

- DeepBench: A simulation package for physical
- benchmarking data
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Software

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Summary

We introduce **DeepBench**, a Python library that employs mechanistic models (i.e., analytic mathematical models) to simulate data that represent physics-related objects and systems: geometric shapes (e.g., polygon), physics objects (e.g., pendulum), and astronomical objects (e.g., elliptical galaxy). These data take the form of images (two-dimensional) or time series (one-dimensional). In contrast to natural image benchmarks and complex physics simulations, these data have simple, direct, numerical, and traceable connections between the input data and the label data. When seeking a quantifiable interpretation, this kind of data is uniquely suitable for developing, calibrating, testing, and benchmarking statistical and machine learning models. Finally, this software includes methods to curate and store these datasets to maximize reproducibility.

Statement of Need

There are multiple open problems and issues that are critical for the machine learning and scientific communities to address - principally, interpretability, explainability, uncertainty quantification, and inductive bias in machine learning models when they are applied to scientific data. Multiple kinds of data sets and data simulation software can be used for developing models and confronting these challenges. These data sets range from natural images and text to multi-dimensional data of physical processes. Indeed, multiple benchmark data and simulation software have been created and inculcated for developing and comparing models. However, these benchmarks are typically limited in significant ways. Natural image data sets comprising images from the real or natural world (e.g., vehicles, animals, landscapes) are widely used in the development of machine learning models. These kinds of data sets tend to be large, diverse, and carefully curated. However, they are not underpinned by or constructed upon physical principles: they cannot be generated by mathematical expressions of formal physical theory, so there is not a robust connection between the data and a quantitative theory. Therefore, these data sets have a severely limited capacity to help address many questions in machine learning models, such as uncertainty quantification. On the other hand, complex physics simulations (e.g., cosmological n-body simulations and particle physics simulators) are accurate, detailed, and based on precise quantitative theories and models. This facilitates studies of interpretability and uncertainty quantification because there is the possibility of linking the simulated data to the input choices through each layer of calculation in the simulator. However, they are relatively small in size and number, and they are computationally expensive to reproduce. In addition, while they are underpinned by specific physical functions, the



- complexity of the calculations makes them challenging as a venue through which to make
- connections between machine learning results and input choices. Complex physics simulations
- have one or more layers of mechanistic models. Mechanistic models are defined with analytic
- functions and equations that describe and express components of a given physical process:
- these are based on theory and empirical observations. In both of these scenarios, it is difficult
- to build interpretable models that connect raw data and labels, and it is difficult to generate
- new data rapidly.
- The physical sciences community lacks sufficient datasets and software as benchmarks for the
- development of statistical and machine learning models. In particular, there currently does 50
- not exist simulation software that generates data underpinned by physical principles and that 51
- satisfies the following criteria:
 - multi-domain
 - multi-purpose
 - fast

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- reproducible
 - extensible
 - based on mechanistic models
- include detailed noise prescriptions.

Related Work

- There are many benchmarks datasets and simulation software widely used for model 61 building in machine learning, statistics, and the physical sciences. First, benchmark
- datasets of natural images include MNIST [@dengMnistDatabaseHandwritten2012c], CIFAR
- [@krizhevskyCIFAR10CanadianInstitute2017a], | Imagenet [@russakovskyImageNetLargeScale2015a]
- Second, there are several large astronomical observation datasets CfA Redshift Survey 65
- [@huchraSurveyGalaxyRedshifts1983], Sloan Digital Sky Survey [@yorkSloanDigitalSky2000],
- and Dark Energy Survey [@abbottDARKENERGYSURVEY]. Third, many n-body cosmology simula-
- tion data sets serve as benchmarks e.g., Millennium [@springelCosmologicalSimulationCode2005], Illustris [@vogelsbergerIntroducingIllustrisProject2014b], EAGLE [@schayeEAGLEProjectSimulat
- Coyote [@heitmannCoyoteUniversePrecision2010], Bolshoi [@klypinDARKMATTERHAL0S2011],
- CAMELS [@villaescusa-navarroCAMELSProjectCosmology2021], Quijote [@villaescusa-
- navarroQuijoteSimulations2020]. Fourth, there have been multiple astronomy 72
- data set challenges that can be considered benchmarks for analysis comparison 73
- PLAsTiCC [@hlozekResultsPhotometricLSST2020a], The Great08 Challenge
- [@bridleHandbookGREAT08Challenge2009a], and the Strong Gravitational Lens Chal-
- lenge [@metcalfStrongGravitationalLens2019c]. Fifth, there are multiple soft-
- ware that generate simulated data for astronomy and cosmology e.g., astropy
- [@theastropycollaborationAstropyCommunityPython2013a], galsim [@roweGalSimModularGalaxy201
- lenstronomy [@birrerLenstronomyMultipurposeGravitational2018a], deeplenstronomy
- [@morganDeeplenstronomyDatasetSimulation2021a], CAMB [@CAMBInfo, @lewisEfficientComputat
 - Pixell [@WelcomePixellDocumentation], SOXs [@SOXSSimulatedObservations]. Fi-
 - nally, particle physics projects use standard codebases for simulations e.g., GEANT
- [@piaGeant4ScientificLiterature2009], GENIE [@andreopoulosGENIENeutrinoMonte2015], 83
- and PYTHIA [@sjostrandPYTHIAEventGenerator2020]. These simulations span wide ranges
- in speed, code complexity, and physical fidelity and detail. Unfortunately, these data and soft-
- ware lack a combination of critical features, including mechanistic models, speed, reproducibility,
- which are needed for more fundamental studies of statistical and machine learning models. The
- work in this paper is most closely related to SHAPES [@wuVisualQuestionAnswering2016a]
- because that work also uses collections of geometric objects with varying levels of complexity
- as a benchmark.



