

SM-01 Module 2

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Contents

1	Module Index	1
1.1	Modules	1
2	Data Structure Index	3
2.1	Data Structures	3
3	File Index	5
3.1	File List	5
4	Module Documentation	7
4.1	General Sky Model Module	7
4.1.1	Detailed Description	11
4.1.2	Define Documentation	11
4.1.2.1	F	11
4.1.2.2	SM_BOOL	11
4.1.2.3	SM_ERROR_ACCES_TXT	11
4.1.2.4	SM_ERROR_BDR_TXT	11
4.1.2.5	SM_ERROR_EIS	11
4.1.2.6	SM_ERROR_EIS_TXT	11
4.1.2.7	SM_ERROR_EXIST_TXT	11
4.1.2.8	SM_ERROR_FAULT_TXT	11
4.1.2.9	SM_ERROR_FOF	11
4.1.2.10	SM_ERROR_FOF_TXT	11
4.1.2.11	SM_ERROR_FOPEN_TXT	11
4.1.2.12	SM_ERROR_IDG_TXT	12
4.1.2.13	SM_ERROR_IDR_TXT	12

4.1.2.14	SM_ERROR_IFE_TXT	12
4.1.2.15	SM_ERROR_IIP_TXT	12
4.1.2.16	SM_ERROR_INVAL_TXT	12
4.1.2.17	SM_ERROR_IO_TXT	12
4.1.2.18	SM_ERROR_IOD_TXT	12
4.1.2.19	SM_ERROR_IOS_TXT	12
4.1.2.20	SM_ERROR_IOV_TXT	12
4.1.2.21	SM_ERROR_ISD	12
4.1.2.22	SM_ERROR_ISD_TXT	12
4.1.2.23	SM_ERROR_ISM	12
4.1.2.24	SM_ERROR_ISM_TXT	12
4.1.2.25	SM_ERROR_LINK_TXT	12
4.1.2.26	SM_ERROR_LOOP_TXT	12
4.1.2.27	SM_ERROR_NAMETOOLONG_TXT	12
4.1.2.28	SM_ERROR_NDA_TXT	12
4.1.2.29	SM_ERROR_NOENT_TXT	12
4.1.2.30	SM_ERROR_NOMEM_TXT	12
4.1.2.31	SM_ERROR_NOSPC_TXT	12
4.1.2.32	SM_ERROR_NOTDIR_TXT	12
4.1.2.33	SM_ERROR_PERM_TXT	13
4.1.2.34	SM_ERROR_ROFS_TXT	13
4.1.2.35	SM_ERROR_SUBROUTINE_TXT	13
4.1.2.36	SM_ERROR_TXTBSY_TXT	13
4.1.2.37	SM_ERROR_UFS_TXT	13
4.1.2.38	SM_ERROR_UNDEF_TXT	13
4.1.2.39	SM_GAS_CONST	13
4.1.2.40	SM_GENERAL_H	13
4.1.2.41	SM_GRAV_ACC	13
4.1.2.42	SM_LENLINE	13
4.1.2.43	SM_MAXLEN	13
4.1.2.44	SM_MAXPAR	14
4.1.2.45	SM_MOL_AIR_DRY	14
4.1.2.46	SM_NMAXERR	14
4.1.2.47	SM_SR_IN_ARCSEC2	14

4.1.2.48	SM_TOL	14
4.1.2.49	T	14
4.1.3	Typedef Documentation	14
4.1.3.1	sm_error_code	14
4.1.3.2	smbool	14
4.1.3.3	smdat	14
4.1.3.4	smgrid	15
4.1.3.5	smparam	15
4.1.3.6	smspec	15
4.1.4	Enumeration Type Documentation	15
4.1.4.1	_sm_error_code_	15
4.1.4.2	SM_BOOL	16
4.1.5	Function Documentation	16
4.1.5.1	sm_basic_access	16
4.1.5.2	sm_basic_chdir	17
4.1.5.3	sm_basic_createdir	17
4.1.5.4	sm_basic_initstring	18
4.1.5.5	sm_basic_interpollin	18
4.1.5.6	sm_basic_isnumber	19
4.1.5.7	sm_basic_mkdir	19
4.1.5.8	sm_basic_replacestring	20
4.1.5.9	sm_basic_rmcntrl	20
4.1.5.10	sm_basic_rmcntrl_inplace	21
4.1.5.11	sm_basic_strtrim	21
4.1.5.12	sm_basic_strtrim_inplace	21
4.1.5.13	sm_basic_terminatestring	22
4.1.5.14	sm_grid_extract	22
4.1.5.15	sm_grid_free	23
4.1.5.16	sm_grid_malloc	23
4.1.5.17	sm_grid_print	24
4.1.5.18	sm_grid_read	24
4.1.5.19	sm_grid_write	25
4.1.5.20	sm_param_check	25
4.1.5.21	sm_param_read	26

4.1.5.22	sm_param_readcheck	26
4.1.5.23	sm_spec_average	27
4.1.5.24	sm_spec_calc	28
4.1.5.25	sm_spec_changegrid	28
4.1.5.26	sm_spec_changetype	29
4.1.5.27	sm_spec_compgrids	29
4.1.5.28	sm_spec_convolve	30
4.1.5.29	sm_spec_convunits	30
4.1.5.30	sm_spec_copy	31
4.1.5.31	sm_spec_create	31
4.1.5.32	sm_spec_extract	32
4.1.5.33	sm_spec_free	32
4.1.5.34	sm_spec_funct	32
4.1.5.35	sm_spec_functnoerr	33
4.1.5.36	sm_spec_interpol	33
4.1.5.37	sm_spec_join	34
4.1.5.38	sm_spec_malloc	34
4.1.5.39	sm_spec_modval	35
4.1.5.40	sm_spec_print	35
4.1.5.41	sm_spec_read	36
4.1.5.42	sm_spec_readcpl	36
4.1.5.43	sm_spec_readfits	37
4.1.5.44	sm_spec_readfitsrange	37
4.1.5.45	sm_spec_readrange	38
4.1.5.46	sm_spec_rebin	39
4.1.5.47	sm_spec_scale	39
4.1.5.48	sm_spec_scalerange	40
4.1.5.49	sm_spec_split	40
4.1.5.50	sm_spec_write	41
4.1.5.51	sm_spec_writecpl	41
4.1.5.52	sm_spec_writefits	42
4.2	Sky Model Module 2	42
4.2.1	Detailed Description	44
4.2.2	Define Documentation	44

4.2.2.1	SM_ERAD	44
4.2.2.2	SM_FILENAMELIST	44
4.2.2.3	SM_LAM_UNIT	44
4.2.2.4	SM_NSIGBIN	44
4.2.2.5	SM_RADMINLAM	44
4.2.2.6	SM_RRSTEP	44
4.2.2.7	SM_SIGMAX	45
4.2.2.8	SM_SKYEMCOMP_H	45
4.2.3	Typedef Documentation	45
4.2.3.1	smparmodel	45
4.2.4	Function Documentation	46
4.2.4.1	sm_comp_airglowcont	46
4.2.4.2	sm_comp_airtovac_single	47
4.2.4.3	sm_comp_airtovac_spec	48
4.2.4.4	sm_comp_alttoairmass	48
4.2.4.5	sm_comp_calcmiescat	49
4.2.4.6	sm_comp_calcmiescat1	49
4.2.4.7	sm_comp_calcrayleighscat	49
4.2.4.8	sm_comp_calcrayleighscat1	50
4.2.4.9	sm_comp_convertlinetab	50
4.2.4.10	sm_comp_convertlinetabo	51
4.2.4.11	sm_comp_extrapolrad	51
4.2.4.12	sm_comp_extrapoltrans	53
4.2.4.13	sm_comp_getlinespec	54
4.2.4.14	sm_comp_getmolspec	56
4.2.4.15	sm_comp_getmoonbedo	57
4.2.4.16	sm_comp_lunskybright	58
4.2.4.17	sm_comp_lunskybright_v	59
4.2.4.18	sm_comp_readlibstruct	62
4.2.4.19	sm_comp_readvarpar	62
4.2.4.20	sm_comp_scalelinetab	63
4.2.4.21	sm_comp_scaletranscurv	64
4.2.4.22	sm_comp_scatstarlight	64
4.2.4.23	sm_comp_skyemcomp	65

4.2.4.24	sm_comp_telem	68
4.2.4.25	sm_comp_vactoir_single	69
4.2.4.26	sm_comp_vactoir_spec	69
4.2.4.27	sm_comp_zodskybright	69
4.2.4.28	sm_etc_calcmodel	71
4.2.4.29	sm_etc_getparams	74
4.2.4.30	sm_etc_parameterlist_find_const	74
4.2.4.31	sm_etc_readfilenames	75
4.2.4.32	sm_etc_splitstring	75
4.2.4.33	sm_etc_writetable	75
5	Data Structure Documentation	77
5.1	_smdat_ Struct Reference	77
5.1.1	Detailed Description	77
5.1.2	Field Documentation	77
5.1.2.1	dflux1	77
5.1.2.2	dflux2	77
5.1.2.3	flux	77
5.1.2.4	lam	78
5.2	_smgrid_ Struct Reference	78
5.2.1	Detailed Description	78
5.2.2	Field Documentation	78
5.2.2.1	nx	78
5.2.2.2	ny	78
5.2.2.3	val	78
5.2.2.4	xpos	78
5.2.2.5	ypos	78
5.3	_smparam_ Struct Reference	79
5.3.1	Detailed Description	79
5.3.2	Field Documentation	79
5.3.2.1	c	79
5.3.2.2	d	79
5.3.2.3	i	79
5.3.2.4	n	79

5.4	<code>_smparmodel_ Struct Reference</code>	79
5.4.1	Detailed Description	80
5.4.2	Field Documentation	81
5.4.2.1	<code>acontname</code>	81
5.4.2.2	<code>airmass</code>	82
5.4.2.3	<code>alpha</code>	82
5.4.2.4	<code>alt</code>	82
5.4.2.5	<code>altmoon</code>	82
5.4.2.6	<code>datapath</code>	82
5.4.2.7	<code>eps</code>	82
5.4.2.8	<code>incl</code>	82
5.4.2.9	<code>lat_ecl</code>	82
5.4.2.10	<code>libpath</code>	82
5.4.2.11	<code>libstruct</code>	82
5.4.2.12	<code>libstruct1</code>	82
5.4.2.13	<code>libstruct2</code>	82
5.4.2.14	<code>linetabname</code>	82
5.4.2.15	<code>lon_ecl</code>	82
5.4.2.16	<code>lunirrname</code>	82
5.4.2.17	<code>miephasename</code>	82
5.4.2.18	<code>moondist</code>	82
5.4.2.19	<code>mscatname</code>	82
5.4.2.20	<code>msolflux</code>	82
5.4.2.21	<code>ncomp</code>	82
5.4.2.22	<code>o3transname</code>	82
5.4.2.23	<code>pwv</code>	82
5.4.2.24	<code>resol</code>	82
5.4.2.25	<code>rho</code>	82
5.4.2.26	<code>rcode</code>	83
5.4.2.27	<code>season</code>	83
5.4.2.28	<code>solspecname</code>	83
5.4.2.29	<code>starspecname</code>	83
5.4.2.30	<code>temp</code>	83
5.4.2.31	<code>time</code>	83

