

N-body simulation

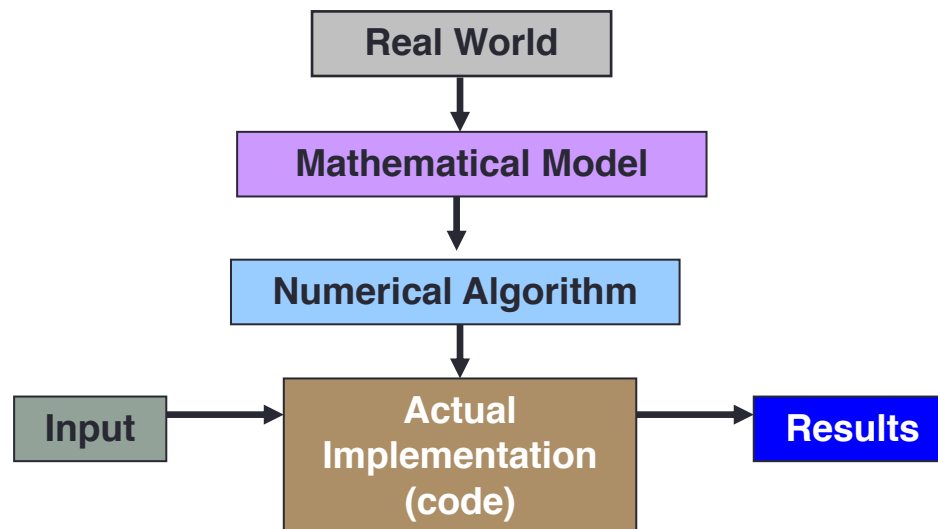
Part 1: introduction

Overview

- What is N-body simulation and what is it used for?
- Underlying physics & mathematical formulation
- General N-body simulation algorithm
- Computational cost and big O notation

What is N-body simulation?

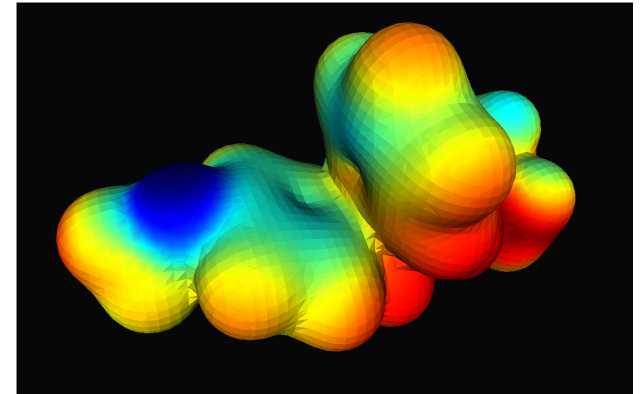
- Particles interact through physical forces
- A mathematical model describe these interactions and how particles move
- A numerical algorithm derived from the mathematical model allows us to simulate the motion of N interacting particles - a dynamical system



Science that uses N-body simulation

- Atomistic simulation / molecular dynamics:

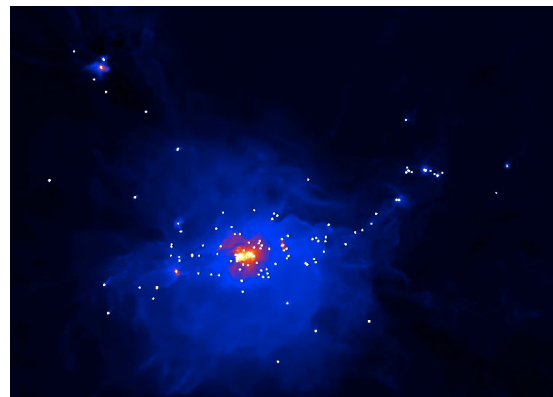
- Biomolecular modelling & simulation
- Computational chemistry
- Materials science



credit RPI/Curt Breneman

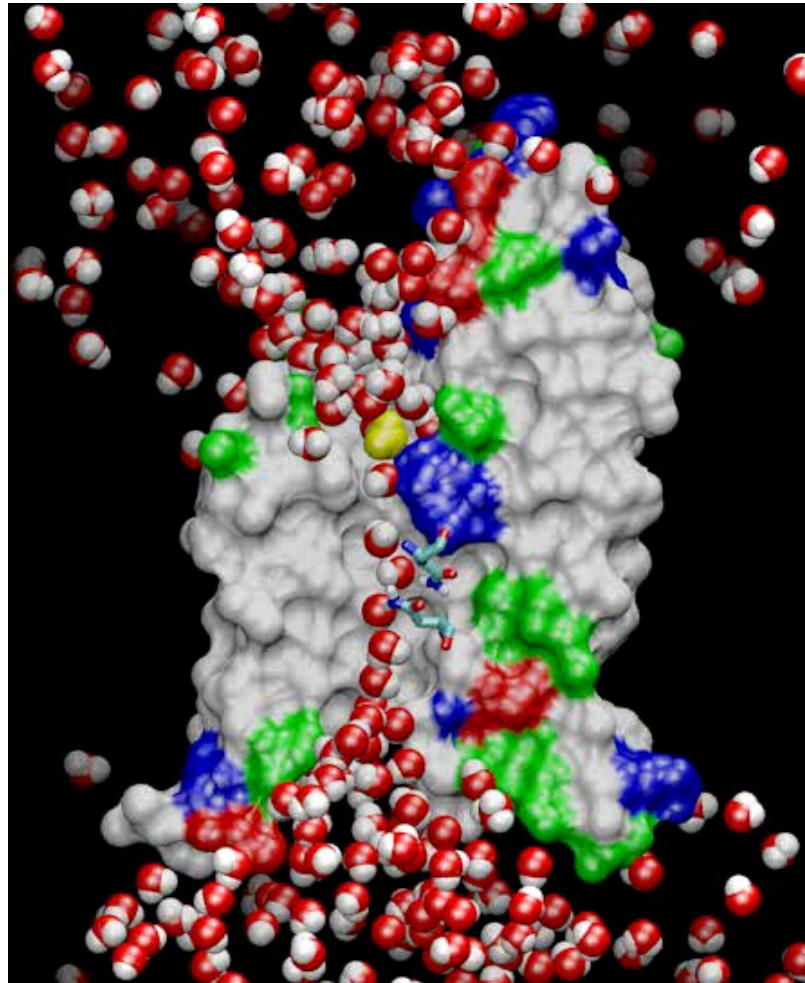
- Astrophysical N-body simulation:

