Integrability and action-angle variables of post-Newtonian binary black holes

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Introduction and theory

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- 1.5PN: action-angles

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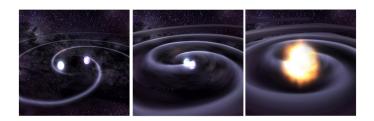
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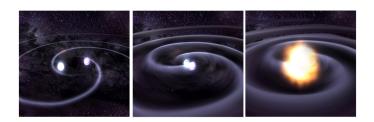
Gravitational waves from binary black holes



 Stellar mass BBHs: gravitational wave (GW) sources of LIGO and LISA.

Image credit: www.eoportal.org

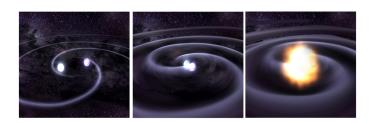
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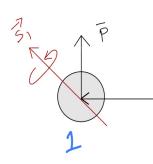
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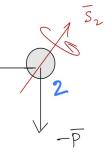
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Starting point: 2PN Hamiltonian due to [Barker, O'Connell-1975]





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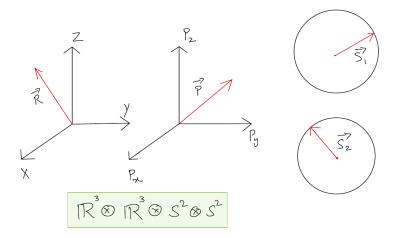
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$$\begin{split} H = & \left(\frac{P^2}{2\mu} - \frac{Gm_1m_2}{R} \right) + \frac{1}{c^2} F_1(R, \vec{P}) + \frac{1}{c^4} F_2(R, \vec{P}) \\ & + \frac{1}{c^2} F_3 \left(\vec{S_1}.\vec{L}, \ \vec{S_2}.\vec{L} \right) + \frac{1}{c^4} F_4 \left(\vec{S_1}.\vec{n}, \ \vec{S_2}.\vec{n}, \ \vec{S_1}.\vec{S_2} \right). \end{split}$$

• With $m=m_1+m_2, \ \mu:=m_1m_2/m$ and $\vec{n}:=\vec{R}/R$, the 2PN Hamiltonian becomes [Barker, O'Connell-1975]

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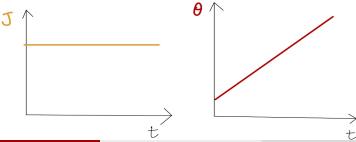
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- $\mathcal{J}_1 = L$, $\mathcal{J}_2 = J$, $\mathcal{J}_3 = J_z$.
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• \mathcal{J}_5 is very lengthy.