5.4.2.32	vac_air	83
5.4.2.33	vardatname	83
5.4.2.34	zodtabname	83
5.5	_smspec_ Struct Reference	83
5.5.1	Detailed Description	84
5.5.2	Field Documentation	84
5.5.2.1	dat	84
5.5.2.2	n	84
5.5.2.3	type	84
6	File Documentation	85
6.1	src/sm_general.c File Reference	85
6.1.1	Detailed Description	87
6.2	src/sm_general.h File Reference	87
6.2.1	Detailed Description	92
6.3	src/sm_scatmoonlight.c File Reference	92
6.3.1	Detailed Description	93
6.3.2	Function Documentation	94
6.3.2.1	sm_scat_calcextcurv	94
6.3.2.2	sm_scat_calcscatmoon	94
6.3.2.3	sm_scat_calcscatspec	95
6.3.2.4	sm_scat_calctau	96
6.3.2.5	sm_scat_createdatatab	96
6.3.2.6	sm_scat_getaerosoldens	97
6.3.2.7	sm_scat_geteffcoldens	97
6.3.2.8	sm_scat_getmiephasefunc	98
6.3.2.9	sm_scat_getmolecdens	98
6.3.2.10	sm_scat_moon	98
6.3.2.11	sm_scat_tau2lam	99
6.4	src/sm_scatmoonlight.h File Reference	100
6.4.1	Detailed Description	101
6.4.2	Define Documentation	102
6.4.2.1	SM_DTAU1	102
6.4.2.2	SM_DTAU2	102

6.4.2.3	SM_DTAU3	102
6.4.2.4	SM_H	102
6.4.2.5	SM_H0_M	102
6.4.2.6	SM_H0_R	102
6.4.2.7	SM_HMAX	102
6.4.2.8	SM_LNSCALE	102
6.4.2.9	SM_N0_M	102
6.4.2.10	SM_N0_R	102
6.4.2.11	SM_NS	103
6.4.2.12	SM_NTAU1	103
6.4.2.13	SM_NTAU2	103
6.4.2.14	SM_NTAU3	103
6.4.2.15	SM_R	103
6.4.2.16	SM_REFTAU0	103
6.4.2.17	SM_SCATMOONLIGHT_H	103
6.4.2.18	SM_SSA_M	103
6.4.3	Function Documentation	103
6.4.3.1	sm_scat_calcextcurv	103
6.4.3.2	sm_scat_calcscatmoon	104
6.4.3.3	sm_scat_calcscatspec	105
6.4.3.4	sm_scat_calctau	105
6.4.3.5	sm_scat_createdatatab	106
6.4.3.6	sm_scat_getaerosoldens	106
6.4.3.7	sm_scat_geteffcoldens	107
6.4.3.8	sm_scat_getmiephasefunc	107
6.4.3.9	sm_scat_getmolecdens	108
6.4.3.10	sm_scat_moon	108
6.4.3.11	sm_scat_tau2lam	109
6.5	src/sm_skyemcomp.c File Reference	109
6.5.1	Detailed Description	111
6.6	src/sm_skyemcomp.h File Reference	112
6.6.1	Detailed Description	114

Chapter 1

Module Index

1.1 Modules

Here is a list of all modules:

General Sky Model Module	7
Sky Model Module 2	42

Chapter 2

Data Structure Index

2.1 Data Structures

Here are the data structures with brief descriptions:

smdat	77
smgrid	78
smparam	79
smparmodel	79
smspec	83

Chapter 3

File Index

3.1 File List

Here is a list of all files with brief descriptions:

src/ sm_general.c	85
src/ sm_general.h	87
src/ sm_scatmoonlight.c	92
src/ sm_scatmoonlight.h	100
src/ sm_skyemcomp.c	109
src/ sm_skyemcomp.h	112

Chapter 4

Module Documentation

4.1 General Sky Model Module

Data Structures

- struct [_smdat_](#)
- struct [_smspec_](#)
- struct [_smparam_](#)
- struct [_smgrid_](#)

Files

- file [sm_general.c](#)
- file [sm_general.h](#)
- file [sm_general.h](#)

Defines

- #define [SM_GENERAL_H](#)
- #define [SM_BOOL](#)
- #define [F](#) [SM_F](#)
- #define [T](#) [SM_T](#)
- #define [SM_NMAXERR](#) 100
- #define [SM_LENLINE](#) 160
- #define [SM_MAXLEN](#) 4000
- #define [SM_MAXPAR](#) 21
- #define [SM_TOL](#) 1e-7
- #define [SM_SR_IN_ARCSEC2](#) 4.254517e+10
- #define [SM_GAS_CONST](#) 8.31447
- #define [SM_MOL_AIR_DRY](#) 0.0289644
- #define [SM_GRAV_ACC](#) 9.80665

- #define [SM_ERROR_FOF](#) SM_ERROR_FOPEN
- #define [SM_ERROR_ISM](#) SM_ERROR_NOMEM
- #define [SM_ERROR_EIS](#) SM_ERROR_SUBROUTINE
- #define [SM_ERROR_ISD](#) SM_ERROR_INSUFF_DATA
- #define [SM_ERROR_FOPEN_TXT](#) "File opening failed"
- #define [SM_ERROR_UFS_TXT](#) "Unexpected file structure"
- #define [SM_ERROR_IFE_TXT](#) "Invalid file name extension"
- #define [SM_ERROR_BDR_TXT](#) "Bad directory"
- #define [SM_ERROR_NDA_TXT](#) "No data"
- #define [SM_ERROR_ISD_TXT](#) "Insufficient data points"
- #define [SM_ERROR_IDG_TXT](#) "Inconsistent data grids"
- #define [SM_ERROR_IDR_TXT](#) "Invalid data range"
- #define [SM_ERROR_IOD_TXT](#) "Invalid order of data points"
- #define [SM_ERROR_IIP_TXT](#) "Invalid input parameter(s)"
- #define [SM_ERROR_IOV_TXT](#) "Invalid object value(s)"
- #define [SM_ERROR_IOS_TXT](#) "Invalid object structure"
- #define [SM_ERROR_SUBROUTINE_TXT](#) "Error in subroutine"
- #define [SM_ERROR_ACCES_TXT](#) "Permission denied"
- #define [SM_ERROR_LOOP_TXT](#) "Too many symbolic links"
- #define [SM_ERROR_NAMETOOLONG_TXT](#) "Pathname too long"
- #define [SM_ERROR_NOENT_TXT](#) "File/dir does not exist"
- #define [SM_ERROR_NOTDIR_TXT](#)
- #define [SM_ERROR_ROFS_TXT](#)
- #define [SM_ERROR_FAULT_TXT](#)
- #define [SM_ERROR_INVALID_TXT](#) "Mode was incorrectly specified"
- #define [SM_ERROR_IO_TXT](#) "I/O error occurred"
- #define [SM_ERROR_NOMEM_TXT](#) "Insufficient memory"
- #define [SM_ERROR_TXTBSY_TXT](#)
- #define [SM_ERROR_EXIST_TXT](#) "File/dir already exists"
- #define [SM_ERROR_NOSPC_TXT](#) "No space left on device"
- #define [SM_ERROR_PERM_TXT](#)
- #define [SM_ERROR_LINK_TXT](#) "Could not create symbolic link"
- #define [SM_ERROR_UNDEF_TXT](#) "Undefined error"
- #define [SM_ERROR_FOF_TXT](#) SM_ERROR_FOPEN_TXT
- #define [SM_ERROR_ISM_TXT](#) SM_ERROR_NOMEM_TXT
- #define [SM_ERROR_EIS_TXT](#) SM_ERROR_SUBROUTINE_TXT

Typedefs

- typedef enum [SM_BOOL](#) smbool
- typedef enum [_sm_error_code_](#) sm_error_code
- typedef struct [_smdat_](#) smdat
- typedef struct [_smspec_](#) smspec
- typedef struct [_smparam_](#) smparam
- typedef struct [_smgrid_](#) smgrid

Enumerations

- enum `SM_BOOL` { `SM_F`, `SM_T` }
- enum `_sm_error_code_` {
`SM_ERROR_FOPEN` = `CPL_ERROR_EOL` + 11, `SM_ERROR_UFS` = `CPL_ERROR_EOL` + 12, `SM_ERROR_IFE` = `CPL_ERROR_EOL` + 13, `SM_ERROR_NDA` = `CPL_ERROR_EOL` + 20,
`SM_ERROR_INSUFF_DATA` = `CPL_ERROR_EOL` + 21, `SM_ERROR_IDG` = `CPL_ERROR_EOL` + 22, `SM_ERROR_IDR` = `CPL_ERROR_EOL` + 23, `SM_ERROR_IOD` = `CPL_ERROR_EOL` + 24,
`SM_ERROR_IIP` = `CPL_ERROR_EOL` + 30, `SM_ERROR_IOV` = `CPL_ERROR_EOL` + 31, `SM_ERROR_IOS` = `CPL_ERROR_EOL` + 32, `SM_ERROR_SUBROUTINE` = `CPL_ERROR_EOL` + 40,
`SM_ERROR_ACCES` = `CPL_ERROR_EOL` + 50, `SM_ERROR_LOOP` = `CPL_ERROR_EOL` + 51, `SM_ERROR_NAMETOOLONG` = `CPL_ERROR_EOL` + 52, `SM_ERROR_NOENT` = `CPL_ERROR_EOL` + 53,
`SM_ERROR_NOTDIR` = `CPL_ERROR_EOL` + 54, `SM_ERROR_ROFS` = `CPL_ERROR_EOL` + 55, `SM_ERROR_FAULT` = `CPL_ERROR_EOL` + 56, `SM_ERROR_INVALID` = `CPL_ERROR_EOL` + 57,
`SM_ERROR_IO` = `CPL_ERROR_EOL` + 58, `SM_ERROR_NOMEM` = `CPL_ERROR_EOL` + 59, `SM_ERROR_TXTBSY` = `CPL_ERROR_EOL` + 60, `SM_ERROR_EXIST` = `CPL_ERROR_EOL` + 61,
`SM_ERROR_NOSPC` = `CPL_ERROR_EOL` + 62, `SM_ERROR_PERM` = `CPL_ERROR_EOL` + 63, `SM_ERROR_BADUSERINPUT` = `CPL_ERROR_EOL` + 70, `SM_ERROR_LINK` = `CPL_ERROR_EOL` + 71,
`SM_ERROR_RFM` = `CPL_ERROR_EOL` + 81, `SM_ERROR_UNDEF` = `CPL_ERROR_EOL` + 80 }

Functions

- `cpl_error_code sm_spec_malloc (smspec *spec, const int size)`
- `cpl_error_code sm_spec_create (smspec *outspec, const double limlam[2], const double dlam)`
- `cpl_error_code sm_spec_read (smspec *spec, const char *filename)`
- `cpl_error_code sm_spec_readrange (smspec *spec, const char *filename, const double limlam[2], const int step)`
- `cpl_error_code sm_spec_readcpl (smspec *spec, const cpl_table *cpltab)`
- `cpl_error_code sm_spec_readfits (smspec *spec, const char *filename)`
- `cpl_error_code sm_spec_readfitsrange (smspec *spec, const char *filename, const double limlam[2], const int step)`
- `cpl_error_code sm_spec_copy (smspec *outspec, const smspec *inspec)`
- `cpl_error_code sm_spec_compgrids (const smspec *spec1, const smspec *spec2)`
- `cpl_error_code sm_spec_join (smspec *spec, const smspec *errfunc, const int errflag)`
- `cpl_error_code sm_spec_split (smspec *spec, smspec *errfunc, const int errflag)`

