

ORBITAL MECHANICS

Today we will finish the theory behind the planet fly-by.

ENTRY CONDITION

$$\underline{\underline{v}}_{\infty} = \underline{\underline{v}}_{p1} + \underline{\underline{v}}_{\infty}^-$$

EXIT CONDITION

$$\underline{\underline{v}}_{\infty}^+ = \underline{\underline{v}}_{p1} + \underline{\underline{v}}_{\infty}^+$$

$$\Delta \underline{\underline{v}}_{\text{fly-by}} = \underline{\underline{v}}_{\infty}^+ - \underline{\underline{v}}_{\infty}^- = \underline{\underline{v}}_{\infty}^+ - \underline{\underline{v}}_{\infty}^- = \Delta \underline{\underline{v}}_{\infty}$$

! Helio reference

↳ planetary reference

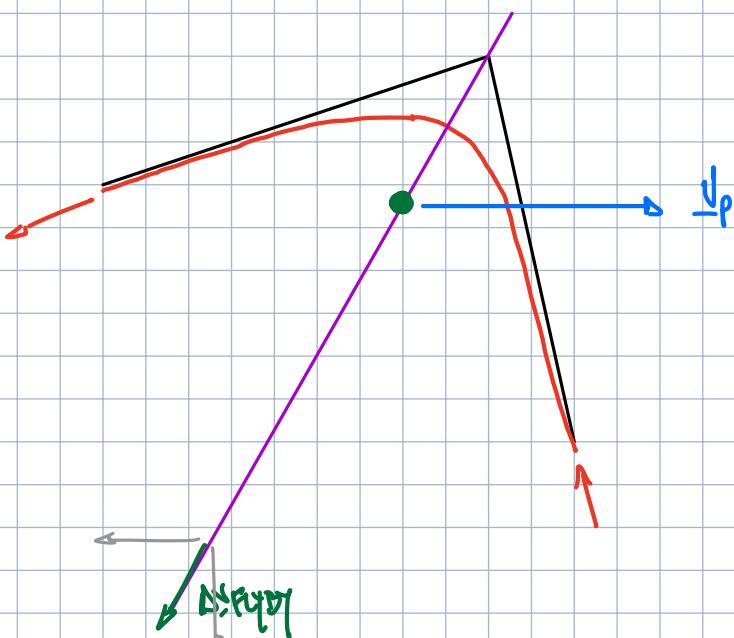
$\Delta \underline{\underline{v}}_{\text{fly-by}}$ → direction of the apse line of the hyperbole & that's why we need to design the hyperbole knowing the $\Delta \underline{\underline{v}}$ I need to give.

$\Delta \underline{\underline{v}}_{\text{fly-by}}$ is directed along the apse line of the HYPERBOLE

$\Delta \underline{\underline{v}}_{\text{fly-by}}$ depends on v_{∞} and β

⇒ Designing the hyperbole so that I can use it as an "heliocentric free maneuver".

LEADING-SIDE HYPERBOLA

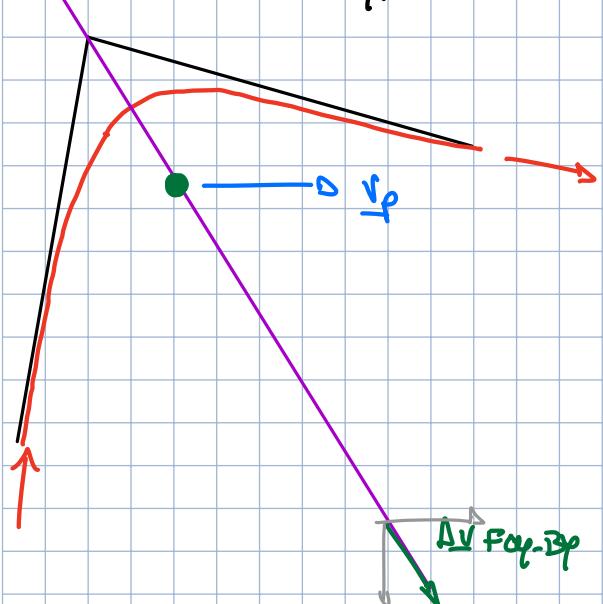


The planet is moving in such a way that the component of $\Delta \underline{\underline{v}}_{\text{fly-by}}$ along the velocity of the planet will be negative.

