max Karl Ernst
Ludwig Planck
1858-1947

Classical → Quantum Mechanics

The density of radiant energy in the cavity per unit wavelength interval, at the wavelength λ , and at the temperature T

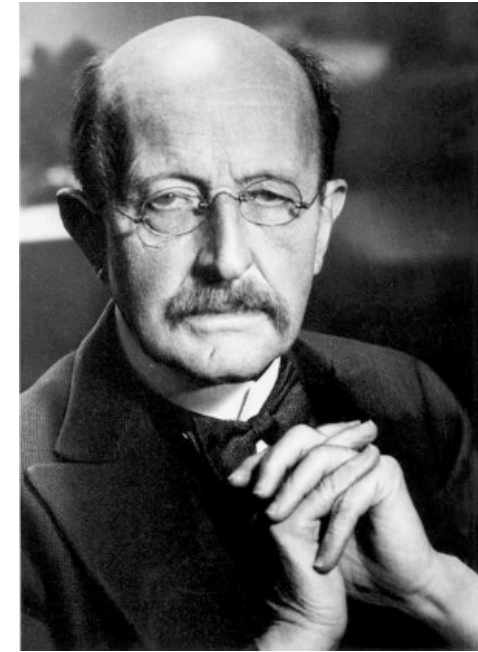
$$E(\lambda, T) = \frac{8\pi hc}{\lambda^5} \times \frac{1}{e^{\frac{hc}{\lambda kT}} - 1}$$

Planck's Radiation Law

**Energy of wave is related with frequency and quantized $E=nh\nu$

*The average energy in a normal mode is of the order of kT .

*The probability of emission of high energy photons with high frequency gets smaller for a given temperature when $h\nu \gg kT$.
There is not enough energy available for emitting these photons.



Max Karl Ernst
Ludwig Planck
1858-1947

Other examples:

Waves behaving as particles

Experiments

1. Photoelectric effect (1902)
2. Compton effect (1922)
3. Pair Production (1932, Anderson)

*Theoretical proposal: Dirac (1928)

Wave-Particle duality

de Broglie postulate (PhD Thesis 1924)

$$\lambda = \frac{h}{p}$$

$$p\lambda = h$$

$$E = h\nu = \frac{h}{T}$$

$$ET = h$$

***Mass of an e- = 9.109383×10^{-28} gm;

Velocity is obtained from the given kinetic energy of 1000 eV

$$p = \sqrt{2m K.E.} = \sqrt{(1000 \text{ eV}) \left(\frac{1.6 \times 10^{-19} \text{ J}}{1 \text{ eV}} \right) 2(9.109383 \times 10^{-31} \text{ kg})}$$
$$= 1.7 \times 10^{-23} \text{ kg} \cdot \text{m/s}$$

$$\lambda = \frac{h}{p}$$
$$= \frac{6.626069 \times 10^{-34} \text{ kg} \cdot \text{m}^2/\text{s}}{1.7 \times 10^{-23} \text{ kg} \cdot \text{m/s}}$$
$$= 3.87 \times 10^{-11} \text{ m}$$
$$= 38.9 \text{ pm}$$

Wave-Particle duality

de Broglie postulate (PhD Thesis 1924)

$$\lambda = \frac{h}{p}$$

$$p\lambda = h$$

***Mass of a tennis ball = 0.11 kg;
Velocity = 44.7 m/s

$$\lambda = \frac{h}{mv} = \frac{6.626069 \times 10^{-34} \text{ kg} \cdot \text{m}^2/\text{s}}{(0.11 \text{ kg})(44.7 \text{ m/s})} = 1.3 \times 10^{-34} \text{ m}$$

$$E = h\nu = \frac{h}{T}$$

$$ET = h$$

Particles behaving as waves

Experiments

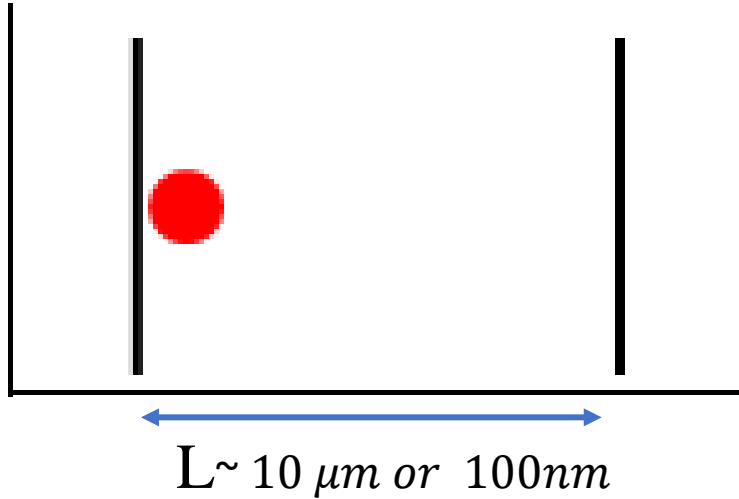
Electron diffraction

Davisson –Germer (USA)
and Thompson (UK) (1927)

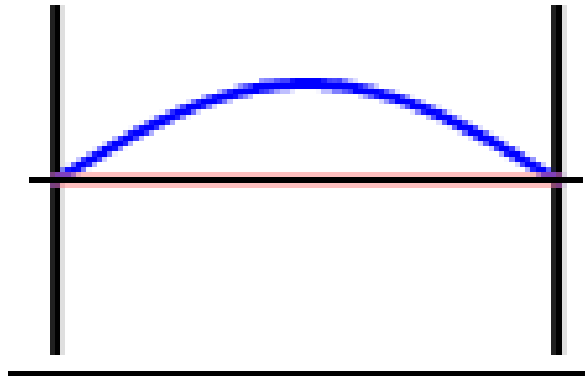
Electron microscope



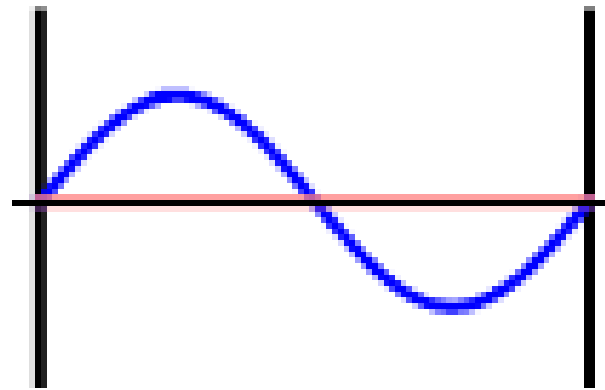
Confining Particle in small box



- If this particle has wave behavior
- The wave has to fit in the length
- Remember, if it is hard wall,
- Amplitude of the wave $= 0$ @ wall



Fundamental wave with $L = \frac{\lambda}{2}$
Ground State of the wave $E = \frac{h^2}{8mL^2}$



1st excited wave with $L = \lambda$

Energy of 1st excited state $E = \frac{2h^2}{8mL^2} = 2E_1$

WF: Wavefunction

- Real part of WF (blue line)
- Imaginary part of WF (red line)

Confining Particle in small box

- Energy of the particle discretized
- Possible energy levels

$$E_1 = h\nu$$

$$E_2 = 2h\nu$$

$$E_3 = 3h\nu$$

⋮

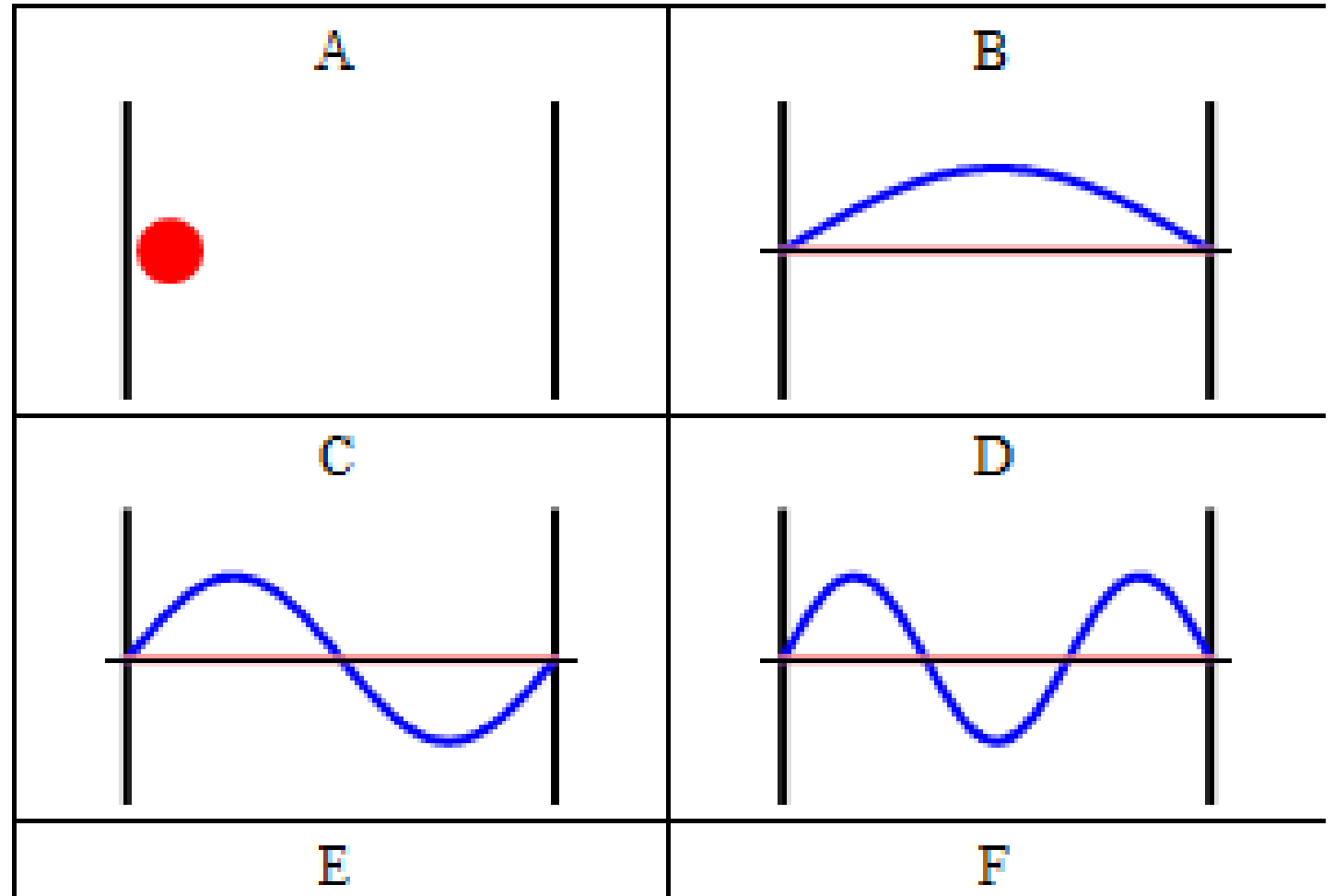
$$E_n = nh\nu$$

$$E_\infty = \infty h\nu$$

Question:

- Are the states are Equally probable??
- Is there any hierarchy?