$$\begin{split} & \text{J} = \text{Jd} \text{J}^{1/2}; \ L = \text{Ld} \text{L}^{1/2}; \ S1 = \text{S1dS1}^{1/2}; \ S2 = \text{S2dS2}^{1/2}; \\ & \text{A1} = \frac{\left(\frac{(3^2-L^2-S1^2-S2^2)}{2} - \frac{\text{serfel}}{\sigma^2}\right)}{\sigma^1 - \sigma^2}; \\ & \text{A2} = \frac{\left(\frac{(3^2-L^2-S1^2-S2^2)}{2} - \frac{\text{serfel}}{\sigma^1}\right)}{\sigma^1 - \sigma^2}; \\ & \text{A2} = \frac{\left(\frac{(3^2-L^2-S1^2-S2^2)}{2} - \frac{\text{serfel}}{\sigma^1}\right)}{\sigma^1 - \sigma^2}; \\ & \sigma^2 = 1 + \frac{9}{16 \times (-1 + \sigma 1)}; \\ & \sigma^1 = \left(1 + \frac{3}{4} \frac{m^2}{m^2}\right); \quad \sigma^2 = \left(1 + \frac{3}{4} \frac{m^1}{m^2}\right); \\ & \text{Seff} = \delta 1 \text{S1} + \delta 2 \text{S2} \\ & \text{cubic} = L^2 \text{S1}^2 \text{S2}^2 - 2 \sigma 1 \sigma 2 (\sigma 1 - \sigma 2) f (f - \Delta 1) (f - \Delta 2) - \\ & L^2 (\sigma 1 - \sigma 2)^2 f^2 - \text{S2}^2 \sigma^2 (f - \Delta 1)^2 - \text{S1}^2 \sigma^2 (f - \Delta 2)^2; \\ & \text{cubic} = \text{Collect}[\text{cubic} \ \# \text{Expand} \ \# \text{Simplify} \ , \ f]; \\ & \text{cubSol} = ((\text{Solve}[\text{cubic} \ \# \text{Expand} \ \# \text{Simplify} \ , \ f]; \\ & \text{cubSol} = ((\text{Solve}[\text{cubic} \ \# \text{e} \theta, f))[1]); \\ & \text{cubSol} = ((\text{Solve}[\text{cubic} \ \# \text{e} \theta, f))[2]); \\ & \text{cubSol} = ((\text{Solve}[\text{cubic} \ \# \text{e} \theta, f))[3]); \\ & \text{f3} = f \ / \ \text{cubSol1}; \\ & \text{f1} = f \ / \ \text{cubSol2}; \\ & \text{f2} = f \ / \ \text{cubSol3}; \\ & \text{f3} = (\text{f2} - \sigma 1) / (f 3 - f 1))^{1/2}; \\ & \text{f1} = \frac{1}{2} \left((\text{SeffdL} + L^2 (\sigma 1 + \sigma 2)) (3 + L) + L (\sigma 1 \text{S1}^2 + \sigma 2 \text{S2}^2 + (\sigma 1 + \sigma 2) (\Delta 2 \sigma 1 - \Delta 1 \sigma 2))); \\ & \text{f2} = \frac{1}{2} \left((\text{SeffdL} + L^2 (\sigma 1 + \sigma 2)) (3 - L) - L (\sigma 1 \text{S1}^2 + \sigma 2 \text{S2}^2 + (\sigma 1 + \sigma 2) (\Delta 2 \sigma 1 - \Delta 1 \sigma 2))); \\ & \text{f1} = L (L - 3) + \Delta 2 \sigma 1 - \Delta 1 \sigma 2; \\ & \text{f2} = L (L - 3) + \Delta 2 \sigma 1 - \Delta 1 \sigma 2; \\ & \text{f2} = \left(-L^2 \text{S1} \ \sigma 1 + \text{J} \ \text{S1}^2 \ \sigma 1 - \text{S1}^3 \ \sigma 1 - \text{S1} \ \Delta 2 \ \sigma 1^2 + \text{J}^2 \ \text{S1} \ \sigma 2 - 2 \ \text{J} \ \text{S1} \ \sigma 2 - 2 \ \text{J} \ \text{S1} \ \sigma 2 - 2 \ \text{J} \ \text{A1} \ \sigma 2^2 - \text{S1} \ \Delta 2 - 2 \ \text{J} \ \text{S1} \ \sigma 2 - 2 \ \text{J} \ \text{A1} \ \sigma 2^2 - \text{S1} \ \Delta 1 \ \sigma 2^2 \right); \\ & \text{f2} = \frac{1}{2} \left(L^2 \text{S1} \ \sigma 1 + \text{J} \ \text{S1}^2 \ \sigma 1 + \text{S1}^3 \ \sigma 1 + \text{S1} \ \Delta 2 \ \sigma 1^2 - \text{J}^2 \ \text{S1} \ \sigma 2 - 2 \ \text{J} \ \text{S1} \ \sigma 2 - 2 \ \text{J} \ \text{A1} \ \sigma 2^2 + \text{S1} \ \Delta 1 \ \sigma 2^2 \right); \\ & \text{f2} = \frac{1}{2} \left(L^2 \text{S1} \ \sigma 1 + \text{J} \ \text{S2}^2 - 2 \ \text{J} \ \text{A1} \ \sigma 1 \sigma 2 + \text{J} \ \text{A1} \ \sigma 2^2$$

$$\alpha 2 = \frac{(\sigma 1 - \sigma 2) (f2 - f1)}{D2 - f1 (\sigma 1 - \sigma 2)};$$