- cpl_error_code [sm_spec_changetype](#) (smspec *spec, const int type)
- cpl_error_code [sm_spec_scalerange](#) (smspec *spec, const double limlam[2], const char op, const double c)
- cpl_error_code [sm_spec_scale](#) (smspec *spec, const char op, const double c)
- cpl_error_code [sm_spec_modval](#) (smspec *spec, const double lam, const char op, const double c)
- cpl_error_code [sm_spec_calc](#) (smspec *spec, const char op, const smspec *opspec)
- cpl_error_code [sm_spec_funct](#) (smspec *spec, const char *funct, const char baselab)
- cpl_error_code [sm_spec_functnoerr](#) (smspec *spec, const char *funct)
- cpl_error_code [sm_spec_convunits](#) (smspec *spec, const double factor, const int lamexp)
- cpl_error_code [sm_spec_changegrid](#) (smspec *spec, const double factor, const char *scale)
- cpl_error_code [sm_spec_average](#) (const smspec *spec, const double limlam[2], double mean[3])
- cpl_error_code [sm_spec_write](#) (const smspec *spec, const char *filename)
- cpl_error_code [sm_spec_print](#) (const smspec *spec)
- cpl_error_code [sm_spec_writecpl](#) (cpl_table *cpltab, const smspec *spec)
- cpl_error_code [sm_spec_writefits](#) (const smspec *spec, const char *filename)
- cpl_error_code [sm_spec_free](#) (smspec *spec)
- cpl_error_code [sm_spec_extract](#) (smspec *outspec, const smspec *inspec, const double limlam[2])
- cpl_error_code [sm_spec_rebin](#) (smspec *outspec, const smspec *inspec)
- cpl_error_code [sm_spec_interpol](#) (smspec *outspec, const smspec *inspec)
- cpl_error_code [sm_spec_convolve](#) (smspec *spec, const int nkpix, const double *kernel)
- cpl_error_code [sm_param_read](#) (FILE *stream, smparam par[])
- cpl_error_code [sm_param_readcheck](#) (FILE *stream, smparam par[], const char *parname, const int npar)
- cpl_error_code [sm_param_check](#) (smparam par[], const char *parname, const int npar)
- cpl_error_code [sm_grid_malloc](#) (smgrid *xy, const int nx, const int ny)
- cpl_error_code [sm_grid_read](#) (smgrid *xy, const char *filename)
- cpl_error_code [sm_grid_extract](#) (const smgrid *xy, const double x0, const double y0, double *outval)
- cpl_error_code [sm_grid_write](#) (const smgrid *xy, const char *filename)
- cpl_error_code [sm_grid_print](#) (const smgrid *xy)
- cpl_error_code [sm_grid_free](#) (smgrid *xy)
- cpl_error_code [sm_basic_chdir](#) (const char *dir)
- cpl_error_code [sm_basic_access](#) (const char *pathname, const int mode)
- cpl_error_code [sm_basic_mkdir](#) (const char *dir, const mode_t mode)
- cpl_error_code [sm_basic_createdir](#) (const char *dir, const mode_t mode)
- void [sm_basic_initstring](#) (char *str, const long n)
- cpl_boolean [sm_basic_isnumber](#) (char *str)
- char * [sm_basic_replacestring](#) (char *instr, char *oldsubstr, char *newsubstr)
- char * [sm_basic_rmcntrl](#) (char *str)

- void `sm_basic_rmcntrl_inplace` (char *str)
- char * `sm_basic_strtrim` (char *str)
- void `sm_basic_strtrim_inplace` (char *str)
- void `sm_basic_terminatestring` (char *str)
- cpl_error_code `sm_basic_interpollin` (const double *x_out, double *y_out, const long n_out, const double *x_ref, const double *y_ref, const long n_ref)

4.1.1 Detailed Description

This module provides basic definitions and functions for both sky model modules (*sm_module1* and *sm_module2*).

4.1.2 Define Documentation

4.1.2.1 #define F SM_F

Boolean false

4.1.2.2 #define SM_BOOL

Boolean type (F = false, T = true)

4.1.2.3 #define SM_ERROR_ACCES_TXT "Permission denied"

4.1.2.4 #define SM_ERROR_BDR_TXT "Bad directory"

4.1.2.5 #define SM_ERROR_EIS SM_ERROR_SUBROUTINE

4.1.2.6 #define SM_ERROR_EIS_TXT SM_ERROR_SUBROUTINE_TXT

4.1.2.7 #define SM_ERROR_EXIST_TXT "File/dir already exists"

4.1.2.8 #define SM_ERROR_FAULT_TXT

Value:

```
"Pathname points outside accessible " \
    "address space"
```

4.1.2.9 #define SM_ERROR_FOF SM_ERROR_FOPEN

4.1.2.10 #define SM_ERROR_FOF_TXT SM_ERROR_FOPEN_TXT

4.1.2.11 #define SM_ERROR_FOPEN_TXT "File opening failed"

4.1.2.12 `#define SM_ERROR_IDG_TXT "Inconsistent data grids"`

4.1.2.13 `#define SM_ERROR_IDR_TXT "Invalid data range"`

4.1.2.14 `#define SM_ERROR_IFE_TXT "Invalid file name extension"`

4.1.2.15 `#define SM_ERROR_IIP_TXT "Invalid input parameter(s)"`

4.1.2.16 `#define SM_ERROR_INVALID_TXT "Mode was incorrectly specified"`

4.1.2.17 `#define SM_ERROR_IO_TXT "I/O error occurred"`

4.1.2.18 `#define SM_ERROR_IOD_TXT "Invalid order of data points"`

4.1.2.19 `#define SM_ERROR_IOS_TXT "Invalid object structure"`

4.1.2.20 `#define SM_ERROR_IOV_TXT "Invalid object value(s)"`

4.1.2.21 `#define SM_ERROR_ISD SM_ERROR_INSUFF_DATA`

4.1.2.22 `#define SM_ERROR_ISD_TXT "Insufficient data points"`

4.1.2.23 `#define SM_ERROR_ISM SM_ERROR_NOMEM`

4.1.2.24 `#define SM_ERROR_ISM_TXT SM_ERROR_NOMEM_TXT`

4.1.2.25 `#define SM_ERROR_LINK_TXT "Could not create symbolic link"`

4.1.2.26 `#define SM_ERROR_LOOP_TXT "Too many symbolic links"`

4.1.2.27 `#define SM_ERROR_NAMETOOLONG_TXT "Pathname too long"`

4.1.2.28 `#define SM_ERROR_NDA_TXT "No data"`

4.1.2.29 `#define SM_ERROR_NOENT_TXT "File/dir does not exist"`

4.1.2.30 `#define SM_ERROR_NOMEM_TXT "Insufficient memory"`

4.1.2.31 `#define SM_ERROR_NOSPC_TXT "No space left on device"`

4.1.2.32 `#define SM_ERROR_NOTDIR_TXT`

Value:

```
"Component used as directory in pathname " \
    "is not a directory"
```


4.1.2.33 #define SM_ERROR_PERM_TXT

Value:

```
"File system does not support creation of " \
    "directories"
```

4.1.2.34 #define SM_ERROR_ROFS_TXT

Value:

```
"Write permission requested for file/dir " \
    "on read-only file system"
```

4.1.2.35 #define SM_ERROR_SUBROUTINE_TXT "Error in subroutine"

4.1.2.36 #define SM_ERROR_TXTBSY_TXT

Value:

```
"Write access requested to executable " \
    "which is being executed"
```

4.1.2.37 #define SM_ERROR_UFS_TXT "Unexpected file structure"

4.1.2.38 #define SM_ERROR_UNDEF_TXT "Undefined error"

4.1.2.39 #define SM_GAS_CONST 8.31447

Universal gas constant [J/(molK)]

4.1.2.40 #define SM_GENERAL_H

4.1.2.41 #define SM_GRAV_ACC 9.80665

Earth-surface gravitational acceleration [m/ s²]

4.1.2.42 #define SM_LENLINE 160

Maximum number of characters per line

4.1.2.43 #define SM_MAXLEN 4000

Maximum number of string characters

4.1.2.44 `#define SM_MAXPAR 21`

Maximum number of space-, tab-, or '='-separated strings per line (see [sm_param_read](#))

4.1.2.45 `#define SM_MOL_AIR_DRY 0.0289644`

Molar mass of dry air [kg/mol]

4.1.2.46 `#define SM_NMAXERR 100`

Maximum number of error codes

4.1.2.47 `#define SM_SR_IN_ARCSEC2 4.254517e+10`

steradians in arcsec²

4.1.2.48 `#define SM_TOL 1e-7`

Required relative accuracy for data comparisons

4.1.2.49 `#define T SM_T`

Boolean true

4.1.3 Typedef Documentation

4.1.3.1 `typedef enum _sm_error_code_ sm_error_code`

Enumeration structure for sky model related error codes

4.1.3.2 `typedef enum SM_BOOL smbool`

4.1.3.3 `typedef struct _smdat_ smdat`

Structure for data points

Parameters

<i>lam</i>	wavelength
<i>flux</i>	flux
<i>dflux1</i>	(lower) flux error
<i>dflux2</i>	upper flux error

4.1.3.4 typedef struct _smgrid_ smgrid

Structure for coordinate grids

Parameters

<i>nx</i>	number of x coordinate values
<i>ny</i>	number of y coordinate values
<i>*xpos</i>	vector of x coordinate values
<i>*ypos</i>	vector of y coordinate values
<i>**val</i>	matrix of data values for the coordinate grid

4.1.3.5 typedef struct _smparam_ smparam

Structure for a parameter value in different data types

Parameters

<i>c</i>	string (maximum length SM_LENLINE)
<i>i</i>	integer number
<i>d</i>	float number of double precision
<i>n</i>	number of value (for arrays; counted backwards)

4.1.3.6 typedef struct _smspec_ smspec

Structure for spectra

Parameters

<i>type</i>	number of columns (1 = no flux, 2 = no errors, 3 = symmetric error, 4 = lower and upper error)
<i>n</i>	number of wavelengths
<i>*dat</i>	vector of data points defined by smdat structure

4.1.4 Enumeration Type Documentation

4.1.4.1 enum _sm_error_code_

Enumeration structure for sky model related error codes

Enumerator:

SM_ERROR_FOPEN
SM_ERROR_UFS
SM_ERROR_IFE
SM_ERROR_NDA

SM_ERROR_INSUFF_DATA
SM_ERROR_IDG
SM_ERROR_IDR
SM_ERROR_IOD
SM_ERROR_IIP
SM_ERROR_IOV
SM_ERROR_IOS
SM_ERROR_SUBROUTINE
SM_ERROR_ACCES
SM_ERROR_LOOP
SM_ERROR_NAMETOOLONG
SM_ERROR_NOENT
SM_ERROR_NOTDIR
SM_ERROR_ROFS
SM_ERROR_FAULT
SM_ERROR_INVALID
SM_ERROR_IO
SM_ERROR_NOMEM
SM_ERROR_TXTBSY
SM_ERROR_EXIST
SM_ERROR_NOSPC
SM_ERROR_PERM
SM_ERROR_BADUSERINPUT
SM_ERROR_LINK
SM_ERROR_RFM
SM_ERROR_UNDEF

4.1.4.2 enum SM_BOOL

Enumerator:

SM_F
SM_T

4.1.5 Function Documentation

4.1.5.1 cpl_error_code sm_basic_access (const char * *pathname*, const int *mode*)

Check for file or directory existence

This function provides a wrapper for the intrinsic function access(). It reports back all access() errors.

INPUT:

Parameters

<i>pathname</i>	name of a file or directory
<i>mode</i>	mode for accessibility checks F_OK = file existence X_OK = file exec permission W_OK = file write permission

OUTPUT:**Returns**

CPL_ERROR_NONE on success, or on failure: SM_ERROR_ROFS, SM_ERROR_INVALID, SM_ERROR_TXTBSY, SM_ERROR_ACCES, SM_ERROR_LOOP, SM_ERROR_NAMETOOLONG, SM_ERROR_NOENT, SM_ERROR_NOTDIR, SM_ERROR_FAULT, SM_ERROR_IO, SM_ERROR_NOMEM, SM_ERROR_UNDEF

ERRORS:

- see access manual for details

4.1.5.2 cpl_error_code sm.basic.chdir (const char * *dir*)

Change working directory

This function provides a wrapper for the intrinsic function chdir(). It reports back all chdir() errors.