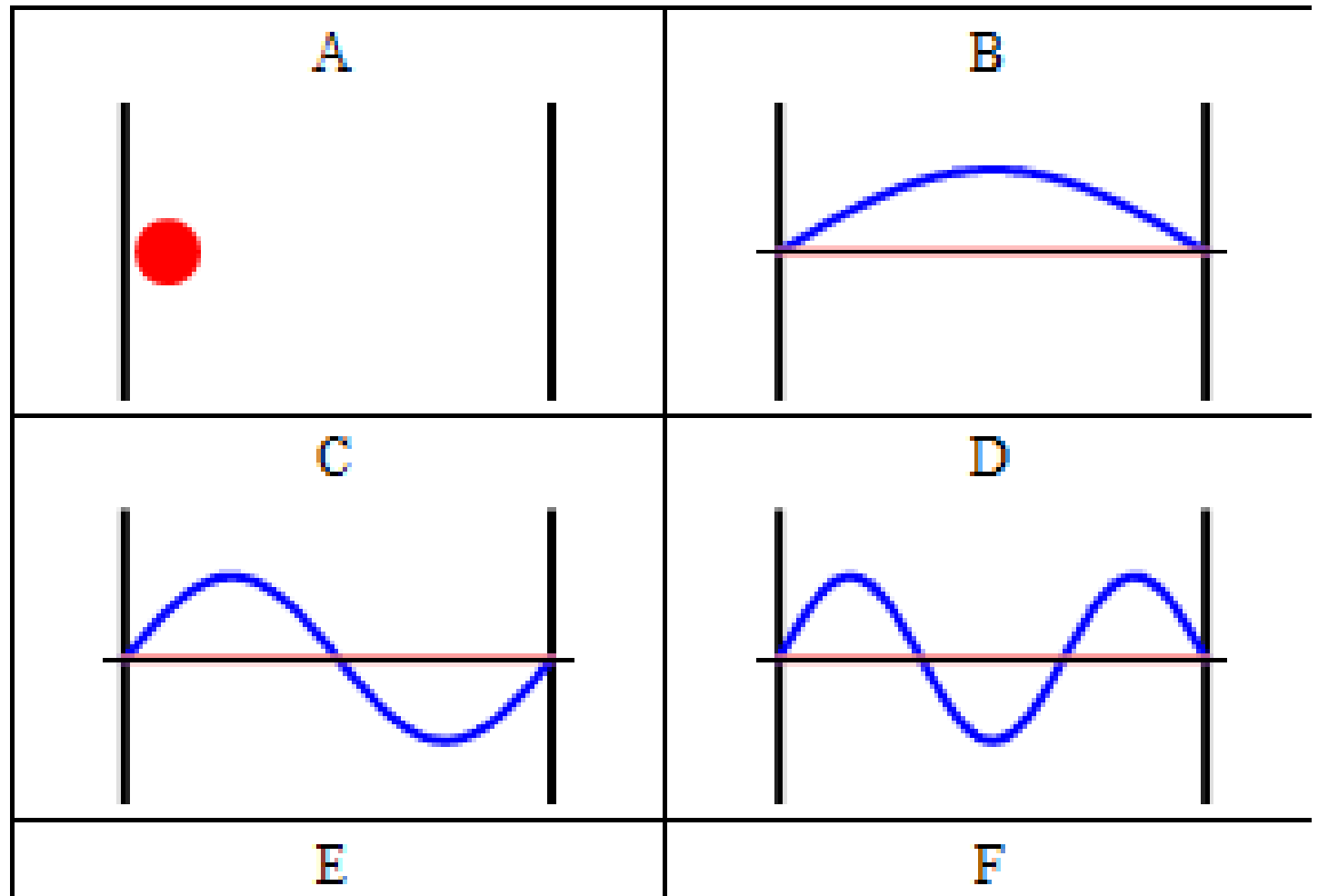
- Nature says: most probable state
→ Ground State
- But, probability of other excited states is finite, but decreases with “n”.



Confining Particle in small box

Look: Ground state wavefunction!

- Probability of finding the particle @box center is maximum
- As we go towards wall, prob. decreases and becomes 0 @wall
- Measurement of position will have uncertainty, but in the range between the wall
- Position uncertainty dictates its canonically conjugate variable
→ momentum uncertainty
(through Hamilton equations)
- This understanding → uncertainty of canonically conjugate variables
→ Uncertainty Relation



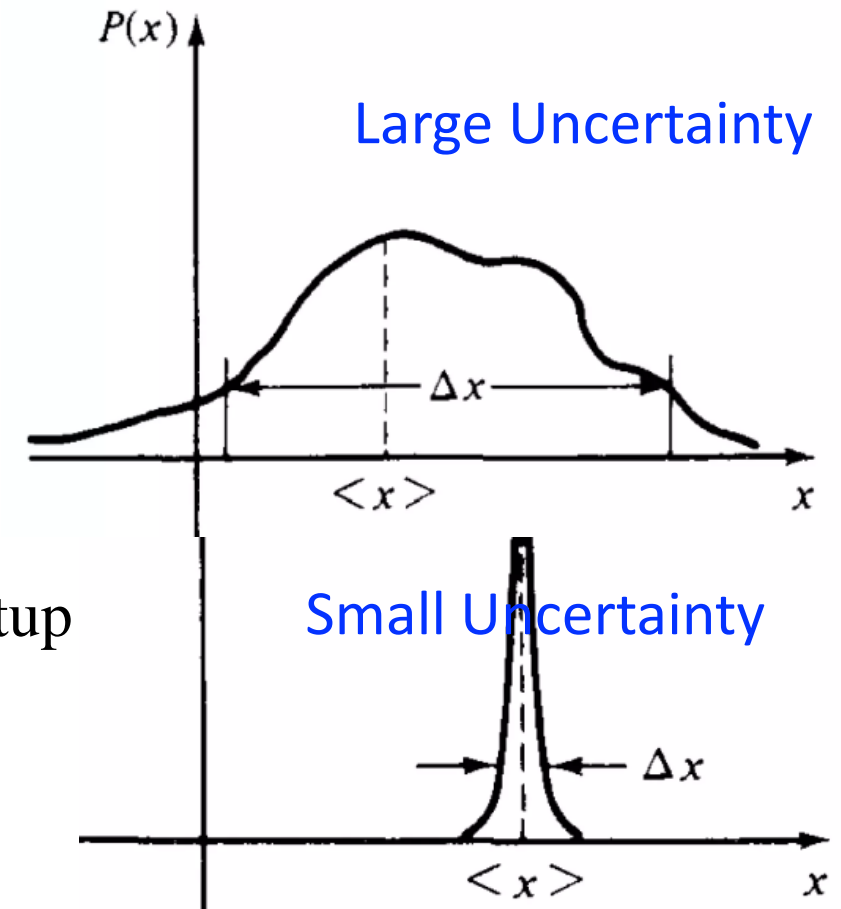
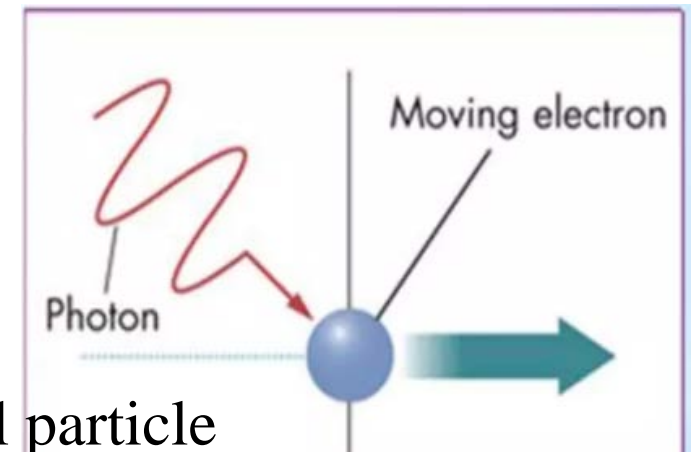
Understanding the Uncertainty

Why do we get uncertainty?

- Measurement requires interaction with probe?
- For very small particle: Interaction changes property of the small particle
- Accurate measurement of position
 - \leftrightarrow requires shorter wavelength of light
 - \leftrightarrow Shorter wavelength of light \rightarrow High energy photon
- Expect high energy or momentum will be transferred from photon to particle to get position accurately.

Uncertainty Definition: $(\Delta x)^2 = \langle x^2 \rangle - \langle x \rangle^2$

- This uncertainty measurement error or inaccuracy of the setup
- You can think about infinitely accurate experimental setup
- Still you can not avoid this quantum uncertainty
- This is fundamental and inherent in nature.





$$p = \frac{h}{\lambda} \Rightarrow$$

Fixed momentum \rightarrow Fixed wavelength

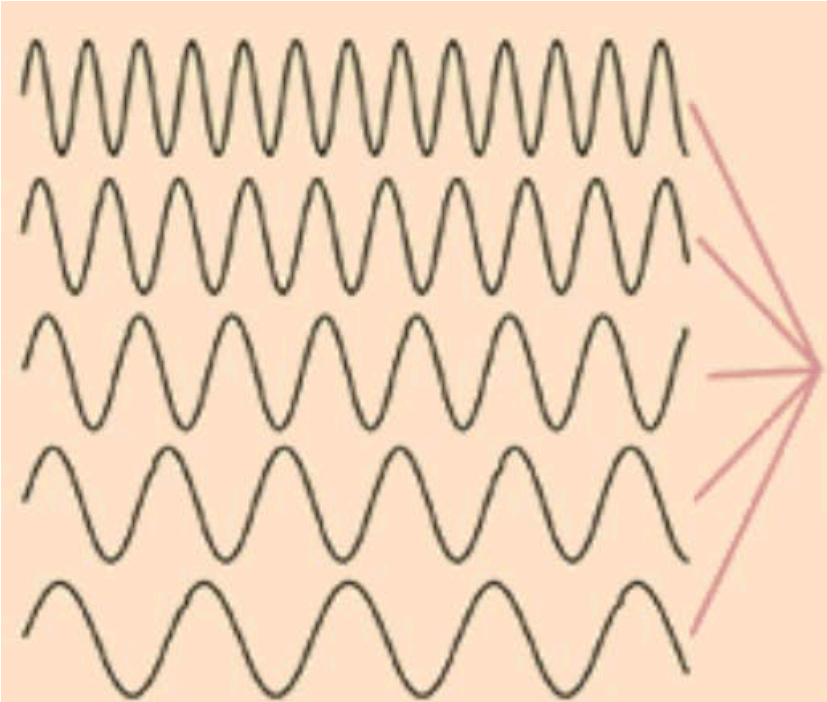
$$\lambda_1 \rightarrow p_1 \Rightarrow$$

$$\lambda_2 \rightarrow p_2 \Rightarrow$$

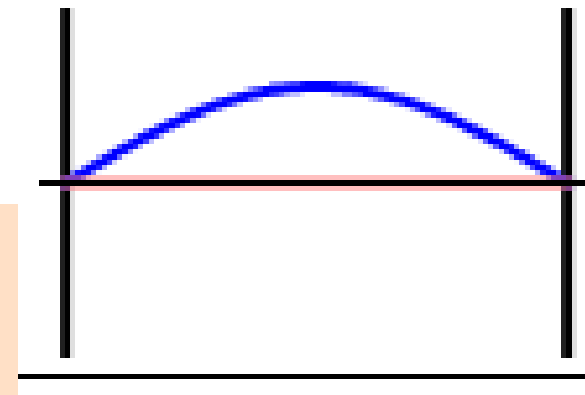
$$\lambda_3 \rightarrow p_3 \Rightarrow$$

$$\lambda_4 \rightarrow p_4 \Rightarrow$$

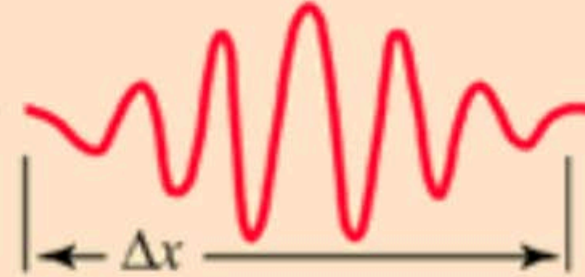
$$\lambda_4 \rightarrow p_4 \Rightarrow$$



Each different wavelength represents a different value of momentum according to the DeBroglie relationship.