$\Delta \underline{\underline{v}}_{\text{fly-by}}$ component along the $\underline{\underline{v}}_p$ direction < 0

\Rightarrow Decrease the heliocentric velocity. $|\underline{v}_{SC}^+| < |\underline{v}_{SC}^-|$

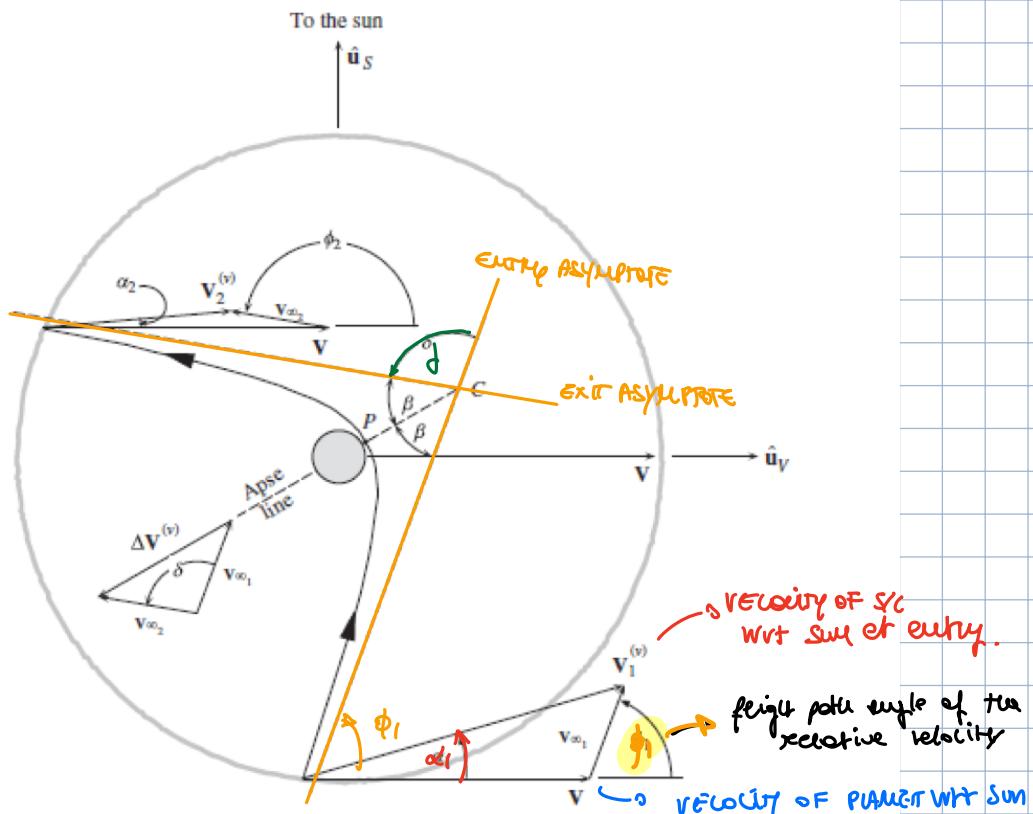
TRAINING-SIDE HYPERBOCA



ΔV_{PUP-BY} COMPONENT ALONG
planet velocity $v_p > 0$
 \Rightarrow INCREASE IN THE HELIOCENTRIC
 $|v_{sc}^+| > |v_{rc}^+|$

HINT: when you pass in front of a planet (leading-side)
the planet sees you \Rightarrow planet "reduces" your velocity
when you pass behind the planet (trailing-side)
the planet doesn't see you \Rightarrow you (src)
"steal" heliocentric energy from the planet.

LEADING SIDE PLANETARY FLY-BY.



Let's define (not needed if we have a fc \rightarrow we have ephemeris)

\hat{u}_v : unit vector in the direction of planet velocity = $\hat{\theta}$ (circular orbit)

\hat{u}_s : unit vector in the direction to the Sun = $-\hat{r}$ (circular orbit)

$\hat{u}_s \perp \hat{u}_v \rightarrow$ as we assume planet circular orbit.