INPUT:**Parameters**

<i>dir</i>	new working directory
------------	-----------------------

OUTPUT:**Returns**

CPL_ERROR_NONE on success, or on failure: SM_ERROR_ACCES, SM_ERROR_LOOP, SM_ERROR_NAMETOOLONG, SM_ERROR_NOENT, SM_ERROR_NOTDIR, SM_ERROR_FAULT, SM_ERROR_IO, SM_ERROR_NOMEM, SM_ERROR_UNDEF

ERRORS:

- see chdir manual for details

4.1.5.3 cpl_error_code sm.basic.createdir (const char * *dir*, const mode_t *mode*)

Create directory including error handling

This function creates a new directory with the specified permissions.

INPUT:

Parameters

<i>dir</i>	name of directory
<i>mode</i>	access permissions

OUTPUT:**Returns**

CPL_ERROR_NONE on success, SM_ERROR_SUBROUTINE else.

ERRORS:

- see mkdir manual for details

NOTE:

- routine uses [sm_basic_mkdir](#)

4.1.5.4 void sm_basic_initstring (char * *str*, const long *n*)

Initialise a string variable.

This function initialises a given string *str* of length *n* with "\0".

INPUT:**Parameters**

<i>n</i>	length of string
----------	------------------

OUTPUT:**Parameters**

<i>str</i>	string
------------	--------

Returns

nothing

4.1.5.5 cpl_error_code sm_basic_interpollin (const double * *x_out*, double * *y_out*, const long *n_out*, const double * *x_ref*, const double * *y_ref*, const long *n_ref*)

Linear interpolation.

This function calculates the interpolated y-values *y_out* at the positions *x_out* with respect to the reference x/y pairs (*x_ref* / *y_ref*).

Note

Points outside the range of the reference vectors will be extrapolated based on the

first / last values in the reference vectors for the low / high end, respectively.

INPUT:

Parameters

<i>x_out</i>	desired output x-spacing
<i>n_out</i>	length of <i>x_out</i> / <i>y_out</i>
<i>x_ref</i>	reference x-spacing
<i>y_ref</i>	reference y-values at <i>x_ref</i>
<i>n_ref</i>	length of <i>x_ref</i> / <i>y_ref</i>

OUTPUT:

Parameters

<i>y_out</i>	requested y-values at <i>x_out</i>
--------------	------------------------------------

Returns

CPL_ERROR_NONE on success, CPL_ERROR_DIVISION_BY_ZERO else

4.1.5.6 `cpl_boolean sm_basic_isnumber (char * str)`

Check if string contains number.

This function checks whether *str* contains a valid number. A leading "+" or "-" is treated as a sign. A single "." is identified as a decimal point. Finally, an "e" is accepted if an exponent follows this letter.

Note

Surrounding spaces are not treated.

INPUT:

Parameters

<i>str</i>	string
------------	--------

OUTPUT:

Returns

CPL_TRUE if *str* is number, CPL_FALSE else.

4.1.5.7 `cpl_error_code sm_basic_mkdir (const char * dir, const mode_t mode)`

Create directory

This function provides a wrapper for the intrinsic function `mkdir()`. It reports back all `mkdir()` errors.

INPUT:**Parameters**

<i>dir</i>	name of directory
<i>mode</i>	access permissions

OUTPUT:**Returns**

CPL_ERROR_NONE on success, or on failure: SM_ERROR_EXIST, SM_ERROR_NOSPC, SM_ERROR_PERM, SM_ERROR_ROFS, SM_ERROR_INVALID, SM_ERROR_TXTBSY, SM_ERROR_ACCES, SM_ERROR_LOOP, SM_ERROR_NAMETOOLONG, SM_ERROR_NOENT, SM_ERROR_NOTDIR, SM_ERROR_FAULT, SM_ERROR_IO, SM_ERROR_NOMEM, SM_ERROR_UNDEF

ERRORS:

- see `mkdir` manual for details

4.1.5.8 char* sm_basic_replacestring (char * *instring*, char * *oldsubstr*, char * *newsubstr*)

Substitutes substring in string

Parameters

<i>instring</i>	input string
<i>oldsubstr</i>	substring to be replaced
<i>newsubstr</i>	new substring

Returns

input string with substituted substring

4.1.5.9 char* sm_basic_rmcntrl (char * *str*)

Remove control characters from string.

This function removes all ASCII control characters from *str* using `iscntrl()`.

INPUT:**Parameters**

<i>str</i>	string.
------------	---------

OUTPUT:

Returns

string.

4.1.5.10 void sm_basic_rmctrl_inplace (char * *str*)

Remove control characters from string.

This function removes all ASCII control characters from *str* using `isctrl()`. In contrast to `sm_basic_rmctrl()`, the action is performed in place, i.e. the input gets overwritten.

INPUT & OUTPUT:

Parameters

<i>str</i>	string
------------	--------

OUTPUT:**Returns**

nothing

4.1.5.11 char* sm_basic_strtrim (char * *str*)

Remove leading and trailing blanks from string.

This function removes all leading and trailing " " characters from *str* using `isspace()`.

INPUT:

Parameters

<i>str</i>	string
------------	--------

OUTPUT:**Returns**

string

4.1.5.12 void sm_basic_strtrim_inplace (char * *str*)

Remove leading and trailing blanks from string.

This function removes all leading and trailing " " characters from *str* using `isspace()`. In contrast to `sm_basic_strtrim()`, the action is performed in place, i.e. the input gets overwritten.

INPUT & OUTPUT:

Parameters

<i>str</i>	string
------------	--------

OUTPUT:**Returns**

nothing

4.1.5.13 void sm_basic_terminatestring (char * *str*)

Put "\0" at end of string.

This function places a "\0" at end of the input string.

Note

No checks are performed for writing beyond the boundary of allocated memory for the input string.

INPUT & OUTPUT:**Parameters**

<i>str</i>	string
------------	--------

OUTPUT:**Returns**

nothing

4.1.5.14 cpl_error_code sm_grid_extract (const smgrid * *xy*, const double *x0*, const double *y0*, double * *outval*)

Extracts value of an [smgrid](#) structure at position (*x0*, *y0*) by means of bilinear interpolation

INPUT:**Parameters**

<i>xy</i>	input smgrid structure
<i>x0</i>	<i>x</i> coordinate position for extraction
<i>y0</i>	<i>y</i> coordinate position for extraction

OUTPUT:**Parameters**

<i>outval</i>	value of smgrid structure at position (<i>x0</i> , <i>y0</i>)
---------------	---

ERRORS:

- NDA: No data
- ISD: Insufficient data points
- IOD: Invalid order of data points
- IIP: Invalid input parameter(s)

4.1.5.15 `cpl_error_code sm_grid_free (smgrid * xy)`

Frees memory occupied by an `smgrid` structure

INPUT:**Parameters**

<code>xy</code>	<code>smgrid</code> structure of size $n_x \times n_y$
-----------------	--

OUTPUT:**Parameters**

<code>xy</code>	<code>smgrid</code> structure of size zero
-----------------	--

ERRORS:

- none

4.1.5.16 `cpl_error_code sm_grid_malloc (smgrid * xy, const int nx, const int ny)`

Allocates memory for an `smgrid` structure of size $n_x \times n_y$. All values are zero.

INPUT:**Parameters**

<code>xy</code>	<code>smgrid</code> structure of size zero
<code>nx</code>	number of x coordinates
<code>ny</code>	number of y coordinates

OUTPUT:**Parameters**

<code>xy</code>	<code>smgrid</code> structure of size $n_x \times n_y$ if there is sufficient memory (otherwise size remains zero)
-----------------	--

ERRORS:

- IIP: Invalid input parameter(s)
- ISM: Insufficient memory

4.1.5.17 `cpl_error_code sm_grid_print (const smgrid * xy)`

Prints `smgrid` structure on stdout

INPUT:

Parameters

<code>xy</code>	<code>smgrid</code> structure
-----------------	-------------------------------

ERRORS:

- none

4.1.5.18 `cpl_error_code sm_grid_read (smgrid * xy, const char * filename)`

Reads `smgrid` structure from ASCII file.

Required file structure:

- Comment lines have to be marked by #.
- Data lines:
 - 1: `nx ny`
 - 2: `xpos (nx values)`
 - 3: `ypos (ny values)`
 - 4-(`nx+3`): `val (ny values each line)`

INPUT:

Parameters

<code>filename</code>	name of input file
-----------------------	--------------------

OUTPUT:

Parameters

<code>xy</code>	output <code>smgrid</code> structure
-----------------	--------------------------------------

ERRORS:

- FOF: File opening failed
- UFS: Unexpected file structure

- ISM: Insufficient memory

4.1.5.19 `cpl_error_code sm_grid_write (const smgrid * xy, const char * filename)`

Writes [smgrid](#) structure to ASCII file or on stdout (indicated by "stdout" as filename)

INPUT:

Parameters

<i>xy</i>	smgrid structure
<i>filename</i>	name of output file or "stdout"

ERRORS:

- FOF: File opening failed

4.1.5.20 `cpl_error_code sm_param_check (smparam par[], const char * parname, const int npar)`

Checks a parameter definition held by an array of [smparam](#) structures. The first string of the read line is expected to be the parameter name. It is checked whether it matches the input name. In the case of a deviation, an error is returned. The routine also checks the number of parameter values (not counting the parameter name), which is also an input parameter of the routine. An error is returned if there is a mismatch. In the case of errors, all entries of the [smparam](#) array are set to the default values "", 0, and 0.0.

INPUT:

Parameters

<i>par</i>	array of SM_MAXPAR smparam structures
<i>parname</i>	name of parameter
<i>npar</i>	number of expected parameter values

OUTPUT:

Parameters

<i>par</i>	smparam array (error: set to the data type default values, otherwise untouched)
------------	---

ERRORS:

- EIS: Error in subroutine (see [sm_param_read](#))
- IOV: Invalid object value(s)

4.1.5.21 `cpl_error_code sm_param_read (FILE * stream, smparam par[])`

Reads a line from a file and returns the space-, tab-, or '='-separated individual strings in an array of [SM_MAXPAR *smparam*](#) structures. A *smparam* structure contains each string converted into integer and double precision floating point numbers. Comment lines marked by # and empty lines are not considered. In this case, the next line with data is read.

INPUT:

Parameters

<i>stream</i>	name of stream (to be defined by fopen command)
<i>par</i>	array of SM_MAXPAR <i>smparam</i> structures (to transfer parameter values)

OUTPUT:

Parameters

<i>par</i>	smparam array that contains the read value(s) as character string ("c"), integer ("i"), or double precision floating point number ("d"). The different data types can be accessed by adding the corresponding suffixes "c", "i", or "d" to the name of the smparam structure. An additional suffix "n" element indicates the number of read values minus the array index, i.e. <i>par</i> [0].n provides the number of values in the read line.
------------	---

ERRORS:

- IIP: Invalid input parameter(s)
- UFS: Unexpected file structure
- ISM: Insufficient memory

In the case of problems, the default values "", 0, and 0.0 are used for the corresponding data types.

4.1.5.22 `cpl_error_code sm_param_readcheck (FILE * stream, smparam par[], const char * parname, const int npar)`

Reads and checks a line from a file and returns the space-, tab-, or '='-separated individual strings in an array of [SM_MAXPAR *smparam*](#) structures. A *smparam* structure contains each string converted into integer and double precision floating point numbers. Comment lines marked by # and empty lines are not considered. In this case, the next line with data is read.

The first string of the read line is expected to be the parameter name. It is checked whether it matches the input name. In the case of a deviation, an error is returned. The routine also checks the number of parameter values (not counting the parameter name), which is also an input parameter of the routine. An error is returned if there is a mismatch.

