# IQM Bonus Project

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#### **Animation Link**

### Improved Animation Link (Post Submission)

## 1 Derivation of the First order correction

In order to plot the wave function a function for the first order wavefunction with first-order corrections must be derived. The Hamiltonian can be expressed as:

$$\hat{\mathbf{H}} = \hat{\mathbf{H}}_0 + \hat{\mathbf{H}}_1 \tag{1}$$

Where  $\hat{H}_0$  is the unperturbed Hamiltonian, and  $\hat{H}_1$  is the perturbing Hamiltonian. This is transformed into:

$$\hat{\mathbf{H}} = \hat{\mathbf{H}}_0 + \lambda \hat{\mathbf{H}} \tag{2}$$

Where  $\lambda$  is used to adjust the strength of the perturbation term. The respective eigenstates and eigenvalues of this Hamiltonian are:

$$|\psi_n\rangle = |\varphi_n^{(0)}\rangle + \lambda |\varphi_n^{(1)}\rangle \tag{3}$$

$$E_n = E_n^{(0)} + \lambda E_n^{(1)} \tag{4}$$

Inserting this in to the equation  $\hat{H} |\psi_n\rangle = E_n |\psi_n\rangle$ , we recieve:

$$(\hat{H}_0 + \lambda \hat{H}_1)(|\varphi_n^{(0)}\rangle + \lambda |\varphi_n^{(1)}\rangle + ...) = (E_n^{(0)} + \lambda E_n^{(1)} + ...)(|\varphi_n^{(0)}\rangle + \lambda |\varphi_n^{(1)}\rangle + ...)$$
 (5)

With the zero-th and first order of this respectively being:

$$\hat{\mathbf{H}}_0 \left| \varphi_n^{(0)} \right\rangle = \mathbf{E}_n^0 \left| \varphi_n^{(0)} \right\rangle \tag{6}$$

$$\hat{H}_0 |\varphi_n^{(1)}\rangle + \hat{H}_1 |\varphi_n^{(0)}\rangle = E_n^0 |\varphi_n^{(1)}\rangle + E_n^1 |\varphi_n^{(0)}\rangle$$
(7)

In order to find the first order correction for the eigenstate, we take the inner product of the first order equation with  $\langle \varphi_k^{(0)} |$  with  $k \neq n$ :

$$\langle \varphi_k^{(0)} | \hat{\mathbf{H}}_0 | \varphi_n^{(1)} \rangle + \langle \varphi_k^{(0)} | \hat{\mathbf{H}}_1 | \varphi_n^{(0)} \rangle = \mathcal{E}_n^{(0)} \langle \varphi_k^{(0)} | \varphi_n^{(1)} \rangle + \mathcal{E}_n^{(1)} \langle \varphi_k^{(0)} | \varphi_n^{(0)} \rangle$$
(8)

Where the second term on the RHS goes to zero, we get:

$$\langle \varphi_k^{(0)} | \varphi_n^{(1)} \rangle = \frac{\langle \varphi_k^{(0)} | \hat{\mathbf{H}} | \varphi_n^{(0)} \rangle}{\mathbf{E}_n^{(0)} - \mathbf{E}_k^{(0)}}, \, \mathbf{k} \neq \mathbf{n}$$
(9)

By using the identity operator on the first order eigenstate  $|\varphi_n^{(1)}\rangle$  we get:

$$|\varphi_n^{(1)}\rangle = \sum_k = |\varphi_n^{(0)}\rangle \langle \varphi_n^{(0)}|\varphi_n^{(1)}\rangle \tag{10}$$

And substituting that into the eigenstate of the Hamiltonian we get the first order eigenstate:

$$|\psi_n\rangle = |\varphi_n^{(0)}\rangle + \lambda \sum_{k \neq n} |\varphi_k^{(0)}\rangle \frac{\langle \varphi_k^{(0)} | \hat{\mathbf{H}}_1 | \varphi_n^{(0)} \rangle}{\mathbf{E}_n^{(0)} - \mathbf{E}_k^{(0)}}$$
 (11)

# 2 Analytical Simplifications and Solutions to the Spherical Schrodinger Equation

With a three-dimensional wavefunction described by the eigenket  $|\psi\rangle$  that is dependent on a spatial coordinate r, the wave function can be written as:

$$\langle \mathbf{r} | \psi \rangle = \langle \mathbf{r} | n, l, m \rangle = \mathbf{R}(\mathbf{r}) \Theta(\theta) \Phi(\phi)$$
 (12)

In spherical coordinates, this wave function can be rewritten:

$$\langle \mathbf{r}, \theta, \phi | \psi \rangle = \langle \mathbf{r}, \theta, \phi | n, l, m \rangle = \mathbf{R}(\mathbf{r}) \mathbf{Y}_{l,m}(\theta, \phi)$$
 (13)