ENTRY OF SOI
notatory cartis

$$\underline{v}_{sc} = \underline{v}_i = [v_i^{sc}]_v \hat{u}_v + [v_i^{sc}]_s \hat{u}_s$$

$$[v_i^{sc}]_v = v_i^{sc} \cos \alpha_1$$

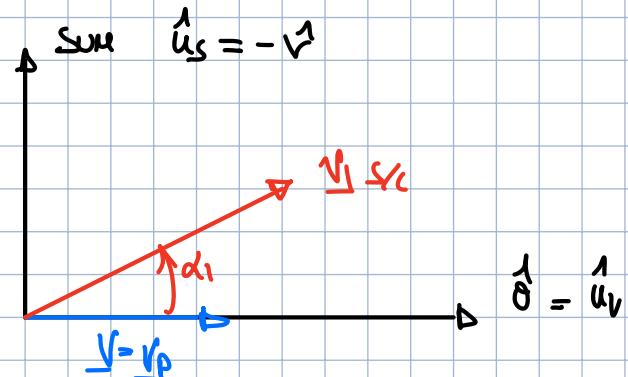
$$[v_i^{sc}]_s = v_i^{sc} \sin \alpha_1$$

α_1 angle between v_i^{sc} and \underline{v} (planet velocity wrt sun)
= flight path angle of v_i^{sc} when it enter the SOI.

$$[v_i^{sc}]_v = v_{\theta 1}$$

$$[v_i^{sc}]_s = -v_{r1}$$

$$\therefore \hat{u}_s = -\hat{r}$$



So we can use r_r and θ

Result

$$v_\theta = \frac{h}{r} (1 + e \cos \theta)$$

$$r_r = \frac{h}{r} (e \sin \theta)$$

Eq (2.34)

Eq (2.33)

$$[v_i^{sc}]_v = v_{\theta 1} = \frac{h_{\text{sun}}}{r_1} (1 + e_1 \cos \theta_1)$$

$$- [v_i^{sc}]_s = v_{r1} = \frac{h_{\text{sun}}}{r_1} (e_1 \sin \theta_1)$$

We are characterizing the velocity of the space craft in the heliocentric pair of ref.

$\left. \begin{matrix} e_1 \\ u_1 \\ \delta_1 \end{matrix} \right\}$ approach trajectory to planet heliocentric leg (green)

$$\underline{v}_p = \underline{v}_p \hat{u}_V$$

$$\underline{v}_p = \sqrt{\frac{\mu_{\text{Sun}}}{r_p}}$$

v_p : planet on circular orbit

② characterize the planetocentric entry velocity.

$$v_{\infty^-} = v_i^{\text{S/C}} - \underline{v}_p$$

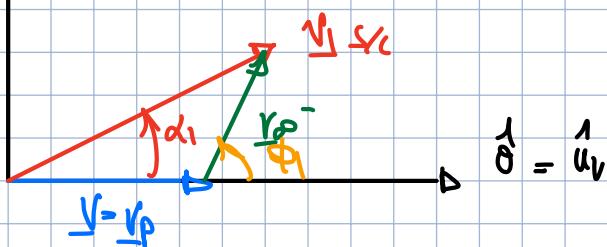
$$\underline{v}_{\infty^-} = [v_{\infty^-}]_V \hat{u}_V + [v_{\infty^-}]_S \hat{u}_S$$

$$[v_{\infty^-}]_V = v_i^{\text{S/C}} \cos \alpha - \underline{v}_p$$

Because we are considering a leading-side fly by.

$$[v_{\infty^+}]_S = v_i^{\text{S/C}} \sin \alpha$$

Sum $\hat{u}_S = -\hat{v}$



$$v_{\infty^-} = \sqrt{v_{\infty^-} \circ v_{\infty^-}} = \sqrt{v_i^{\text{S/C}}^2 + v_p^2 - 2 v_i^{\text{S/C}} v_p \cos \alpha}, \quad (6.4)$$

by solving the triangle of velocities

We can compute v_{∞^-} given $v_i^{\text{S/C}}$ (satellite entry velocity) and planet entry velocity (v_p) and the flight path angle α .

We can compute everything about the hyperbola.

