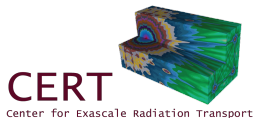


A High-Order Low-Order Algorithm with Exponentially-Convergent Monte Carlo for Thermal Radiative Transfer

Simon Bolding, Matt Cleveland, and Jim Morel

4 August 2016



We are interested in modeling thermal radiation transport in the high energy density physics regime

Modeling materials under extreme conditions

Temperatures $\mathcal{O}(10^6)$ K or more

Photon radiation transports through a material

Significant **energy** may be exchanged

We want to improve efficiency of Monte Carlo calculations

e.g., inertial confinement fusion, supernovae, et. al.

Our method has been applied to a simplified model:
the 1D frequency-integrated radiative transfer equations

Energy balance equations for radiation and material.

Radiation intensity $I(x, \mu, t)$, material temperature $T(x, t)$

$$\frac{1}{c} \frac{\partial I}{\partial t} + \mu \frac{\partial I}{\partial x} + \sigma_t I(x, \mu, t) = \frac{1}{4\pi} \sigma_a a c T^4,$$

$$C_v \frac{\partial T(x, t)}{\partial t} = \sigma_a \phi(x, t) - \sigma_a a c T^4$$

Equations are **nonlinear** and may be tightly coupled

Absorption opacity (σ_a) can be a strong function of T

Typically solved with implicit Monte Carlo (IMC)

which linearizes the system over a time step

Our high-order low-order (HOLo) method improves on several drawbacks of IMC

Standard IMC

Large **statistical noise** possible

Effective scattering can make MC very expensive

Linearization can cause **non-physical** results (maximum principle violations)

Reconstruction of linear emission shape limits artificial energy propagation

HOLo Method

ECMC is **very efficient** for TRT problems

MC solution has **no scattering**

Fully **implicit** time-discretization and LO solution **resolves nonlinearities**

Linear-discontinuous FE for $T(x)$ preserving equilibrium diffusion limit

A HOLO Algorithm for Thermal Radiative Transfer



Overview of algorithm

Derivation of the LO equations

Exponentially Convergent MC High-Order Solver

Computational Results

Extensions and Improvements on Algorithm

A HOLO Algorithm for Thermal Radiative Transfer



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Solve a non-linear low-order system with high-order angular correction from efficient MC simulations

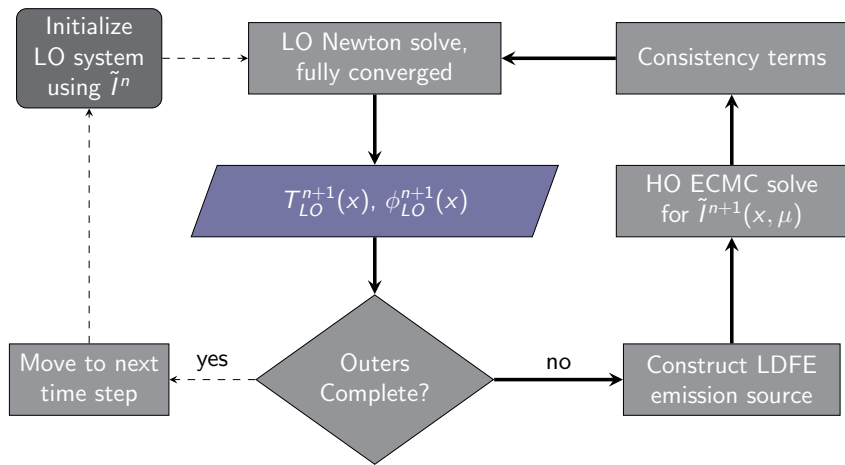
The **LO system** is space-angle moment equations, on a fixed finite-element (FE) spatial mesh

- ▶ Reduced dimensionality in angle
allows for solution with Newton's method
- ▶ **Output:** linear-discontinuous $\phi(x)$ and $T(x)$
Construct LDFE scattering and emission source

The **HO system** is a pure-absorber transport problem

- ▶ Solved with exponentially-convergent MC (ECMC)
for *efficient* reduction of statistical noise
- ▶ **Output:** consistency terms

Iterations between the HO and LO systems
can be performed each time step



A HOLO Algorithm for Thermal Radiative Transfer



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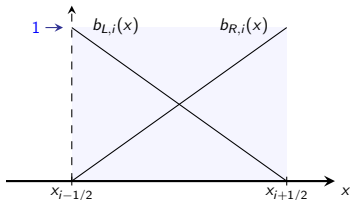
Computational Results

