Milky Way Rotation Curve

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Methods & Assumptions

R ≤ 30 kiloparsecs

Bulge

 $M = 10^10$ solar masses

1≤ R ≤ 30 kiloparsecs

Disk, approximated as a disk

 $M = 10^11 \text{ solar masses}$

 $1 \le R \le 10$ kiloparsecs

Halo, approximated as a sphere

 $M = 10^12 \text{ solar masses}$

 $1 \le R \le 30$ kiloparsecs

Calculations

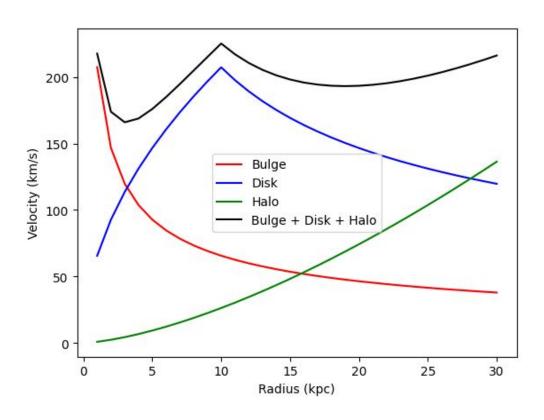
Orbital velocity = $\sqrt{(GM/R)}$

Enclosed mass of disk = π * R² * density

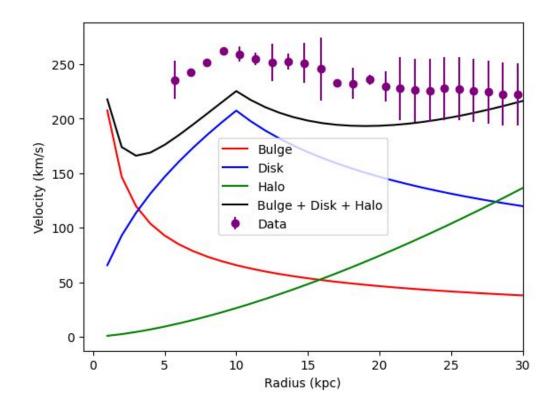
Volume of sphere = $4/3 * \pi * R^3$

Enclosed mass of sphere = $4/3 * \pi * R^3 *$ density

The Graph



The Graph (With Data)



Works Cited

Carignan, Claude, et al. "The extended HI rotation curve and mass distribution of M31." The Astrophysical Journal, vol. 641, no. 2, 30 Mar. 2006, https://doi.org/10.1086/503869.