DeepBench Software

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The DeepBench software simulates data for analysis tasks that require precise numerical calculations. First, the simulation models are fundamentally mechanistic - based on relatively simple analytic mathematical expressions, which are physically meaningful. This means that for each model, the number of input parameters that determine a simulation output is small (<10 for most models). These elements make the software fast and the outputs interpretable - conceptually and mathematically relatable to the inputs. Second, DeepBench also includes methods to precisely prescribe noise for inputs, which are propagated to outputs. This permits studies and the development of statistical inference models that require uncertainty quantification, which is a significant challenge in modern machine learning research. Third, the software framework includes features that permit a high degree of reproducibility: e.g., random seeds at every key stage of input, a unique identification tag for each simulation run, tracking and storage of metadata (including input parameters) and the related outputs. Fourth, the primary user interface is a YAML configuration file, which allows the user to specify every aspect of the simulation - e.g., types of objects, numbers of objects, noise type, and number of classes. This feature - which is especially useful when building and studying complex models like deep learning neural networks - permits the user to incrementally decrease or increase the complexity of the simulation with a high level of granularity.

DeepBench has the following features:

- Exact reproducibility
- Noise and error propagation
- Mechanistic modeling
- Physical sciences-based modeling
- Computational efficiency
- Simulations relevant to multiple domains
- Outputs of varying dimensions
- Readily extensible to new physics and outputs

Primary Modules

- Geometry objects: two-dimensional images generated with matplotlib [@hunterMatplotlib2DGrap The shapes include N-sided polygons, arcs, straight lines, and ellipses. They are solid, filled or unfilled two-dimensional shapes with edges of variable thickness.
- Physics objects: one-dimensional profiles for two types of implementations of pendulum dynamics: one using Newtonian physics, the other using Hamiltonian.
- Astronomy objects: two-dimensional images generated based on radial profiles of typical astronomical objects. The star object is created using the Moffat distribution provided by the AstroPy [@theastropycollaborationAstropyCommunityPython2013a] library. The spiral galaxy object is created with the function used to produce a logarithmic spiral [@ringermacherNewFormulaDescribing2009a]. The elliptical Galaxy object is created using the Sérsic profile provided by the AstroPy library. Two-dimensional models are representations of astronomical objects commonly found in data sets used for galaxy morphology classification.
- Image: two-dimensional images that are combinations and/or concatenations of Geometry
 or Astronomy objects. The combined images are within matplotlib meshgrid objects.
 Sky images are composed of any combination of Astronomy objects, while geometric
 images comprise individual geometric shape objects.
- Collection: Provides a framework for producing module images or objects at once and storing all parameters that were included in their generation, including exact noise levels, object hyper-parameters, and non-specified defaults.



All objects also come with the option to add noise to each object. For Physics objects – i.e., the pendulum – the user may add Gaussian noise to parameters: initial angle θ_0 , the pendulum length L, the gravitational acceleration g, the planet properties $\Phi=(M/r^2)$, and Newton's gravity constant G. Note that $g=G*\Phi=G*M/r^2$: all parameters in that relationship can receive noise. For Astronomy and Geometry Objects, which are images, the user can add Poisson or Gaussian noise to the output images. Finally, the user can regenerate the same noise using the saved random seed.

Example Outputs

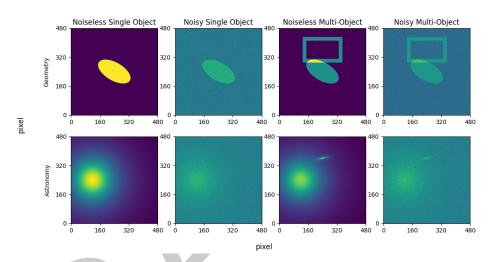


Figure 1: Example outputs of **DeepBench**, containing shapes, and astronomy objects. Variants include a single object, a noisy single object, two objects, and two noisy objects.

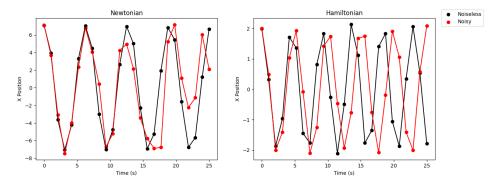


Figure 2: Example physics simulations from **DeepBench**. Pendulums show noisy and non-noisy variants of the Newtonian (left) and Hamiltonian (right) mathematical simulations.

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References

