

Charge Conservation and Higher Moments of Charge Fluctuations

Scott Pratt

*Department of Physics and Astronomy and National Superconducting Cyclotron Laboratory
Michigan State University, East Lansing, MI 48824 USA*

Rachel Steinhorst

*Department of Physics and Astronomy
Michigan State University, East Lansing, MI 48824 USA*

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Higher moments of distributions of net charge and baryon number in heavy-ion collisions have been proposed as signals of fundamental QCD phase transitions. In order to better understand background processes for these observables, models are presented which enable one to gauge the effects of local charge conservation, Bose symmetrization, and volume fluctuations. The models consist of a thermal model superimposed onto a simple parameterization of collective flow, known as a blast-wave, with emission being consistent with individual canonical ensembles. The spatial extent of local charge conservation is parameterized by a patch volume. The sensitivity of third and fourth order moments, skewness and kurtosis, to the patch volume and other parameters, including beam energy and average net baryon multiplicity is explored. Comparisons with STAR data show that a significant part of the observed non-Poissonian fluctuations in net-proton fluctuations are explained by charge-conservation, but that measurements of the STAR collaboration for fluctuations of net electric charge qualitatively differ from expectations of the models presented here.

I. INTRODUCTION

The fluctuation of conserved charges is a standard means by which to investigate and classify phase transitions. At the critical point correlation lengths diverge, which results in peaks in charge fluctuations as one approaches the critical point. For systems with first-order phase transitions, fluctuations turn into phase separation and fluctuations are no longer intensive quantities. The growth of fluctuations becomes increasingly dramatic as one considers progressively higher-order fluctuations. In a volume V , fluctuations of a charge Q can be defined as

$$\mathcal{M}_N \equiv \frac{1}{V} \langle (Q - \bar{Q})^N \rangle = \frac{1}{V} \sum_n P_n (n - \bar{n})^N, \quad (1)$$

when particles have unit charge. The measure is increasingly sensitive to the tails of the multiplicity distribution, P_n , as N increases. The free energy, $F \sim a_2(n - \bar{n})^2 + a_3(n - \bar{n})^3 \dots$, is minimized for $n = \bar{n}$, but the quadratic part vanishes at the critical point, $a_2 \rightarrow 0$, which allow the fluctuations to grow and be dominated by higher-order terms. The shape of the tails of the distribution are then profoundly altered.

The properties of the QCD transition, deconfinement and the restoration of chiral symmetry, are not well understood at finite baryon density. There exists the possibility that this transition is a true phase transition, with a critical point at several times nuclear density and with a critical temperature close to the pion mass. If this is the case, it begs the question as to whether the conditions for phase separation or for critical phenomena can be reproduced in the laboratory. Heavy ion collisions at high energy, measured at the Relativistic Heavy Ion Collider (RHIC) or at the LHC, can produce mesoscopic regions at temperatures of a few hundred MeV, which is well above the expectations for a critical temperature, and densities of several times nuclear matter density, which might or might not be sufficiently high to investigate the phase transition.

High-energy heavy ion collisions are characterized by strong explosive collective flow. Measurement is confined to the outgoing asymptotic momenta, but because of strong flow, correlations in coordinate space manifest themselves as correlations in relative momentum. Thus, measurements of correlations binned by relative momentum or charge fluctuations within some defined region of momentum space serve as surrogates for the corresponding observables in coordinate space. Indeed, measurements of charge and baryon number fluctuations have been performed at RHIC. By adjusting the beam energy of the colliding nuclei, experiments at RHIC have explored conditions at which novel phase phenomena might occur. Fluctuations of electric charge and baryon number have been especially popular. An initial scan of beam energies [2–4, 8, 9, 13–18] was rather inconclusive, but measurements with greatly improved statistics are currently being undertaken and analyzed. The main thrust of these studies is to search for evidence of a QCD phase transition with a critical point at finite baryon density, with the hope that the phenomena one expects for an idealized equilibrated system [44, 45] becomes manifest in the measured debris of heavy-ion collisions.

In addition to the finite system size (event multiplicities might number in the thousands), the novel states of matter created in heavy-ion collisions persist for $\lesssim 10$ fm/c. This severely limits the degree to which phases can separate or to which critical fluctuations can grow. This limits the degree to which conserved charges can separate from one another. For example, if a strange and an anti-strange quanta are produced together their separation is limited by diffusion. To justify thermal models of fluctuations of conserved charge based on equilibration, sufficient time is required for particles to enter and exit the defining volume.

The goal of this paper is to gauge the degree to which charge-balance correlations affect higher-order correlations. In addition to making it difficult for phases to separate or for critical correlations to emerge, conservation also represents its own source of correlation, which needs to be understood as a potential source of background before making firm arguments to have observed phenomena related to phase transitions. It is well known that charge-balance correlations are readily measurable at the two particle level, $N = 2$ in Eq. (49), and that they explain the bulk of the $N = 2$ fluctuation measurement. However, their impact on $N = 3, 4$ fluctuations has not been investigated in great detail. For instance, the four-particle measure of correlation,

$$C_4 = \mathcal{M}_4 - 3\mathcal{M}_2^2, \quad (2)$$

which is based on cumulants, subtracts much of the contribution to \mathcal{M}_4 coming from purely two-particle correlations. However, charge conservation can involve multiple particles, and the degree to which a cumulant-based measure, like the kurtosis, is affected by charge conservation is not fully understood. Relations based on a uniform acceptance probability and for a single type of charge were worked out in [31], which provide significant insight into how higher-order correlations are affected by charge conservation. The goal of this study is to extend such ideas to a more realistic picture, which takes into account the conservation of all three types of charge (baryon number, electric charge and strangeness) and applies a more realistic model of experimental acceptance and efficiency. The interplay of charge conservation with chemical equilibrium, decays and Bose statistics are all considered.

To understand the role of chemical equilibrium and decays, a model is presented which creates small volumes in which the net charge, B, Q and S , is fixed at some value. Even if the net charges are all zero, charged particles exist

but in combinations that conserve the net charge. Theoretical methods for exact calculation of the canonical ensemble and a method for Monte Carlo generation of statistically independent events are presented here. The method lends itself to including decays and accounting for experimental acceptance and efficiency. The physical picture of treating small volumes as independent patches was previously done for calculation of charge balance functions in [26, 28], and was also applied in [32]. Following the terminology in [32], we sometimes refer to these sub-volumes as patches. The method presented here creates perfectly independent samplings, and can generate billions of such patches within a few hours. This enables highly accurate calculations of higher moments with minimal numeric costs.

After a brief review of cumulants and the definitions of skewness and kurtosis in Sec. II, the method for exact solution of the canonical ensemble describing a multi-component, multi-charge hadron gas is presented in Sec. III. These techniques extend those used for canonical ensembles used to study isospin fluctuations of a hadron gas [34], nuclear fragmentation [33], the level density of a Fermi gas [22], and the effect of restricting a quark-gluon plasma (QGP) to having fixed charge, including being in an overall color singlet [35]. Exact methods for calculating correlations up to fourth-order are presented. Unfortunately, to include realistic acceptance effects and complex decays, the exact expressions are no longer viable. However, as shown in Sec. IV, the exact expressions show how sample events can be generated. Each event, defined by a set of particles and momenta, is generated with perfect independence from the others, and perfectly reproduces the canonical expressions from Sec. III. Section V extends the previous sections to show how Bose correlations can be included. Section VI considers the case of a single type of charge and a uniform efficiency, i.e. the probability of any particle being recorded is set to some fixed value. Much of this discussion repeats what is said in [?], and is included for completeness. This provides for a physical discussion of how charge conservation, volume fluctuations, chemical equilibrium, decays, clustering, and Bose corrections should affect higher moments.

The heart of this study is presented in Sec. VII. Here, the patches are assigned collective velocities consistent with the collective flow deduced from heavy-ion collisions. A canonical sampling of particles are generated from each patch, followed by a simulation of their decays. The particles were then overlaid onto the acceptance of the STAR detector at RHIC. Each patch is uncorrelated with any other patch, so moments can be calculated by averaging over the independent contributions from the patches. Results are displayed alongside results from the STAR Collaboration. The size and sign of the fluctuations of the net-proton distributions are consistent with observations, but the calculations of the net charge distributions differs qualitatively from STAR observations. A detailed discussion of the lessons derived from this study is presented in Sec. VIII.

II. CUMULANTS, SKEWNESS AND KURTOSIS

For the manuscript to be more self-contained, a brief review of cumulants and the definition of skewness and kurtosis is presented here.

Cumulants of a charge distribution are defined by

$$C_1 = \langle Q \rangle, \quad (3)$$

$$C_2 = \langle (Q - \overline{Q})^2 \rangle, \quad (4)$$

$$C_3 = \langle (Q - \overline{Q})^3 \rangle, \quad (5)$$

$$C_4 = \langle (Q - \overline{Q})^4 \rangle - C_2^2, \quad (6)$$

where Q is the net charge. Here, Q might refer to baryon number, to strangeness, or to the electric charge measured in units of e .

Rather than showing the cumulants C_n , ratios are presented to help minimize trivial dependences on system size. The skewness, S , is a measure of the third moment,

$$S = \frac{C_3}{C_2^{3/2}}. \quad (7)$$

This definition has the advantage in being dimensionless, but it does not become independent of volume in the limit of large volumes. Thus, it is more common to consider the ratio

$$S\sigma = S\sqrt{C_2} = \frac{C_3}{C_2}, \quad (8)$$

which becomes an intensive measure in the limit of larger volumes. The kurtosis is a measure of four-particle correlations,

$$K = \frac{C_4}{C_2^2}, \quad (9)$$

but instead of K , one typically chooses

$$K\sigma^2 = \frac{C_4}{C_2}, \quad (10)$$

where $\sigma^2 = C_2$, to find an intensive measure of the fluctuation. For a measure to be intensive, it should be independent of volume in the large-volume limit.