Localization of Wave



Position uncertainty

$$\langle \Delta x \rangle$$

$$\Rightarrow \langle \Delta p_x \rangle$$

\rightarrow Mom. uncertainty

Superposition of different wavelengths is necessary to localize the position. A wider spread of wavelengths contributes to a smaller Δx .

Heisenberg's Uncertainty Relations

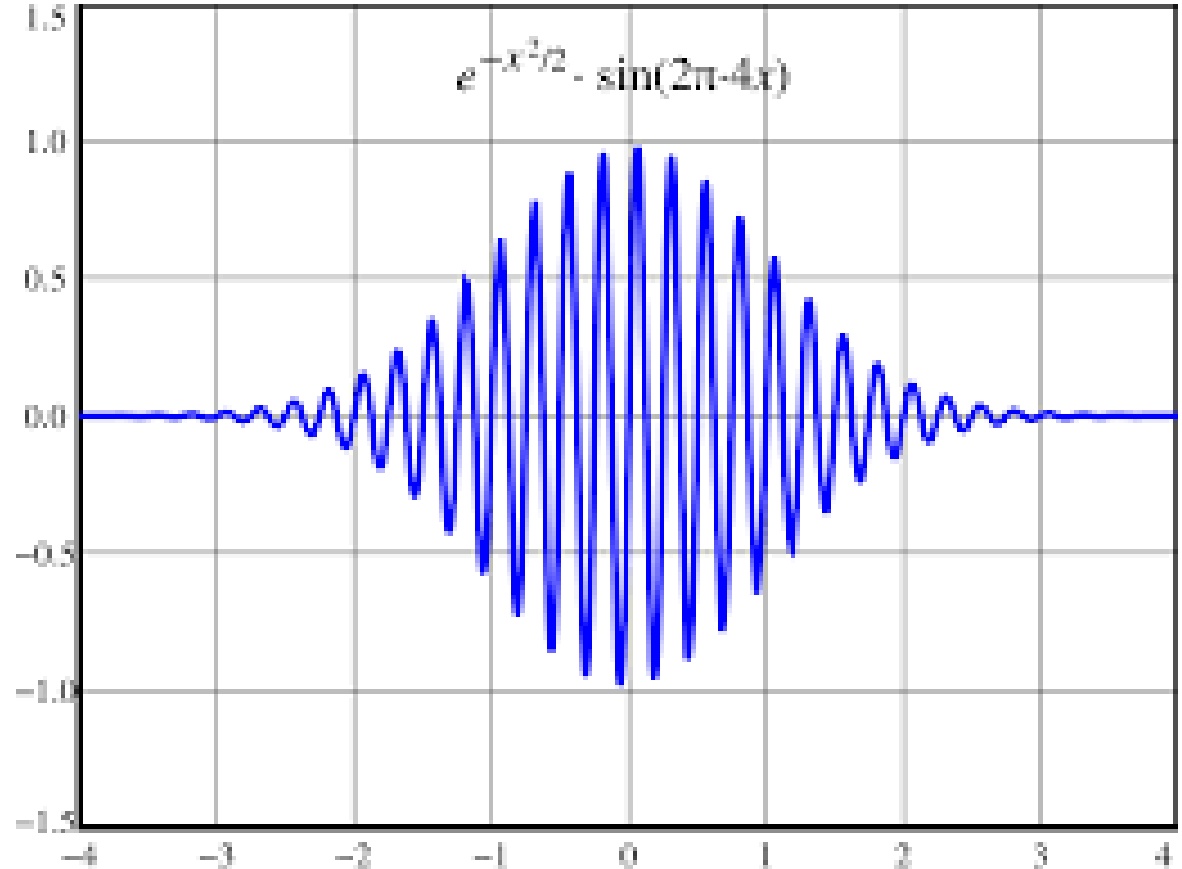
Consider, Gaussian wave packets: $\langle A \rangle = \int_{-\infty}^{\infty} y^*(x) A(x) y(x) dx$

$$\langle \Delta x \rangle \langle \Delta p_x \rangle \geq \frac{\hbar}{2};$$

$$\langle \Delta z \rangle \langle \Delta p_z \rangle \geq \frac{\hbar}{2};$$

$$\langle \Delta y \rangle \langle \Delta p_y \rangle \geq \frac{\hbar}{2};$$

$$\langle \Delta x \rangle \langle \Delta p_x \rangle \sim \frac{\hbar}{2}; \text{ etc}$$



Minimum Uncertainty

Single-slit diffraction: Δx in slit-width, changes Δp_x

This principle is of great importance in understanding many phenomena

Electrons confined in atoms

Max. movement span of electron: Diameter of atom

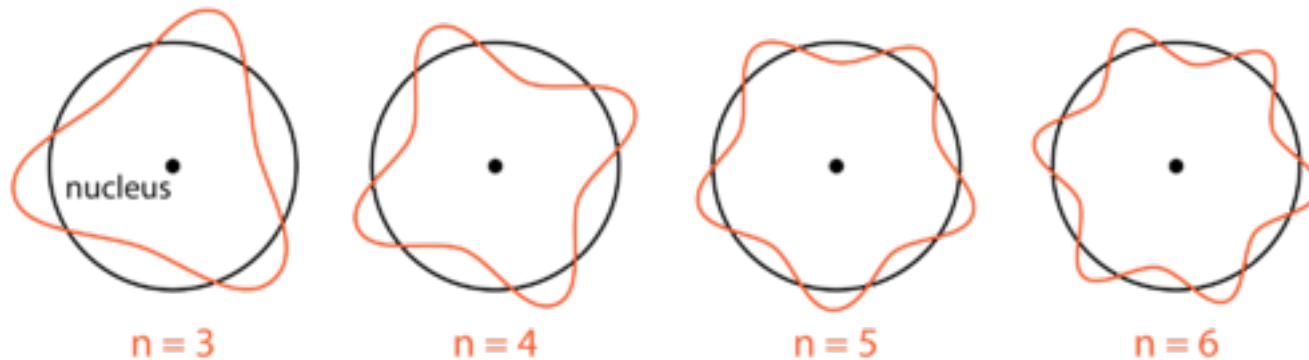
$$x \sim 10^{-10} m \quad \left| \quad \Delta p \sim \hbar / \Delta x \right.$$

$$x \sim \Delta x \quad \left| \quad \Delta p \sim p \right.$$

Electron Cloud inside atom



@<https://mipt.ru/en/>



Standing Wave of electron orbital

@scienceready.com.au/

Kinetic Energy of Electron in Atom

$$E = p^2 / 2m = \hbar^2 / (2m(\Delta x)^2)$$
$$= (10^{-34})^2 / (2 \times 10^{-30} \times (10^{-10})^2)$$

J	eV
----------	-----------

We know, **Hydrogen atom ionisation potential** : 13.6 eV

This is in combination with **Nuclear attraction potential energy**

This principle is of great importance in understanding many phenomena

Hydrogen atom: Ground state

$$E = \frac{\hbar^2}{2m(\Delta r)^2} - \frac{e^2}{4\pi\epsilon_0\Delta r}$$

$$\frac{\partial E}{\partial \Delta r} = -\frac{\hbar^2}{m(\Delta r)^3} + \frac{e^2}{4\pi\epsilon_0(\Delta r)^2} = 0$$

$$\Rightarrow \Delta r = \frac{4\pi\epsilon_0}{me^2} \hbar^2$$

$$\Delta r = \frac{(1.05 \times 10^{-34})^2}{9 \times 10^9 \times 9.1 \times 10^{-31} \times (1.6 \times 10^{-19})^2}$$

$$= 0.528 \times 10^{-10} \text{ m} = 0.528 \text{ \AA} \quad \text{Bohr radius}$$

Minimum energy

$$\Delta r = \frac{4\pi\epsilon_0}{me^2}$$

This principle is of great importance in understanding many phenomena
particles confined in nucleus

Size of the nucleus

$$E = p^2 / 2m = \hbar^2 / (2m(\Delta x)^2)$$

$$= (10^{-34})^2 / (2 \times 4 \times 1.67 \times 10^{-27} \times 4 \times (10^{-15})^2)$$

$$= 2 \times 10^{-13} \text{ J} \sim 1 \text{ MeV}$$

This principle is of great importance in understanding many phenomena

Harmonic oscillator: Ground state

$$E = \frac{p_x^2}{2m} + \frac{1}{2}kx^2$$

$$x \sim \Delta x$$

$$\Delta p_x \sim p_x \sim \hbar/2\Delta x$$

$$E = \frac{\hbar^2}{8m(\Delta x)^2} + \frac{1}{2}k(\Delta x)^2$$

$$\frac{\partial E}{\partial \Delta x} = -\frac{\hbar^2}{4m(\Delta x)^3} + k(\Delta x) = 0$$

$$\Rightarrow (\Delta x)^2 = \frac{\hbar}{2\sqrt{mk}}$$

Zero point energy

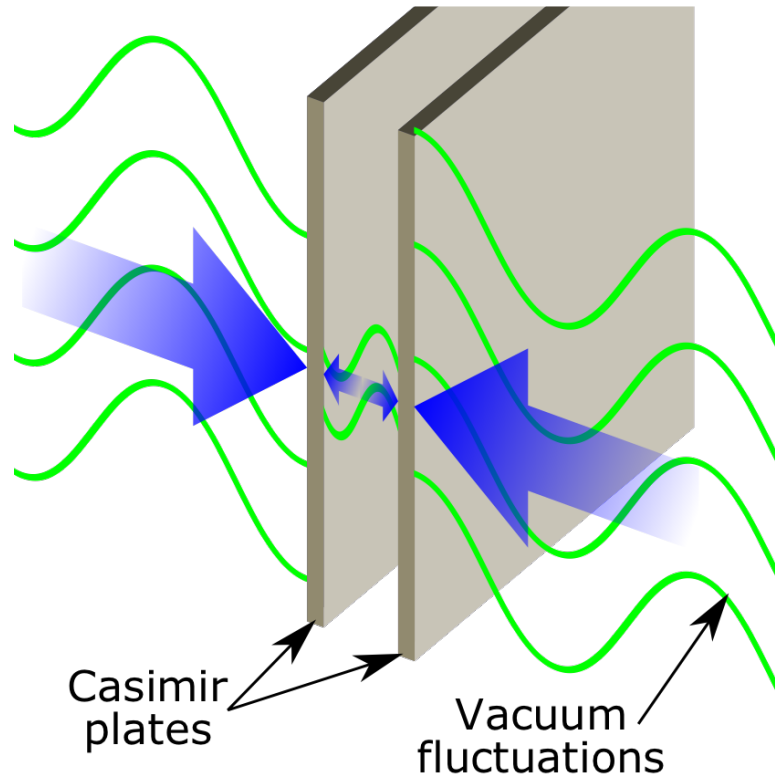
Consequences of Zero Point Energy

- We found **consequence of Heisenberg uncertainty** is the zero point energy
- We had earlier calculated radius of hydrogen atom using the uncertainty relation
- Other way to say, zero point energy decides the span of electron