INPUT:

Parameters

<i>stream</i>	name of stream (to be defined by fopen command)
<i>par</i>	array of SM_MAXPAR smparam structures (to transfer parameter values)
<i>parname</i>	name of parameter
<i>npar</i>	number of expected parameter values

OUTPUT:**Parameters**

<i>par</i>	smparam array that contains the read value(s) as character string ("c"), integer ("i"), or double precision floating point number ("d"). The different data types can be accessed by adding the corresponding suffixes "c", "i", or "d" to the name of the smparam structure. An additional suffix "n" element indicates the number of read values minus the array index, i.e. <i>par</i> [0].n provides the number of values in the read line.
------------	---

ERRORS:

- EIS: Error in subroutine (see [sm_param_read](#))
- IOV: Invalid object value(s)

In the case of problems, the default values "", 0, and 0.0 are used for the corresponding data types.

4.1.5.23 `cpl_error_code sm_spec_average (const smspec * spec, const double limlam[2], double mean[3])`

Calculates mean of a spectrum (and its error function if given) in the given wavelength range

INPUT:**Parameters**

<i>spec</i>	input spectrum
<i>limlam</i>	lower and upper wavelength limit for operation (Take HUGE_VAL for unconstrained limits.)

OUTPUT:**Parameters**

<i>mean</i>	mean values for spectrum (and lower and upper error function if given)
-------------	--

ERRORS:

- NDA: No data
- IIP: Invalid input parameter(s)

4.1.5.24 `cpl_error_code sm_spec_calc (smspec * spec, const char op, const smspec * opspect)`

Performs addition, subtraction, multiplication, division, or equalisation of two spectra with error functions. Moreover, a spectrum can be used as a power of another spectrum. Error propagation (adding of squared errors) is considered. For asymmetric errors positive fluxes are required for multiplicative operations.

The two subtraction operators '-' and '_' differ by the order of the spectra, i.e., '-' and '_' imply $x_1 - x_2$ and $x_2 - x_1$, respectively. The two division operators '/' and '|' differ by the order of the spectra, i.e., '/' and '|' imply x_1/x_2 and x_2/x_1 , respectively.

INPUT:

Parameters

<i>spec</i>	input smspec structure
<i>op</i>	operation: '+', '-', '_', '*', '/', ' ', '^', or '='
<i>opspect</i>	smspec structure to modify spec (identical wavelengths needed!)

OUTPUT:

Parameters

<i>spec</i>	modified spectrum (no modification in the case of errors)
-------------	---

ERRORS:

- IIP: Invalid input parameter(s)
- CPL: Division by zero [warning]
- CPL: Access out of range [warning]

4.1.5.25 `cpl_error_code sm_spec_changegrid (smspec * spec, const double factor, const char * scale)`

Modifies the wavelengths of an [smspec](#) structure by multiplying a factor and/or performing $\log_{10}(x)$ or 10^x .

INPUT:

Parameters

<i>spec</i>	input smspec structure
<i>factor</i>	wavelengths are multiplied by this value (<i>c</i>).
<i>scale</i>	wavelength scale, options: <ul style="list-style-type: none"> • "log": logarithmic ($\log_{10}(cx)$) • "exp": exponential ($c 10^x$) \rightarrow reverses "log" option • "lin": linear (cx)

OUTPUT:**Parameters**

<i>spec</i>	spectrum with modified wavelength scale (no modification in the case of errors)
-------------	---

ERRORS:

- IDR: Invalid data range
- IIP: Invalid input parameter(s)

4.1.5.26 `cpl_error_code sm_spec.changetype (smspec * spec, const int type)`

Changes spectrum type (1 = no flux, 2 = no errors, 3 = symmetric error, 4 = lower and upper error). For changes to 1 or 2 the obsolete columns are set to zero. For a change from 4 to 3 the lower and upper errors are made symmetric by simple averaging. For a change from 1-3 to 4 or 3 it is ensured that the lower and upper errors are equal (using the lower errors as reference). Only for a change from 4 to 4 nothing is done.

INPUT:**Parameters**

<i>spec</i>	spectrum of arbitrary type
<i>type</i>	spectrum type (see smspec)

OUTPUT:**Parameters**

<i>spec</i>	spectrum of given <i>type</i>
-------------	-------------------------------

ERRORS:

- IIP: Invalid input parameter(s)

4.1.5.27 `cpl_error_code sm_spec.compgrids (const smspec * spec1, const smspec * spec2)`

Tests agreement of wavelength grids of two smspec structures. Returns $\neq 0$ for disagreement.

INPUT:**Parameters**

<i>spec1</i>	first spectrum
<i>spec2</i>	second spectrum

ERRORS:

- IDG: Inconsistent data grids

4.1.5.28 `cpl_error_code sm_spec_convolve (smspec * spec, const int nkpix, const double * kernel)`

Convolution of `smspec` structure with given kernel. Errors are convolved in the same way as fluxes → not suitable for statistical noise.

Note

The centre of the convolution function is shifted by -0.5 pixels for an even number of kernel pixels.

INPUT:**Parameters**

<i>spec</i>	input spectrum (<code>smspec</code> structure)
<i>nkpix</i>	number of kernel pixels
<i>kernel</i>	vector of kernel values (required: sum of all values = 1)

OUTPUT:**Parameters**

<i>spec</i>	convolved input spectrum
-------------	--------------------------

ERRORS:

- IIP: Invalid input parameter(s)

4.1.5.29 `cpl_error_code sm_spec_convunits (smspec * spec, const double factor, const int lamexp)`

Modifies the flux of an `smspec` structure by multiplying a *factor* times λ^{lamexp} .

INPUT:**Parameters**

<i>spec</i>	input <code>smspec</code> structure
<i>factor</i>	constant factor
<i>lamexp</i>	power for wavelengths

OUTPUT:

Parameters

<i>spec</i>	modified spectrum
-------------	-------------------

ERRORS:

- IDR: Invalid data range

4.1.5.30 `cpl_error_code sm_spec_copy (smspec * outspec, const smspec * inspec)`

Copies [smspec](#) structure

INPUT:**Parameters**

<i>inspec</i>	input spectrum
---------------	----------------

OUTPUT:**Parameters**

<i>outspec</i>	output spectrum
----------------	-----------------

ERRORS:

- none

4.1.5.31 `cpl_error_code sm_spec_create (smspec * outspec, const double limlam[2], const double dlam)`

Initialisation of a wavelength grid that is characterised by constant bin size. The spectrum type 3 is set (see [smspec](#)).

INPUT:**Parameters**

<i>limlam</i>	lower and upper limit of wavelength range
<i>dlam</i>	bin size

OUTPUT:**Parameters**

<i>outspec</i>	smspec structure that provides wavelengths, fluxes (= 0.), and flux errors (= 0.) of a spectrum
----------------	---

ERRORS:

- IIP: Invalid input parameter(s)

4.1.5.32 `cpl_error_code sm_spec_extract (smspec * outspec, const smspec * inspec, const double limlam[2])`

Extracts subspectrum in a given wavelength range of the input spectrum (:: smspec structure)

INPUT:

Parameters

<i>inspec</i>	input spectrum
<i>limlam</i>	wavelength limits for extraction

OUTPUT:

Parameters

<i>outspec</i>	extracted spectrum
----------------	--------------------

ERRORS:

- IIP: Invalid input parameter(s)

4.1.5.33 `cpl_error_code sm_spec_free (smspec * spec)`

Frees memory occupied by an [smspec](#) structure

INPUT:

Parameters

<i>spec</i>	smspec structure with n data points
-------------	---

OUTPUT:

Parameters

<i>spec</i>	smspec structure with zero data points
-------------	--

ERRORS:

- none

4.1.5.34 `cpl_error_code sm_spec_funct (smspec * spec, const char * funct, const char baselab)`

Applies exponential or logarithmic function. Names of functions equal those of `<math.h>` excepting "exp10", "expmag", and "logmag" which correspond to 10^x , $10^{-0.4x}$, and $-2.5 \log_{10}(x)$, respectively. Correction of resulting error asymmetry by means of the recipe $\frac{1}{2}(f(x+dx) - f(x-dx))$ for spectrum type 3 (see [smspec](#)).

INPUT:**Parameters**

<i>spec</i>	input smspec structure
<i>funct</i>	"exp" or "log"
<i>baselab</i>	'e', 'd' (10), or 'm' ($10^{-0.4}$)

OUTPUT:**Parameters**

<i>spec</i>	modified spectrum if function is valid
-------------	--

ERRORS:

- IIP: Invalid input parameter(s)
- CPL: Access out of range [warning]

4.1.5.35 cpl_error_code sm_spec_funcnoerr (smspec * spec, const char * funct)

Applies mathematical function to flux. A possible error function is not changed. Names of functions equal those of `<math.h>` excepting "exp10", "expmag", and "logmag" which correspond to 10^x , $10^{-0.4x}$, and $-2.5 \log_{10}(x)$, respectively.

INPUT:**Parameters**

<i>spec</i>	input smspec structure
<i>funct</i>	"acos", "asin", "atan", "cos", "sin", "tan", "cosh", "sinh", "tanh", "exp", "exp10", "expmag", "log", "log10", "logmag", "sqrt", "ceil", "fabs", "floor"

OUTPUT:**Parameters**

<i>spec</i>	modified spectrum if function is valid
-------------	--

ERRORS:

- IIP: Invalid input parameter(s)
- CPL: Access out of range [warning]

4.1.5.36 cpl_error_code sm_spec_interpol (smspec * outspec, const smspec * inspec)

Linear interpolation of [smspec](#) structure for conversion to finer wavelength grid. Optimal for sparse, unbinned data.

INPUT:**Parameters**

<i>outspec</i>	desired wavelength grid
<i>inspec</i>	wavelengths and fluxes (per wavelength unit) of input spectrum

OUTPUT:**Parameters**

<i>outspec</i>	interpolated input spectrum
----------------	-----------------------------

ERRORS:

- none

4.1.5.37 `cpl_error_code sm_spec_join (smspec * spec, const smspec * errfunc, const int errflag)`

Uses the flux of a second [smspec](#) structure to fill the error vector of the first spectrum. If two error columns are used, lower or upper error has to be specified and the function has to be called for either case (starting with the lower error).

Note

Pre-existing error data in spec will be overwritten.

INPUT:**Parameters**

<i>spec</i>	spectrum without errors (2 columns)
<i>errfunc</i>	error function (2 columns)
<i>errflag</i>	error flag (0 = symmetric error, 1 = lower error, 2 = upper error)

OUTPUT:**Parameters**

<i>spec</i>	spectrum with errors (3 columns)
-------------	----------------------------------

ERRORS:

- IIP: Invalid input parameter(s)

4.1.5.38 `cpl_error_code sm_spec_malloc (smspec * spec, const int size)`

Allocates memory for an [smspec](#) structure of given size. Assumes symmetric errors (see [smspec](#)).

INPUT:**Parameters**

<i>spec</i>	smspec structure (no memory allocated)
<i>size</i>	number of data points

OUTPUT:**Parameters**

<i>spec</i>	smspec structure with allocated memory (all values = 0)
-------------	---

ERRORS:

- ISM: Insufficient memory

4.1.5.39 `cpl_error_code sm_spec_modval (smspec * spec, const double lam, const char op, const double c)`

Modifies the flux of a given wavelength belonging to an [smspec](#) structure by adding, subtracting, multiplying, or dividing a constant. Moreover, the flux can be raised to a constant power or the flux of the selected wavelength can be substituted.

The two subtraction operators '-' and '_' differ by the order of spectrum and constant, i.e., '-' and '_' imply $x - c$ and $c - x$, respectively. The two division operators '/' and '|' differ by the order of spectrum and constant, i.e., '/' and '|' imply x/c and c/x , respectively.