The measures $S\sigma$ and $K\sigma^2$ approach simple values in the limit that the distributions would be Poissonian. For Poissonian emission the observation of a charge in one region of momentum space is uncorrelated with the emission into any other space. Thus, particles are correlated only with themselves. If charges appear only in integral positive units, one can apply the usual expression for the Poissonian moments where the mean is η ,

$$C_1 = \bar{n} = \eta \quad (11)$$

$$C_2 = \langle (n - \bar{n})^2 \rangle = \eta, \quad (12)$$

$$C_3 = \langle (n - \bar{n})^3 \rangle = \eta = C_1, \quad (13)$$

$$C_4 = \langle (n - \bar{n})^4 \rangle - 3C_2^2 = \eta. \quad (14)$$

If there exist both positive and negative charges, the distribution of the net charge can be derived by convoluting the two distributions. Convoluting two Poissonians results in a Skellam distribution. If the mean number of positives is η_+ and the mean number of negatives is η_- , the distribution of net charge for a Skellam distribution, $Q = n_+ - n_-$, yields the following cumulants

$$C_1 = \eta_+ - \eta_-, \quad (15)$$

$$C_2 = \eta_+ + \eta_-, \quad (16)$$

$$C_3 = \eta_+ - \eta_- = C_1, \quad (17)$$

$$C_4 = \eta_+ + \eta_- = C_2. \quad (18)$$

Thus, if charges are produced in an uncorrelated fashion in increments of either ± 1 , the skewness and kurtosis become

$$S\sigma = \frac{C_3}{\sigma^2} = \frac{C_1}{\sigma^2}, \quad (19)$$

$$K\sigma^2 = \frac{C_4}{\sigma^2} = 1, \quad (20)$$

where $\sigma^2 \equiv \langle (Q - \bar{Q})^2 \rangle = \eta_+ + \eta_-$. Sec. VI considers C_3/C_1 rather than $S\sigma$, as that ratio is unity in the limit that emission is uncorrelated. However, because experimental analyses have often concentrated on the two ratios above, higher order correlations in Sec. VII, where calculations are compared to STAR results, will be presented in terms of these ratios.

Moments depend on the efficiency α with which particles are measured. In the limit of vanishing efficiency all distributions of positives or of negatives tend to become Poissonian [19], and the distribution of the net charge will thus become Skellam. This can be understood by having the probability of observing a positive charge being $\alpha \rightarrow 0$ and the probability of observing two charges being proportional to α^2 being negligible. This assumption would fall through if charges are measured in pairs, i.e. multiple charges on each measured particle, but for final-state hadrons the charges are only ± 1 . As $\alpha \rightarrow 0$, the moments are dominated by the probability of observing either zero charges or a single charge, and the moments for counting particles of a given charge are

$$\langle n^m \rangle = \alpha, \quad (21)$$

for all $m > 0$. It is then easy to see that $\langle (n - \bar{n})^m \rangle = \alpha$, which quickly leads to the result that as $C_4/C_2 = C_3/C_1 = 1$.

If the net charge is fixed, a non-perfect efficiency is required to produce fluctuations. For fixed charge, the efficiency divides into two sets, the measured and non-measured. Each set fluctuates equally, but oppositely, relative to the mean. Thus, all the even moments of net charge will have an even reflection symmetry about an efficiency of $1/2$, and the odd moments will have an odd symmetry,

$$C_n(1/2 + \delta\alpha) = \begin{cases} C_n(1/2 - \delta\alpha), & n = 2, 4, 6 \dots \\ -C_n(1/2 - \delta\alpha), & n = 3, 5, 7 \dots \end{cases} \quad (22)$$

III. RECURSIVE TECHNIQUES FOR GENERATING CANONICAL PARTITION FUNCTIONS

For non-interacting particles the canonical partition function can be calculated exactly, or at least to the level that all partitions of $A \leq A_{\max}$ hadrons are taken into account, with the exact solution being reached with $A_{\max} = \infty$.

For our case, we conserve three quantities: the electric charge Q , the baryon number B and the strangeness S . For states i with energies E_i , the partition function,

$$Z(Q, B, S) = \sum_{i, Q_i=Q, B_i=B, S_i=S} e^{-\beta E_i}, \quad (23)$$

where Q_i, B_i and S_i are the discrete values of the conserved quantities for the state i , can be calculated recursively. The function $Z_A(Q, B, S)$ refers to the subset of states with A hadrons,

$$Z(Q, B, S) = \sum_{A \geq 0} Z_A(Q, B, S). \quad (24)$$

The recursive procedure begins with

$$Z_{A=0}(0, 0, 0) = 1, \quad (25)$$

the canonical partition function of the vacuum. The contribution for a given A , $Z_A(Q, B, S)$, can be written as

$$Z_A(Q, B, S) = \frac{1}{A} \sum_h z_h Z_{A-1}(Q - q_h, B - b_h, S - s_h), \quad (26)$$

where z_h is the single-particle partition function for hadron species h , which has charges q_h, b_h and s_h . This was proved in [33], and can be understood by realizing that one can count all the ways to arrange A hadrons with a given charge by considering all the ways to arrange one hadron multiplied by all the ways to arrange the remaining hadrons. To avoid double counting, a factor of $1/A$ is applied. For a fixed charge the probability to have A hadrons is,

$$P(A) = \frac{Z_A(Q, B, S)}{\sum_A Z_A(Q, B, S)} = \frac{Z_A(Q, B, S)}{Z(Q, B, S)}. \quad (27)$$

In practice, the sum over A is cut off at some A_{\max} , but in our studies here that cutoff is made large enough that contributions to Z for $A > A_{\max}$ are negligible. Thus, once one builds the partition function from $A = 0$ to A_{\max} one has the partition function for all Q, B, S .

Once the partition function is calculated one can also calculate the multiplicities and moments of observing specific species. For example, the multiplicity of species h in a system with charge Q, B, S is

$$\langle N_h \rangle = z_h \frac{Z(Q - q_h, B - b_h, S - s_h)}{Z(Q, B, S)}. \quad (28)$$

This also provides expressions for the various charges, e.g.,

$$\langle Q \rangle = \sum_h q_h z_h \frac{Z(Q - q_h, B - b_h, S - s_h)}{Z(Q, B, S)}. \quad (29)$$

Spectra can also be calculated

$$\frac{dN_h}{d^3p} = \frac{(2s_h + 1)\Omega}{(2\pi\hbar)^3} e^{-E_h(p)} \frac{Z(Q - q_h, B - b_h, S - s_h)}{Z(Q, B, S)}. \quad (30)$$

Higher moments can also be extracted,

$$\langle N_h N_{h'} \rangle = \delta_{hh'} z_h \frac{Z(Q - q_h, B - b_h, S - s_h)}{Z(Q, B, S)} \quad (31)$$

$$+ z_h z_{h'} \frac{Z(Q - q_h - q_{h'}, B - b_h - b_{h'}, S - s_h - s_{h'})}{Z(Q, B, S)}. \quad (32)$$

It is straightforward to extend this expression to the fluctuation of charges.

These expressions can also be extended to consider non-additive conservation laws. Net isospin conservation of a hadron gas was invoked in [34], i.e. restricting the states to being in an iso-singlet. Quark-gluon states restricted to being in both an iso-singlet and a color singlet are addressed in [35]. Bose and Fermi corrections are discussed in Sec. V.

For the case of a single kind of charge, one can see how the recursive method above yields the same result as what one would expect by writing down the partition function for a system with $(A - Q)/2$ negative charges and $(A + Q)/2$ positive charges as shown in [31],

$$Z_{A,Q} = \begin{cases} \frac{z^A}{[(A-Q)/2]![(A+Q)/2]!}, & A - Q \text{ is even,} \\ 0, & A - Q \text{ is odd.} \end{cases}, \quad (33)$$

where z is the partition function of a single charge. One can readily see that this is consistent with the recurrence relations,

$$Z_{A,Q} = \frac{z}{A} \{Z_{A-1,Q-1} + Z_{A-1,Q+1}\} \quad (34)$$

$$\begin{aligned} &= \frac{z}{A} \left\{ \frac{z^{A-1}}{[(A-Q-2)/2]![(A+Q)/2]!} + \frac{z^{A-2}}{[(A-Q)/2]![(A+Q-2)/2]!} \right\} \\ &= \frac{z}{A} \left\{ \frac{z^{A-1}(A-Q)/2}{[(A-Q)/2]![(A+Q)/2]!} + \frac{z^{A-2}(A+Q)/2}{[(A-Q)/2]![(A+Q)/2]!} \right\} \\ &= \frac{z^A}{[(A-Q)/2]![(A+Q)/2]!}. \end{aligned} \quad (35)$$

This result is also equivalent to expectations based on setting reaction rates equal. If one assumes that pairs are created with some rate β , and that they are destroyed with some rate $\alpha N_+ N_-$, where N_+ and N_- are the number of positive and negative charges, $N_+ + N_- = A$. Setting the rates equal,

$$\alpha \frac{(A-Q)(A+Q)}{4} Z_{A,Q} = \beta Z_{A-2,Q}. \quad (36)$$

One can see that if one chooses $\beta/\alpha = z$ that Eq.s (36) and (33) are consistent. If the net charge is zero, the result is even simpler,

$$\begin{aligned} P(A|Q) &= \frac{Z_{A,Q}}{\sum_A Z_{A,Q}}, \\ P(A|Q=0) &= \frac{z^A}{[(A/2)!]^2} \left\{ \sum_{A=\text{even}} \frac{z^A}{[(A/2)!]^2} \right\}^{-1}. \end{aligned} \quad (37)$$

Aside from the assumptions that Q is fixed and that there exists only one kind of charge, Eq.(36) also requires that Bose and Fermi quantum statistical corrections are negligible, and that only unit charges exist. Despite these shortcomings, this picture is useful in that it allows one to see how multiplicity fluctuations are affected by charge conservation in a simple picture.