□ In 1934 Paul Dirac and W. Heisenberg: **Vacuum Fluctuation**

- Vacuum does not mean empty.
 - $(\Delta E)(\Delta t) \sim \hbar \Rightarrow$ Within Δt time, e^- & e^+ can be generated & destroyed in vacuum \Rightarrow Sea of transient charges always present : Energy of Vacuum randomly changes temporarily
- \Rightarrow Consequence, we can not measure bare charge: e^+ cloud around test e^-
- \Rightarrow Experimentally verified theory: Physical charge of e^- matches with theoretically evaluated value up to 10 decimal places

□ In 1948 Hendrik Casimir showed:
one consequence of the zero-point field is an attractive force between two uncharged, perfectly conducting parallel plates, the so-called **Casimir effect.**



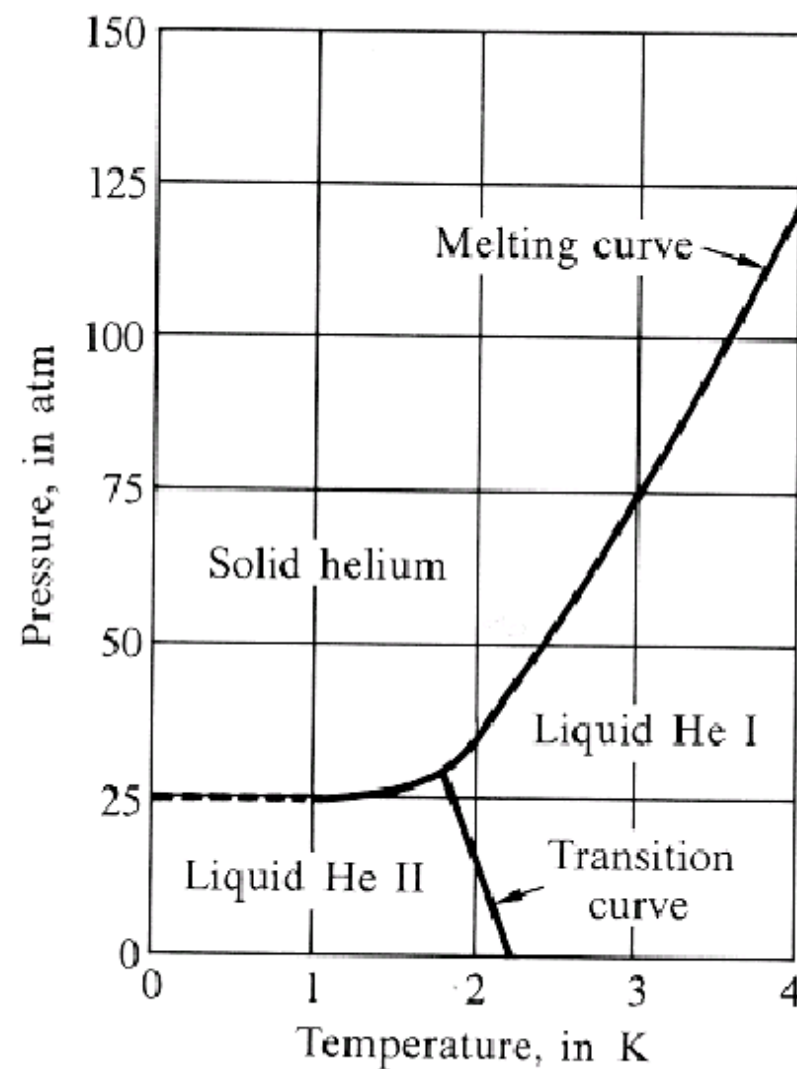
Consequences of Zero Point Energy

Helium atom:

- It remains **liquid** @ absolute zero temperature.
- The large vibrational zero-point energy is responsible for this


For Fermion Gas:

- **Degeneracy Pressure**



Dynamics of Particle in Quantum Mechanics

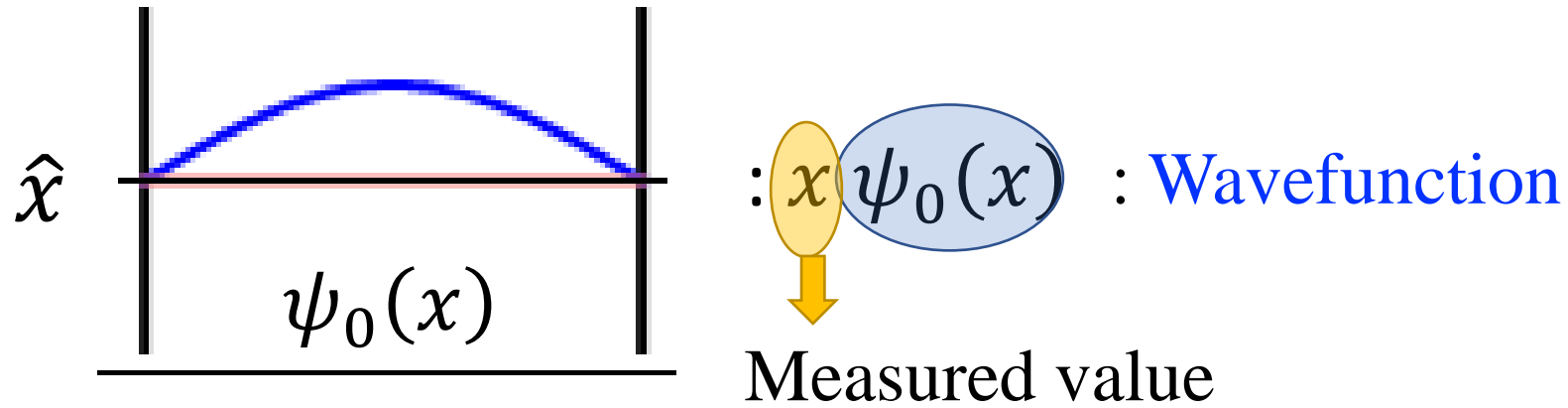
Dynamics of Particle

	Newtonian Mechanics	Lagrangian Mechanics	Hamiltonian Mechanics	Quantum Mechanics
Principle	Force = $m \times accl^n$	Principle of Least Action	Reformulation of Lagrangian Mechanics (in Phase space coordinates (x,p))	Hamilton's Equation of motion
Driving Equation	$m\ddot{x} = F$	Lagrange's Eq. $\frac{d}{dt} \left\{ \frac{\partial L}{\partial \dot{x}} \right\} = \frac{\partial L}{\partial x}$ $L = K.E - P.E.$	$\dot{p} = -\frac{\partial H}{\partial x}, \dot{x} = \frac{\partial H}{\partial p}$ $H = K.E. + P.E.$ 	x and p are not simple dynamical variables. They are self-adjoint Canonical conjugate operators

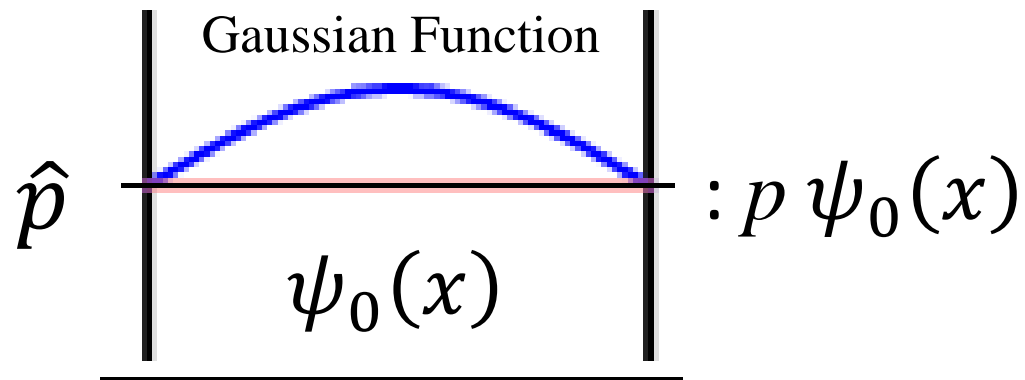

 \hat{x} & \hat{p} : Operate on states representing the system

Example of states representing system: Energy states of particle in box

Position Operator



Momentum Operator: Canonically conjugate operator (Fourier transform dual)



Consider Plane wave

$$\psi(x, t) = e^{\frac{i}{\hbar}(px - Et)}$$

$$\frac{\partial \psi(x, t)}{\partial x} = \frac{ip}{\hbar} e^{\frac{i}{\hbar}(px - Et)} = \frac{ip}{\hbar} \psi.$$

$$\hat{p} = -i\hbar \frac{\partial}{\partial x}$$

Time Evaluation of Quantum System

Momentum operator: $\hat{p}_x \equiv -i\hbar \frac{\partial}{\partial x}$

Total energy operator: $\hat{E} \equiv i\hbar \frac{\partial}{\partial t}$

Now the total energy of a particle is just the sum of the kinetic energy and the potential energy:

$$E = p^2/2m + V$$

In “operator” form this becomes:

$$i\hbar \frac{\partial \psi}{\partial t} = \frac{1}{2m} (-i\hbar)^2 \frac{\partial^2 \psi}{\partial x^2} + V\psi$$

Rearranging we get
Schrödinger's equation:
(one dimensional version)

$$\boxed{-\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V\psi = i\hbar \frac{\partial \psi}{\partial t}}$$

Schrödinger's equation

Things to note from Schrödinger's eqn.

$$-\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V\psi = i\hbar \frac{\partial \psi}{\partial t}$$

- (i) The equation is linear in ψ that is there are no terms like ψ^2 or $\left(\frac{\partial \psi}{\partial x}\right)^2$.
- (ii) The equation is homogeneous, that is there are no terms independent of ψ

Taken together these features mean that if ψ is a solution to the equation, then so is $c\psi$ where c is any complex number.

This implies that any linear combination of solutions is also a solution

Differences between Schrödinger and classical wave equations

$$-\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V\psi = i\hbar \frac{\partial \psi}{\partial t}$$

K.E. + P.E. = Total Energy

$$\frac{\partial^2 y}{\partial x^2} = \frac{1}{v^2} \frac{\partial^2 y}{\partial t^2}$$

Schrödinger Wave Equation	Classical Wave Equation
Eqn. derived for total energy	Eqn. derived from force
First order derivative w.r.t. time	Second order derivative w.r.t time
Complex Equation	Real Equation
→ ψ must be complex	→ y is a displacement : Real

Interpretation of the wavefunction

What does $\psi(x, t)$ tell us?

c.f. Young's slit experiment: high probability of detecting particle at bright fringes, low probability at dark fringes

- for light expect probability of detecting a photon to be proportional to the *intensity* of E.M. wave.

Phase of $\psi(x, t)$ is not observable.

The probability of finding a particle in the range x to $x+dx$ at time t is proportional to $|\psi(x, t)|^2 dx$

Define: $P(x, t) dx \propto |\psi(x, t)|^2 dx$

Need to normalise probabilities:

$$\int_{-\infty}^{\infty} P(x) dx = 1$$

(**one particle!**)

Normalisation of wavefunction

$$P(x,t) dx. \propto |\psi(x,t)|^2 dx$$

To get rid of proportional sign we need a constant of proportionality: $P(x,t) dx. = A |\psi(x,t)|^2 dx.$

Eliminate A by using the normalisation: $\int_{-\infty}^{\infty} P(x) dx = 1$

$$A \int_{-\infty}^{\infty} |\psi(x,t)|^2 dx = 1$$

It follows that:
$$A = \frac{1}{\int_{-\infty}^{\infty} |\psi(x,t)|^2 dx}$$

If we normalise the wavefunction so that

$$\int_{-\infty}^{\infty} |\psi(x,t)|^2 dx = 1$$

we can define

$$P(x,t) = |\psi(x,t)|^2$$

Time independent Schrödinger equation

Time-dependent, Schrödinger equation:
$$-\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V(x) \psi = i \hbar \frac{\partial \psi}{\partial t}$$

Suppose time-independent potential, then we expect the total energy to be a constant.