INPUT:**Parameters**

<i>spec</i>	input smspec structure
<i>lam</i>	wavelength (accuracy ruled by SM_TOL)
<i>op</i>	operation: '+', '-', '*', '/', '^', '=', '_', or ' '
<i>c</i>	constant (double precision float)

OUTPUT:**Parameters**

<i>spec</i>	modified spectrum (no modification in the case of errors)
-------------	---

ERRORS:

- see [sm_spec_scalerange](#)

4.1.5.40 `cpl_error_code sm_spec_print (const smspec * spec)`

Prints [smspec](#) structure on stdout

INPUT:**Parameters**

<i>spec</i>	input smspec structure
-------------	--

ERRORS:

- none

4.1.5.41 `cpl_error_code sm_spec_read (smspec * spec, const char * filename)`

Fills [smspec](#) structure by data read from a file. Each line of the file must consist of wavelength and flux. The presence of one or two optional error columns is detected automatically. Header lines are allowed if they are marked by #.

INPUT:**Parameters**

<i>filename</i>	name of data file
-----------------	-------------------

OUTPUT:**Parameters**

<i>spec</i>	output spectrum with read values
-------------	----------------------------------

ERRORS:

- FOF: File opening failed
- UFS: Unexpected file structure
- CPL: Access out of range [warning]

4.1.5.42 `cpl_error_code sm_spec_readcpl (smspec * spec, const cpl_table * cpltab)`

Reads [smspec](#) structure from CPL table with one to four columns. The column IDs "lam", "flux", "dflux" (for symmetric errors), "dflux1" (lower error), and "dflux2" (upper error) are mandatory.

INPUT:**Parameters**

<i>cpltab</i>	CPL table with one to four columns named "lam", "flux", "dflux"/"dflux1", or "dflux2"
---------------	---

OUTPUT:

Parameters

<i>spec</i>	spectrum of structure smspec
-------------	--

ERRORS:

- NDA: No data
- IOS: Invalid object structure
- CPL: Access out of range [warning]

4.1.5.43 `cpl_error_code sm_spec_readfits (smspec * spec, const char * filename)`

Reads [smspec](#) structure from FITS file by means of CPL

INPUT:**Parameters**

<i>filename</i>	name of FITS file (extensions "fits" or "mt")
-----------------	---

OUTPUT:**Parameters**

<i>spec</i>	spectrum of structure smspec
-------------	--

ERRORS:

- IFE: Invalid file name extension
- FOF: File opening failed
- UFS: Unexpected file structure

4.1.5.44 `cpl_error_code sm_spec_readfitsrange (smspec * spec, const char * filename, const double limlam[2], const int step)`

Reading of a spectrum from FITS file in a given wavelength range by means of CPL. The wavelength range of the extracted spectrum tends to be wider than the interval between the provided limits and depends on the given step size.

INPUT:**Parameters**

<i>filename</i>	name of FITS file (extensions "fits" or "mt")
<i>limlam</i>	wavelength limits for extraction
<i>step</i>	step in FITS table rows for search of given wavelength range (optimum: square root of total number of rows)

OUTPUT:**Parameters**

<i>spec</i>	spectrum of structure smspec
-------------	--

ERRORS:

- IIP: Invalid input parameter(s)
- IFE: Invalid file name extension
- FOF: File opening failed
- UFS: Unexpected file structure

4.1.5.45 `cpl_error_code sm_spec_readrange (smspec * spec, const char * filename, const double limlam[2], const int step)`

Reading of a spectrum from file in a given wavelength range. The wavelength range of the extracted spectrum tends to be wider than the interval between the provided limits and depends on the given step size. The presence of one or two optional error columns is detected automatically. Header lines are allowed if they are marked by #.

INPUT:**Parameters**

<i>filename</i>	name of file with wavelength and flux (and error) columns
<i>limlam</i>	wavelength limits for extraction
<i>step</i>	step in file lines for search of given wavelength range (optimum: square root of total number of lines)

OUTPUT:**Parameters**

<i>spec</i>	spectrum of structure smspec
-------------	--

ERRORS:

- IIP: Invalid input parameter(s)
- FOF: File opening failed
- UFS: Unexpected file structure
- CPL: Access out of range [warning]

4.1.5.46 `cpl_error_code sm_spec_rebin (smspec * outspec, const smspec * inspec)`

Rebins [smspec](#) structure by using bins with constant fluxes. Conserves integral of fluxes. Errors are rebinned in the same way as fluxes → not suitable for statistical noise.

INPUT:**Parameters**

<i>outspec</i>	desired wavelength grid
<i>inspec</i>	wavelengths and fluxes (per wavelength unit) of input spectrum

OUTPUT:**Parameters**

<i>outspec</i>	rebinned input spectrum
----------------	-------------------------

ERRORS:

- none

4.1.5.47 `cpl_error_code sm_spec_scale (smspec * spec, const char op, const double c)`

Modifies the flux of an [smspec](#) structure by adding, subtracting, multiplying, or dividing a constant. Moreover, the flux can be raised to a constant power or the full spectrum can be given a constant flux value.

The two subtraction operators '-' and '_' differ by the order of spectrum and constant, i.e., '-' and '_' imply $x - c$ and $c - x$, respectively. The two division operators '/' and '|' differ by the order of spectrum and constant, i.e., '/' and '|' imply x/c and c/x , respectively.

INPUT:**Parameters**

<i>spec</i>	input smspec structure
<i>op</i>	operation: '+', '-', '*', '/', '^', '=', '_', or ' '
<i>c</i>	constant (double precision float)

OUTPUT:**Parameters**

<i>spec</i>	modified spectrum (no modification in the case of errors)
-------------	---

ERRORS:

- see [sm_spec_scalerange](#)

4.1.5.48 `cpl_error_code sm_spec_scalerange (smspec * spec, const double limlam[2], const char op, const double c)`

Modifies the flux of an [smspec](#) structure in a given wavelength range by adding, subtracting, multiplying, or dividing a constant. Moreover, the flux can be raised to a constant power or the selected wavelength range can be given a constant flux value.

The two subtraction operators '-' and '_' differ by the order of spectrum and constant, i.e., '-' and '_' imply $x - c$ and $c - x$, respectively. The two division operators '/' and '|' differ by the order of spectrum and constant, i.e., '/' and '|' imply x/c and c/x , respectively.

INPUT:

Parameters

<i>spec</i>	input smspec structure
<i>limlam</i>	lower and upper wavelength limit for operation (Take HUGE_VAL for unconstrained limits.)
<i>op</i>	operation: '+', '-', '_', '*', '/', ' ', '^', or '='
<i>c</i>	constant (double precision float)

OUTPUT:

Parameters

<i>spec</i>	modified spectrum (no modification in the case of errors)
-------------	---

ERRORS:

- IIP: Invalid input parameter(s)
- CPL: Division by zero [warning]
- CPL: Access out of range [warning]

4.1.5.49 `cpl_error_code sm_spec_split (smspec * spec, smspec * errfunc, const int errflag)`

Writes error function of an [smspec](#) structure to another spectrum. The error vector of the former spectrum is filled with zero. If two error columns are used, lower or upper error has to be specified and the function has to be called for either case.

INPUT:

Parameters

<i>spec</i>	spectrum with errors (3 columns)
-------------	----------------------------------

OUTPUT:

Parameters

<i>spec</i>	spectrum without errors (2 columns)
-------------	-------------------------------------

<i>errfunc</i>	error function (2 columns)
<i>errflag</i>	error flag (0 = symmetric error, 1 = lower error, 2 = upper error)

ERRORS:

- IIP: Invalid input parameter(s)

4.1.5.50 `cpl_error_code sm_spec_write (const smspec * spec, const char * filename)`

Writes `smspec` structure to ASCII file or on stdout (indicated by "stdout" as filename)

INPUT:**Parameters**

<i>spec</i>	input <code>smspec</code> structure
<i>filename</i>	name of output file or "stdout"

ERRORS:

- FOF: File opening failed
- NDA: No data
- IOV: Invalid object value(s)

4.1.5.51 `cpl_error_code sm_spec_writecpl (cpl_table * cpltab, const smspec * spec)`

Writes `smspec` structure to CPL table

INPUT:**Parameters**

<i>spec</i>	input <code>smspec</code> structure
-------------	-------------------------------------

OUTPUT:**Parameters**

<i>cpltab</i>	output CPL table
---------------	------------------

ERRORS:

- NDA: No data
- IOV: Invalid object value(s)

4.1.5.52 `cpl_error_code sm_spec_writfits (const smspec * spec, const char * filename)`

Writes `smspec` structure to FITS file by means of CPL

INPUT:

Parameters

<i>spec</i>	input <code>smspec</code> structure
<i>filename</i>	name of output file

ERRORS:

- none

4.2 Sky Model Module 2

Data Structures

- struct `_smparmodel_`

Files

- file `sm_scatmoonlight.c`
- file `sm_scatmoonlight.h`
- file `sm_skyemcomp.c`
- file `sm_skyemcomp.h`
- file `sm_skyemcomp.h`

Defines

- `#define SM_SKYEMCOMP_H`
- `#define SM_FILENAMESLIST "sm_filenames.dat"`
- `#define SM_RRSTEP 1000`
- `#define SM_LAM_UNIT 1e-6`
- `#define SM_RADMINLAM 1.3 * 1e-6 / SM_LAM_UNIT`
- `#define SM_ERAD 6371.`
- `#define SM_SIGMAX 4.`
- `#define SM_NSIGBIN 20`

Typedefs

- typedef struct `_smparmodel_ smparmodel`

Functions

- cpl_error_code [sm_etc_calcmmodel](#) (cpl_table *skytable, const cpl_parameterlist *params)
- cpl_error_code [sm_etc_getparams](#) (smparmodel *modelpar, const cpl_parameterlist *params)
- const cpl_parameter * [sm_etc_parameterlist_find_const](#) (const cpl_parameterlist *self, const char *name, cpl_type type)
- cpl_error_code [sm_etc_splitstring](#) (cpl_array *outval, char *instring)
- cpl_error_code [sm_etc_readfilenames](#) (smparmodel *modelpar, const cpl_parameterlist *params)
- cpl_error_code [sm_etc_writetable](#) (cpl_table *skytable, const smspec *radiance, const smspec *transmission)
- cpl_error_code [sm_comp_skyemcomp](#) (smspec *radiance, smspec *transmission, const smparmodel modelpar)
- cpl_error_code [sm_comp_extrapoltrans](#) (smspec *spec, const smparmodel modelpar, const double lim[2])
- double [sm_comp_altoairmass](#) (const double alt)
- cpl_error_code [sm_comp_getmolspec](#) (smspec *spec, const cpl_parameterlist *libfilepar, const double lim[2])
- cpl_error_code [sm_comp_readlibstruct](#) (cpl_parameterlist *libfilepar, const smparmodel modelpar, const char *spectype)
- double [sm_comp_vactoir_single](#) (const double lam)
- cpl_error_code [sm_comp_vactoir_spec](#) (smspec *spec)
- double [sm_comp_airtovac_single](#) (const double lam)
- cpl_error_code [sm_comp_airtovac_spec](#) (smspec *spec)
- double [sm_comp_calcrayleighscat1](#) (const double lam)
- cpl_error_code [sm_comp_calcrayleighscat](#) (smspec *rscattrans)
- double [sm_comp_calcmiescat1](#) (const double lam)
- cpl_error_code [sm_comp_calcmiescat](#) (smspec *mscattrans)
- cpl_error_code [sm_comp_scaletranscurv](#) (smspec *trans, const smparmodel modelpar)
- cpl_error_code [sm_comp_lunskybright](#) (smspec *spec, const smparmodel modelpar, const smspec *solspec, const smspec *molabstrans, const smspec *rscattrans, const smspec *mscattrans)
- cpl_error_code [sm_comp_lunskybright_v](#) (smspec *spec, const smparmodel modelpar, const smspec *solspec, const smspec *molabstrans, const smspec *rscattrans, const smspec *mscattrans)
- cpl_error_code [sm_comp_getmoonbedo](#) (smspec *albedo, const smparmodel modelpar)
- cpl_error_code [sm_comp_scatsstarlight](#) (smspec *spec, const smparmodel modelpar, const smspec *molabstrans)
- cpl_error_code [sm_comp_zodskybright](#) (smspec *spec, const smparmodel modelpar, const smspec *solspec, const smspec *molabstrans, const smspec *rscattrans, const smspec *mscattrans)
- cpl_error_code [sm_comp_tele](#) (smspec *spec, const smparmodel modelpar)
- cpl_error_code [sm_comp_extrapolrad](#) (smspec *spec, const smspec *trans, const smparmodel modelpar, const double lim[2])