Extensions and Improvements on Algorithm

The LO equations are formed as *consistently* as possible with spatial and angular moments of TRT equations

The time discretization is backward Euler
for both the HO and LO equations

Spatial moments are weighted with FE basis functions:



$$\langle \cdot \rangle_{L,i} = \frac{2}{h_i} \int_{x_{i-1/2}}^{x_{i+1/2}} b_{L,i}(x)(\cdot) dx$$

Half-range integrals reduce angular dimensionality

$$I^+(x) = 2\pi \int_0^1 I(x, \mu) d\mu$$

Apply moments to the TRT equations
and manipulate to form **angular consistency terms**

For example, apply $\langle \cdot \rangle_{L,i}^+$ to a streaming term:

$$\begin{aligned} \left\langle \mu \frac{\partial I}{\partial x} \right\rangle_L^+ &= -\frac{2}{h_i} \{\mu I\}_{i-1/2}^+ + \frac{1}{h_i} [\langle \mu I \rangle_{L,i}^+ + \langle \mu I \rangle_{R,i}^+] \\ &= -\frac{2}{h_i} \frac{\{\mu I\}_{i-1/2}^+}{I_{i-1/2}^+} I_{i-1/2}^+ + \frac{1}{h_i} \frac{\langle \mu I \rangle_{L,i}^+}{\langle I \rangle_{L,i}^+} \langle I \rangle_{L,i}^+ + \frac{1}{h_i} \frac{\langle \mu I \rangle_{R,i}^+}{\langle I \rangle_{R,i}^+} \langle I \rangle_{R,i}^+ \end{aligned}$$

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Now, close the final moment equations:

- Approximate consistency terms with \tilde{I}_{HO}^{n+1} from previous HO solve
- Eliminate remaining face unknowns with LD spatial closure
 $T^4(x)$ and $T(x)$ also assumed LD

A HOLO Algorithm for Thermal Radiative Transfer



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Exponentially Convergent MC High-Order Solver

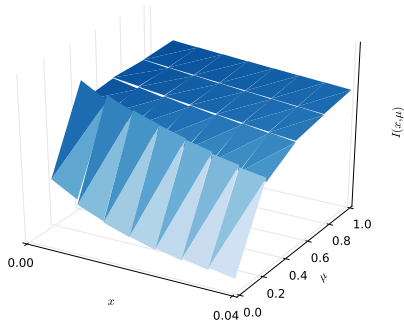
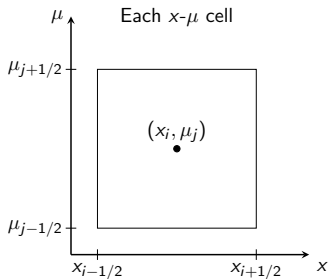
Computational Results

Extensions and Improvements on Algorithm

We use a **projection** $\tilde{I}(x, \mu)$ onto a space-angle LDFE mesh to represent the solution

local volumetric tallies

$$\tilde{I}_{ij}(x, \mu) = I_a + \frac{2}{h_x} I_x (x - x_i) + \frac{2}{h_\mu} I_\mu (\mu - \mu_i)$$



We apply the ECMC algorithm to the **pure-absorber** HO transport equation

$$\left[\mu \frac{\partial}{\partial x} + \left(\sigma_t + \frac{1}{c \Delta t} \right) \right] I^{n+1} = \frac{1}{4\pi} \left[\sigma_a a c (T_{LO}^{n+1})^4 + \sigma_s \phi_{LO}^{n+1} \right] + \frac{\tilde{J}^n}{c \Delta t}$$
$$\mathbf{L} I^{n+1} = q$$

For each batch m :

- ▶ Evaluate residual source: $r^{(m)} = q - \mathbf{L} \tilde{J}^{n+1,(m)}$
- ▶ Estimate $\epsilon^{(m)} = \mathbf{L}^{-1} r^{(m)}$ via **MC simulation**
- ▶ Update solution: $\tilde{J}^{n+1,(m+1)} = \tilde{J}^{n+1,(m)} + \tilde{\epsilon}^{(m)}$

Our HO system allows for straight-forward variance reduction

$I^n(x, \mu)$ is often an **excellent** estimate of $I^{n+1}(x, \mu)$
No MC sampling from thermal equilibrium regions

Histories stream without collision
along path s , weight reduces as $w(s) = w_0 e^{-\sigma_t s}$

Use cell-wise systematic sampling for $|r^{(m)}|$ source
Particularly effective in thick cells