Example: $V(x, t) \sim$ Potential due to oscillating dipole $\rightarrow V(x) \sim$ due to constant dipole

“Separation of variables”: $\psi(x, t) = u(x) f(t)$

Insert into Schrödinger equation:

$$-\frac{\hbar^2}{2m} f(t) \frac{\partial^2 u(x)}{\partial x^2} + V(x) u(x) f(t) = i \hbar u(x) \frac{\partial f(t)}{\partial t}$$

Divide through by $u(x) f(t)$:

$$-\frac{\hbar^2}{2m} \frac{1}{u(x)} \frac{\partial^2 u(x)}{\partial x^2} + V(x) = i \hbar \frac{1}{f(t)} \frac{\partial f(t)}{\partial t}$$

LHS

RHS

Only depends on x

Only depends on t

Time independent Schrödinger equation

$$-\frac{\hbar^2}{2m} \frac{1}{u(x)} \frac{\partial^2 u(x)}{\partial x^2} + V(x) = i\hbar \frac{1}{f(t)} \frac{\partial f(t)}{\partial t}$$

- The only way a function of x = a function of t , $\forall x, t \Rightarrow$ L.H.S.=R.H.S. constant.
- Call the constant E (we will show later this is the total energy).

RHS becomes:

$$i\hbar \frac{\partial f(t)}{\partial t} = E f(t)$$

Integrating: $\int \frac{\partial f}{f} = \frac{E}{i\hbar} \int \partial t \longrightarrow \boxed{f(t) = f(0) \exp(-i E t / \hbar)}$

LHS becomes:

$$\boxed{-\frac{\hbar^2}{2m} \frac{\partial^2 u(x)}{\partial x^2} + V(x) u(x) = E u(x)}$$

This is called **the time-independent Schrödinger equation**

ψ has a solution of the form: $\psi(x, t) = u(x) \exp(-i E t / \hbar)$

Probability density is
independent of time:

$$|\psi(x, t)|^2 = |u(x)|^2$$

'Solving' the Schrödinger equation:

What do we want? Usually, the allowed energy levels.

First: Assume a solution - e.g. $u(x) = A \sin(kx)$ (TISE)

Second: Substitute solution into Schrödinger equation.

That gives relationship between E (and V) and k

Third: Use boundary conditions (e.g., u and du/dx continuous)
to solve for allowed values of k (e.g. in terms of well size, a)

Fourth: Use earlier relationship between E (and V) and k to
obtain allowed values of E

Example of solution of T.I.S.E:
Particle in a box: infinite potential well

$$-\frac{\hbar^2}{2m} \frac{\partial^2 u(x)}{\partial x^2} + V(x)u(x) = E u(x)$$

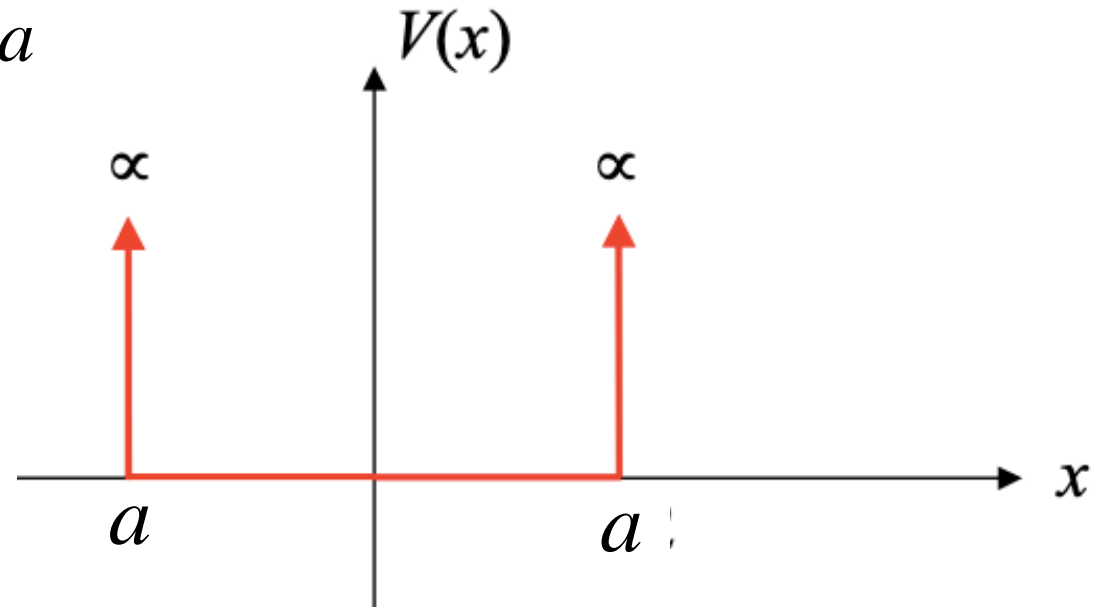
Box in one dimension with walls at $-a$ and $+a$

$$V(x) = 0 \quad \text{for } |x| \leq a,$$

$$V(x) = \infty \quad \text{for } |x| > a$$

For $|x| \leq a$, T.I.S.E becomes:

$$-\frac{\hbar^2}{2m} \frac{\partial^2 u(x)}{\partial x^2} = E u(x)$$



Boundary condition: $u(x)$ must vanish for $|x| > a$,



Types of solution (I)

$$-\frac{\hbar^2}{2m} \frac{\partial^2 u(x)}{\partial x^2} = E u(x)$$

A possible solution is: $u(x) = A \sin(kx)$

Check:

$$\frac{\partial u}{\partial x} = Ak \cos(kx)$$

$$\frac{\partial^2 u}{\partial x^2} = -Ak^2 \sin(kx)$$

Insert into T.I.S.E.:
$$\frac{\hbar^2 k^2}{2m} A \sin(kx) = E A \sin(kx)$$

Therefore it is a solution as long as:
$$E = \frac{\hbar^2 k^2}{2m}$$

Boundary condition: $u(x) = 0$ for $|x| > a$

implies $\sin(kx) = 0$ for $|x| = a$. True if $ka = m\pi$ (m integer)

Types of solution (II)

$$-\frac{\hbar^2}{2m} \frac{\partial^2 u(x)}{\partial x^2} = E u(x)$$

Another type of solution is: $u(x) = B \cos(kx)$

Check: $\frac{\partial u}{\partial x} = -Bk \sin(kx)$

$$\frac{\partial^2 u}{\partial x^2} = -Bk^2 \cos(kx)$$

Insert into T.I.S.E.: $\frac{\hbar^2 k^2}{2m} B \cos(kx) = E B \cos(kx)$

Therefore it is a solution as long as: $E = \frac{\hbar^2 k^2}{2m}$

Boundary condition: $u(x) = 0$ for $|x| > a$ implies

$\cos(kx) = 0$ for $|x| = a$. **True if** $ka = m\pi/2$ (m odd integer)

Summary of solutions:

$$u(x) = A \sin(kx)$$

$$u(x) = B \cos(kx)$$

$$ka = m\pi \quad (m \text{ integer})$$

$$ka = n\pi/2 \quad (n \text{ odd integer})$$

Equivalently: $ka = n\pi/2$

(n even integer)

$$n = 2, 4, 6, 8, \dots$$

$$k = \frac{n\pi}{2a}$$

$$n = 1, 3, 5, 7, \dots$$

$$E = \frac{\hbar^2 k^2}{2m} = \frac{\hbar^2 \pi^2 n^2}{8ma^2}$$

Energy levels for a particle in a box

$$E = \frac{\hbar^2 k^2}{2m} = \frac{\hbar^2 \pi^2 n^2}{8ma^2} \quad n = 1, 2, 3, 4, 5, 6, 7, 8, \dots \quad \square$$

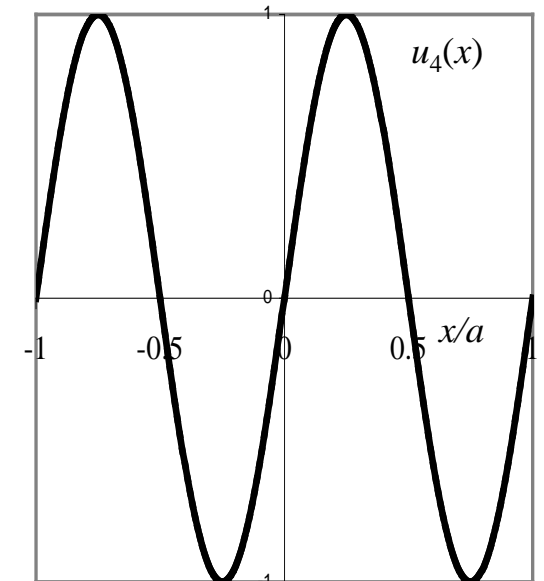
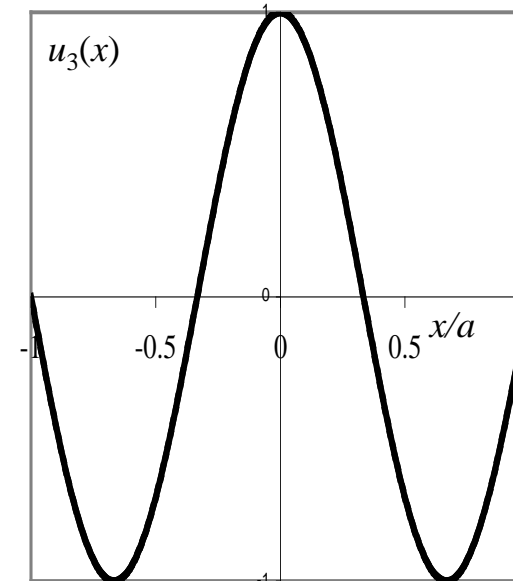
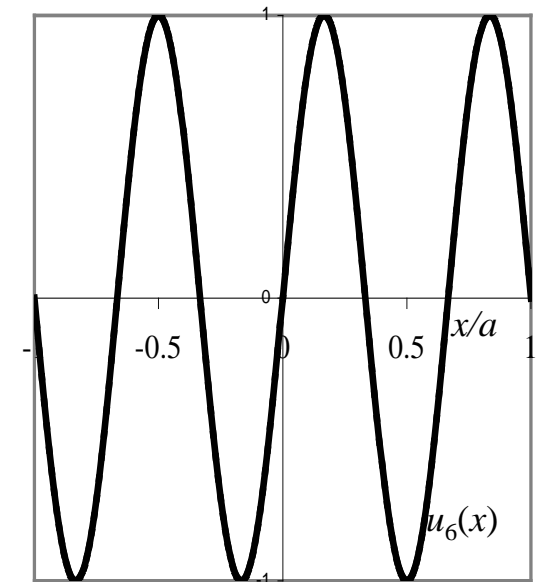
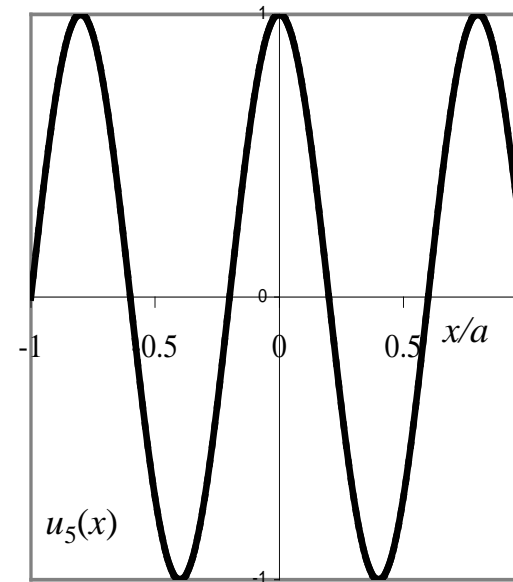
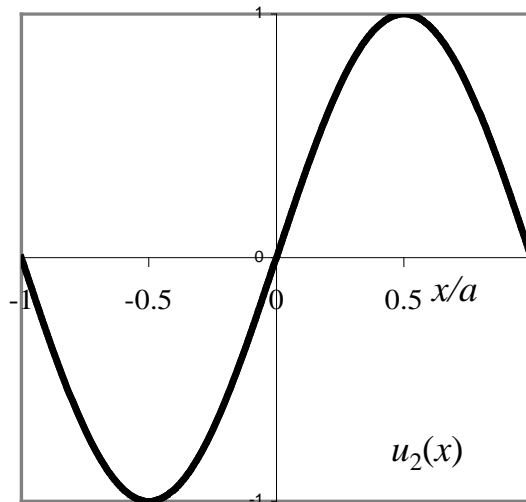
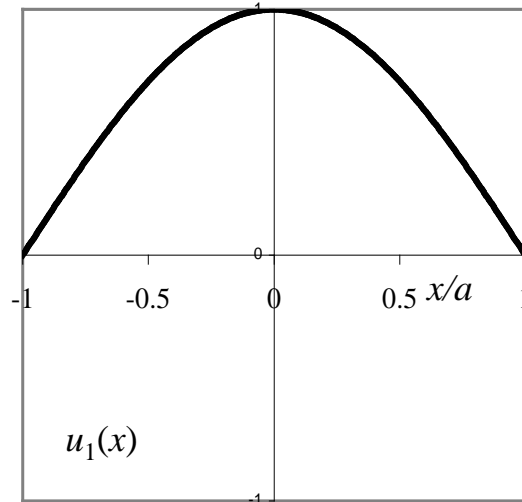
Solutions for energy are called energy eigenvalues