The wave function can be solved with the spherical Schrodinger Equation:

$$\frac{-\hbar^2}{2\mu} \left[ \frac{1}{r^2} \frac{\partial}{\partial r} (r^2 \frac{\partial \psi}{\partial r}) + \frac{1}{r^2 \sin(\theta)} \frac{\partial}{\partial \theta} (\sin(\theta) \frac{\partial \psi}{\partial \theta}) \frac{1}{r^2 \sin^2 \theta} \frac{\partial^2 \varphi}{\partial \psi^2} \right] - \frac{e^2}{4\pi\epsilon_0 r} \psi = E\psi \tag{14}$$

The following separated differential equations were used to solve for the radial wave function, azimuthal wave equation, and polar wave function respectively:

$$\frac{\mathrm{d}}{\mathrm{dr}}(\mathrm{r}^2 \frac{\mathrm{dR}}{\mathrm{dr}}) + \frac{2\mu \mathrm{r}^2}{\hbar^2} (\mathrm{E} + \frac{e^2}{4\pi\epsilon_0 \mathrm{r}}) \mathrm{R} - \mathrm{AR} = 0$$
(15)

$$\frac{\sin(\theta)}{\Theta} \frac{\mathrm{d}}{\mathrm{d}\theta} (\sin(\theta) \frac{\mathrm{d}\Theta}{\mathrm{d}\theta}) + A\sin^2(\theta) - \mathrm{B} = 0 \tag{16}$$

$$\frac{1}{\Phi} \frac{\mathrm{d}^2 \Phi}{\mathrm{d}\phi} + \mathrm{B} = 0 \tag{17}$$

 $Y_{l,m}(\theta,\phi)$  is expressed as:

$$Y_{l,m}(\theta,\phi) = \frac{-1^{l}}{2^{l} l!} \sqrt{\frac{(2l+1)(l+m)!}{4\pi(l-m)!}} e^{im\phi} \frac{1}{\sin^{m}(\theta)} \frac{d^{l-m}}{d(\cos^{m}(\theta))} \sin^{2l}(\theta)$$
(18)

The radial equation R(r) has solutions given by the **Laguerre polynomials**. The azimuthal component (provided by the **Legendre polynomials**) and polar components can be combined to form the **Spherical Harmonics**. In order to calculated the first order energy shift caused by a time-independent perurbative Hamiltonian, we will make use of the unique properties of these equations. To properly determine the wave function of a perturbed Hydrogen atom in the ground state, an inner product will have to be computed with all of the possible quantum states an electron can be found in (described by n, l, and m). In order to simplify the the wave function analytically, we can first look to one particular component within  $Y_{l,m}(\theta,\phi)$ :  $(-1)^l$ . This phase factor introduces a unique effect on a given complex conjugate of  $Y_{l,m}(\theta,\phi)$ , where the inner product of two states with the same l state will vanish. This can be shown by applying the parity operator to the azimuthal-polar wave function:

$$\hat{\Pi}(Y_{l,m}(\theta,\phi)) = (-1)^l Y_{l,m}(\theta,\phi) \tag{19}$$

With this state of **definite parity**, the integration of an inner product between two states with the same quantum number l will result in zero because the inner product would be an odd function. Therefore, our first simplification can made with the following condition:

$$l \neq l' \tag{20}$$

Additionally, another simplification can be made due to the relationship between the pertubative Hamiltonian  $\hat{H}_1$  and the orbital angular momentum operator about the z-axis  $\hat{L}_z$ . Because the perturbative Hamiltonian acts along the z-axis, it does not commute with this angular momentum operator because  $\hat{L}_z$  only acts along the XY-plane:

$$[\hat{\mathbf{H}}_1, \hat{\mathbf{L}}_z] = 0 \tag{21}$$

As a result, the eigenvalue associated with  $\hat{L}_z$  must be the same for two different wave functions in order for the Stark Effect to have a non-vanishing effect on the wave function with a given quantum number m. Thus the second simplification can be made with the condition:

$$m = m' (22)$$

Since we have that m=0 for non-vanishing states, we find that only states  $|k \neq 1, 1, 0>$  have a non-vanishing value for  $\langle k, l, m | \hat{\mathbf{H}}_1 | 1, 0, 0 \rangle$ , by the following property (note:  $\mathbf{P}_{l,0} = \mathbf{P}_l$ , i.e. the L-th m=0 associated Legendre Polynomial is just the L-th Legendre Polynomial: We can make the substitution  $x=\cos(\theta)$  and we find the left hand side becomes:

$$\int_0^{\pi} P_L(\cos(\theta)) \cos(\theta) P_N(\cos(\theta)) \sin(\theta) d\theta$$

which matches the form of the integral taken to find  $\langle k, l, m | \hat{H}_1 | 1, 0, 0 \rangle$  for the polar term involving Legendre Polynomials, and thus this term vanishes for  $l \neq 1$ . Thus we have further restricted the non-vanishing terms and we have that the non-vanishing terms will be:

$$|k \neq 1, 1, 0\rangle$$
.