IV. GENERATION OF UNCORRELATED SAMPLE EVENTS

Complicated experimental acceptances are difficult to incorporate into expressions for the moments. It is then easiest to generate entire events via Monte Carlo, and filter the events through the acceptance. The Monte Carlo procedure involves choosing a hadron proportional the number of ways the system might have such hadron, i.e. a product of the partition function of the individual hadron multiplied by the partition function of the remainder. The procedure becomes:

1. Calculate and store the partition function, $Z_A(Q, B, S)$, up to some size $A \leq A_{\max}$ for all Q, B, S that might ultimately couple back to a given $A = A_{\max}$ for the given total values Q, B, S .
2. For total charge Q, B, S , choose the number of hadrons A proportional to $Z_A(Q, B, S)/Z(Q, B, S)$, where $Z(Q, B, S) = \sum_{A \leq A_{\max}} Z_A(Q, B, S)$.
3. Choose a hadron h proportional to the probability $z_h Z_{A-1}(Q-q_h, B-b_h, S-s_h)/Z_A(Q, B, S)$. If Bose degeneracy is to be taken into account this procedure is slightly modified as described in Sec. V.
4. Choose the momentum proportional to the thermal weight $e^{-E_p/T}$.
5. Repeat (3,4) but with A, Q, B, S being replaced by $A-1, Q-q_h, B-b_h, S-s_h$. The procedure is finished when $A=0$.

Bose effects for pions can be included by altering the second and third steps above. This is addressed in Sec. V.

Storing the partition function can require substantial memory for larger A_{\max} because the indices Q, B and S must also vary over a range of order $\pm A_{\max}$, so memory usage roughly scales with A_{\max}^4 . Because one is usually interested in calculations with total charge near zero, one can ignore partition functions for charges that cannot couple back to the fixed overall charge at A_{\max} . Once A exceeds $A_{\max}/2$, the calculations here cutoff values of Q, B and S that

could not ultimately affect the $Q = B = S = 0$ partition function for $A = A_{\max}$. Even with this savings, partition functions with $A_{\max} = 250$ could require approximately 13 GB of memory, and need on the order of 10 minutes to calculate on a single processor. For $A_{\max} = 125$, less than a GB of memory was needed and partition functions could be calculated in less than a minute. For hadron gases at temperatures of 150 MeV, $A_{\max} = 250$ was readily sufficient for patch volumes $\lesssim 700 \text{ fm}^3$. If multiple patch volumes are to be explored for the same temperature, computational time can also be saved by realizing that the partition functions scale as Ω^A . Thus, if one performs a calculation for some initial volume Ω_0 , scaling can provide results for new volumes with minimal computation.

Once the partition function is calculated, event generation is remarkably fast. The time to generate an event scales linearly with the volume, or equivalently, linearly with the average number of particles generated. Running sufficient events to generate a million individual particles can be accomplished within a few seconds on a single CPU. Unlike Metropolis methods where events are modified by considering small changes to existing events, such as in [32], each event in this method is independent as long as the random numbers are without correlation.

V. BOSE AND FERMI STATISTICS

Including Bose and Fermi statistics into the recursive relations for partition functions is straight-forward, and was shown in [22, 34]. The method is related to that used for calculating the effects of multi-boson interference for pion interferometry [41]. In a fixed volume the partition function can be first treated as the usual procedure of accounting for n identical particles being in different single-particle states. This includes the $1/n!$ term to account for the fact that particles are indistinguishable, i.e. the Gibbs paradox. If m_ℓ indistinguishable particles are in the same single-particle state ℓ , one must correct the weight by a factor of $m_\ell!$ for each level, which can also be thought of as the analog of the symmetrized relative wave function with all the momenta being equal. For fermions, the weight becomes $(-1)^{\ell-1} m_\ell$. As demonstrated in [22], the recurrence relation to the partition function then becomes

$$Z_A(Q, B, S) = \frac{1}{A} \sum_h \sum_n Z_{A-n}(Q - nq_h, B - nb_h, S - ns_h) z_{h,n} (\pm 1)^{n-1}, \quad (38)$$

where $z_{h,n}$ is the partition function for n particles in some level,

$$z_{h,n} = \sum_\ell e^{-n\beta\epsilon_\ell}, \quad (39)$$

where ℓ refers to single-particle levels of energy ϵ_ℓ . The ± 1 refers to bosons or fermions. For hadron gases in the high temperature environments of relativistic heavy ion collisions, only pions have a non-negligible correction from quantum degeneracy. The correction of order n for any level ℓ is of the order $e^{-\beta\epsilon_\ell}$ lower than the previous term. This factor is largest for zero momentum, and for pions becomes $e^{-\beta m}$, where m is the pion mass. For the zero-momentum level at a temperature of 150 MeV, the factor is $e^{-m/T} \approx 0.4$, and as the system cools the factor falls slightly [40]. For a more characteristic thermal momentum the factor is ≈ 0.1 . For heavier particles the factor is always small. For example, for a zero-temperature ρ meson the factor is a fraction of a percent.

Given that symmetrization is only being applied to pions, which are bosons, one can incorporate these corrections into the Monte Carlo procedure outlined in Sec. IV. For fermions this might be problematic because of the negative weights coming from the $(-1)^{n-1}$ factors in Eq (38), but fortunately this is unnecessary because the degeneracy of fermions is negligible in the systems considered here. For pions, the algorithm is adjusted by treating each value of n as being a different species, with charges nq_h and with the partition function calculated with a reduced temperature $T \rightarrow T/n$. If one picks such a species in step 3 of the algorithm, n pions are generated, all with the same momentum. For a finite system, the pions would have a small relative momentum of order the inverse system size.

It is well known that bosonic effects can broaden multiplicity distributions, consistent with negative binomial distributions [24, 25], making them super-Poissonian. One of the goals of this study is to discern how bosonic statistics alter the skewness and kurtosis.

VI. MODELS WITH UNIFORM EFFICIENCY AND FIXED CHARGE

Even for a volume of fixed charge, finite efficiency and acceptance leads to non-zero fluctuations. The degree to which these fluctuations affect the skewness and kurtosis was worked out in [31] for emission of a fixed charge where the probability of any charge being observed was a constant α . If α were zero or unity, there would be no fluctuations, and because the charge on those particles that are not observed must fluctuate exactly opposite to the charge that is observed, the even moments must be symmetric about $\alpha = 1/2$, and the odd moments must be anti-symmetric. One of the most important results of [31] is that for fixed Q and α the ratios of cumulants depend only on α and the variance and mean of the underlying multiplicity distribution. Even though Q is fixed, the net number of charged particles M can fluctuate.

From [31], the probability that M charged particles with total charge Q will result in a measured charge $q = n_+ - n_-$ due to a uniform efficiency α is the convolution of two binomial distributions

$$P(q|M, Q) = \sum_{n_+=0}^{(M+Q)/2} \sum_{n_-=0}^{(M-Q)/2} \frac{[(M+Q)/2]![(M-Q)/2]!}{[(M+Q)/2 - n_+]![(M-Q)/2 - n_-]!n_+!n_-!} \alpha^{n_++n_-} (1-\alpha)^{M-n_+-n_-}. \quad (40)$$

After some tedious algebra, one can find the cumulants for fixed multiplicity,

$$\begin{aligned} C'_1 &= \bar{q} = \alpha Q, \\ C'_2 &= \langle (q - \bar{q})^2 \rangle_M = \alpha(1-\alpha)M, \\ C'_3 &= \langle (q - \bar{q})^3 \rangle_M = \alpha(1-\alpha)(1-2\alpha)Q, \\ C'_4 &= \langle (q - \bar{q})^4 \rangle_M - 3\langle (q - \bar{q})^2 \rangle_M^2 \\ &= \alpha(1-\alpha) - 6\alpha^2(1-\alpha)^2 M. \end{aligned} \quad (41)$$

Here, the primes emphasize that the averages $\langle \dots \rangle_M$ denote that they consider only those events with fixed base multiplicities M . The even moments are all linear in M , while the odd moments are linear in Q . It might appear that due to the linearity one might replace M with $\langle M \rangle$ for a fluctuating base multiplicity, but the second term in the fourth cumulant

$$-3 \sum_M P(M) \langle (q - \bar{q})^2 \rangle_M^2 \neq -3 \langle (q - \bar{q})^2 \rangle^2. \quad (42)$$

Instead,

$$\langle (q - \bar{q})^2 \rangle^2 = \alpha^2(1-\alpha)^2 \bar{M}^2. \quad (43)$$

This provides a contribution to C_4 and the cumulants and their ratios become [31],

$$\begin{aligned} C_1 &= \alpha Q, \\ C_2 &= \alpha(1-\alpha)\bar{M}, \\ C_3 &= \alpha(1-\alpha)(1-2\alpha)Q, \\ C_4 &= \alpha(1-\alpha)\bar{M} - 6\alpha^2(1-\alpha)^2\bar{M} \\ &\quad + 3\alpha^2(1-\alpha)^2\langle (M - \bar{M})^2 \rangle, \\ \frac{C_2}{C_1} &= (1-\alpha)\frac{\bar{M}}{Q}, \\ \frac{C_3}{C_1} &= (1-\alpha)(1-2\alpha), \\ \frac{C_3}{C_2} &= (1-2\alpha)\frac{Q}{\bar{M}}, \\ \frac{C_4}{C_2} &= 1 + 3\alpha(1-\alpha)(\omega_M - 2), \\ \omega_M &\equiv \frac{\langle (M - \bar{M})^2 \rangle}{\bar{M}}. \end{aligned} \quad (44)$$

The relative variance of the multiplicity of the base distribution, ω_M , is unity for a Poissonian distribution. In that case C_3/C_1 and C_4/C_2 fall below unity for non-zero α . The assumptions for this relation are that the charge and volume are fixed, that the efficiency is uniform, and that Bose and Fermi symmetrization is ignored. This relation is important as it implies that if one understands the efficiency and the second moment of the base multiplicity distribution, one can construct a baseline of cumulant ratios, and attribute any deviation from the baseline as due to fluctuations in Q , which is precisely the goal.