Wavefunctions for particle in a box

$$u(x) = B \cos(n\pi x/2a) \quad u(x) = A \sin(n\pi x/2a)$$

$$n = 1, 3, 5, 7, \dots \quad n = 2, 4, 6, 8, \dots$$



Solutions for wavefunctions are called **eigenfunctions**

Normalisation

$u(x) = A \sin(n\pi x/2a)$ or $B \cos(n\pi x/2a)$ Values of A, B ?

Normalisation condition: $\int_{-\infty}^{\infty} |\psi(x,t)|^2 dx = 1$ \leftarrow As we know one particle is present @ $x \in [-\infty, \infty]$

In the present case we only need to integrate between $-a$ and $+a$ since $u(x)$ vanishes outside this range:

$$\int_{-a}^a |\psi(x,t)|^2 dx = \int_{-a}^a A^2 \sin^2\left(\frac{n\pi x}{2a}\right) dx = \frac{A^2}{2} [x]_{-a}^a = A^2 a = 1$$

➤ It follows that: $A = \frac{1}{\sqrt{a}}$.

➤ Similarly it can be shown that $B = \frac{1}{\sqrt{a}}$

Takes care time evolution of wavefunction; making Sure # of particle is not changing over time

Normalised eigenfunctions:

$$\psi_n(x,t) = u_n(x) \exp(-i E_n t / \hbar) = \frac{1}{\sqrt{a}} \cos\left(\frac{n\pi x}{2a}\right) \exp\left(-\frac{i E_n t}{\hbar}\right)$$

$n = 1, 3, 5, \dots$

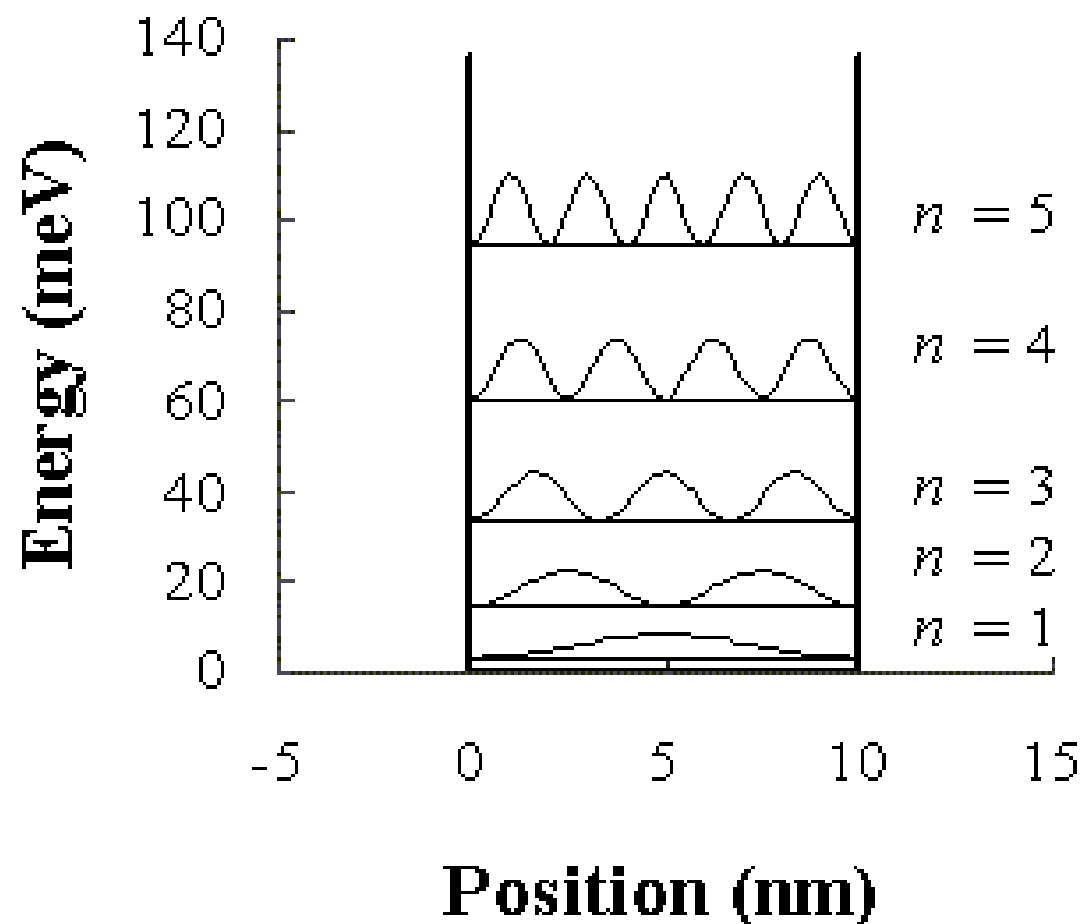
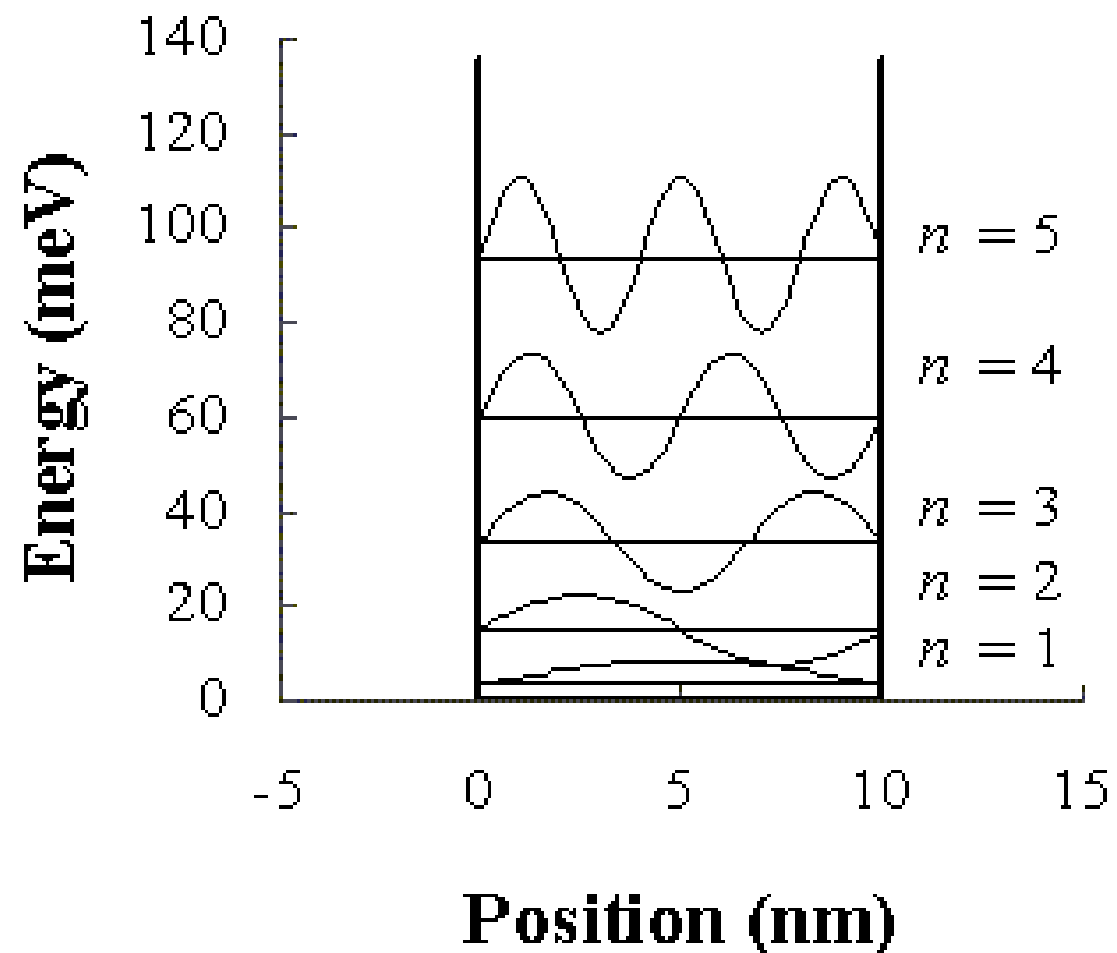
$n = 2, 4, 6, \dots$

$$\psi_n(x,t) = \frac{1}{\sqrt{a}} \sin\left(\frac{n\pi x}{2a}\right) \exp\left(-\frac{i E_n t}{\hbar}\right)$$



Probability density $P(x, t) = |\psi(x, t)|^2$

For first five eigenfunctions for particle in a box of width 10 nm





Zero point energy

The lowest energy state for a particle in a box is $E_1 = \frac{\hbar^2 \pi^2}{8ma^2}$

Classically, Lowest possible energy should be zero.

➤ Why can't the energy be zero?

Heisenberg uncertainty relation $\Delta p \Delta x \approx \hbar$

Particle is confined in box, so $\Delta x \sim a$.