$$\alpha 1S = \frac{(-\sigma 1) (f2 - f1)}{D1S + f1 \sigma 1};$$

$$\alpha 2S = \frac{(-\sigma 1) (f2 - f1)}{D2S + f1 \sigma 1};$$

$$\Delta Y = EllipticK[\beta^2];$$

$$\Delta \lambda 1 = \frac{4 \Delta Y}{(A (f3 - f1))^{1/2}};$$

$$\Delta\lambda 2 = \left(\frac{-1}{2\,\,\mathrm{JdJ}^{1/2}}\right) \left(\frac{4}{(\mathrm{A}\,(\mathrm{f3-f1}))^{1/2}}\right) \left(\frac{\mathrm{B1}\,\,\mathrm{EllipticPi}\left[\alpha\mathrm{1}\,,\,\beta^2\right]}{\mathrm{D1-f1}\,(\sigma\mathrm{1-}\sigma\mathrm{2})} + \frac{\mathrm{B2}\,\,\mathrm{EllipticPi}\left[\alpha\mathrm{2}\,,\,\beta^2\right]}{\mathrm{D2-f1}\,(\sigma\mathrm{1-}\sigma\mathrm{2})}\right) \quad ;$$

$$\Delta\lambda3 = \left(\frac{-1}{2\;\text{LdL}^{1/2}}\right) \left(\left(\frac{4}{\left(\text{A}\;(\text{f3-f1})\right)^{1/2}}\right) \left(\frac{\text{B1 EllipticPi}\left[\alpha 1\;,\;\beta^2\right]}{\text{D1-f1}\left(\sigma 1-\sigma 2\right)} - \frac{\text{B2 EllipticPi}\left[\alpha 2\;,\;\beta^2\right]}{\text{D2-f1}\left(\sigma 1-\sigma 2\right)}\right) - \frac{1}{2} \left(\frac{1}{2}\left(\frac{1}{2}\right)^{1/2}\right) \left(\frac{1}{2}\left(\frac{1}{2}\right)^{1$$

(SeffdL + (
$$\Delta$$
1 - Δ 2) σ 1 σ 2 + LdL (σ 1 + σ 2)) $\frac{\Delta \lambda 1}{LdL^{1/2}}$);

$$\Delta\lambda 4 = \left(\frac{-1}{2\,\text{SldS1}^{1/2}}\right) \left(\left(\frac{4}{\left(\text{A}\,(\text{f3}-\text{f1})\right)^{1/2}}\right) \left(-\frac{\text{B1S EllipticPi}\left[\alpha\text{1S}\,,\,\beta^2\right]}{\text{D1S}+\text{f1}\,\sigma\text{1}} + \frac{\text{B2S EllipticPi}\left[\alpha\text{2S}\,,\,\beta^2\right]}{\text{D2S}+\text{f1}\,\sigma\text{1}}\right) + \frac{\text{SldS1}^{1/2}\left(\sigma\text{2}-\sigma\text{1}\right)\Delta\lambda\text{1}}{\text{SldS1}^{1/2}\left(\sigma\text{2}-\sigma\text{1}\right)\Delta\lambda\text{1}}\right);$$

 $\Delta\lambda 5 = \Delta\lambda 4 /. \sigma 1 \rightarrow \sigma 2 /. S1dS1 \rightarrow S1DS1 /. S2dS2 \rightarrow S2DS2 /. S1DS1 \rightarrow S2dS2 /. S2DS2 \rightarrow S1dS1;$

$$J5 = \frac{1}{\pi} \left(SeffdL \Delta \lambda 1 + J^2 \Delta \lambda 2 + L^2 \Delta \lambda 3 + S1^2 \Delta \lambda 4 + S2 \Delta \lambda 5 \right);$$

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- Other examples: Radius of convergence of Taylor series, Fermat's last theorem.

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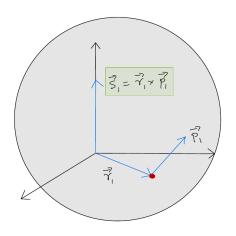
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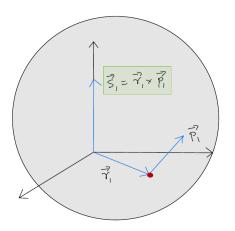
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- Numerical check: matches with numerical solution [github.com/sashwattanay/BBH-PN-Toolkit].

A pictorial mnemonic for fictitious variables



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Essentially, trade the spherical manifold in favor of a Cartesian manifold.

2PN: Integrable or non-integrable?