A reasonable value of α can be taken from balance function analyses. If a charge is observed, there should be a balancing charge emitted nearby, and detected with probability α . This should correspond to the integrated strength of the charge balance function, which for electric charge is in the neighborhood of 0.4. Of course, α is not a constant. If a charge is observed in the center of the detector, its balancing charge has a better chance of being observed than for one observed near the periphery of the acceptance. Nonetheless, for the purposes of roughly setting expectations, this suggests that C_4/C_2 and C_3/C_1 can significantly differ from unity. Next, we consider how various effects, aside from fluctuating the base charge Q , might push ω_M to be either super-Poissonian, $\omega_M > 1$ or sub-Poissonian. We discuss the effects of volume fluctuations, equilibrated charge distributions, decays and Bose condensation.

A. Volume Fluctuations

Heavy ion experiments measure collisions spanning a range of impact parameters. Even for a fixed impact parameter, energy deposition might significantly vary depending on how many nucleons actually collided or how many jets were produced. If more energy is deposited into a fixed volume, it might expand further before it hadronizes, resulting in larger final volumes. Experimental analyses attempt to minimize these fluctuations by constraining a given fluctuation measurement to a specific centrality bin, where “centrality” might be defined by multiplicity, transverse energy, or energy deposition in a forward calorimeter. To reduce auto-correlation centrality measurements are usually constrained to particles other than those used to construct the moments. Nonetheless, it is inevitable that a range of initial conditions is explored within any centrality bin, and one might thus expect the base multiplicity distribution to broaden. If the patch volumes are fixed at some constant value, volume fluctuations can be thought of as a fluctuation in the number of patches.

It is tempting to expect that the ratio of cumulants would be independent of the number of patches, and thus impervious to volume fluctuations. Independently, each cumulant, C_n , scales linearly with the number of patches because there are no cross correlations between patches. Thus, if $P_p(N)$ is the probability of having N patches, and if c_n is the cumulant for a single patch, the cumulant for the overall system is

$$C_n = \sum_N P_p(N) N c_n, \quad (45)$$

and the ratio of cumulants is

$$\begin{aligned} \frac{C_n}{C_m} &= \frac{\sum_N P_p(N) N c_n}{\sum_M P_p(M) M c_m} \\ &= \frac{c_n}{c_m}. \end{aligned} \quad (46)$$

Thus, the ratio of cumulants is independent of the number of patches, and thus independent of the overall volume, but the ratio can still depend on the volume of an individual patch.

Unfortunately, although the ratios of cumulants are independent of the overall volume, they are not impervious to fluctuations of the overall volume, or equivalently, to fluctuations of the number of patches [?]. To illustrate this, one can consider a system with a distribution $P_p(N)$ as in Eq. (45), where the average number of patches is \bar{N} . For $n > 1$ the cumulants are expressed in terms of $\delta Q = Q - \bar{Q}$. Because \bar{Q} varies with the number of patches we rewrite it as

$$\begin{aligned} \delta Q &= \delta Q_N + \delta N \bar{Q}_1, \\ \delta Q_N &\equiv Q - N \bar{Q}_1, \end{aligned} \quad (47)$$

where $\bar{Q}_1 = c_1$ is the average charge in a single patch. Here δQ_N is the fluctuation of the charge for a specific number N of patches, and $\delta N = N - \bar{N}$. Inserting these definitions into the expressions for the cumulants,

$$\begin{aligned} C_1 &= \kappa_1 c_1, \\ C_2 &= \sum_N P_p(N) \langle (\delta Q_N + \delta N \bar{Q}_1)^2 \rangle = \kappa_1 c_2 + \kappa_2 c_1^2, \\ C_3 &= \sum_N P_p(N) \langle (\delta Q_N + \delta N \bar{Q}_1)^3 \rangle = \kappa_1 c_3 + \kappa_3 c_1^3, \\ C_4 &= \sum_N P_p(N) \langle (\delta Q_N + \delta N \bar{Q}_1)^4 \rangle - 3C_2^2 \\ &= \kappa_1 c_4 + 3\kappa_2 c_2^2 + 6\kappa_3 c_2 c_1^2 + \kappa_4 c_1^4, \end{aligned} \quad (48)$$

where κ_n are the cumulants of $P_p(N)$,

$$\begin{aligned} \kappa_1 &= \sum_N N P_p(N) = \bar{N}, \\ \kappa_2 &= \sum_N (N - \bar{N})^2 P_p(N), \\ \kappa_3 &= \sum_N (N - \bar{N})^3 P_p(N), \\ \kappa_4 &= \sum_N (N - \bar{N})^4 P_p(N) - 3\kappa_2^2. \end{aligned} \quad (49)$$

If the number of patches are fixed, i.e. the overall volume does not fluctuate, then $\kappa_{n>1} = 0$ and each cumulant satisfies $C_n = \bar{N}c_n$. The ratio of cumulants then cancels the factor \bar{N} . Once the volume fluctuates the ratios of cumulants of the charge distribution depends on the ratio of cumulants of $P_p(N)$. Methods have been constructed to eliminate the dependence volume fluctuations [36–39]. These methods are built on the assumption that the observable used to identify the volume is correlated to the fluctuating charge only through the fact that they both scale with the volume. If the phase space used for the centrality measure is clearly distinct of the phase space over which charge is measured, this should be a good approximation.

Equations (48) and (49) are built on the assumption that the average charge scales linearly with the number of patches. This differs from the assumptions going into Eq. (44), where it was assumed that charge was fixed within the overall volume. If the total charge in the volume is absolutely fixed despite fluctuations in the number of patches, then volume fluctuations affect the answer in that the relative variance of the multiplicity distribution, ω_M , is increased by volume fluctuations as described in Eq. (44).

B. Chemically Equilibrated Distribution

One can also calculate ω_M for an equilibrated system of a single type of charge, where the measurement directly samples the equilibrated distribution. This ignores decays, which in later stages proceed without the regeneration needed to maintain equilibrium. The results of the equilibrated canonical ensemble in Eq. (37) can be used to find ω_M , and thus generate the moments using Eq. (44). For large average multiplicities, \bar{M} , the value of ω_M approaches unity, the Poissonian limit. For small \bar{M} it approaches two. This is expected, because for a fixed charge, you can only add particles pair-wise. In fact, as will be shown in the next section, if one generated a Poissonian number of pairs, the relative variance would be $\omega_M = 2$ for all multiplicities. With ω_M near unity for any system of significant size the values of C_4/C_2 and C_3/C_1 stay below unity as shown in Fig. 1.

C. Decays

In addition to the results from the canonical ensemble, a system of uncorrelated neutral particles that decays to charged particles is also presented to illustrate the difference in how charge conservation manifests itself in fluctuation observables depending on whether the charges arise directly from an equilibrated system, or whether they arise from the decays of uncorrelated neutral particles. Here, the multiplicity of the uncorrelated neutral particles is assumed to be Poissonian. If each neutral particle decays into N_{bodies} charged particles, where the charges are ± 1 , the charged particle multiplicity distributions become

$$\begin{aligned} \langle M_{\text{ch}} \rangle &= N_{\text{bodies}} \bar{M}_0, \\ \langle (M_{\text{ch}} - \bar{M}_{\text{ch}})^2 \rangle &= N_{\text{bodies}} \langle (M_0 - \bar{M}_0)^2 \rangle, \\ \omega_M &= \frac{\langle (M_{\text{ch}} - \bar{M}_{\text{ch}})^2 \rangle}{\bar{M}_{\text{ch}}} \\ &= N_{\text{bodies}} \frac{\langle (M_0 - \bar{M}_0)^2 \rangle}{\bar{M}_0}. \end{aligned} \tag{50}$$

Fig. 1 illustrates how this picture affects C_4/C_2 . In this case, because ω depends only on N_{bodies} and does not change with multiplicity or system size, the cumulants are also independent of multiplicity. For $N_{\text{bodies}} = 2$ the cumulant ratios does not even depend on the efficiency. For $N_{\text{bodies}} > 2$ the ratio C_4/C_2 exceeds unity, so if charge creation proceeded through the creation and decay of neutral clusters it would be easy to generate large values of C_4/C_2 . This has been discussed at length in [42].

In an equilibrated system decays and recombination have equal rates. However, as the system decouples recombination stops and decays proceed until only stable hadrons remain. Very few decays proceed via more than one charged pair. During the hadronic phase the number of charged particles nearly doubles due to these decays. Thus, if the systems do equilibrate, then decay after chemical freeze-out, one would expect the ratio C_4/C_2 to lie somewhere between the value for the canonical ensemble in Fig. 1 and the value for pure decays with $N_{\text{bodies}} = 2$. STAR's experimental results for net proton fluctuations indeed satisfy this expectation, but their measured C_4/C_2 for net charge fluctuations well exceed unity. This discrepancy could be due to volume fluctuations or the decay of larger clusters, or perhaps some shortcoming in this overly simple picture, e.g. the neglect of other conservation laws or Bose condensation.