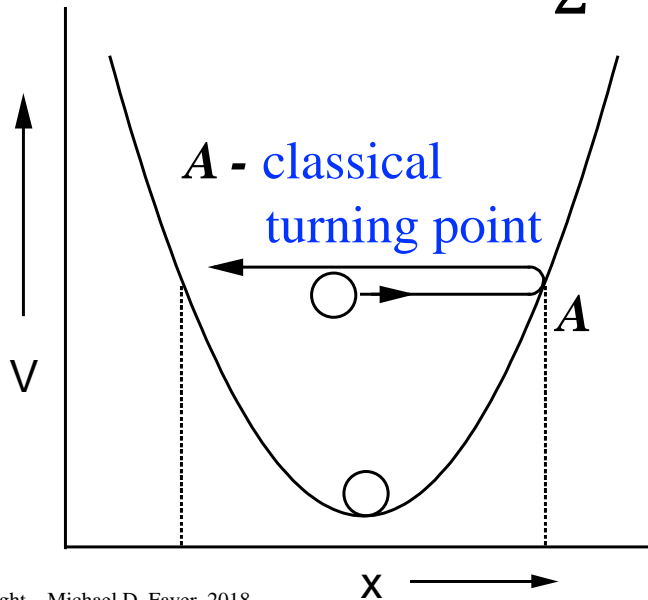
Must be an uncertainty in momentum $\Delta p \approx \frac{\hbar}{a}$

Since momentum cannot be zero (static), minimum energy must be of order

$$E_{\min} \approx \frac{(\Delta p)^2}{2m} \approx \frac{\hbar^2}{2ma^2} \quad (\text{Qualitative Agreement with } E_1!!)$$

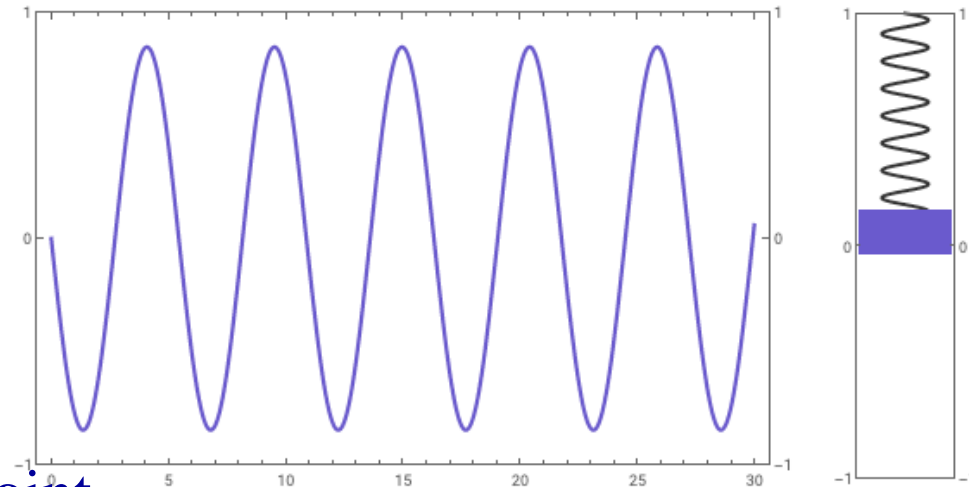
Particle in Harmonic Oscillator: Quantum Mechanics point of View

Potential : $V(x) = \frac{1}{2}kx^2$ Classical Mechanics: $m\ddot{x} = -kx$



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Real Solution: $x(t) = A \cos kx$ with energy $E = \frac{1}{2}kA^2$
A can take on any value \rightarrow Energy is continuous.



➤ Classical particle can never be past turning point.

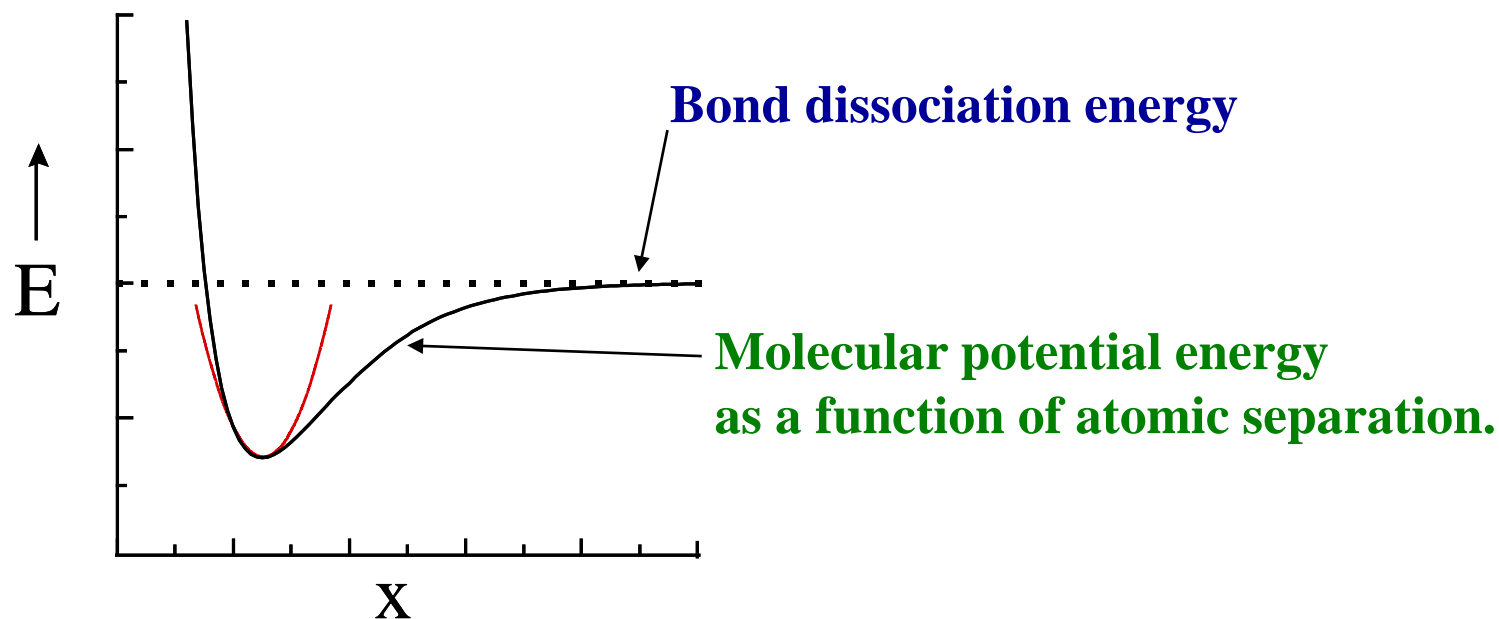
➤ Turning point K.E. zero; potential energy max.

➤ Particle can be stationary at bottom of well,
know position, $x = 0$; know momentum, $p = 0$.

$$\therefore \Delta x \Delta p = 0$$

Example: Quantum Harmonic Oscillator

Simplest model of molecular vibrations



- **Bonds between atoms act as "springs".**
- **Near bottom of molecular potential well, molecular potential approximately parabolic**
Harmonic Oscillator.

One Dimensional Quantum Harmonic Oscillator in the Schrödinger Representation

$$H|\psi\rangle = E|\psi\rangle \quad \text{Where, } H = -\frac{\hbar^2}{2m} \frac{d^2}{dx^2} + \frac{1}{2} kx^2$$

$$\triangleright \frac{d^2\psi(x)}{dx^2} + \frac{2m}{\hbar^2} [E - (2\pi^2 m \nu^2) x^2] \psi(x) = 0$$

Define

$$\alpha = 2\pi m \nu / \hbar$$

$$\lambda = \frac{2mE}{\hbar^2}$$

$$\frac{d^2\psi(x)}{dx^2} + (\lambda - \alpha^2 x^2) \psi(x) = 0$$

Find $\psi(x)$!!

Case: 1 $\alpha^2 x^2 \gg \lambda$ Therefore, λ can be dropped.

$$\frac{d^2\psi}{dx^2} = \alpha^2 x^2 \psi$$

$$\frac{d^2\psi}{dx^2} = \alpha^2 x^2 e^{\pm \frac{\alpha}{2} x^2} \pm \alpha e^{\pm \frac{\alpha}{2} x^2}$$

This is negligible compared to the 1st term as $x \rightarrow \infty$.

Must obey Born Conditions

- finite everywhere
- single valued
- Continuous
- first derivative continuous

Two solutions

$$e^{-\frac{\alpha}{2}x^2}$$

This is O.K. at
 $x = \pm\infty$

Therefore, large x solution is

$$\psi(x) = e^{-\frac{\alpha}{2}x^2}$$



***How to derive $f(x)$?

$$e^{+\frac{\alpha}{2}x^2}$$

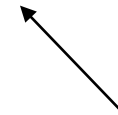
*Remember $\alpha = \frac{2\pi m v}{\hbar} \geq 0$

This blows up at $x = \pm\infty$

Not finite everywhere.

For all x

$$\psi(x) = e^{-\frac{\alpha}{2}x^2} f(x)$$



Must find this.

$$\boxed{\frac{d^2 \psi(x)}{d x^2} + (\lambda - \alpha^2 x^2) \psi(x) = 0} \longrightarrow \psi(x) = e^{-\frac{\alpha}{2} x^2} f(x)$$

Need second derivative in Schrödinger equation

$$\frac{d^2 \psi(x)}{d x^2} = e^{-\frac{\alpha}{2} x^2} (\alpha^2 x^2 f - \alpha f - 2\alpha x f' + f'')$$

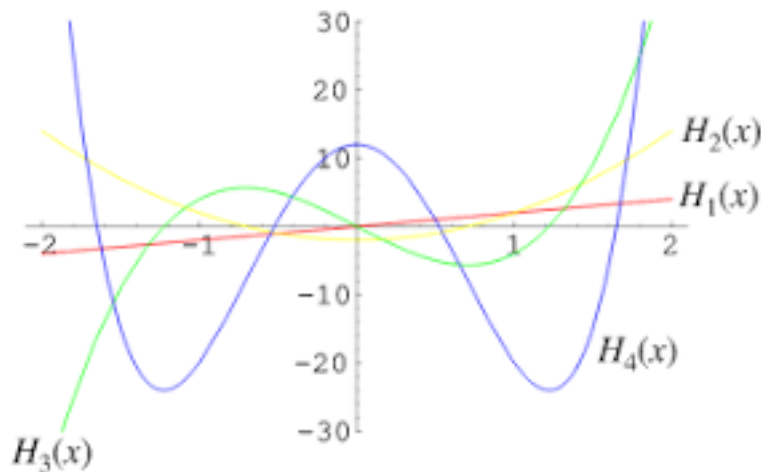
With $f' = \frac{d f}{d x}$ and $f'' = \frac{d^2 f}{d x^2}$

Substitute $\frac{d^2 \psi(x)}{d x^2}$ and $\psi(x)$ **into the original equation**

and divide by $e^{-\frac{\alpha}{2} x^2}$ gives $f'' - 2\alpha x f' + (\lambda - \alpha) f = 0$

- Equation only in f .
- Solve for f and have $\psi(x)$

Solution will be Hermite polynomial



Quantization of Energy

If there are a finite number of terms in the series for $\underline{H}(\gamma)$, wavefunction does not blow up. Goes to zero at infinity.

$e^{-\gamma^2/2} \gamma^n$ The exponential goes to zero faster than γ^n blows up.

To make series finite, truncate by choice of $\lambda \rightarrow$ One can show $\lambda = (2n + 1)\alpha, n = \text{integer}$

Therefore, $\lambda = \frac{2mE}{\hbar^2} = (2n + 1) 2\pi m v / \hbar$
definition of λ definition of α

$$E_n = \left(n + \frac{1}{2} \right) h\nu \quad n \text{ is the quantum number}$$

$n = 0$ $E_0 = 1/2 h\nu$ Lowest energy, not zero. Called zero point energy.

Energy levels equally spaced by $h\nu$.