$$\int_{-1}^{1} x P_{L}(x) P_{N}(x) dx = \begin{cases} \frac{2(L+1)}{(2L+1)(2L+3)} & \text{for } N = L+1\\ \frac{2L}{(2L-1)(2L+1)} & \text{for } N = L-1 \end{cases}$$

Figure 1: Legendre Polynomial Identity (https://mathworld.wolfram.com/LegendrePolynomial.html)

To calculate the perturbative effect on the ground state wavefunction, we used the equation for first-order corrections:

$$|\psi_n\rangle = |\varphi_n^{(0)}\rangle + \lambda \sum_{k \neq n} |\varphi_k^{(0)}\rangle \frac{\langle \varphi_k^{(0)} | H_1 | \varphi_n^{(0)}\rangle}{E_n^{(0)} - E_k^{(0)}}$$

The ground state  $|\phi_n^{(0)}\rangle$  is given by  $|1,0,0\rangle$ . In addition, the first-order Hamiltonian is defined as  $e|E|\hat{z}$ . So, the ground state correction is

$$e|E|\sum_{k\neq 1}\sum_{l,m}|k,l,m\rangle \frac{\langle k,l,m|\,\hat{z}\,|1,0,0\rangle}{E_1^{(0)}-E_k^{(0)}}$$

Based on the simplifications from earlier, we know that l=1 and m=0 for the terms that don't disappear.

Next, we project both sides onto the position bra, using P to denote only the perturbative term:

$$P\langle r, \theta, \phi | \psi_n \rangle = e|E| \sum_{k \neq 1} \langle r, \theta, \phi | k, 1, 0 \rangle \frac{\langle k, 1, 0 | \hat{z} | 1, 0, 0 \rangle}{E_1^{(0)} - E_k^{(0)}}$$

Since the spherical harmonics depend only on l and m, all the perturbative wavefunctions must have the same spherical harmonics. So, we can integrate the spherical harmonics and place those outside the sum as prefactors.

Polar part  $(\Theta_{00} = \frac{1}{\sqrt{2}}, \Theta_{10} = \frac{\sqrt{6}}{2} \cos \theta)$ :

$$\int_0^\pi \Theta_{10}(\theta) \cos(\theta) \Theta_{00}(\theta) \sin(\theta) d\theta = \frac{\sqrt{3}}{2} \int_0^\pi \cos^2(\theta) \sin(\theta) d\theta = \frac{1}{\sqrt{3}}$$

Azimuthal part  $(\Phi_0 = \frac{1}{\sqrt{2\pi}})$ :

$$\int_0^{2\pi} (\frac{1}{\sqrt{2\pi}})^2 d\phi = 1$$

So, with our spherical harmonic prefactor of  $\frac{1}{\sqrt{3}}$ , we can express the term in a radial integral:

$$\langle k, 1, 0 | \hat{z} | 1, 0, 0 \rangle = \langle k, 1, 0 | r \cos(\theta) | 1, 0, 0 \rangle = \frac{1}{\sqrt{3}} \int_0^\infty R_{k,1}(r) R_{1,0}(r) r^3 dr$$

## 3 Plotting the Perturbation Effect

We plotted the perturbed wavefunctions with respect to a changing electric field. To find the maximum value of the electric field that should be used to fully capture the effects, we set the potential energies of the non-perturbed and of the first order effects equal to each other at twice the Bohr radius:

$$\frac{e^2}{4\pi\epsilon_0(2a_0)} = e|\mathbf{E}|(2a_0)$$

We say that  $r_c$  is the distance from the x-y plane where the Hamiltonians are similar:

$$\frac{H_1(r_c)}{H_0(r_c)} \sim 1$$

since  $H_1 = e|E|r$  with  $|E| \propto 1/r_c^2$  and  $H_0 \propto 1/r$ , then

$$\frac{H_1(r)}{H_0(r)} \sim \frac{r^2}{r_c^2}$$

Thus, the first order Stark Effect will be about 1/4 the magnitude of the unperturbed Hamiltonian at the Bohr Radius. This makes sense, and is the reason the radius was set to  $2a_0$ . We expect the electron to move opposite the direction of the electric field which is what our model predicts.

The final animation of the first-order correction to the Hydrogen atom is linked <u>here</u>.

EDIT: An improved animation created after the original submission can be found here.

Our code used for the animation is linked https://github.com/schwas3/iqm-bonus-problem.

# 4 References

Solution to the Hydrogen Atom Hyperphysics Hydrogen Atom Wolfram - Laguerre Polynomials Wolfram - Legendre Polynomials