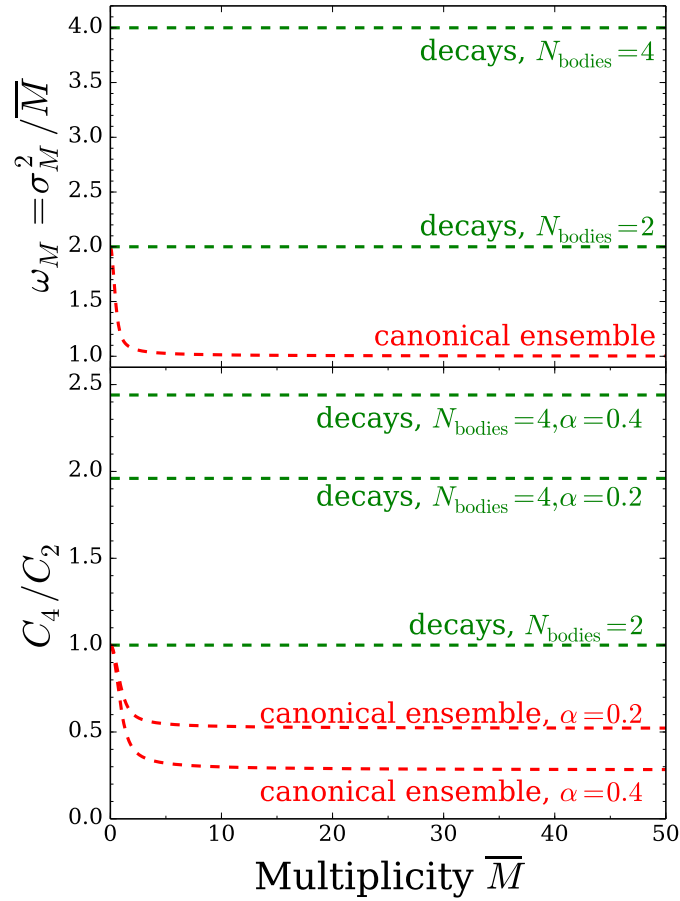


FIG. 1. The relative variance of the charged multiplicity distribution is shown for three cases in the upper panel for a system carrying only a single type of \pm unit charge. For a neutral equilibrated system ω approaches the Poissonian limit for higher mean multiplicities, \overline{M} . This contrasts to a system where one has a Poissonian distribution of neutral particles which all decay into pairs. In this case $\omega = 2$ and if the number of charged particles coming from the decay, N_{bodies} , of the neutral decay were n , one would have $\omega = n$. Using Eq. (44) from [?] the ratio C_4/C_2 is plotted for the cases above for different fixed efficiencies α . For $\omega < 2$, C_4/C_2 falls below unity, and for $\omega > 2$ the ratio exceeds unity. The difference from unity is strongest for α near $1/2$, and the result is symmetric for reflecting α about $1/2$. For α approaching zero or unity $C_4/C_2 \rightarrow 1$.

D. Bose correlations

It has long been understood that Bose correlations induce super-Poissonian fluctuations. Here, we illustrate how Bose effects combine with charge conservation to determine C_4/C_2 . Partition functions for a gas of positive and negative pions, $m_\pi = 139.57 \text{ MeV}/c^2$, were considered to be kinetically equilibrated but with a chemical potential enforcing a fixed average density. If a gas is created at chemical equilibrium one expects $\mu = 0$, but if it cools while maintaining a fixed number of pions, and if the pion number is fed by decays, a non-zero value of μ is required. At decoupling one expects kinetic temperatures to fall near 100 MeV and the pion chemical potential to grow to perhaps as high as 75 MeV [40]. This estimate can be understood by the fact that the phase space occupancies should stay roughly constant for fixed entropy per particle for an expansion at fixed entropy, which suggests a chemical potential of approximately 50 MeV by the time the system cools to 100 MeV. Decay products can then further feed the phase space occupancy. Pion condensation occurs when this chemical potential reaches the pion mass. In the absence of Bose effects, this requires roughly doubling the phase space density of pions as compared to the expectations above. If such conditions were realized Bose effects could result in super-radiance [1], which should be accompanied by large multiplicity fluctuations.

As a function of the effective chemical potential μ , the partition function for a pion gas in a volume of 500 fm^3 and at a temperature of 100 MeV was calculated from Eq. (38). For μ in the range of 75 MeV the relative variance of the charged multiplicity distribution is $\omega_M \approx 1.2$, which modestly increases C_4/C_2 as shown in Fig. 2 for a fixed efficiency of $\alpha = 0.3$. More dramatic results for C_4/C_2 are not expected unless the chemical potential reaches within a



FIG. 2. Fluctuations for a canonical ensemble at fixed charge, $Q = 0$, with an effective chemical potential, μ , applied to adjust the net pion number in a volume of 500 fm^3 at a temperature of 100 MeV . The relative variance of the multiplicity distribution (dashed red line) and the ratio C_4/C_2 of the net-charge distribution (solid green line) grow dramatically as μ approaches the pion mass. Super-radiant effects can take place once μ reaches $m_\pi = 139.57 \text{ MeV}$. Heavy-ion collisions are expected to decouple with effective chemical potentials near 75 MeV , which is well below the onset of large fluctuations.

few MeV of the pion mass. As discussed above, this is not expected. However, if the number of pion sources fluctuated wildly from one event to another, and if there were some events with twice the number of sources emitting into the same phase space, super-radiance might occur in some small fraction of the events. Such behavior would strongly contradict expectations based on chemical equilibrium.

E. Hadron Gas

Thus far, all the simple examples presented in this section considered a system with one type of conserved charge, but in actuality a hadron gas obeys the conservation of three types of charge: baryon number, strangeness and electric charge. This invalidates the use of the recurrence relation of Eq. (44) and requires the application of Eq. (26), or if Bose corrections are included, Eq. (38). Here, we calculate the canonical ensemble, $Z_A(Q, B, S)$, using the recurrence relation, then apply the Monte Carlo techniques of Sec. IV to generate sample sets of particles. The calculation is based on a large number, ~ 300 , of hadron resonances listed in [30]. After generating the particles, unstable resonances are decayed. For this section, a uniform efficiency α is assumed, independent of species, momenta, or whether the products came from a weak decay (charged pions and kaons were not decayed).

Emission is assumed to come from sub-volumes, or patches, of fixed size. Because each volume is independent, there are no correlations between the various sub-volumes. Also, because cumulants scale linearly with the number of sub-volumes, ratios of cumulants can depend only on the patch volume, not on the overall volume. However, if the number of patches fluctuates, the result would be modified by volume fluctuations as shown in Sec. VIA.

Here, the sensitivity to three parameters is investigated. The three parameters are the patch volume, the baryon density, and the fixed efficiency. When one parameter is varied, the other two are set at default values. The default is that the temperature is set at $T = 150 \text{ MeV}$, the baryon density is set to $\rho_B = 0$, the efficiency is $\alpha = 0.3$ and the patch volume is 200 fm^3 . The electric charge density is set to half of the given baryon density. After these sensitivities are studied, calculations at finite baryon density are modified so that the net baryon and electric charges responsible for the non-zero charge density are allowed to fluctuate across sub-volumes according to a Poissonian distribution. This is motivated by the fact that charges transported far away from mid-rapidity at early times cause this non-zero baryon density, so local charge conservation should play little role in the distribution of net charge across patches.

The dependence on patch size is exhibited in Fig. 3. For small patches the emission of baryons is discouraged because each baryon must be accompanied by the existence of an anti-baryon. Thus, the Boltzmann factor, $e^{-M/T}$, for the mass is compounded by the necessity of having a second accompanying massive particle. Once the volumes become larger, and the mean number of baryon pairs exceeds unity, the observation of a baryon might be accompanied by seeing one fewer baryon in the remainder of the system, rather than seeing an extra anti-baryon. In fact, this is exactly what happens in a canonical ensemble in the large volume limit. If a baryon is observed in some small amount of phase space, the mean number of baryons observed in the remainder of the system is reduced by one half. The mean number of anti-baryons then increases by one half. It is expected that the dependence of cumulant ratios on

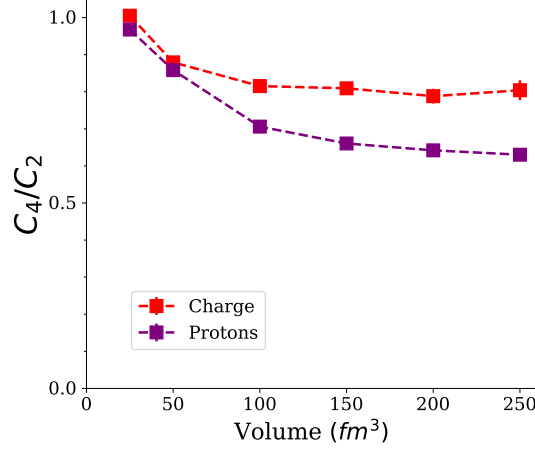


FIG. 3. The ratio of cumulants is shown for a $\rho_B = 0$ system as a function of the patch volume. For small patches a baryon is always accompanied by an extra anti-baryon, whereas for large patches the observation of a baryon is as likely to reduce the chance of observing an extra baryon as it is to enhance the observation of anti-baryons.

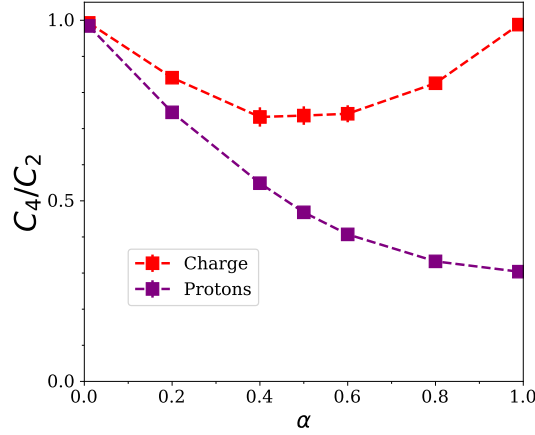


FIG. 4. As the fixed acceptance probability α approaches zero, distribution become Poissonian and the ratios, C_4/C_2 , approach unity. Even for perfect acceptance, the fluctuations are non-zero for a multi-charge system because the conserved charges carried by a proton can be balanced by an array of other species, and charge within the proton and anti-proton sector is not conserved. Thus, the ratio is not symmetric about $\alpha = 0.5$ as one might infer from Eq. (44), because protons and anti-protons only represent a fraction of the species that carry electric charge and baryon number.

patch size should disappear as the patch size increases beyond the threshold for the mean baryon density to be unity. Figure ?? displays this behavior by displaying the ratio C_4/C_2 as a function of patch volume. Because the mean multiplicity for charged particles exceeds unity at a smaller volume, the ratio levels off faster for net charge than for net protons. Qualitatively, the same behavior was seen for a system with a single type of charge in Fig. 1.