CHARACTERISE THE HYPERBOLA

$$\text{eq (4.21)} \quad u = r_p \sqrt{\frac{v_\infty^2 + 2\frac{1}{r_p}}{\mu}}$$

$$\text{eq (4.21)} \quad e = 1 + \frac{r_p v_\infty^2}{\mu}$$

μ of planet!

r_p, e, v_∞ on hyperbola

so planetocentric ref.

① angle between v_∞ and v_p

$$\phi_1 = \tan^{-1} \left(\frac{[r_\infty]_S}{[v_\infty]_V} \right) = \tan^{-1} \left(\frac{\frac{v_{S/C}}{v_p \sin \alpha_1}}{\sqrt{\frac{v_{S/C}}{v_p \cos \alpha_1} - 1}} \right)$$

$$\boxed{\phi_1 = \tan^{-1} \frac{\frac{v_{S/C}}{v_p \sin \alpha_1}}{\sqrt{\frac{v_{S/C}}{v_p \cos \alpha_1} - 1}}} \quad (4.42)$$

② CHARACTERISE THE HYPERBOLA AT EXIT

$$\phi_2 = \phi_1 + \delta$$

\hookrightarrow turn angle.

$$\text{Recall } \delta = 2 \sin^{-1} \frac{1}{e} \quad \text{eq (4.2)}$$

We have characterize the velocity at exit.

$$v_\infty^+ = v_\infty \cos \phi_2 \hat{u}_V + v_\infty \sin \phi_2 \hat{u}_S$$

$$v_\infty = v_\infty^+ = v_\infty^- \rightarrow \text{In modulus.}$$

③ CHARACTERISE THE HELIOCENTRIC EXIT CONDITIONS

$$\underline{v}_2 = \underline{v}_{S/C}^+ = \underline{v}_p + \underline{v}_\infty^+$$

\hookrightarrow wrt's

$$\underline{v}_2^{sc} = \underline{v}_{sc} = [\underline{v}_2^{sc}]_V \hat{u}_V + [\underline{v}_2^{sc}]_S \hat{u}_S$$

We can compute \underline{v}_2^{sc} given v_∞ , v_p , $\phi_2 = \phi_1 + \delta$

$$[\underline{v}_2^{sc}]_V = v_p + r_\infty \cos \phi_2 = v_\infty$$

$$[\underline{v}_2^{sc}]_S = r_\infty \sin \phi_2 = -v_{r_2}$$

(4.4b)

We can obtain the parameters of the heliocentric exit leg.

$$h_2 = R_p v_\infty \rightarrow h_1 \text{ of hel's leg after fly-by}$$

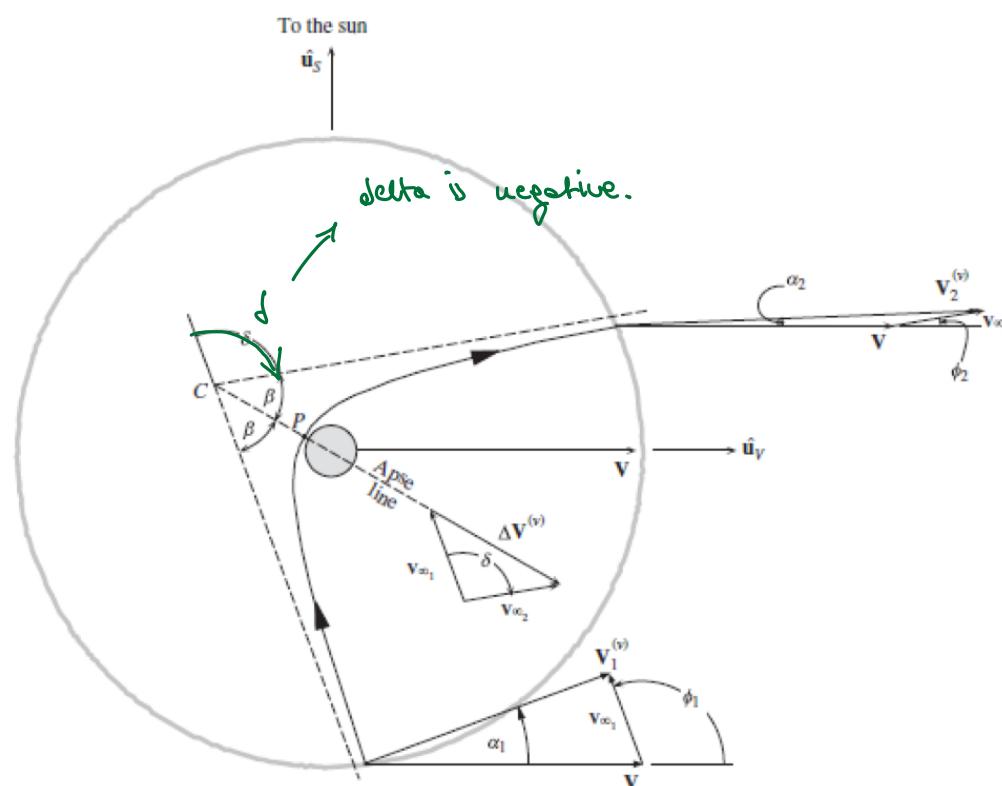
$$\left\{ \begin{array}{l} R_p = \frac{R_2^2}{\mu_{\text{sum}}} \frac{1}{1 + e_2 \cos \phi_2} \\ v_{r_2} = \frac{\mu_{\text{sum}}}{h_2} e_2 \sin \phi_2 \end{array} \right.$$