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- See the Introduction of [gr-qc:0511009] and [2012.06586] for details.

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- The non-exact nature of integrability ⇒ the tension b/w the two camps.

The fourth commuting constant of motion

With the definitions:

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The 4th commuting constant is

$$\begin{split} & \mathbf{L^2} - \epsilon \left[\frac{(m_2 \ P^i S_{1i} + m_1 \ P^i S_{2i})^2}{m_1^2 \ m_2^2} + \frac{2G(m_2 \ R^i S_{1i} + m_1 \ R^i S_{2i})^2}{(m_1 + m_2)(R^i R_i)^{3/2}} \right. \\ & + \left. \left(\frac{P^i P_i}{m_1 m_2} - \frac{2Gm_1 m_2}{(m_1 + m_2)\sqrt{R^i R_i}} \right) S_{1a} S_2^a \right]. \end{split}$$

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$$\begin{split} & \vec{S}_{\text{eff}} \cdot \vec{L} + \frac{1}{2} \left(S_{1}^{a} S_{2a} \right) + \frac{\epsilon \left(P^{a} S_{1a} \right)^{2}}{m_{1}^{2}} + \frac{3 m_{2} \epsilon \left(P^{a} S_{1a} \right)^{2}}{4 m_{1}^{3}} - \frac{2 G m_{2}^{2} \epsilon \left(R^{a} S_{1a} \right)^{2}}{\left(m_{1} + m_{2} \right) \left(R_{a} R^{a} \right)^{3/2}} \\ & - \frac{3 G m_{2}^{3} \epsilon \left(R^{a} S_{1a} \right)^{2}}{2 m_{1} \left(m_{1} + m_{2} \right) \left(R_{a} R^{a} \right)^{3/2}} + \frac{3 \epsilon \left(P^{a} S_{1a} \right) \left(P^{a} S_{2a} \right)}{4 m_{1}^{2}} \\ & + \frac{3 \epsilon \left(P^{a} S_{1a} \right) \left(P^{a} S_{2a} \right)}{4 m_{2}^{2}} + \frac{2 \epsilon \left(P^{a} S_{1a} \right) \left(P^{a} S_{2a} \right)}{m_{1} m_{2}} + \frac{3 m_{1} \epsilon \left(P^{a} S_{2a} \right)^{2}}{4 m_{2}^{3}} \\ & + \frac{\epsilon \left(P^{a} S_{2a} \right)^{2}}{m_{2}^{2}} - \frac{3 G m_{1}^{2} \epsilon \left(R^{a} S_{1a} \right) \left(R^{a} S_{2a} \right)}{2 \left(m_{1} + m_{2} \right) \left(R_{a} R^{a} \right)^{3/2}} \\ & - \frac{4 G m_{1} m_{2} \epsilon \left(R^{a} S_{1a} \right) \left(R^{a} S_{2a} \right)}{\left(m_{1} + m_{2} \right) \left(R_{a} R^{a} \right)^{3/2}} - \frac{3 G m_{2}^{2} \epsilon \left(R^{a} S_{1a} \right) \left(R^{a} S_{2a} \right)}{2 \left(m_{1} + m_{2} \right) \left(R_{a} R^{a} \right)^{3/2}} \\ & - \frac{2 G m_{1}^{2} \epsilon \left(R^{a} S_{2a} \right)^{2}}{\left(m_{1} + m_{2} \right) \left(R_{a} R^{a} \right)^{3/2}} - \frac{3 G m_{1}^{3} \epsilon \left(R^{a} S_{2a} \right)^{2}}{2 m_{2} \left(m_{1} + m_{2} \right) \left(R_{a} R^{a} \right)^{3/2}}. \end{split}$$

Conclusions

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Afterthoughts: (1) Nature is complex; elaborate math unavoidable (2) Using classical mechanics to do GW research.

• Find action-angles at 2PN using canonical pert. theory. Add 2.5PN radiation reaction.

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- Redo EOB for spinning systems by matching PN and EMRI action angles.



Thank you! Questions?

Refs:

- Papers: 2012.06586, 2110.15351, 2210.01605, 2110.09608.
- Lecture notes: 2206.05799
- Mathematica package: github.com/sashwattanay/BBH-PN-Toolkit