The sensitivity to the acceptance probability, α , is displayed in Fig. 4. As α approaches zero, distributions tend to become Poissonian, because the mean number of particles is only zero or unity. Eq. (44) describes similar behavior for a system of a single type of charge. For net charge, the ratio is symmetric about $\alpha = 0.5$. This is because, since net charge is conserved, any charges not measured fluctuate exactly the same as those that are measured. The same property is apparent in Eq. (44). However, the distribution of net protons considers only protons and anti-protons, so it does not represent the entirety of particles that carry baryon number or electric charge. Thus, that distribution is not symmetric about $\alpha = 0.5$ and fluctuations are large even for $\alpha \rightarrow 1$.

Figure 5 presents C_4/C_2 and C_3/C_1 as a function of baryon density. The patch volume is fixed at 200 fm³ and the densities chosen correspond to a fixed number, 4, 8, 12, \dots , of baryons. The choice is made to display C_3/C_1 rather than C_3/C_2 because C_3/C_1 would be unity for uncorrelated emission. The ratios all decrease, moving further from unity, as baryon density increases. The densities displayed cover the range of what might be reached in heavy-ion systems at the point when temperatures fall below 150 MeV.

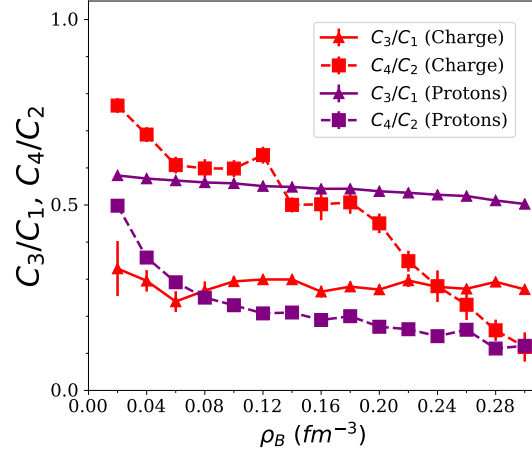


FIG. 5. The ratios C_4/C_2 and C_3/C_1 are shown for net charge and for net baryon number as a function of the baryon density. The ratios fall increasingly below the Poissonian limit of unity as the baryon number is increased. For this calculation the net baryon charge was fixed at $B = \rho_B V$ and the net electric charge was fixed at $Q = B/2$.

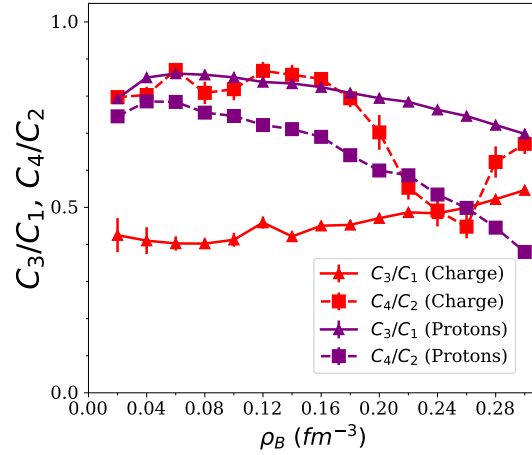


FIG. 6. The ratios C_4/C_2 and C_3/C_1 are shown for net charge and for net baryon number as a function of the baryon density. Calculations differ from those in Fig. 5 in that the net charge fluctuates according to a Poissonian distribution. The ratios rise relative to those in the fixed-charge case, but they remain below unity.

For the calculation illustrated in Fig. 5, the net baryon number, electric charge, and strangeness were all fixed. Because of local charge conservation and the limits of diffusion, balancing charges created after the collision are constrained to stay within the same neighborhood. This constraint motivated the choice of patches with finite volumes. However, even for measurements at mid-rapidity, non-zero charges can arise due to the transfer of baryon number and electric charge from the projectile and target rapidities. This transfer happens either due to hard collisions or diffusion of baryon number from higher baryon density regions. These intruder charges are not typically balanced by charges within the acceptance of the detector, and thus their fluctuation should be considered separately. Figure 6 shows how C_4/C_2 and C_3/C_1 behave as a function of baryon density, if the non-zero baryon number and electric charge are chosen to fluctuate within the patch according to a Poissonian distribution. As expected, the ratios rise as compared to the fixed case shown in Fig. 5, but they do not exceed unity. Discerning these fluctuations of the actual net charges is of particular interest. These results suggest that the starting point for the ratios C_4/C_2 and C_3/C_1 is below unity for random uncorrelated net charges accompanied by an ensemble of baryon charges.

VII. BLAST WAVE MODEL WITH A FULL HADRON GAS AND COMPARISON TO STAR RESULTS

The calculations of the previous section were based on a simple picture, where each sub-volume emitted particles whose probability of being observed was uniform, denoted by α . In practice, this probability depends on where the sub-volume is located within the overall reaction volume. A sub-volume in the region with spatial rapidity, $|\eta_s| > 1$, emits particles that are only rarely detected in by detectors that specialize in mid-rapidity measurements. Even for a sub-volume centered at mid-rapidity, thermal motion can spread some of its emission to rapidities outside the acceptance. This is especially true when the experiments narrow their acceptance. For example, STAR's acceptance nominally covers ± 1 units of pseudo-rapidity, but the acceptance for identified particles is confined to ± 0.9 units. For real rapidity, the effective acceptance is narrowed further, due to the fact that the true rapidity y is less than the pseudo-rapidity η . This difference magnifies for massive, or slower particles. In fact, STAR analyses of identified particles often make cuts and only consider particles with true rapidities $-0.5 < y < 0.5$. Of course, even for particles within the rapidity acceptance, the particles must exceed some minimum transverse momentum and the efficiency for being detected is imperfect. For particles identified only by charge, the efficiencies are typically $\gtrsim 80\%$, and for identified particles the efficiency falls by another few tens of percent.

Collective flow plays a critical role. First, longitudinal flow is what allows the measurement of rapidity to correlate to a measurement in coordinate-space rapidity. In the Bjorken model of boost invariant flow, the mapping is simple in that a fluid element with spatial rapidity η_s moves with a rapidity η_s , and particles emitted from a sub-volume with rapidity η_s have rapidities $y \approx \eta_s$. The mapping is smeared by the thermal motion. Collective radial flow and cooling combine to better align η_s and y . The thermal spread for pions is ≈ 0.6 units of rapidity, while that for protons is ≈ 0.25 .

The size of the sub-volume clearly drives the result. If the sub-volume is small, there is an enhanced probability that a charge and its balancing charge will both be identified and cancel one another. If the sub-volume extends over a large rapidity range, the effects of charge balance are minimized. The extent of the sub-volume in the transverse direction is less important. For sub-volumes at the edge of the fireball, which have more collective velocity, there is a modest increase of having balancing charges both pushed into the acceptance, and avoid the low p_t cutoff, but that sensitivity is modest compared to the extent of the sub-volume in spatial rapidity. The transverse size also plays a role, as illustrated in the previous session, in that the canonical suppression is important for smaller overall volumes.

The intent of this study is to overlay canonical sub-volumes onto a blast-wave model of collective flow and to filter the calculations with the STAR acceptance and efficiency. This should be sufficiently realistic to make meaningful comparisons to STAR data. The three main parameters for the blast-wave model typically describe the radial flow u_\perp , the kinetic freeze-out temperature, T_k , and the temperature of chemical freeze-out, T_c . The chemical freeze-out temperature, T_c , is chosen to fit relative particle yields, while T_k and U_\perp are determined by simultaneously fitting the spectra of species with varying mass, typically π, K, p . A variety of parameterizations exist, such as having the velocity increase linearly from the origin, or the transverse rapidity, or having a sharp cutoff in radius vs. assuming a Gaussian profile. Depending on the choice, the value of T_k and u_\perp vary, but are typically in the neighborhood of $T_k \approx 100$ MeV and $u_\perp \approx 0.6$. For increasing multiplicities, the reaction volumes can expand and cool further, which leads to modestly increased values of u_\perp and modestly decreased values of T_k for either more central or for more energetic collisions. For this study the chemical freeze-out temperatures were taken from [], which extracted T_c for a variety of beam energies. For this section, Bose statistics were ignored. Over 300 hadron species were included in the analysis. The flow and collective velocities were

Here η_0 is a uniformly distributed random value, with the uniformity being motivated by the assumption of boost invariance. The parameter σ_η describes how far particles created from the same sub-volume may have separated by the time of emission. For increasingly larger values of σ_η , the chance that any two observed particles are correlated decreases. One goal of this section is to evaluate the sensitivity of fluctuations to Ω_c and σ_η . For this specific blast-wave model, each sub-volume was assigned a radial transverse velocity according to a Gaussian,

$$\frac{dN}{d^2u_\perp} \approx e^{-(u_x^2 + u_y^2)/2u_\perp^2}. \quad (51)$$

Each sub-volume was also assigned a mean longitudinal rapidity. In the Bjorken picture the distribution of such rapidities would be uniform, but in order to better express the reality that the beams were not of infinite energy, the distribution of sub-volume rapidities was assigned with a Gaussian weight, with the width dependent on the beam energy so as to crudely describe the rapidity distribution. The width was chosen as

$$\sigma_{\text{beam}} = \sqrt{(0.4y_{\text{beam}})^2 - \sigma_\eta^2}, \quad (52)$$

where y_{beam} is the rapidity of the beam in the center-of-mass. With this choice, the overall rapidity width of the distribution is within 10% of the experimental results from STAR reported in [43].