e_1, e_2, δ heliocentric orbit

R_p = planet radius wrt Sun

Note fly-by is considered as impulsive manoeuvre (as seen from Sun) so R_p is constant during fly-by.

The same procedure is valid for trailing edge fly-by.



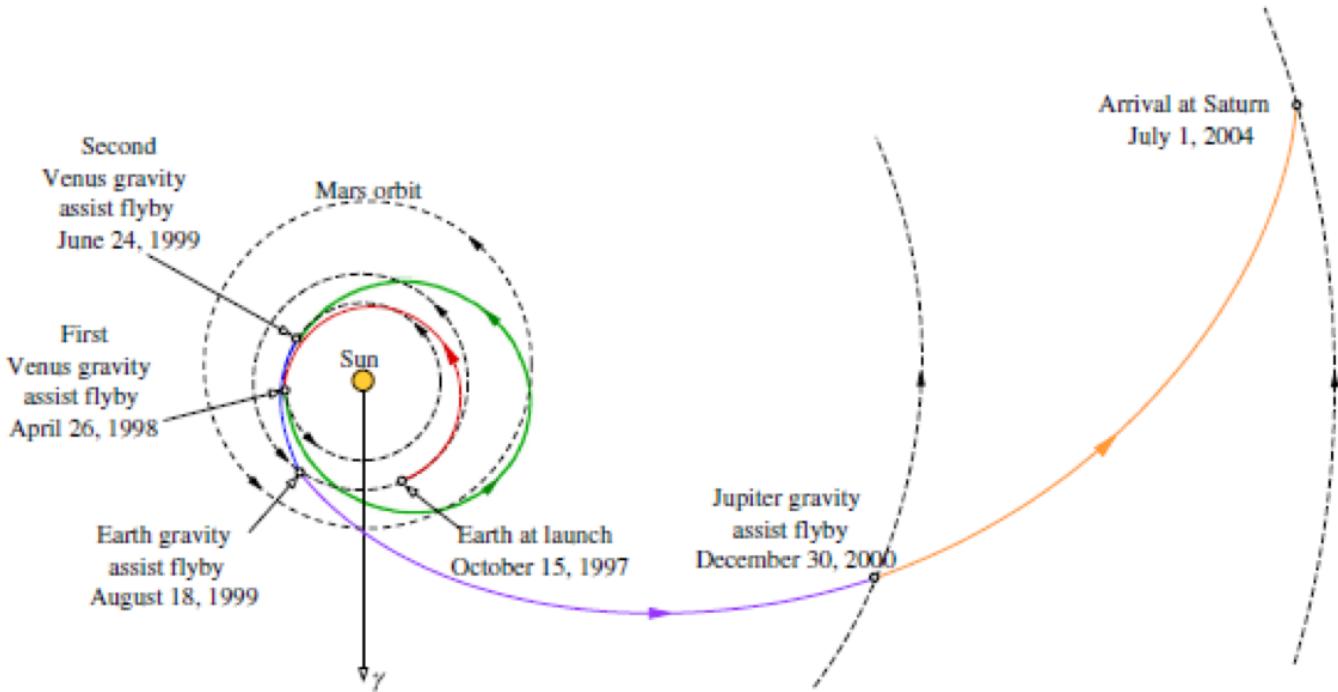
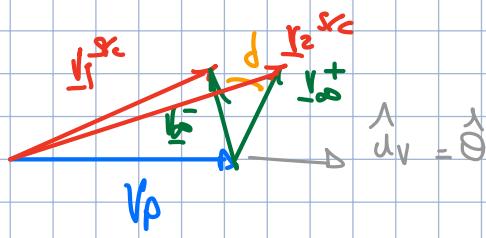