- cpl_error_code `sm_comp_getlinespec` (`smspec *spec`, const `smparmodel modelpar`, const `smspec *rscattrans`, const `smspec *mscattrans`)
- cpl_error_code `sm_comp_readvarpar` (`cpl_table *varpar`, const `smparmodel modelpar`)
- cpl_error_code `sm_comp_scalelinetab` (`cpl_table *linedat`, const `smspec *linetab`, const `cpl_table *varpar`)
- cpl_error_code `sm_comp_convertlinetab` (`smspec *spec`, const `cpl_table *linedat`)
- cpl_error_code `sm_comp_convertlinetabo` (`smspec *spec`, const `cpl_table *linedat`)
- cpl_error_code `sm_comp_airglowcont` (`smspec *spec`, const `smparmodel modelpar`, const `smspec *molabstrans`, const `smspec *rscattrans`, const `smspec *mscattrans`)

4.2.1 Detailed Description

This module provides functions for the computation of the sky model for the ESO ETC.

4.2.2 Define Documentation

4.2.2.1 `#define SM_ERAD 6371.`

average earth radius in km

4.2.2.2 `#define SM_FILENAMELIST "sm_filenames.dat"`

List of data paths and file names

4.2.2.3 `#define SM_LAM_UNIT 1e-6`

μm

4.2.2.4 `#define SM_NSIGBIN 20`

Minimum number of bins for σ width of airglow lines

4.2.2.5 `#define SM_RADMINLAM 1.3 * 1e-6 / SM_LAM_UNIT`

Minimum wavelength for extraction of radiance \rightarrow faster code

4.2.2.6 `#define SM_RRSTEP 1000`

Line jumps for wavelength searches in LBLRTM/RFM library spectra

4.2.2.7 #define SM_SIGMAX 4.

Extension of Gaussian for airglow lines in σ

4.2.2.8 #define SM_SKYEMCOMP_H

4.2.3 Typedef Documentation

4.2.3.1 typedef struct _smparmodel_ smparmodel

Structure for sky emission parameters

Parameters

<i>alt</i>	altitude of object above horizon [0,90]
<i>airmass</i>	line-of-sight airmass
<i>alpha</i>	separation of sun and moon as seen from earth [0,180]
<i>rho</i>	separation of moon and object [0,180]
<i>altmoon</i>	altitude of moon above horizon [-90,90]
<i>moondist</i>	distance to moon (mean distance = 1; [0.945,1.055])
<i>lon_ecl</i>	heliocentric ecliptic longitude of object [-180,180]
<i>lat_ecl</i>	ecliptic latitude of object [-90,90]
<i>ncomp</i>	number of emissivity/temperature pairs (\leq SM_MAXPAR)
<i>eps[]</i>	emissivity (factor for conversion black body \rightarrow grey body)
<i>temp[]</i>	temperature in K
<i>msolflux</i>	monthly-averaged solar radio flux [s.f.u.]
<i>season</i>	period of the year (1: Dec/Jan, ..., 6: Oct/Nov.; 0: entire year)
<i>time</i>	period of the night (x/3 of night, x = 1,2,3; 0: entire night)
<i>vac_air</i>	vac[uum] or air wavelengths
<i>pwv</i>	precipitable water vapour in mm (-1: bimonthly mean)
<i>rtcode</i>	radiative transfer code L(BLRTM) or R(FM) for molecular spectra
<i>resol</i>	resolution of molecular spectra in library (crucial for run time)
<i>libpath</i>	directory path for library of molecular spectra
<i>libstruct</i>	name of file containing the structure of the selected LBLRTM/RFM library
<i>libstruct1</i>	name of structure file for time-dependent library
<i>libstruct2</i>	name of structure file for PWV-dependent library
<i>datapath</i>	directory path for input data files
<i>solspec-</i> <i>name</i>	solar spectrum
<i>lunirname</i>	file for lunar irradiance model parameters
<i>miephase-</i> <i>name</i>	file for Mie scattering phase functions
<i>mscatname</i>	file for multiple scattering correction factors
<i>o3transname</i>	file for UV/optical ozone transmission
<i>starspec-</i> <i>name</i>	mean spectrum of scattered starlight
<i>zodtabname</i>	V-brightness of zodiacal light [$10^{-8} \text{ W m}^{-2} \mu\text{m}^{-1} \text{ sr}^{-1}$]

<i>linetabname</i>	sky line table
<i>vardatname</i>	file for airglow lines scaling parameters
<i>acontname</i>	file for airglow continuum (scaling parameters included)
<i>incl</i>	rules inclusion of sky emission components <ul style="list-style-type: none"> • format: "xxxxxxx" where x = "Y" (yes) or x = "N" (no) • position: <ul style="list-style-type: none"> – 1: scattered moonlight – 2: scattered starlight – 3: zodiacal light – 4: thermal emission by telescope/instrument – 5: molecular emission of lower atmosphere – 6: sky emission lines of upper atmosphere – 7: airglow continuum (residual continuum)

4.2.4 Function Documentation

4.2.4.1 `cpl_error_code sm_comp_airglowcont (smspec * spec, const smparmodel modelpar, const smspec * molabstrans, const smspec * rscattrans, const smspec * mscattrans)`

Reads airglow continuum spectrum (with uncertainties) and scales it depending on the emission layer width (related to airmass), the monthly-averaged solar flux in sfu, the season of the year, and the time of the day. Moreover, the change of continuum emission by molecular absorption, Rayleigh scattering, and aerosol extinction is considered by provided transmission curves. For the scattering curves the effective optical depth is reduced by airmass-dependent recipes which were obtained by 3D atmospheric scattering calculations for extended airglow emission.

INPUT:

Parameters

<i>spec</i>	desired wavelength grid
<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)
<i>molabstrans</i>	transmission curve from radiative transfer code (scaled to airmass = 1)
<i>rscattrans</i>	transmission curve for Rayleigh scattering
<i>mscattrans</i>	transmission curve for Mie scattering

OUTPUT:

Parameters

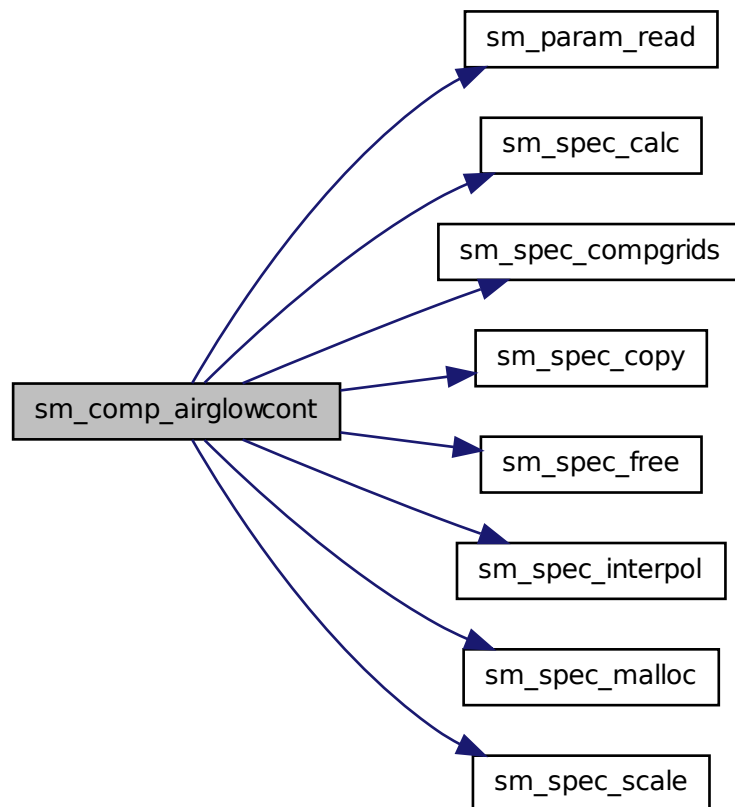
<i>spec</i>	airglow continuum
-------------	-------------------

ERRORS:

- NDA: No data

- IDG: Inconsistent data grids
- FOF: File opening failed
- UFS: Unexpected file structure
- IIP: Invalid input parameter(s)
- ISM: Insufficient memory

Here is the call graph for this function:



4.2.4.2 double sm_comp_airtovac_single (const double *lam*)

Converts air wavelength [μm] to vacuum wavelength by using the formula of Edlen (1966)

INPUT:**Parameters**

<i>lam</i>	air wavelength in μm
------------	---------------------------------

RETURN:

- vacuum wavelength in μm

ERRORS:

- none

4.2.4.3 cpl_error_code sm_comp_airtovac_spec (smspec * spec)

Converts spectrum of air wavelengths to spectrum of vacuum wavelengths by using the formula of Edlen (1966)

INPUT:**Parameters**

<i>spec</i>	spectrum with air wavelengths
-------------	-------------------------------

OUTPUT:**Parameters**

<i>spec</i>	spectrum with vacuum wavelengths
-------------	----------------------------------

ERRORS:

- none

4.2.4.4 double sm_comp_altoairmass (const double alt)

Calculates airmass from altitude angle in deg using the formula of Rozenberg (1966). The maximum airmass is 40.

INPUT:**Parameters**

<i>alt</i>	altitude angle in deg
------------	-----------------------

RETURN:

- airmass

ERRORS:

- none

4.2.4.5 `cpl_error_code sm_comp_calcmiescat (smspec * mscattrans)`

Calculates transmission curve for Mie (aerosol) extinction by using the parametrisation given by Patat (2011) and a cut in the optical depth for wavelengths below $0.4\mu\text{m}$

INPUT:**Parameters**

<i>mscattrans</i>	desired wavelength grid in μm
-------------------	--

OUTPUT:**Parameters**

<i>mscattrans</i>	Mie extinction curve (transmission at zenith)
-------------------	---

ERRORS:

- none

4.2.4.6 `double sm_comp_calcmiescat1 (const double lam)`

Calculates optical depth my Mie (aerosol) extinction for given wavelength. The parametrisation of Patat (2011) is used and a cut in the optical depth for wavelengths below $0.4\mu\text{m}$ is applied.

INPUT:**Parameters**

<i>lam</i>	wavelength in μm
------------	-----------------------------

RETURN:

- optical depth at zenith for given wavelength

ERRORS:

- none

4.2.4.7 `cpl_error_code sm_comp_calcrayleighscat (smspec * rscattrans)`

Calculates Rayleigh scattering transmission curve by using the parametrisation given by Liou (2002, P. 352) and the mean pressure profile of Paranal

INPUT:**Parameters**

<i>rscattrans</i>	desired wavelength grid in μm
-------------------	--

OUTPUT:**Parameters**

<i>rscattrans</i>	Rayleigh scattering curve (transmission at zenith)
-------------------	--

ERRORS:

- none

4.2.4.8 double sm_comp_calcrayleighscat1 (const double *lam*)

Calculates optical depth by Rayleigh scattering for given wavelength. The parametrisation of Liou (2002, P. 352) and the mean pressure profile of Paranal are used.