- Orbital dynamics
- Formation & evolution of planetary systems
- Cosmology



Matthew Bate, U. Exeter

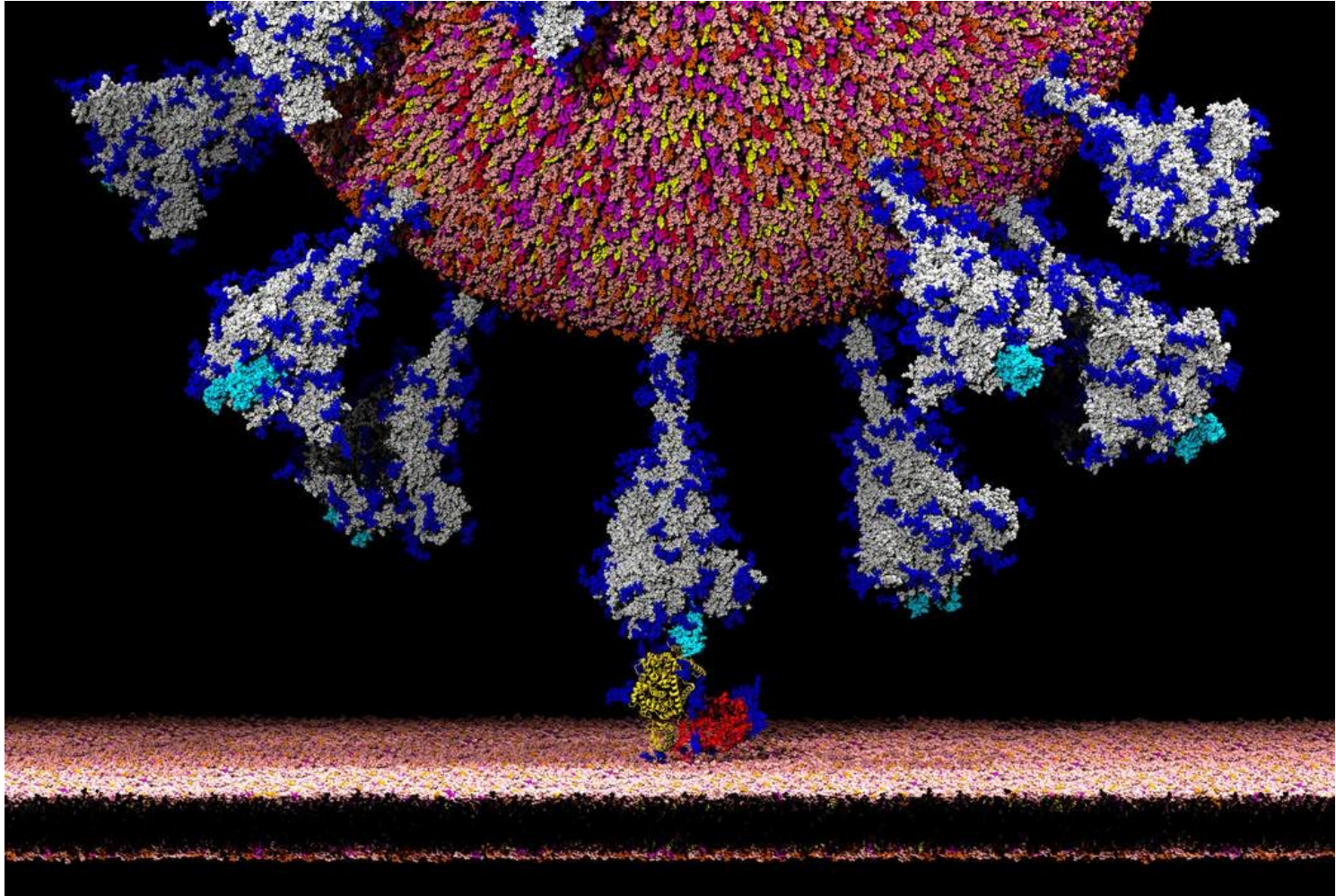
Water permeation in an aquaporin



Star formation: Density



SARS-CoV-2



What is a particle?