FIGURE 8.24

Cassini's 7-year mission to saturn.

The problem of the trajectory interplanetary design is that they can get really complicated \rightarrow really fast.

INPUT/TOOLS

- planet ephemeris
- Lambert solver
- gravity / fly-by model

We do not know how many fly by we'll do when and with which planet.

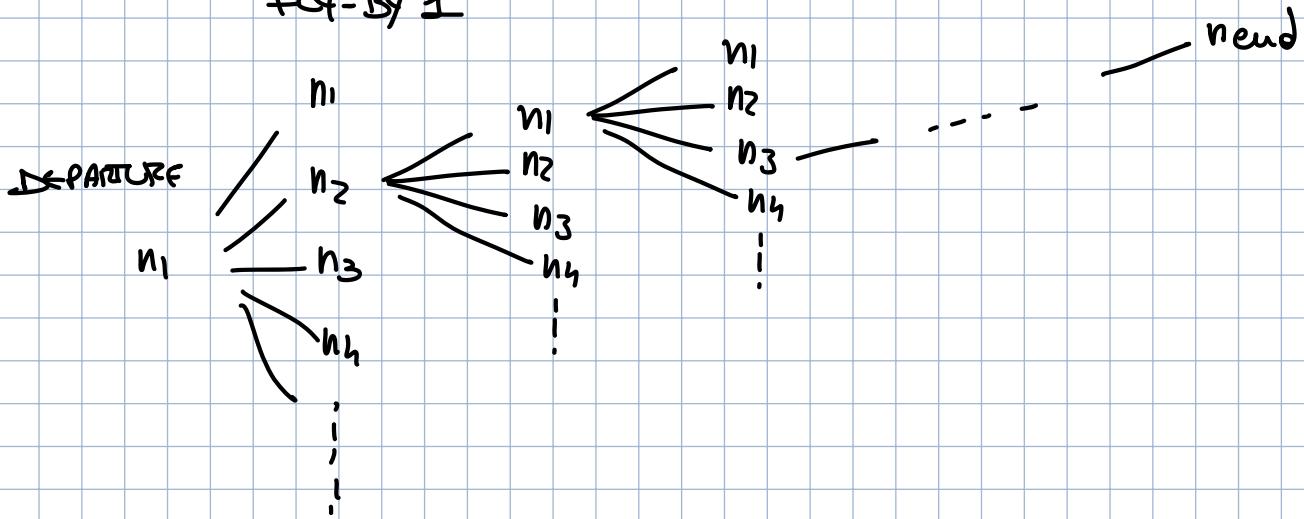
First problem

SEQUENCE SELECTION PROBLEM : sequence of fly-by.
 COMBINATORIAL PROBLEM

$n_1 \dots n_{\text{planet}}$ --- n_{end}
 fly-by

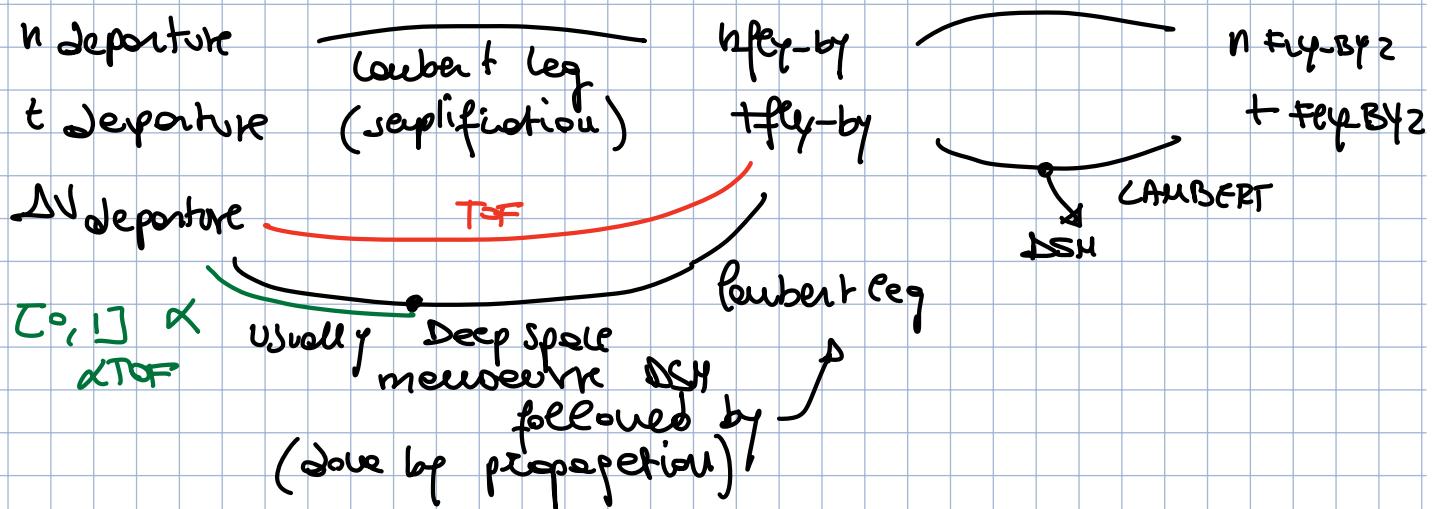
integer programming problem

Fly-By 1



So I need to find a sequence of Problem

Given a sequence.



$$TOF = t_{\text{fly-by}} - t_{\text{departure}}$$

$$x : \begin{cases} t_{\text{departure}} \\ \end{cases}$$

$$\alpha_1, \Delta V_{\text{DSN}}^1(3)$$

$$n_{\text{fly-by}} \\ t_{\text{fly-by}}$$