- ▶ Particles in each x - μ cell $\propto |r^{(m)}|$ in cell
- ▶ Set minimum n for cells
except for cells in thermal equilibrium

A HOLO Algorithm for Thermal Radiative Transfer



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Derivation of the LO equations

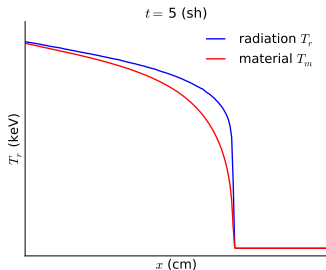
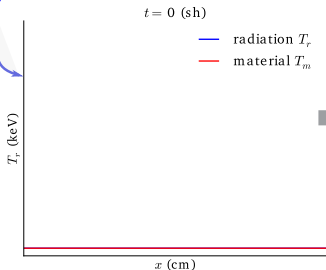
Exponentially Convergent MC High-Order Solver

Computational Results

Extensions and Improvements on Algorithm

We will test our method with several standard **Marshak Wave** problems

constant radiation
boundary source



Results show radiation temperature $T_r = \sqrt[4]{\phi/ac}$

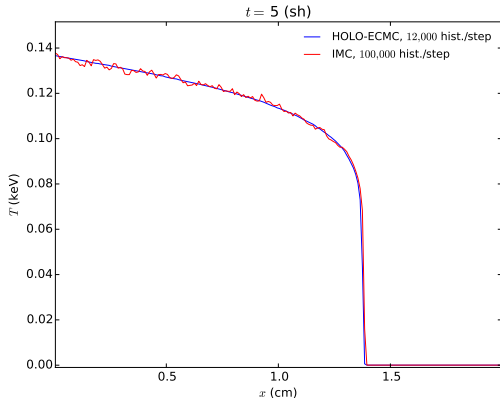
Implementation specifics for results are given below:

- ▶ Written as stand-alone C++ code
- ▶ One HO solve per time step (two LO solves)
 - ▶ *each HO solve* has 3 ECMC batches
no mesh refinement
- ▶ $\text{FOM} = \frac{1}{\|s^2\| N_{\text{total}}}$, normalized to IMC result

The HOLO method produces significantly less noise than IMC for a typical Marshak Wave: **FOM=145**

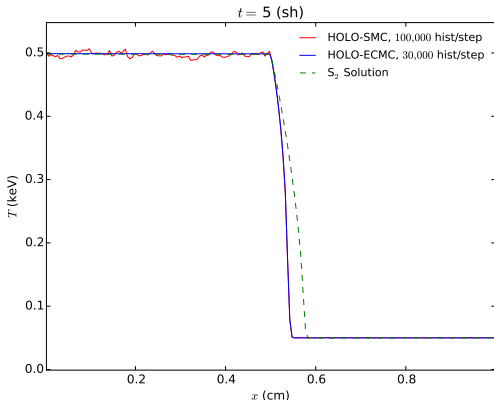
► $\sigma_a \propto T^{-3}$

- Transient solution after 5 shakes
200 x cells and 4 μ cells (for ECMC)



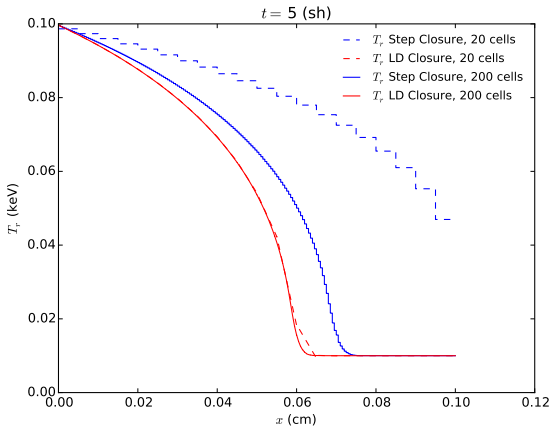
ECMC is more efficient than standard MC as a HO solver

- ▶ Left half is optically thin ($\sigma=0.2 \text{ cm}^{-1}$), right half is thick ($\sigma_a=2000 \text{ cm}^{-1}$). 8 μ cells
- ▶ Different HO solvers: **ECMC** (FOM=10,000), **standard MC** (FOM=0.46), and an S_2 solution



The LDFE discretization for the LO equations preserves the equilibrium diffusion limit