- Asteroid, planet, star (astrophysical n-body simulation)
 - To model solar systems, globular star clusters, colliding galaxies, ...
- Atom (atomistic simulation, molecular dynamics)
 - To model liquids, solids (e.g. crystals), gases, molecules, viruses
- Lots of things inbetween (coarse-grained / multi-scale simulation)
 - molecule
 - protein
 - small volume of gas, liquid, or solid (10^x atoms)

Interactions

- Gravity (astrophysical simulation)
- Electrostatic (atomistic simulation)
- Use simple approximate descriptions of forces
 - *Accurate enough* for simulation purpose
 - *Fast enough* to compute in useful time
- Include approximate quantum effects where essential
 - ‘Traditional’ N-body simulation uses Newton’s Laws – classical physics
 - Describing motion of electrons requires full quantum physics

Example: gravitational force

- Newton's Universal Law of Gravitation
 - Force between two bodies labelled i and j (in vector form)

Gravitational Constant

Masses of the two bodies

$$\vec{F}_{ij} = G \frac{m_i m_j}{|\vec{r}_i - \vec{r}_j|^2} \times \frac{\vec{r}_i - \vec{r}_j}{|\vec{r}_i - \vec{r}_j|}$$

Separation of the two bodies, squared

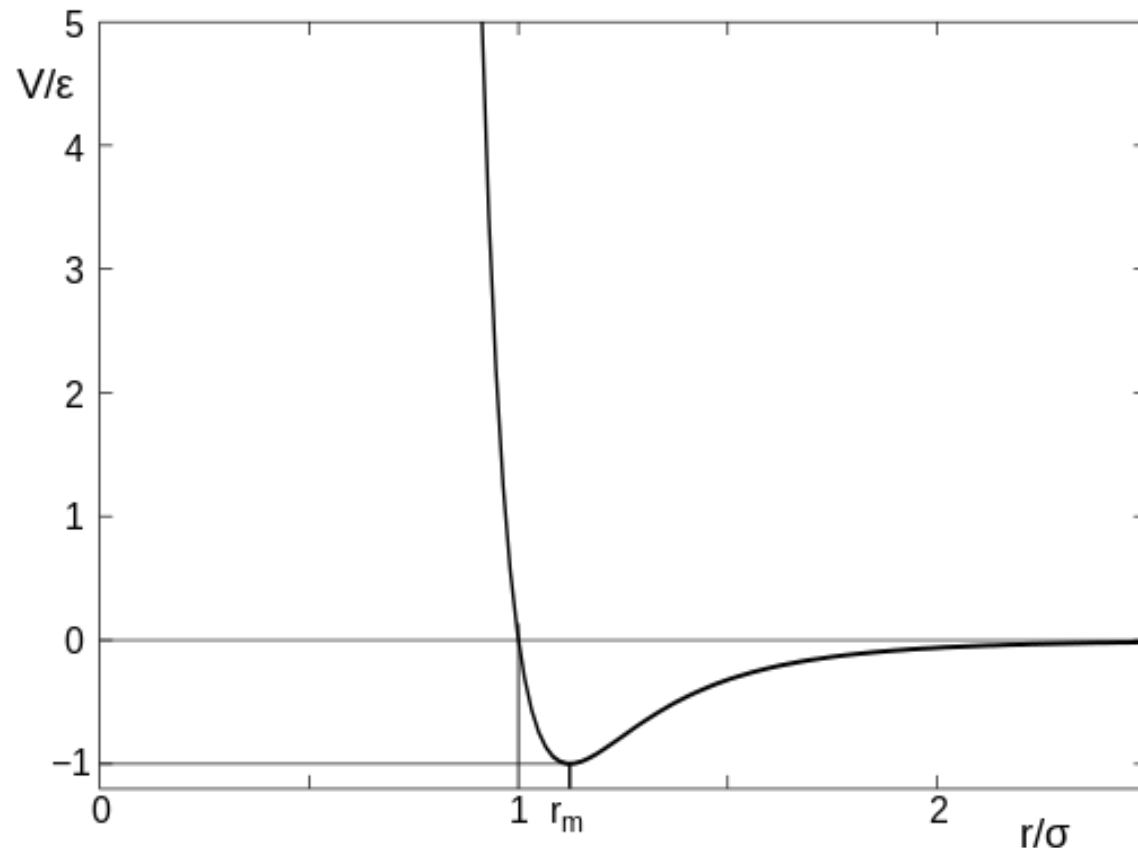
Unit vector in
direction of
resultant force

Example: Lennard-Jones potential

- Electrostatic attraction ($1/r^2$ like gravity) dominates between atoms further apart than r_m
- Repulsion dominates between atoms closer together than r_m

$$V(r_{ij}) = 4\epsilon_{ij} \left(\left[\frac{\sigma_{ij}}{r_{ij}} \right]^{12} - \left[\frac{\sigma_{ij}}{r_{ij}} \right]^6 \right) \quad r_m = 2^{\frac{1}{6}} \sigma$$

$$F = -\frac{dV}{dr}$$



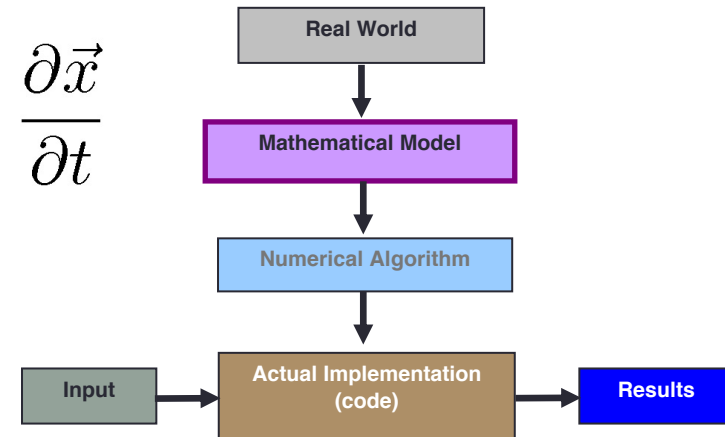
Why N-body simulations?

