

Second-Order Discretization in Space and Time for Radiation-Hydrodynamics

Simon Bolding^a, Joshua Hansel^a, Jarrod D. Edwards^b, Jim E. Morel^a,
Robert B. Lowrie^c

^a*Department of Nuclear Engineering, 337 Zachry Engineering Center, TAMU 3133,
Texas A&M University, College Station, Texas, 77843*

^b*Phenomenology and Sensor Science Department, Sandia National Laboratory,
Albuquerque, NM*

^c*Computational Physics Group CCS-2, Los Alamos National Laboratory, P.O. Box 1663,
MS D413, Los Alamos, NM 87545*

1. Introduction

In this work, we derive, implement, and test a new IMEX scheme for solving the equations of radiation hydrodynamics that is second-order accurate in both space and time. We consider a RH system that combines a 1-D slab model of compressible fluid dynamics with a grey radiation S_2 model, given by:

$$\frac{\partial \rho}{\partial t} + \frac{\partial}{\partial x} (\rho u) = 0, \quad (1a)$$

$$\frac{\partial}{\partial t} (\rho u) + \frac{\partial}{\partial x} (\rho u^2) + \frac{\partial}{\partial x} (p) = \frac{\sigma_t}{c} F_{r,0}, \quad (1b)$$

$$\frac{\partial E}{\partial t} + \frac{\partial}{\partial x} [(E + p) u] = -\sigma_a c (aT^4 - E_r) + \frac{\sigma_t u}{c} F_{r,0}, \quad (1c)$$

$$\frac{1}{c} \frac{\partial \psi^+}{\partial t} + \frac{1}{\sqrt{3}} \frac{\partial \psi^+}{\partial x} + \sigma_t \psi^+ = \frac{\sigma_s}{4\pi} c E_r + \frac{\sigma_a}{4\pi} a c T^4 - \frac{\sigma_t u}{4\pi c} F_{r,0} + \frac{\sigma_t}{\sqrt{3}\pi} E u, \quad (1d)$$

$$\frac{1}{c} \frac{\partial \psi^-}{\partial t} - \frac{1}{\sqrt{3}} \frac{\partial \psi^-}{\partial x} + \sigma_t \psi^- = \frac{\sigma_s}{4\pi} c E_r + \frac{\sigma_a}{4\pi} a c T^4 - \frac{\sigma_t u}{4\pi c} F_{r,0} - \frac{\sigma_t}{\sqrt{3}\pi} E u, \quad (1e)$$

where ρ is the density, u is the velocity, $E = \frac{\rho u^2}{2} + \rho e$ is the total material energy density, e is the specific internal energy density, T is the material

temperature, E_r is the radiation energy density,

$$E_r = \frac{2\pi}{c} (\psi^+ + \psi^-) , \quad (2)$$

F_r is the radiation energy flux,

$$F_r = \frac{2\pi}{\sqrt{3}} (\psi^+ - \psi^-) \quad (3)$$

and $F_{r,0}$ is an approximation to the comoving-frame flux,

$$F_{r,0} = F_r - \frac{4}{3} E_r u . \quad (4)$$

Note that if we multiply Eqs. ((1d)) and ((1e)) by 2π and sum them, we obtain the radiation energy equation:

$$\frac{\partial E_r}{\partial t} + \frac{\partial F_r}{\partial x} = \sigma_a c (aT^4 - E_r) - \frac{\sigma_t u}{c} F_{r,0} , \quad (5a)$$

and if we multiply Eq. (1d) by $\frac{2\pi}{c\sqrt{3}}$, multiply Eq. (1e) by $-\frac{2\pi}{c\sqrt{3}}$ and sum them, we get the radiation momentum equation:

$$\frac{1}{c^2} \frac{\partial F_r}{\partial t} + \frac{1}{3} \frac{\partial E_r}{\partial x} = -\frac{\sigma_t}{c} F_{r,0} . \quad (5b)$$

Equations (1a) through (1e) are closed in our calculations by assuming an ideal equation of state (EOS):

$$p = \rho e (\gamma - 1) , \quad (6a)$$

$$T = \frac{e}{C_v} , \quad (6b)$$

where γ is the adiabatic index, and C_v is the specific heat. However, our method is compatible with any valid EOS.

2. Linearization of Equations

Consider the case of the non-linear system to be solved for Crank Nicolson over a time step from t_n to t_{n+1} . The changes to the non-linear system for the predictor and corrector time steps will only effect the choice of Δt , the end time state, and the known source terms on the right hand side from previous states in time. The non-linear equations to be solved in this case are

$$\frac{\mathcal{E}^{n+1} - \mathcal{E}}{\Delta t} = \quad (7)$$

$$\begin{aligned} \frac{E^{n+1} - E^*}{\Delta t} = & -\frac{1}{2} [\sigma_a c (aT^4 - \mathcal{E})]^{n+1,k+1} - \frac{1}{2} [\sigma_a c (aT^4 - \mathcal{E})]^n \\ & - \frac{1}{2} \left[\sigma_t \frac{u}{c} \left(\frac{4}{3} \mathcal{E} u - \mathcal{F} \right) \right]^{n+1,k} + \frac{1}{2} \left[\sigma_t \frac{u}{c} \left(\frac{4}{3} \mathcal{E} u - \mathcal{F} \right) \right]^n \end{aligned} \quad (8)$$

To simplify the algebra, define a source term for all the known, lagged values

$$Q_E^{k,n} = \frac{1}{2} [\sigma_a (a(T^n)^4 - \mathcal{E})]^n - \frac{1}{2} \left[\sigma_t \frac{u}{c} \left(\frac{4}{3} \mathcal{E} u - \mathcal{F} \right) \right]^{n+1,k} \quad (9)$$

We linearize the Planckian function about some temperature near T^{n+1} , denoted T^k , giving

$$(T^{n+1,k+1})^4 = (T^k)^4 + \frac{4(T^k)^3}{\rho c_v^k} (e^{n+1,k+1} - e^k) \quad (10)$$

and substitute this into Eq. (8). We then solve for the quantity $e^{n+1,k+1} - e^k$ through algebraic manipulation, as follows