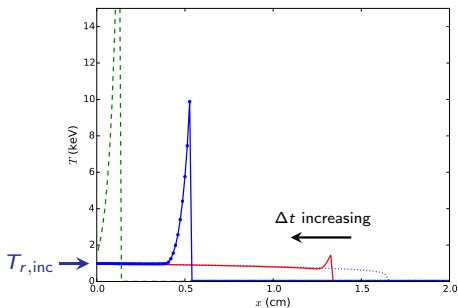
- Large, constant σ_a and small c_v



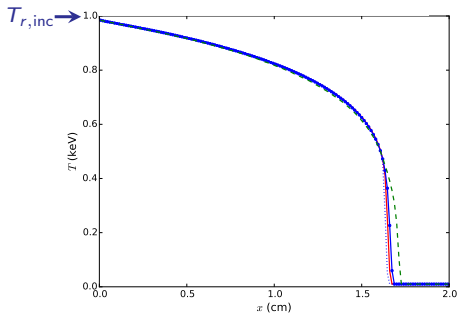
Our HOLO method preserves the maximum principle with sufficient nonlinear convergence

- ▶ **Material temperatures** plotted; all simulations end at $t = 0.1$ sh
 $\sigma_a \propto T^{-3}$, c_v small, $\Delta t \in [10^{-4}, 10^{-2}]$ sh
- ▶ LO Newton iterations required damping

IMC T_m



HOLO T_m



A HOLO Algorithm for Thermal Radiative Transfer



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Exponentially Convergent MC High-Order Solver

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We need a way to resolve issues when the LDFF representation of the intensity is negative

Negative intensities can occur in optically thick cells
Mesh refinement is of minimal use

$\tilde{I}_{HO}(x, \mu)$ must be positive for consistency terms
to produce a physical, stable LO solution

Independent fix up for LO solution
E.g., lumping or preserving balance with floored $\phi(x)$

Can add source δ to produce a positive projection \tilde{l}_{pos} such that \tilde{l}_{pos} satisfies the latest residual equation

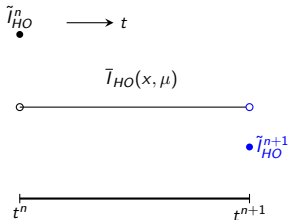
Produce \tilde{l}_{pos} by scaling $x - \mu$ moments equally,
to estimate source for the next iteration

$$\begin{array}{lcl} \mathbf{L}\tilde{l}^{(m)} = q - r^{(m)} & \longrightarrow & \delta^{(m+1)} = \mathbf{L} \left(\tilde{l}^{(m)} - \tilde{l}_{pos}^{(m)} \right) \\ \mathbf{L}\tilde{l}_{pos}^{(m)} = q - r^{(m)} + \delta^{(m+1)} & & q \rightarrow q + \delta^{(m+1)} \end{array}$$

We can delay error stagnation

Investigating alternative positive projection of l

The time variable can be included in the ECMC trial space with a consistent LO closure



Include $\frac{1}{c} \frac{\partial}{\partial t} (\cdot)$ in transport operator \mathbf{L}
with $T(x)$ still implicit

Sample and track in time

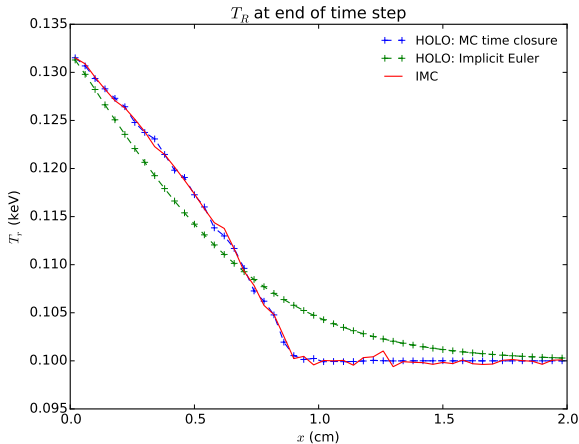
Tally time-average error projection and at t^{n+1}

In **LO equations**, parameterize ϕ_{LO}^{n+1}
in terms of **time-averaged** unknowns

$$\langle \phi \rangle_{L,i}^{n+1} = 2 \gamma_{L,i}^{HO} \overline{\langle \phi \rangle}_{L,i} - \langle \phi \rangle_{L,i}^n$$

The time-closure parameters preserve MC time integration in the LO solution

- ▶ Material is near-void, so temperature uncouples
take 3 large time steps and compare T_R^{n+1}
- ▶ 300,000 histories/step



A couple other improvements I didn't have time to discuss

The HO solution can be used to estimate the spatial closure

by including face tallies in ECMC

Source iteration with diffusion synthetic acceleration
as a solution to the LO equations

A HOLO Algorithm for Thermal Radiative Transfer

ECMC is very efficient for TRT simulations
and fits well in global HOLO context

The LO system can resolve nonlinearities
with bounded angular consistency terms

Moving forward need to extend to higher spatial
dimensions
main hurdle to overcome is infrastructure

M

ore details in: S.R. Bolding, M. Cleveland, and J.E. Morel.
A HOLO Algorithm with ECMC for Radiative Transfer. NS&E:
M&C 2015 Special Issue, 2016. Accepted.