Classical mechanics

Mathematical model underlying N-body simulation:
differential equations expressing Newton's Laws of Motion
(i.e. classical mechanics):

$$\vec{F} = m\vec{a} \quad \vec{a} = \frac{\partial \vec{v}}{\partial t} = \frac{\partial^2 \vec{x}}{\partial t^2} \quad \vec{v} = \frac{\partial \vec{x}}{\partial t}$$

These apply for each of the N particles



If we define forces and initial conditions (positions & velocities),
can **in principle** solve these equations

Real problems are difficult to solve

- Simple problem: apple falling (1 body)
 - Can solve exactly analytically
 - Actually a simplified 2-body problem (apple & Earth)
- Less simple: planet orbiting a star (2 bodies)
 - Can still be solved exactly
- Three body problem
 - No analytic solution *in general*
- Most real problems too complex for exact analytical solution
 - Applies to general n-body problems, where $n \geq 3$



➔ **need to solve numerically using a computer**

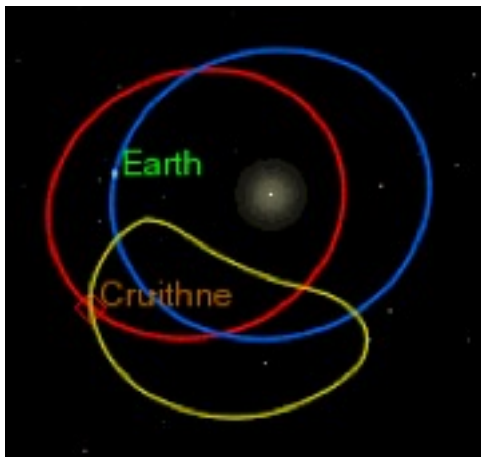
Earth's “second moon”

“3753 Cruithne”

- First observed 10 Oct 1986
- Near-Earth asteroid, 5km diameter

Orbit not determined until 1997

- Co-orbits the Sun with Earth



Complicated ‘kidney bean’ orbit:

- Not possible to predict by solving equations of motion exactly
- Required simulation that includes other planets: many-body problem

Solving differential equations numerically

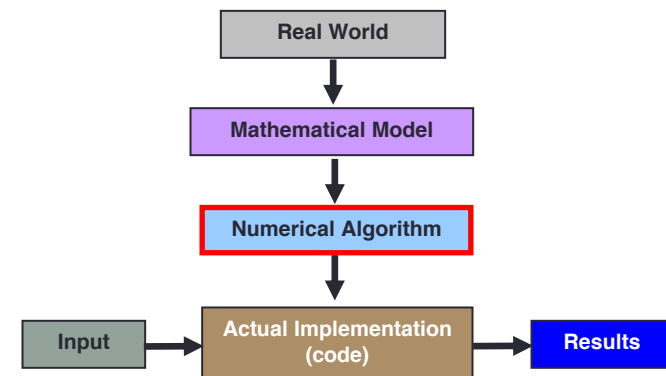
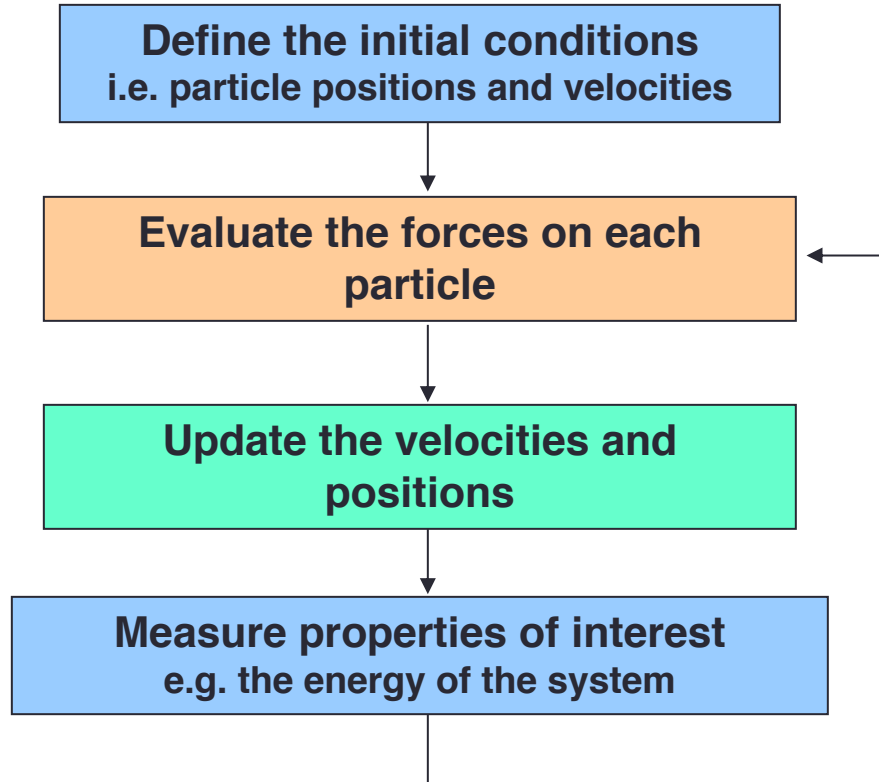
- Mathematical model:

$$\vec{F} = m\vec{a} \quad \vec{a} = \frac{\partial \vec{v}}{\partial t} = \frac{\partial^2 \vec{x}}{\partial t^2} \quad \vec{v} = \frac{\partial \vec{x}}{\partial t}$$

- Differential equations are continuous, computers are digital
- **Discretisation** converts differential equations into forms suitable for numerical solution (details next lecture)
- Solution computed by **time integration**:
 - estimate $F \Rightarrow a \Rightarrow v \Rightarrow x \Rightarrow F$ over small discrete time increments
 - repeat to simulate motion of particles over longer time