$$\alpha_2, \Delta V_{\text{DSN}}^2(3) \\ n_{\text{fly-by}} 2 \\ t_{\text{fly-by}} 2$$

↳ optimization vector

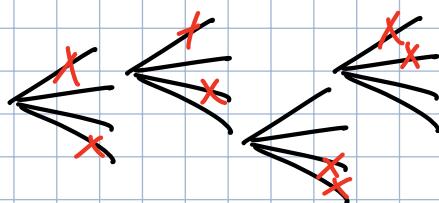
Every transfer we add we add off at the optimisation vector

$$\text{find } x = \left\{ t_{\text{dep}}^1, \alpha_1, \Delta V_{\text{DSH}}^1, t_{\text{FB1}}^1, \alpha_2, \Delta V_{\text{DSH}}^2, \dots \right\}$$

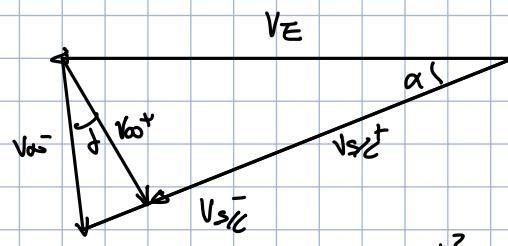
$$\min_x \sum_{\text{TOT}} \Delta V$$

INTEGER - CONTINUOUS OPTIMISATION PB

There are a lot of algorithm →
and they do not consider bad branches
—branch and bound



→ Check at website: global trajectory optimisation competition.



$$V_E = 23.$$

$$V_a^{+2} = V_E^2 + V_{sC}^{+2} - 2V_E V_{sC}^{+} \cos\alpha$$

$$V_{sC}^{+2} + \underbrace{(-2V_E \cos\alpha)}_A V_{sC}^{+} + \underbrace{(V_E^2 - V_a^{+2})}_B = 0$$

$$V_{sC}^{+} = \frac{-A \pm \sqrt{A^2 - 4B}}{2}$$

23.752
 23.709