A HOLO Algorithm for Thermal Radiative Transfer

For more details see:

S.R. Bolding, M. Cleveland, and J.E. Morel. A High-Order Low-Order Algorithm with Exponentially-Convergent Monte Carlo for Thermal Radiative Transfer. Nuclear Science & Engineering: M&C 2015 Special Issue, 2016. Accepted.

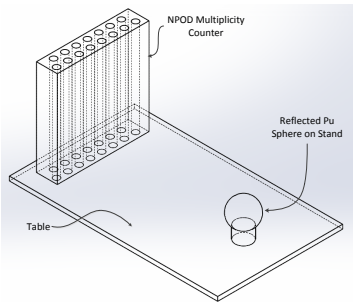
Simulations of Neutron Multiplicity Experiments with Perturbations to Nuclear Data

Simon Bolding and C.J. Solomon

4 August 2016



Multiplicity experiments were performed at Sandia for validating subcritical simulations



*Not to scale

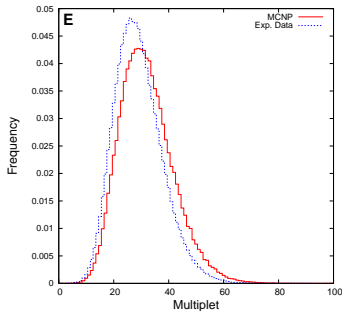
Experimental Parameters

- ▶ 94% ^{239}Pu sphere
- ▶ 5 Different HDPE shells
From none to 3.0 cm HDPE

Experiments repeated w/ ^{252}Cf

Solomon modeled experiments
with MCNP

MCNP5 multiplicity simulations showed discrepancy with experiments for Pu but not for ^{252}Cf



Pu with 3.0-cm HDPE reflector

Previous work by Mattingly [2010]

- Caused by ^{239}Pu nuclear data
- Adjusted energy-integrated $\bar{\nu}$
mean # of neutrons/fission

ENDF-VII raised $\bar{\nu}$ for ^{239}Pu
to match k_{eff} benchmarks

- $\bar{\nu}$ is $\sim 2\sigma$ above measured data
for $E < 1.5$ MeV

Can we reduce discrepancy in multiplicity distributions without significantly altering k_{eff} ?

Perform energy-dependent perturbations of $\bar{\nu}(E)$ in ^{239}Pu
Random samples drawn from ENDF-VII.1 covariance data

Compare experimental and simulated multiplicity dist.
and a k_{eff} benchmark (Jezebel)

Compare $\bar{\nu}(E)$ results to uniform shifts of microscopic cross sections

LANL Python verification library was modified heavily to generate energy-dependent $\bar{\nu}$ samples

1. Generate a correlated sample of $\bar{\nu}(E)$
 - ▶ Assumed multivariate Gaussian
with group-averaged covariances
2. Modify continuous $\bar{\nu}(E)$ data in **ACE** file
3. Perform all MCNP simulations
with modified ACE data

A cost function provides a measure of inaccuracy for each data realization

Reduced χ^2 values for the 5 multiplicity experiments and criticality benchmark

$$\chi_{\text{red,mult},m}^2 = \frac{1}{N_{\text{bins}} - 1} \sum_{i=1}^{N_{\text{bins}}} \frac{(P_i^{\text{exp}} - P_i^{\text{mcnp}})^2}{\sigma^2(P_i^{\text{exp}}) + \sigma^2(P_i^{\text{mcnp}})}$$

Equally weight χ^2 values in a cost function
A lower score indicates higher accuracy

$$\text{Cost} = \sum_{m=1}^5 \chi_{\text{red,mult},m}^2 + \chi_{\text{red},k_{\text{eff}}}^2$$

Multiplicity and k_{eff} simulations were performed for 500 unique realizations of $\bar{\nu}$ data