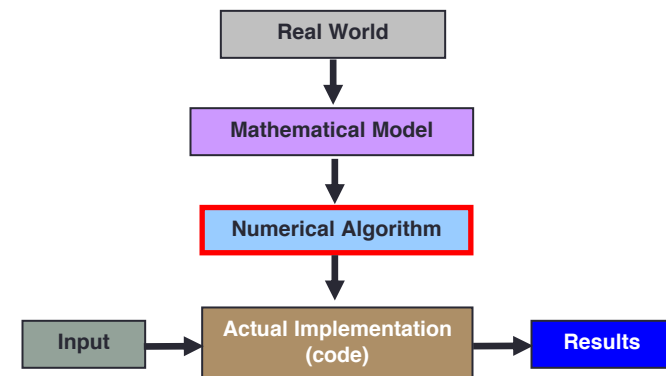
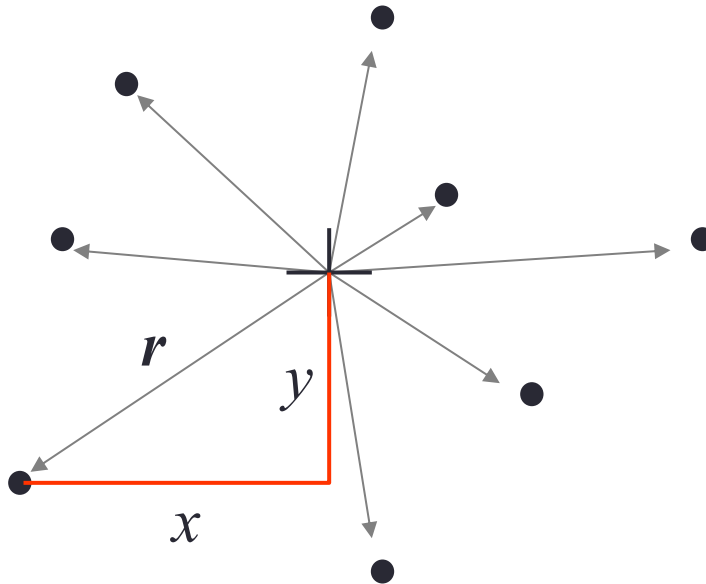
General N-body algorithm

General N-body simulation algorithm



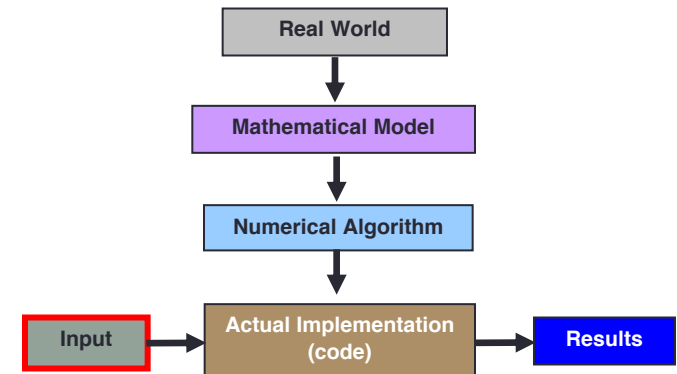
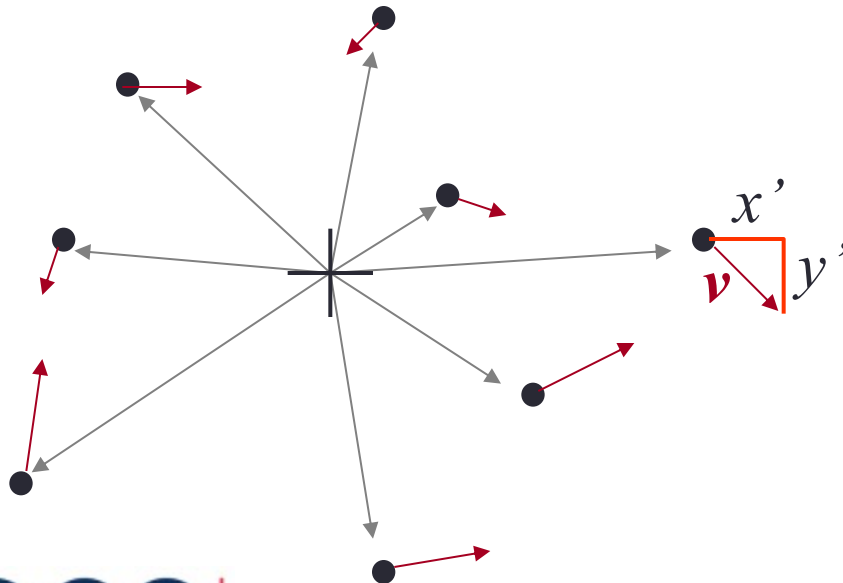
Particle representation: positions

- Positions stored as arrays of real numbers
 - e.g. in 2-D, vector position \mathbf{r} of $i = 1$ to N particles stored as $x[i]$ and $y[i]$ displacements:



Initial Conditions

- We need initial positions and velocities
 - position \mathbf{r} stored as x and y displacements,
 - e.g. miles, metres.
 - velocity \mathbf{v} stored as x' and y' “speeds”:
 - units of displacement per unit time, e.g. mph, m/s.



