

Figure 12. A measure of r-process strength across our parameter space, using a wave frequency of 2×10^3 rad s⁻¹ and $Y_{e,eq} = 0.48$, and using the shock heating prescription from equation (13). Two regimes of r-process production emerge: a region of high wave and neutrino luminosities across all masses, driven by shock heating; and a region of moderate wave and neutrino luminosities driven by the wave stress, which becomes significant at larger neutron star masses.