Trial	Cost	$\chi^2_{k_{\text{eff}}}$
$\bar{\nu}$ -1.14%	164.24	33.66
303	197.07	4.18
55	267.9	0.01
Original	426.86	0.27

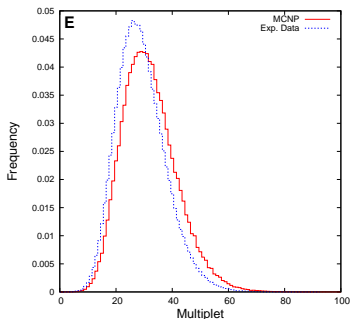
MCNP criticality test suite performed for best data which includes 39 criticality benchmarks w/ ^{239}Pu :

Trial	<i>RMSD</i>
$\bar{\nu}$ -1.14%	1.23%
303	0.51%
Original	0.49%

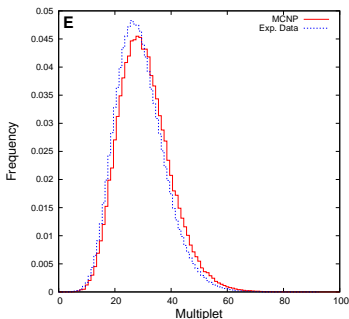
Energy-dependent $\bar{\nu}$ perturbations improved all 5 multiplicity distributions

- Plots for best data realization and 3.0 cm HDPE case

Original $\bar{\nu}$ Data



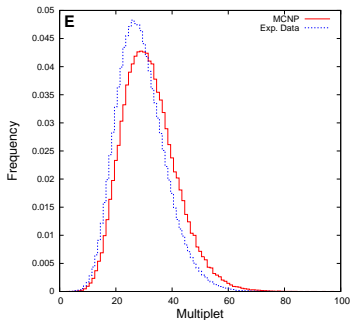
Trial 303: Lowest Cost



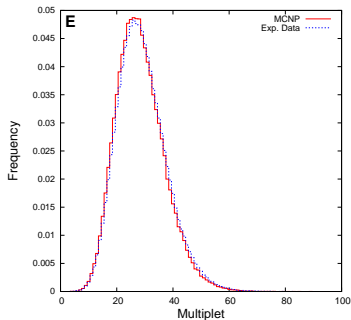
- Best data set reduced **bias** in 1st and 2nd moments, averaged over all 5 simulations, by $\sim 35\%$

Adjusting the fission cross section **uniformly**
showed good correction to multiplicity simulations

Original σ_f Data



σ_f decreased 1.5%



- High accuracy for all simulations: $\sum_{i=1}^5 \chi_{red,mult,m}^2 = 14.6$
- k_{eff} is **not preserved**: $\chi_{k_{eff}}^2 = 22.6$

Simulations of Multiplicity Experiments with Nuclear Data Perturbations

Energy-dependent $\bar{\nu}$ perturbations **reduced inaccuracies**
in multiplicity while **preserving** k_{eff}

- ▶ Majority of cross-correlation terms $\mathcal{O}(10^{-4})$ or less
- ▶ σ_f may need more investigation
not sensitive to capture cross section

Subcritical simulations should be considered
in validation of nuclear data

Covariance sampling methodology for nuclear data was
developed and demonstrated

- ▶ Ideally sample all cross sections and $\bar{\nu}$ simultaneously

Simulations of Neutron Multiplicity Experiments with Nuclear Data Perturbations

S.R. Bolding¹, C.J. Solomon²

¹ *Texas A&M University, College station, TX*

² *Los Alamos National Laboratory, Los Alamos, NM*

4 August 2016



Backup Slides

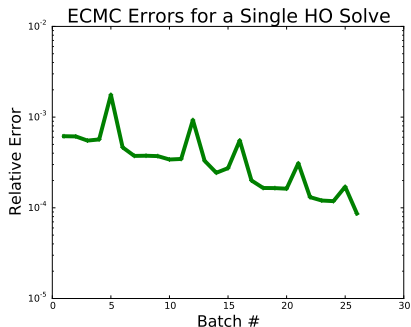
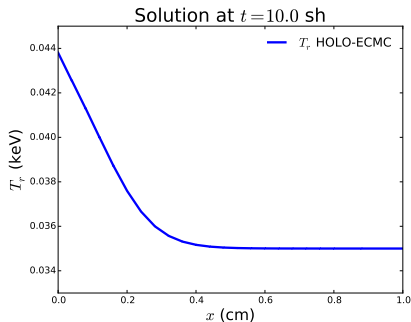
S.R. Bolding¹, C.J. Solomon²