The calculations of this section are filtered through a simplified model of the STAR detector's acceptance and efficiency. For unidentified particles, pseudo-rapidities are required to be less than 1.0 and transverse momenta are constrained to be above 150 MeV/c. An efficiency of 0.8 is applied. For identified particles, additional cuts are added. Rapidities are required to be less than 0.9 and transverse momenta are cut off above 1.5 GeV/c. The efficiency is lowered to 0.6. In order to compare results to measurement from STAR the baryon densities and chemical freeze-out temperatures were mapped to beam energy according to the analysis of [?], which extracted chemical freeze-out temperatures and chemical potentials by considering ratios of particle yields. The parameter, u_\perp , which controls the transverse radial flow was chosen along with the kinetic breakup temperature, T_k , to simultaneously fit the mean transverse momenta of both the protons and pions. Decays were simulated, and the products of weak decays (aside from pions or charged kaons) were included in the analysis. Undoubtedly, a more realistic model of the acceptance might change the ratios, but given that these are ratios, and that the overall efficiency and acceptances are not wildly off, a more rigorous model of the acceptance is unlikely to change the result by more than a few percent.

Figure ?? illustrates how results are sensitive to the size of the sub-volume. For a volume of 100 fm³ the average number of hadrons is several dozen. For very small sub-volumes, where in a grand canonical ensemble the typical number of charges would be zero or one, the thermodynamic cost of having a second charge to balance the first charge reduces the probability of having any charges, and thus lowers both the multiplicities and the moments. For larger sub-volumes, the thermodynamic cost vanishes as the system might have had fewer balancing charges of the same sign, rather than an extra balancing charge of the opposite sign. The characteristic volume for the disappearance of canonical suppression is the volume where the mean number of pairs exceeds unit. For charged particle this sets in at around 10 fm³, but for baryons the characteristic volume is closer to 100 fm³ because of their being heavier and fewer. Thus, the baryon moments for volumes of 100 fm³ and 200 fm³ differ more noticeably. The time for a fluid element to expand and cool to thermal equilibrium tends to be on the order of 5 fm/c, less time on the periphery, at a lower beam energy or at less centrality, and more time for a fluid element at the center, at higher beam energy, or in a more central collision. The maximum transverse distance a charge can travel is 5 fm, but if charge moves diffusively, the separation should be significantly less. As discussed below, charge can separate further along the beam axis, and because that is not well understood, the size of the sub-volume carries a large uncertainty. Anywhere from 50 fm³ to a few hundred fm³ might be reasonable.

Charge can spread further in the longitudinal direction, but that distance depends sensitively on when the charges are created. Matter thermalizes at a point with large collective velocity gradients along the beam axis. If the motion is diffusive, the separation along the beam axis depends logarithmically on the ratio of the final time to the initial creation time. Thus, if a pair are created at 0.2 fm/c, there separation for times, $0.2 < \tau < 1.0$ fm/c are as important as the additional separation they gain from $1.0 < \tau < 5.0$ fm/c. For this reason the size of charge spread, σ_η , in spatial rapidity might be anywhere between a few tenths of a unit of rapidity to a full unit. Figure ?? shows the sensitivity of the moments to this parameter. For large σ_η the observation of a charge is less likely to influence the observation of a second charge, which is similar to having a lower efficiency. For low efficiencies one expects the behavior to behave more like a Poissonian, and that C_4/C_2 and C_3/C_1 would be closer to unity. Indeed this is the case.

Finally, we compare to STAR data. Figure ?? provides two comparisons. In the upper panel, each sub-volume has the same fixed charge, with the net electric charge being set to half the net baryon charge. The ratios C_4/C_2 and C_3/C_2 are both below the experimental results. The difference for net charge is especially striking. In the lower panel, the same procedure is followed with one exception. Rather than keeping the net charge fixed at the same value for each sub-volume, the baryon number and electric charge are assigned values according to Poissonian distributions, where the means are chosen to match the same fixed values as applied in the upper panel. Physically, this picture is based on having two distinct sources of charged particles. The first source comes from baryons and charge that originated from the target and projectile, and were moved into the mid-rapidity region. Given that those charges are displaced by several units of rapidity, their charge balance is likely accounted for well outside the range of the measurement, and it is not unreasonable to think of their contribution to each sub-volume as being random. The second source of charge would be that coming from strings, radiation, and decays. The charges from these processes are not able to transport large distances, hence the phenomenological picture based on sub-volumes. As expected, the ratios displayed in the lower panel are higher than those of the upper panel, for which charges were fixed to the same value for each sub-volume.

For net-baryon fluctuations, where the ratio C_4/C_2 from STAR is mainly below unity, it seems that charge conservation is the main source of non-Poissonian fluctuation. The experimental results seem higher, but statistical uncertainties make it difficult to make strong conclusions. In contrast, the net-charge fluctuations are far above the calculations of this paper, with the ratios roughly a factor of two higher than the calculations. It is difficult to see what missing physics might explain the discrepancy. If large volume fluctuations were responsible, one would expect the baryon fluctuations to also differ more substantially. As seen in Sec. ??, Bose effects for pions would be unlikely to change the result more than 10%. Phase separation would be expected to lead to significant fluctuations in baryon number, and it might be responsible for some of the model/data mismatch of the baryon fluctuations. However, it is

more difficult to see how phase separation would lead to larger net-charge fluctuations than net-baryon fluctuations.

A principal impediment to interpreting the STAR results is in understanding the role of volume fluctuations. This could be better constrained by simultaneously investigating multiplicity fluctuations with the centrality bins and kinematic cuts used for net-charge and net-baryon fluctuations. Net-charge and net-baryon fluctuations should also not be investigated independently, as adjusting parameters to explain net-baryon number behavior while not reproducing net-charge fluctuations suggests that important physical considerations are still missing from the modeling.

VIII. SUMMARY

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- [1] N. Abdelwahab *et al.* [STAR], Phys. Rev. C **92**, no.2, 021901 (2015) doi:10.1103/PhysRevC.92.021901 [arXiv:1410.5375 [nucl-ex]].
 - [2] A. Sarkar [STAR], Nucl. Phys. A **931**, 796-801 (2014) doi:10.1016/j.nuclphysa.2014.10.042
 - [3] J. Xu [STAR], J. Phys. Conf. Ser. **779**, no.1, 012073 (2017) doi:10.1088/1742-6596/779/1/012073 [arXiv:1611.07132 [hep-ex]].
 - [4] J. Thäder [STAR], Nucl. Phys. A **956**, 320-323 (2016) doi:10.1016/j.nuclphysa.2016.02.047 [arXiv:1601.00951 [nucl-ex]].
 - [5] A. Sarkar [STAR], EPJ Web Conf. **60**, 13013 (2013) doi:10.1051/epjconf/20136013013
 - [6] X. Luo [STAR], PoS **CPOD2013**, 019 (2013) doi:10.22323/1.185.0019 [arXiv:1306.3106 [nucl-ex]].
 - [7] N. R. Sahoo [STAR], DAE Symp. Nucl. Phys. **57**, 766-767 (2012)
 - [8] X. Luo [STAR], Nucl. Phys. A **904-905**, 911c-914c (2013) doi:10.1016/j.nuclphysa.2013.02.163 [arXiv:1210.5573 [nucl-ex]].
 - [9] T. J. Tarnowsky [STAR], J. Phys. G **38**, 124054 (2011) doi:10.1088/0954-3899/38/12/124054 [arXiv:1106.6110 [nucl-ex]].
 - [10] X. F. Luo [STAR], J. Phys. Conf. Ser. **316**, 012003 (2011) doi:10.1088/1742-6596/316/1/012003 [arXiv:1106.2926 [nucl-ex]].
 - [11] N. R. Sahoo [STAR], [arXiv:1101.5125 [nucl-ex]].
 - [12] N. R. Sahoo [STAR], DAE Symp. Nucl. Phys. **56**, 1028-1029 (2011)
 - [13] M. Aggarwal *et al.* [STAR], Phys. Rev. Lett. **105**, 022302 (2010) doi:10.1103/PhysRevLett.105.022302 [arXiv:1004.4959 [nucl-ex]].
 - [14] A. Chatterjee [STAR], PoS **CORFU2018**, 164 (2019) doi:10.22323/1.347.0164 [arXiv:1904.01302 [hep-ex]].
 - [15] J. Adam *et al.* [STAR], Phys. Rev. C **100**, no.1, 014902 (2019) doi:10.1103/PhysRevC.100.014902 [arXiv:1903.05370 [nucl-ex]].
 - [16] L. Adamczyk *et al.* [STAR], Phys. Rev. Lett. **113**, 092301 (2014) doi:10.1103/PhysRevLett.113.092301 [arXiv:1402.1558 [nucl-ex]].
 - [17] D. McDonald [STAR], Nucl. Phys. A **904-905**, 907c-910c (2013) doi:10.1016/j.nuclphysa.2013.02.162 [arXiv:1210.7023 [nucl-ex]].
 - [18] D. McDonald, Ph.D. thesis, Rice University, <https://drupal.star.bnl.gov/STAR/files/mcdonaldthesis.pdf> (2013).
 - [19] A. Bzdak and V. Koch, Phys. Rev. C **86**, 044904 (2012) doi:10.1103/PhysRevC.86.044904 [arXiv:1206.4286 [nucl-th]].
 - [20] S. Cheng and S. Pratt, Phys. Rev. C **67**, 044904 (2003) doi:10.1103/PhysRevC.67.044904 [arXiv:nucl-th/0207051 [nucl-th]].
 - [21] S. Pratt and J. Ruppert, Phys. Rev. C **68**, 024904 (2003) doi:10.1103/PhysRevC.68.024904 [arXiv:nucl-th/0303043 [nucl-th]].
 - [22] S. Pratt, Phys. Rev. Lett. **84**, 4255-4259 (2000) doi:10.1103/PhysRevLett.84.4255 [arXiv:nucl-th/9905055 [nucl-th]].
 - [23] S. Pratt and S. Das Gupta, Phys. Rev. C **62**, 044603 (2000) doi:10.1103/PhysRevC.62.044603 [arXiv:nucl-th/9903006 [nucl-th]].
 - [24] P. Carruthers, E. M. Friedlander, C. C. Shih and R. M. Weiner, Phys. Lett. B **222**, 487-492 (1989) doi:10.1016/0370-2693(89)90350-X
 - [25] P. Carruthers and C. C. Shih, Phys. Lett. B **127**, 242-250 (1983) doi:10.1016/0370-2693(83)90884-5
 - [26] S. Schlichting and S. Pratt, Phys. Rev. C **83** (2011), 014913 doi:10.1103/PhysRevC.83.014913 [arXiv:1009.4283 [nucl-th]].
 - [27] S. Pratt, S. Schlichting and S. Gavin, Phys. Rev. C **84** (2011), 024909 doi:10.1103/PhysRevC.84.024909 [arXiv:1011.6053 [nucl-th]].
 - [28] S. Schlichting and S. Pratt, [arXiv:1005.5341 [nucl-th]].
 - [29] D. Oliinychenko, S. Shi and V. Koch, [arXiv:2001.08176 [hep-ph]].
 - [30] M. Tanabashi *et al.* [Particle Data Group], Phys. Rev. D **98** (2018) no.3, 030001 doi:10.1103/PhysRevD.98.030001