The Simulation Domain: Boundaries

- Examples

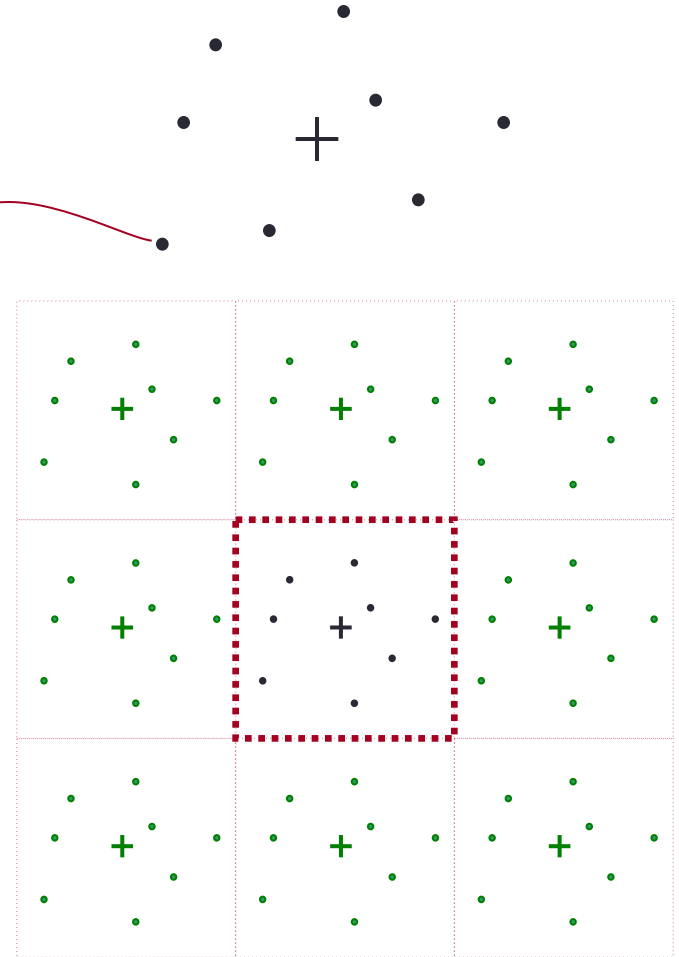
- open:

- particles can go anywhere

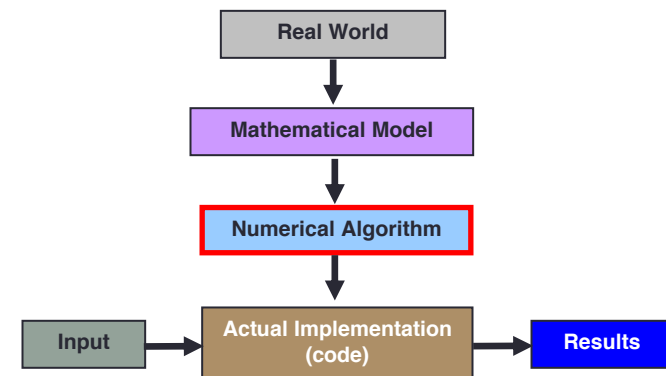
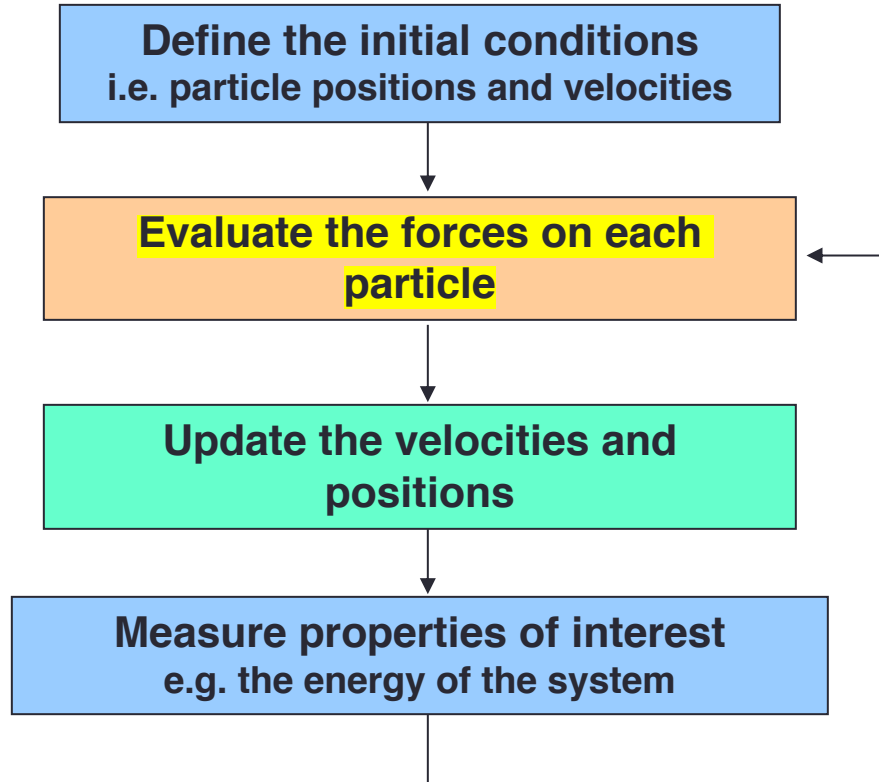


- periodic:

- particles stay in a fixed volume, by wrapping around the edges.
 - The system ‘sees’ images of itself tiled out in all directions.
 - suitable for simulating “bulk” behaviour inside of many systems (e.g. gas, biomolecule in water)