Energy Levels

Correspond to each n : $E_n = \left(n + \frac{1}{2}\right) h\nu$

Wavefunctions

$$\psi_n(x) = N_n e^{-\frac{\gamma^2}{2}} H_n(\gamma)$$

$$\gamma = \sqrt{\alpha} x \quad \alpha = 2\pi m\nu / \hbar$$

$$N_n = \left\{ \left(\frac{\alpha}{\pi} \right)^{\frac{1}{2}} \frac{1}{2^n n!} \right\}^{\frac{1}{2}} \quad \text{normalization constant}$$

Ground state ($n = 0$) $\psi_0(x) = \left(\frac{\alpha}{\pi} \right)^{\frac{1}{4}} e^{-\frac{\alpha}{2} x^2} = \left(\frac{\alpha}{\pi} \right)^{\frac{1}{4}} e^{-\frac{\gamma^2}{2}}$

This is a Gaussian.
Minimum uncertainty.

Hermite Polynomials

$$H_0(\gamma) = 1\gamma^0$$

$$H_1(\gamma) = 2\gamma$$

$$H_2(\gamma) = 4\gamma^2 - 2\gamma^0$$

$$H_3(\gamma) = 8\gamma^3 - 12\gamma$$

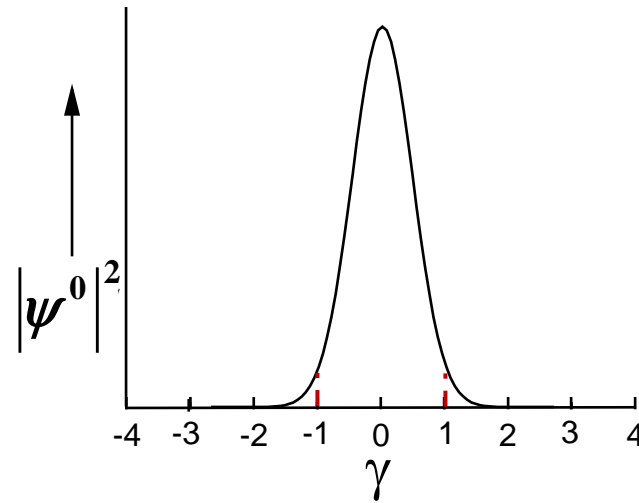
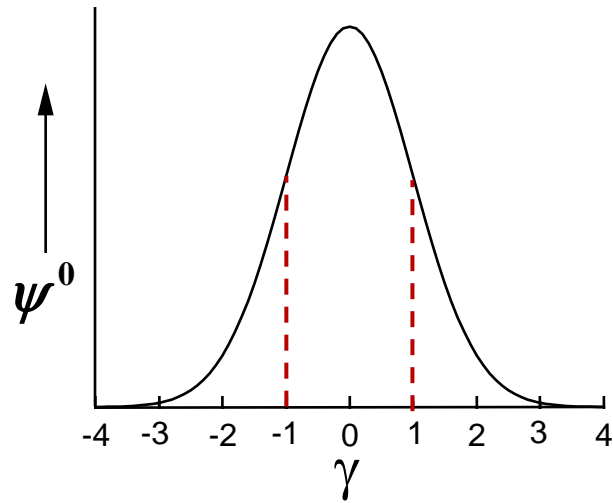
$$H_4(\gamma) = 16\gamma^4 - 48\gamma^2 + 12\gamma^0$$

$$H_5(\gamma) = 32\gamma^5 - 160\gamma^3 + 120\gamma$$

$$H_6(\gamma) = 64\gamma^6 - 480\gamma^4 + 720\gamma^2 - 120\gamma^0$$

**** Given an oscillator $0 \leq \alpha \rightarrow$ Large Number**

Ground State of Quantum Harmonic Oscillator



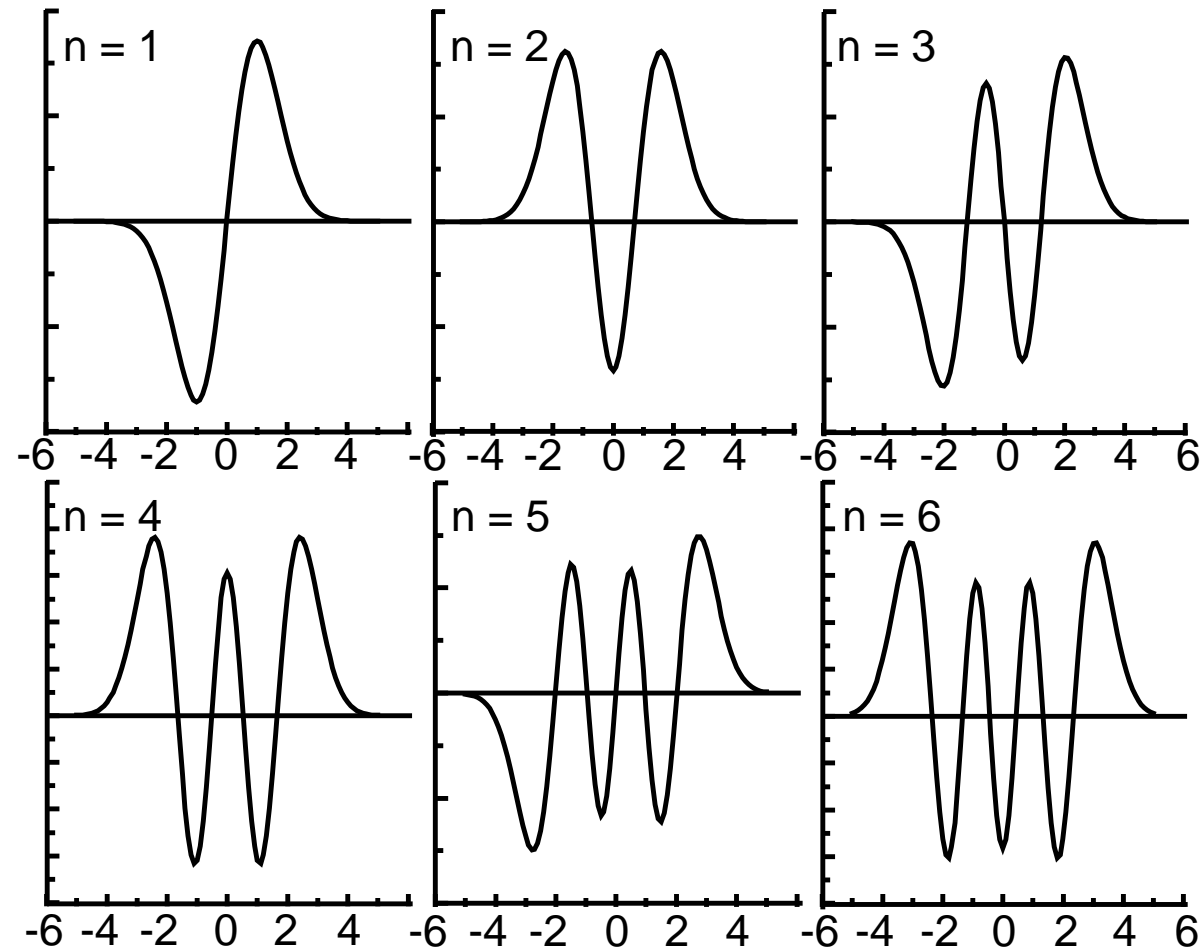
Classical turning points

$$\begin{array}{ccc} \uparrow & & \uparrow \\ 1/2 kx^2 & = & 1/2 h\nu \\ \text{potential} & & \text{total} \\ \text{energy} & & \text{energy} \end{array}$$

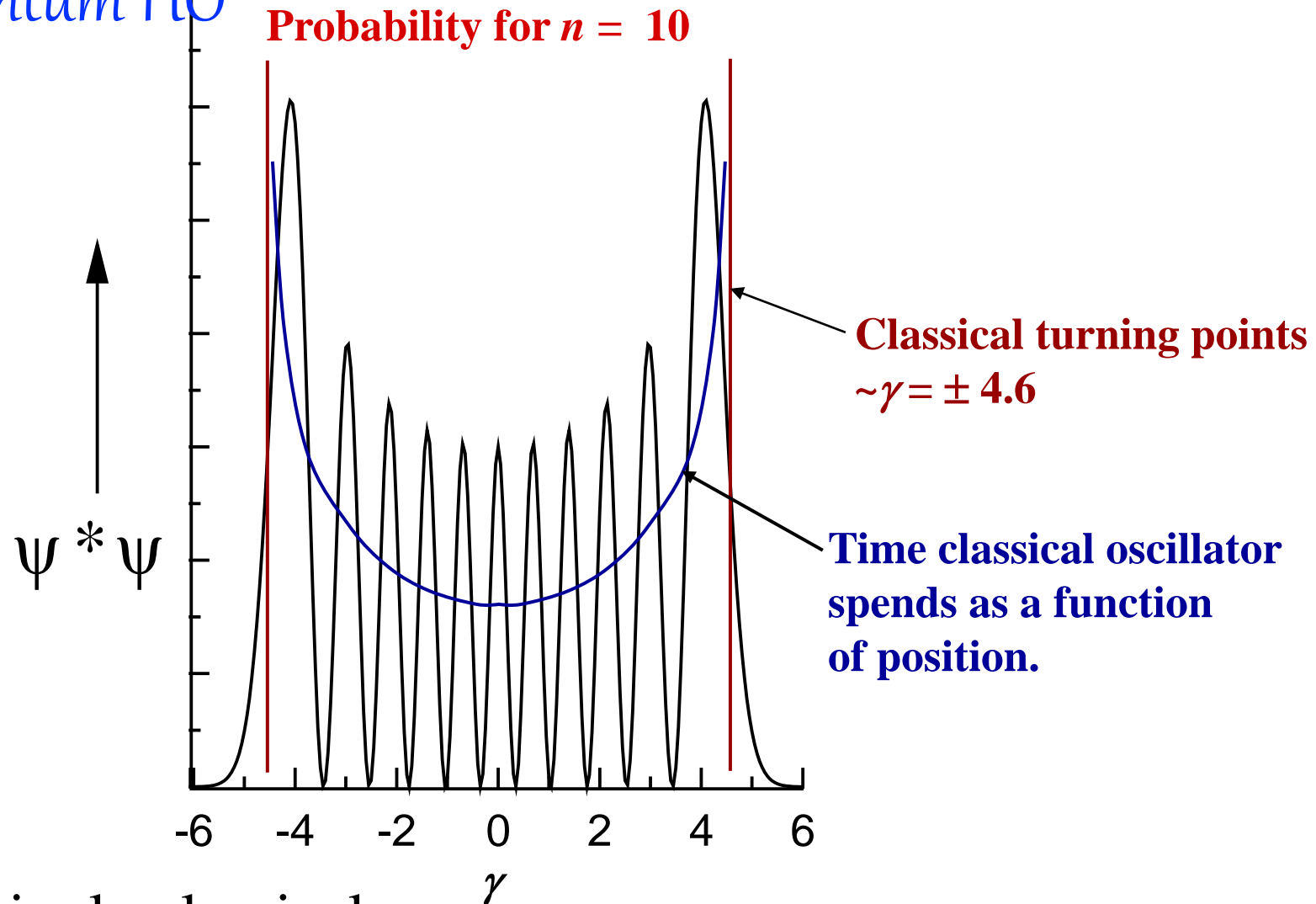
$$\begin{aligned} \therefore x^2 &= \frac{h\nu}{k} \\ x &= \pm \sqrt{h\nu / k} = \pm \gamma \end{aligned}$$

classical turning points - wavefunction extends into classically forbidden region.

More wavefunctions - larger n , more nodes



Comparison of Quantum HO With Classical HO



Looks increasingly classical.

For **large object**, nodes so closely spaced because n very large that can't detect nodes.