



## Model Order Reduction of Rarefied Gases Using Neural Networks

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## Outline

### Introduction

### The BGK-Model

### Sod's shock tube

### Proper Orthogonal Decomposition (POD)

### Neural Networks

### Results

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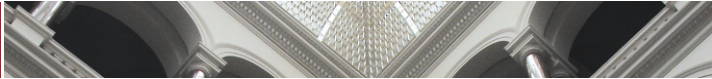
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## Governing equations

- The Boltzmann equation approximated by  $\mathbf{Q}$  the BGK operator as a source term with

$$\partial_t f + v \partial_x f = \overbrace{\frac{1}{\tau} (M_f - f)}^{\mathbf{Q}} \quad (1)$$

- The equilibrium solution is a Maxwellian distribution  $\mathbf{M}_f$  with

$$M_f = \frac{\rho(x, t)}{(2\pi RT(x, t))^{\frac{3}{2}}} \exp\left(-\frac{(v - u(x, t))^2}{2RT(x, t)}\right) \quad (2)$$

- The duration to evolve into equilibrium is given by the relaxation time  $\tau$  with

$$\tau^{-1} = \frac{\rho(x, t) T^{1-\nu}(x, t)}{Kn} \quad (3)$$

- The rarefaction level is defined over the Knudsen number  $\mathbf{Kn}$  with

$$Kn = \frac{\lambda}{l} \quad (4)$$

1

<sup>1</sup>PhysRev.94.511.





## Knudsen number

- Solution is  $f(x, v, t)$  in 1D and  $f(x, y, v_x, v_y, t)$  in 2D and  $f(x, y, z, v_x, v_y, v_z, t)$  in 3D

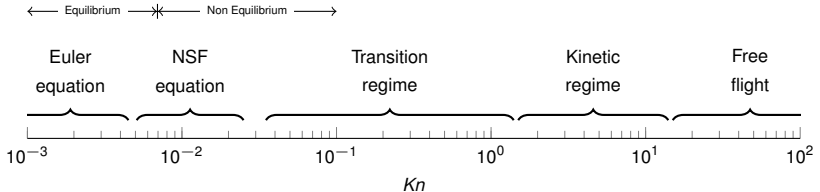


Figure: Partitioning of  $Kn$ , the Knudsen number, into levels of rarefaction.

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<sup>2</sup>NumaKUL.





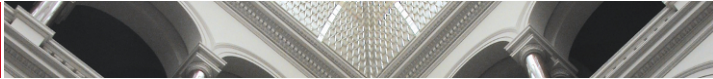
## Discretization in space and velocity space in 1D

- Space and time discretization considering a uniform grid i.e.  
 $x_j = j\Delta x$  and  $j \in \mathbb{Z}$ ,  $v_k = k\Delta v$  and  $k \in \mathbb{Z}$ ,  $t^i = i\Delta t$  and  $t \in \mathbb{N}$ ,
- is leading from the full PDE to a set of ODE's in time

$$\partial_t f_{j,k} = -(v_k)_1 D_x f|_{j,k}(t) + \frac{1}{\tau} (M_{f_{j,k}}(t) - f_{j,k}(t)). \quad (5)$$

- $KJ$  first-order differential equations need to be evaluated when  $K$  and  $J$  are the number of gridpoints in space and velocity space.
- In 3D there are  $K^3 J^3$  first-order differential equations.
- The discretization in velocity space requires the computation of the moments of  $f$ .





## Moments/ Expected values of $f$

- By multiplying the collision invariants  $\Phi(v) = [1, v, \frac{1}{2}v^2]$  with  $f$  and integrating in velocity space  $v$  the moments are obtained.
- The first moment/ the Density is

$$\rho(x, t) = \int f \, dv, \quad (6)$$

- the second moment/ the Momentum is

$$\rho(x, t)u(x, t) = \int vf \, dv, \quad (7)$$

- the third moment/ the kinetic Energy is

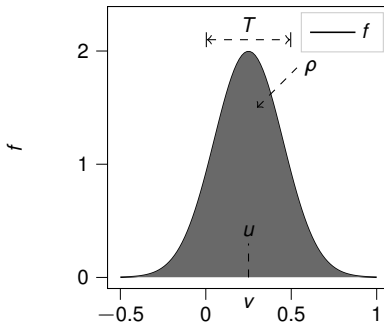
$$E(x, t) = \int \frac{1}{2}v^2 f \, dv. \quad (8)$$







## Moments/ Expected values of $f$



**Figure:** Illustration of the linkage between the macroscopic quantities of the gas flow and the distribution function  $f$ .



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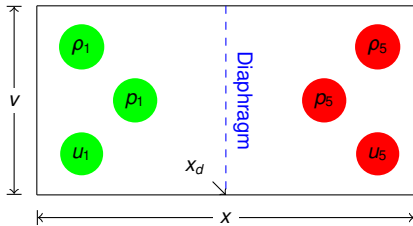
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## Sod's shock tube



**Figure:** Problem setup of Sod's shock tube for the BGK model in 1D.

- Test case for numerical schemes solving
- **non-linear hyperbolic conservation laws** in gas dynamics (Gary A. Sod in 1978)
- **Idea:**
  - Solve problem analytically (Rankine-Hugoniot jump conditions)
  - Solve problem numerically
  - Compare resolution of discontinuities





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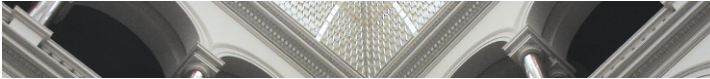
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## ToDo

- **ToDo** schreiben
- **ToDo** abarbeiten