¹ *Texas A&M University, College station, TX*

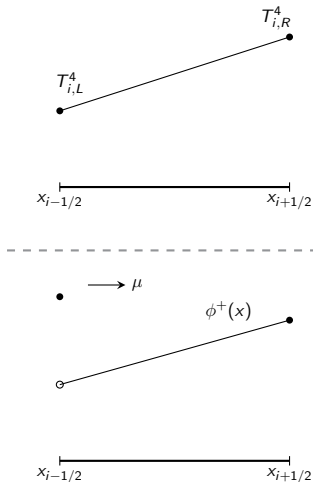
² *Los Alamos National Laboratory, Los Alamos, NM*



Exponential convergence can be maintained if the LDFFE mesh resolves the solution reasonably



We can close equations with HO angular information and a linear-discontinuous (LD) spatial discretization



1. Assume $T(x)$ and $T^4(x)$ are LD

2. A lagged $\tilde{I}_{\text{HO}}^{n+1}$ is used to evaluate consistency terms

3. Eliminate $\phi_{i+1/2}^\pm$ with LD closure preserving equi. diff. limit

$$\phi_{i+1/2}^+ = 2\langle\phi\rangle_{R,i}^+ - \langle\phi\rangle_{L,i}^+$$

4. Global system solved with Newton's method and lagged **implicit opacities**
Energy is always conserved

We need a way to resolve issues when the LDFF representation of the intensity is negative

Negative intensities can occur in optically thick cells
Mesh refinement is of minimal use

$\tilde{I}_{HO}(x, \mu)$ must be positive for consistency terms
to produce a physical, stable LO solution

Independent fix up for LO solution

E.g., lumping or preserving balance with floored $\phi(x)$

Can add source δ to produce a positive projection \tilde{l}_{pos} such that \tilde{l}_{pos} satisfies the latest residual equation

Produce \tilde{l}_{pos} by scaling $x - \mu$ moments equally,
to estimate source for the next iteration

$$\begin{array}{ll} \mathbf{L}\tilde{l}^{(m)} = q - r^{(m)} & \longrightarrow \delta^{(m+1)} = \mathbf{L} \left(\tilde{l}^{(m)} - \tilde{l}_{pos}^{(m)} \right) \\ \mathbf{L}\tilde{l}_{pos}^{(m)} = q - r^{(m)} + \delta^{(m+1)} & q \rightarrow q + \delta^{(m+1)} \end{array}$$

We can delay error stagnation

Investigating alternative positive projection of l

Solving LO System with Newton's Method

- Linearization:

$$\underline{B}(T^{n+1}) = \underline{B}(T^*) + (T^{n+1} - T^*) \left. \frac{\partial \underline{B}}{\partial t} \right|_{t^*}$$

- Modified system

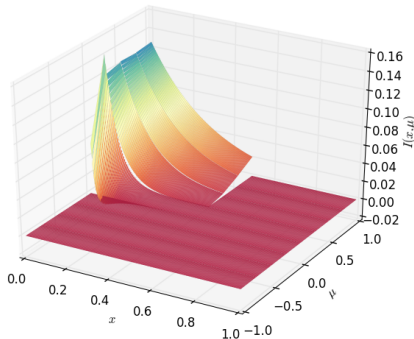
$$[\mathbf{D}(\mu^\pm) - \sigma_a^*(1 - f^*)] \underline{\Phi}^{n+1} = f^* \underline{B}(T^*) + \frac{\underline{\Phi}^n}{c\Delta t}$$

$$\hat{\mathbf{D}} \underline{\Phi}^{n+1} = \underline{Q}$$

$$f = \left(1 + \sigma_a^* c \Delta t \frac{4aT^{*3}}{\rho c_v} \right)^{-1} \quad T_i^* = \frac{T_{L,i}^* + T_{R,i}^*}{2}$$

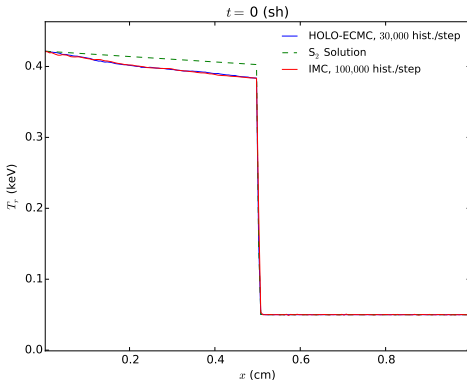
- Equation for T^{n+1} based on linearization that is conservative
- Converge T^{n+1} and $\langle \phi \rangle$ with Newton Iterations

