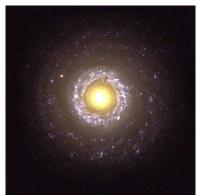
# **Astronomy from 4 Perspectives: the Dark Universe**

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# **Questions: Dark matter and galaxy rotation curves**

# 1. Orientations of galaxies

Think about how the galaxy should be orientated to be observed? Here are some pictures as example:



(a) Edge on galaxy NGC 7742 Inclination angle  $i = 0^{\circ}$ Image credit: Hubble Heritage Team (AURA/STScI/NASA/ESA)



dromeda as seen in infrared. Inclination angle  $i \approx 13^{\circ}$ Image credit: NASA/JPL-Caltech/UCLA



(b) Our galactic neighbour An- (c) The almost edge-on sombrero galaxy. Inclination angle  $i \approx 90^{\circ}$ Image credit: Carsten Frenzl

#### 2. Galactic Rotation curves

Calculate the radial velocities from the measured wavelengths and plot them over the distance from the galaxy center. use  $\lambda_0 = 21.106 \,\text{Å}$  and  $1 \, \text{pc} = 3.1 \cdot 10^{16} \, \text{m}$ .

λin Å	Radius R in Mpc	$v_{\text{rotation}}$ in $\frac{km}{s}$
21.1195	1	
21.1130	2	
21.1173	5	
21.1194	7	
21.1201	10	
21.1208	15	
21.1211	20	
21.1215	22	
21.1213	25	

## 3. Circular obits

Derive for circular orbits the formula for the velocity v in dependence of the distance r. Assume a radially symmetric mass distribution.

#### 4. Velocity of planets

Assuming circular orbits, compute the velocities of the planets in our solar system. Plot the resulting rotation curve v over r.

### 5. Expected Rotation curve

Formulate an expectation for the rotation curve of the Milky Way, assuming the mass in the bulge to be  $1.6 \cdot 10^{10} \, M_\odot$  and the disk to be  $4 \cdot 10^{10} \, M_\odot$ 

**6. Observed rotation curve** The observed rotation curves of spiral galaxies are of the following form: This cannot be explained by visible mass alone. Assuming that dark matter is the source of the

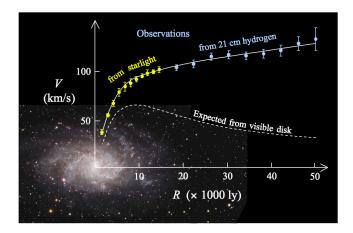


Figure 2: Expected and observed rotation curve of M33. Image credit: Steffania Deluca

difference between the observed and the predicted rotation curves, please calculate the mass of the dark matter depending on the velocities v(r) and  $v_{axis}(r)$ 

**7.** *Dark matter distribution* To find out how dark matter is distributed throughout a spiral galaxy, please consider a simple rotation curve consisting of a linear and a constant branch. Assume a spherically symmetric mass distribution of the form

$$\mu(r) \propto r^k$$

For the mass use the formula

$$M(r) = 4\pi \int_{r_0}^{r} \mu(\rho) \rho^2 d\rho$$

- (a) Please calculate the mass of the bulge in dependence of k
- (b) Using the formula for v from Task 3 and the result from a), please determine the exponent k for the bulge. Calculate the mass of the bulge in dependence of r and the complete mass  $M_B$  of the bulge.
- (c) To determine k for the halo (v = const.), consider the total mass to be composed of the mass of the bulge  $M_B$  and the mass of the halo  $M_H$ .

$$M(r) = M_R + M_H(r)$$

Calculate the mass outside of the bulge with the integral for the mass. Determine the exponent *k* using the results of b) and Task 3. Find a formula for the mass of the halo in dependence of *r* 

(d) Compare the rotation curve of the bulge to the rotation curve of a rigid body.