- [31] O. Savchuk, R. V. Poberezhnyuk, V. Vovchenko and M. I. Gorenstein, Phys. Rev. C **101** (2020) no.2, 024917 doi:10.1103/PhysRevC.101.024917 [arXiv:1911.03426 [hep-ph]].
- [32] D. Oliinychenko, S. Shi and V. Koch, arXiv:2001.08176 [hep-ph].
- [33] S. Pratt and S. Das Gupta, Phys. Rev. C **62**, 044603 (2000) doi:10.1103/PhysRevC.62.044603 [nucl-th/9903006].
- [34] S. Cheng and S. Pratt, Phys. Rev. C **67**, 044904 (2003) doi:10.1103/PhysRevC.67.044904 [arXiv:nucl-th/0207051 [nucl-th]].
- [35] S. Pratt and J. Ruppert, Phys. Rev. C **68**, 024904 (2003) doi:10.1103/PhysRevC.68.024904 [arXiv:nucl-th/0303043 [nucl-th]].
- [36] E. Sangaline, [arXiv:1505.00261 [nucl-th]].
- [37] M. Gorenstein and M. Gazdzicki, Phys. Rev. C **84** (2011), 014904 doi:10.1103/PhysRevC.84.014904 [arXiv:1101.4865 [nucl-th]].
- [38] M. Gazdzicki, M. Gorenstein and M. Mackowiak-Pawlowska, Phys. Rev. C **88** (2013) no.2, 024907 doi:10.1103/PhysRevC.88.024907 [arXiv:1303.0871 [nucl-th]].
- [39] V. V. Begun, M. I. Gorenstein and K. Grebieszko, J. Phys. G **42** (2015) no.7, 075101 doi:10.1088/0954-3899/42/7/075101 [arXiv:1409.3023 [nucl-th]].
- [40] C. Greiner, C. Gong and B. Muller, Phys. Lett. B **316** (1993), 226-230 doi:10.1016/0370-2693(93)90317-B [arXiv:hep-ph/9307336 [hep-ph]].
- [41] S. Pratt, Phys. Lett. B **301** (1993), 159-164 doi:10.1016/0370-2693(93)90682-8
- [42] A. Bzdak, V. Koch, D. Oliinychenko and J. Steinheimer, Phys. Rev. C **98** (2018) no.5, 054901 doi:10.1103/PhysRevC.98.054901 [arXiv:1804.04463 [nucl-th]].
- [43] C. E. Flores [STAR], Nucl. Phys. A **956** (2016), 280-283 doi:10.1016/j.nuclphysa.2016.05.020
- [44] M. A. Stephanov, K. Rajagopal and E. V. Shuryak, Phys. Rev. Lett. **81**, 4816-4819 (1998) doi:10.1103/PhysRevLett.81.4816 [arXiv:hep-ph/9806219 [hep-ph]].
- [45] M. A. Stephanov, K. Rajagopal and E. V. Shuryak, Phys. Rev. D **60**, 114028 (1999) doi:10.1103/PhysRevD.60.114028 [arXiv:hep-ph/9903292 [hep-ph]].
- [46] S. Pratt, E. Sangaline, P. Sorensen and H. Wang, Phys. Rev. Lett. **114**, 202301 (2015).
- [47] S. Pratt, W. P. McCormack and C. Ratti, Phys. Rev. C **92**, 064905 (2015).
- [48] J. E. Bernhard, J. S. Moreland, S. A. Bass, J. Liu and U. Heinz, Phys. Rev. C **94**, no. 2, 024907 (2016).
- [49] J. E. Bernhard, P. W. Marcy, C. E. Coleman-Smith, S. Huzurbazar, R. L. Wolpert and S. A. Bass, Phys. Rev. C **91**, no. 5, 054910 (2015).
- [50] K. M. Burke *et al.* [JET Collaboration], Phys. Rev. C **90**, no. 1, 014909 (2014).
- [51] Y. He, L. G. Pang and X. N. Wang, arXiv:1808.05310 [hep-ph].
- [52] Y. Xu, J. E. Bernhard, S. A. Bass, M. Nahrgang and S. Cao, Phys. Rev. C **97**, no. 1, 014907 (2018).
- [53] M. Greif, J. A. Fotakis, G. S. Denicol and C. Greiner, Phys. Rev. Lett. **120**, 242301 (2018).
- [54] S. Pratt, arXiv:1908.01053 [nucl-th].
- [55] S. Borsanyi, Z. Fodor, S. D. Katz, S. Krieg, C. Ratti and K. Szabo, JHEP **1201**, 138 (2012).
- [56] G. Aarts, C. Allton, A. Amato, P. Giudice, S. Hands and J. I. Skullerud, JHEP **1502**, 186 (2015).
- [57] A. Amato, G. Aarts, C. Allton, P. Giudice, S. Hands and J. I. Skullerud, Phys. Rev. Lett. **111**, no. 17, 172001 (2013).
- [58] G. Policastro, D. T. Son and A. O. Starinets, JHEP **0209**, 043 (2002).
- [59] J. Casalderrey-Solana and D. Teaney, Phys. Rev. D **74**, 085012 (2006).
- [60] J. Ghiglieri, G. D. Moore and D. Teaney, JHEP **1803**, 179 (2018).
- [61] M. Greif, C. Greiner and G. S. Denicol, Phys. Rev. D **93**, no. 9, 096012 (2016) Erratum: [Phys. Rev. D **96**, no. 5, 059902 (2017)]
- [62] J. Hammelmann, J. M. Torres-Rincon, J. B. Rose, M. Greif and H. Elfner, Phys. Rev. D **99**, no. 7, 076015 (2019).
- [63] S. Pratt and C. Plumberg, to appear in Phys. Rev. C, arXiv:1812.05649 [nucl-th].
- [64] C. Shen, Z. Qiu, H. Song, J. Bernhard, S. Bass and U. Heinz, Comput. Phys. Commun. **199**, 61 (2016)
- [65] S. Pratt, J. Kim and C. Plumberg, Phys. Rev. C **98**, no. 1, 014904 (2018).
- [66] H. Wang, Ph.D. Thesis, arXiv:1304.2073 [nucl-ex].
- [67] B. I. Abelev *et al.* [STAR Collaboration], Phys. Lett. B **690**, 239 (2010).
N. Li *et al.* [STAR Collaboration],
- [68] J. Adams *et al.* [STAR Collaboration], Phys. Rev. Lett. **90**, 172301 (2003).
- [69] M. M. Aggarwal *et al.* [STAR Collaboration], Phys. Rev. C **82**, 024905 (2010).
- [70] B. Abelev *et al.* [ALICE Collaboration], Phys. Lett. B **723**, 267 (2013).
- [71] C. Alt *et al.* [NA49 Collaboration], Phys. Rev. C **76**, 024914 (2007).
- [72] L. Adamczyk *et al.* [STAR Collaboration], Phys. Rev. C **94**, no. 2, 024909 (2016).
- [73] L. Adamczyk *et al.* [STAR Collaboration], Phys. Rev. C **88**, no. 6, 064911 (2013).
- [74] B. I. Abelev *et al.* [STAR Collaboration], Phys. Rev. Lett. **103**, 251601 (2009).
- [75] S. A. Bass, P. Danielewicz and S. Pratt, Phys. Rev. Lett. **85**, 2689 (2000).
- [76] S. Pratt, Phys. Rev. Lett. **108**, 212301 (2012).
- [77] Y. Pan and S. Pratt, Phys. Rev. C **89**, no. 4, 044911 (2014).
- [78] J. Steinheimer, J. Aichelin, M. Bleicher and H. Stöcker, Phys. Rev. C **95**, no. 6, 064902 (2017).
- [79] J. Steinheimer, J. Aichelin and M. Bleicher, Phys. Rev. Lett. **110**, no. 4, 042501 (2013).