The angular flux for the two material problem is difficult to resolve near $\mu = 0$



Two Material Problem, comparison in optically thin region

- Plot of radiation temperature after 10 time steps



Backup slide with timing results

Forming the LO System

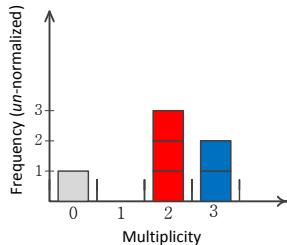
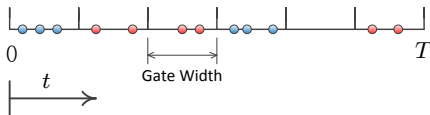
- Taking moments of TE yields 4 equations, per cell i , e.g.

$$\begin{aligned}
 & -2\mu_{i-1/2}^{n+1,+} \phi_{i-1/2}^{n+1,+} + \{\mu\}_{L,i}^{n+1,+} \langle \phi \rangle_{L,i}^{n+1,+} + \{\mu\}_{R,i}^{n+1,+} \langle \phi \rangle_{R,i}^{n+1,+} + \\
 & \left(\sigma_t^{n+1} + \frac{1}{c\Delta t} \right) h_i \langle \phi \rangle_{L,i}^{n+1,+} - \frac{\sigma_s h_i}{2} \left(\langle \phi \rangle_{L,i}^{n+1,+} + \langle \phi \rangle_{L,i}^{n+1,-} \right) \\
 & = \frac{h_i}{2} \langle \sigma_a^{n+1} a c T^{n+1,4} \rangle_{L,i} + \frac{h_i}{c\Delta t} \langle \phi \rangle_{L,i}^{n,+}, \quad (1)
 \end{aligned}$$

- Cell unknowns: $\langle \phi \rangle_{L,i}^+$, $\langle \phi \rangle_{R,i}^+$, $\langle \phi \rangle_{L,i}^-$, $\langle \phi \rangle_{R,i}^-$, T_L , T_R
- Need angular consistency terms and spatial closure (LD)

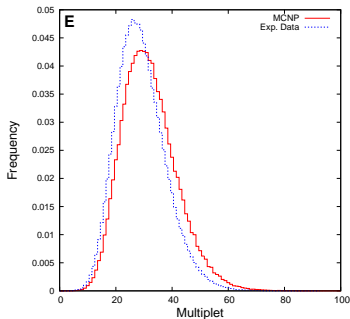
Neutron multiplicity distributions provide passive multiplication information about a fissionable system

○ = Detected Neutron

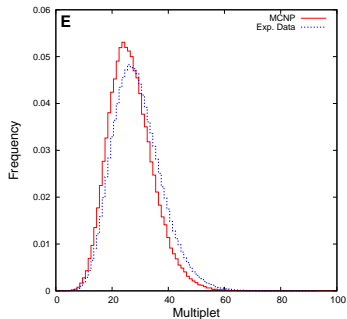


Energy-Integrated $\bar{\nu}$ Shift – 3.0 cm HDPE reflector

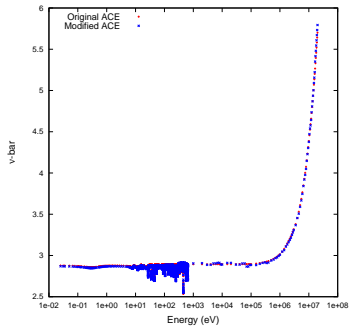
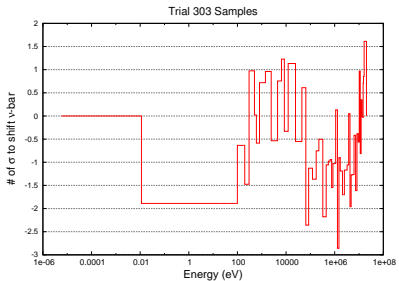
Original $\bar{\nu}$ Data



-1.14% $\bar{\nu}$ Data



The best $\bar{\nu}$ data (trial 303)



Fractional shifts to cross sections were made for comparison

Adjusted cross section uniformly at all energies
compensated with σ_{tot} or σ_{el}

Increasing capture cross section was not effective
relative to variance in data

Scaling fission cross section 1.5% ($< 1\sigma$)
improved multiplicity distributions

- ▶ Adjust elastic scattering (σ_{el}) to compensate change in σ_f , for $E > 1$ keV
- ▶ Better improvement than *uniform* scaling of $\bar{\nu}$