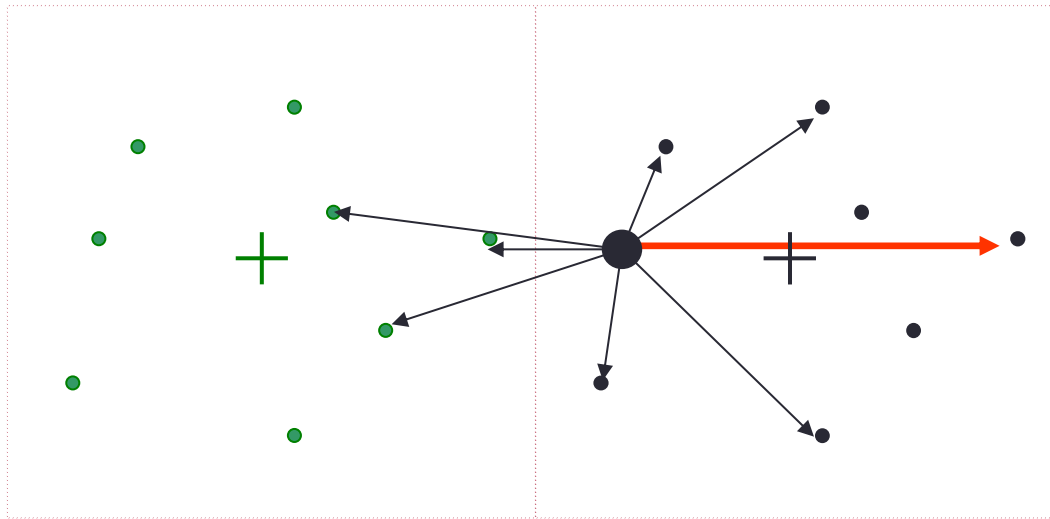


General N-body simulation algorithm



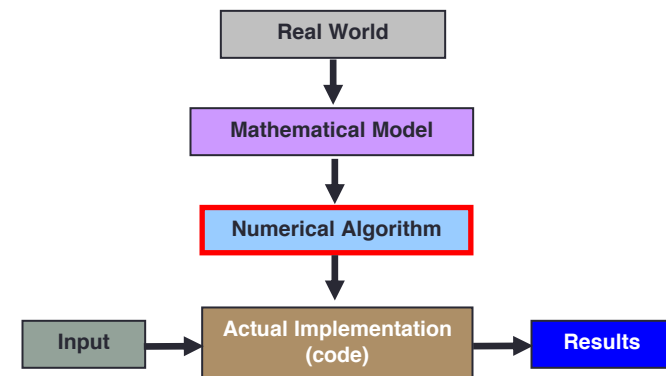
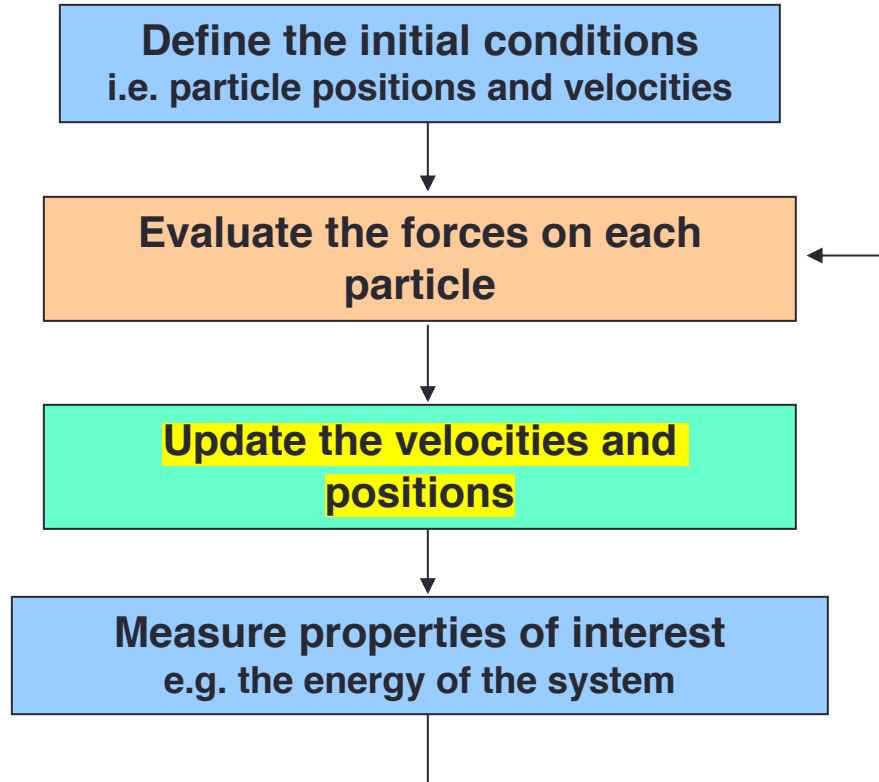
Evaluating the total force on each particle

- For each particle, numbered i from 1 to N
 - add up all forces acting upon that particle due to all other particles.
 - when using periodic boundary conditions, we usually use the nearest image of the other particles.



- particles should not interact directly with (images of) themselves

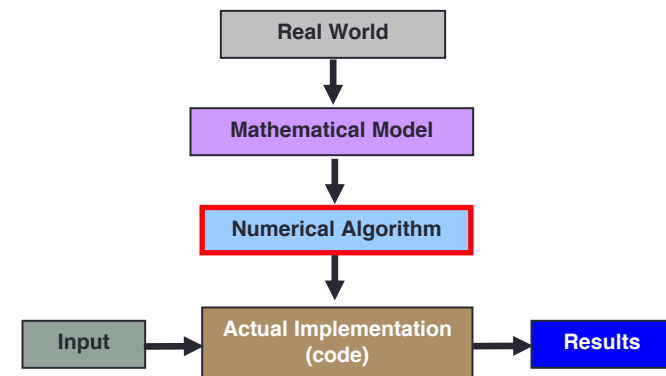
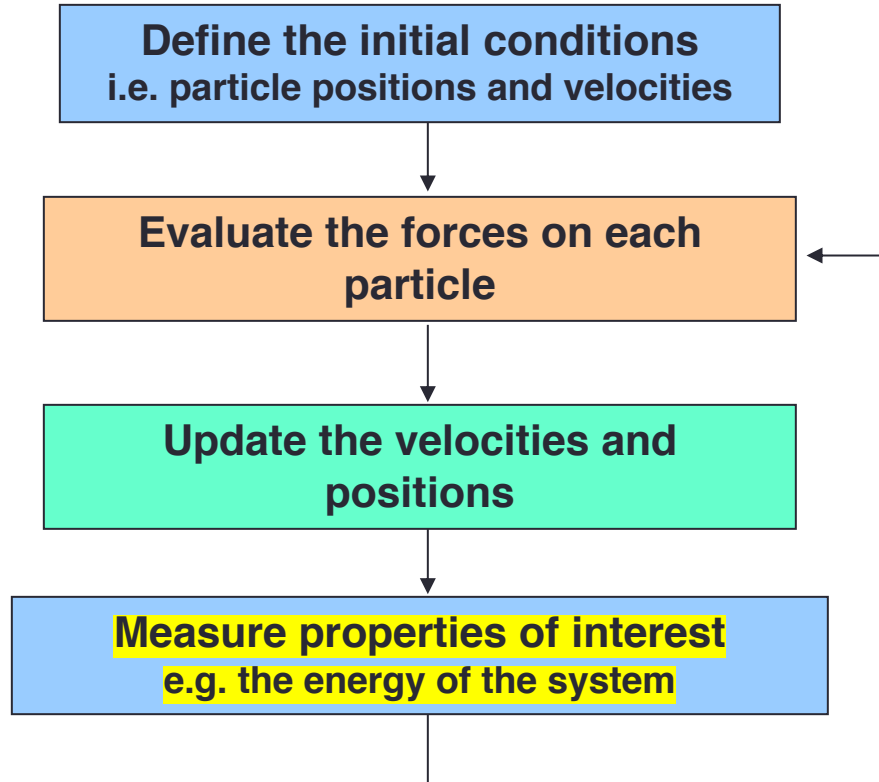
General N-body simulation algorithm



Updating Positions and Velocities

- Update positions and velocities using the discretised forms of Newton's equations of motion and the forces on each particle computed in the previous step
- Many possible ways of doing this
 - Discretisation and time integration schemes trade off speed and accuracy in different ways
 - Consider some of these methods in next lecture & in practical

General N-body simulation algorithm



Measuring properties of interest

- Kinetic energy:
$$E_K(t) = \frac{1}{2} \sum_i m_i \vec{v}_i(t)^2$$
- Potential energy:
$$E_P(t) = \sum_{\langle i,j \rangle} U(|\vec{r}_j(t) - \vec{r}_i(t)|)$$

–
($\sum_{\langle i,j \rangle}$ means sum over all pairs)
- Also relevant when we consider errors and accuracy

Computational Cost

Computational cost of force calculations

If each person in this (virtual) room of N people interacts with each other person (“pair-wise”), how many conversations are happening in total?

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$$= \binom{N}{2} = \frac{N!}{2!(N-2)!}$$

Computational cost of force calculations

Suppose each pairwise force calculation requires execution of q operations

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Time cost T of performing all force calculations is then:

$$T = \binom{N}{2} qt$$

Asymptotic analysis of computational cost

Big O notation

Asymptotic analysis of computational cost

- Want to understand how the runtime of an algorithm scales as a key parameter changes (typically grows)
- Estimating this dependence for very large parameter values: ***asymptotic analysis***
- Asymptotic analysis allows us to
 - Predict which parts of an algorithm dominate runtime in different scenarios
 - Compare the expected relative performance of different algorithms

Big O notation

- Big-O notation specifies upper bound on the quantity of interest (typically runtime, could be memory requirements):

$f(N) = O(g(N)) \rightarrow$ for some sufficiently large N
it is true that $f(N) \leq c * g(N)$, where c is a constant

- e.g. $T = O(N^3)$ implies that for large enough N the runtime scales at most (i.e. no worse than) as N^3
- In practice the value of c can be crucial: algorithm that scales worse asymptotically can, for small N , still be much faster

Computational cost of force calculations

- Without writing any code, could perform asymptotic analysis of force calculation runtime based on $T = \binom{N}{2} qt$

$$\text{Since } \binom{N}{2} = \frac{N!}{2!(N-2)!} = \frac{N(N-1)}{2} = \frac{N^2}{2} - \frac{N}{2}$$

$$\text{and } \lim_{N \rightarrow \infty} \binom{N}{2} = \lim_{N \rightarrow \infty} \frac{N^2}{2}$$

we can conclude that $T = O(N^2)$

Computational cost of force calculations

For N-body simulation the naive serial calculation is

```
for (i=0;i<N;i++){  
    for (j=0;j<N;j++){  
        if (i != j) {  
            force[i] += force(i,j);  
        }  
    }  
}
```

This scales as $O(N^2)$

Asymptotic analysis of computational cost

- Consider: how does the overall runtime of the following sequential matrix-matrix multiplication $\mathbf{A} \times \mathbf{B}$ scale as a function of the linear size N of two square matrices \mathbf{A} , \mathbf{B} ?

```
for (i=0; i<N; i++){  
    for (j=0; j<N; j++){  
        for (k=0; k<N; k++){  
            c[i][j] += a[i][k]*b[k][j];  
        }  
    }  
}
```

Summary

- N-body simulations compute motion of particles interacting through forces
 - Need to store particle positions and velocities, compute forces, evolve equation(s) of motion
- Evaluating total forces on all atoms is an $O(N^2)$ calculation
 - Asymptotic scaling of computational cost: number of particles doubles \rightarrow runtime quadruples – this is why force calculation is typically the most time-consuming part of an N-body simulation
- Next lecture (and planetary orbits practical):
 - Discretisation / time integration schemes
 - Accuracy and stability
 - N-body methods: force calculation schemes