INPUT:**Parameters**

<i>lam</i>	wavelength in μm
------------	-----------------------------

RETURN:

- optical depth at zenith for given wavelength

ERRORS:

- none

4.2.4.9 cpl_error_code sm_comp_convertlinetab (smspec * *spec*, const cpl_table * *linedat*)

Converts table of emission lines with central wavelength, line flux, flux uncertainty, and line width into line spectrum with given wavelength grid. Gaussian line shape is assumed.

INPUT:**Parameters**

<i>spec</i>	desired wavelength grid as smspec structure
<i>linedat</i>	CPL table with line data

OUTPUT:

Parameters

<i>spec</i>	sky line spectrum
-------------	-------------------

ERRORS:

- NDA: No data
- ISD: Insufficient data points

4.2.4.10 `cpl_error_code sm_comp_convertlinetabo (smspec * spec, const cpl_table * linedat)`

Converts table of emission lines with central wavelength, line flux, flux uncertainty, and line width into line spectrum with given wavelength grid.

INPUT:**Parameters**

<i>spec</i>	desired wavelength grid as smspec structure
<i>linedat</i>	CPL table with line data

OUTPUT:**Parameters**

<i>spec</i>	sky line spectrum
-------------	-------------------

ERRORS:

- NDA: No data
- ISD: Insufficient data points

4.2.4.11 `cpl_error_code sm_comp_extrapolrad (smspec * spec, const smspec * trans, const smparmodel modelpar, const double lim[2])`

Extracts and rebins library LBLRTM/RFM molecular radiance spectrum and extrapolates it to the requested airmass, which is computed by means of the formula of Rozenberg (1966). For the optical depth dependent extrapolation, the corresponding transmission curve as provided by [sm_comp_extrapoltrans](#) is used.

INPUT:**Parameters**

<i>spec</i>	desired wavelength grid
<i>trans</i>	transmission curve belonging to radiance spectrum (scaled to airmass = 1)
<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)
<i>lim</i>	minimum and maximum wavelength extraction limits

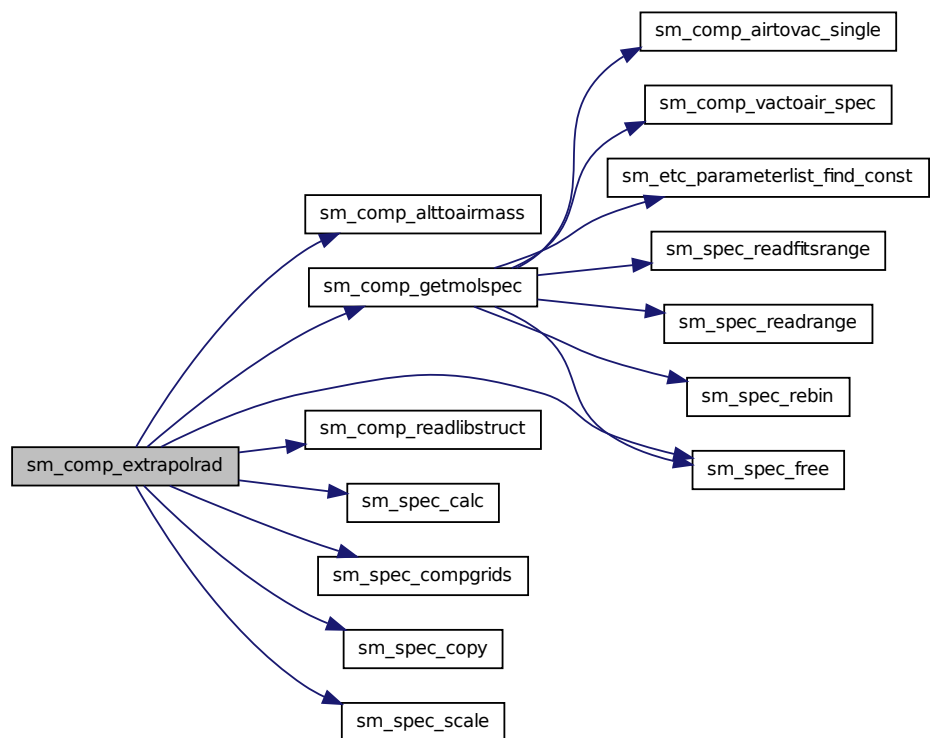
OUTPUT:**Parameters**

<i>spec</i>	extracted, rebinned, and scaled LBLRTM/RFM radiance spectrum
-------------	--

ERRORS:

- NDA: No data
- IOD: Invalid order of data points
- IIP: Invalid input parameter(s)
- IDG: Inconsistent data grids

Here is the call graph for this function:



4.2.4.12 `cpl_error_code sm_comp_extrapoltrans (smspec * spec, const smparmodel modelpar, const double lim[2])`

Extracts and rebins library LBLRTM/RFM molecular transmission spectrum and extrapolates it to airmass = 1. For the airmass calculations, the formula of Rozenberg (1966) is used.

INPUT:

Parameters

<i>spec</i>	desired wavelength grid
<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)
<i>lim</i>	minimum and maximum wavelength extraction limits

OUTPUT:

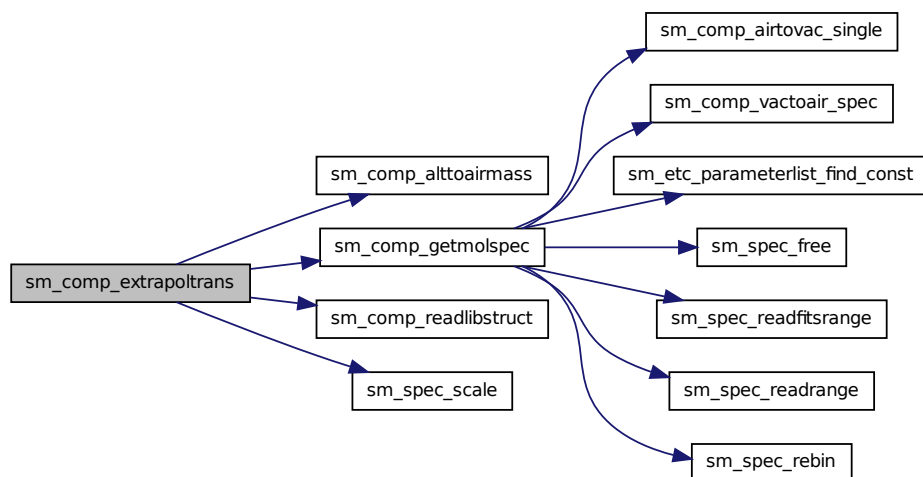
Parameters

<i>spec</i>	extracted, rebinned, and scaled LBLRTM/RFM transmission curve
-------------	---

ERRORS:

- NDA: No data
- IOD: Invalid order of data points
- IIP: Invalid input parameter(s)

Here is the call graph for this function:



4.2.4.13 `cpl_error_code sm_comp_getlinespec (smspec * spec, const smparmodel
modelpar, const smspec * rscattrans, const smspec * mscattrans)`

Converts table of airglow emission lines (see Hanuschik 2003 and Rousselot et al. 2000) to spectrum ([smspec](#) structure) by applying wavelength-dependent correction factors which depend on object altitude, emission layer height, solar radio flux, period of the year, and period of the night (see [sm_comp_readvarpar](#)). The wavelengths are converted from vacuum to air if required. Moreover, the change of line intensity by molecular absorption, Rayleigh scattering, and aerosol extinction is considered. The former correction is realised by a precalculated library of lists of line-specific transmission correction factors for each radiative transfer code transmission spectrum. The latter corrections are performed by provided extinction curves for which the effective optical depth is reduced by airmass-dependent recipes which were obtained by 3D atmospheric scattering calculations for extended airglow emission.

INPUT:

Parameters

<i>spec</i>	desired wavelength grid
<i>modelpar</i>	airglow parameters (see typedef of smparmodel)
<i>rscattrans</i>	transmission curve for Rayleigh scattering
<i>mscattrans</i>	transmission curve for Mie scattering

OUTPUT:

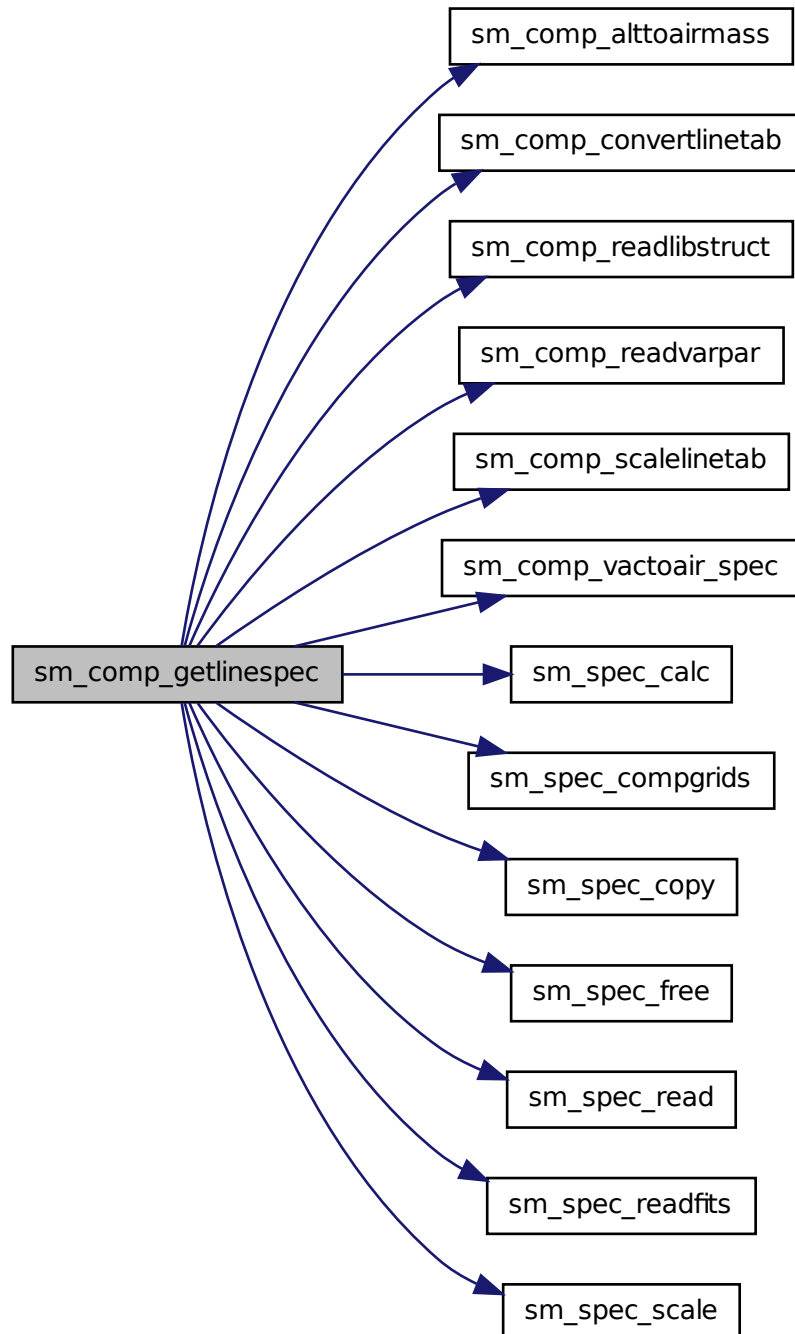
Parameters

<i>spec</i>	airglow emission line spectrum
-------------	--------------------------------

ERRORS:

- IIP: Invalid input parameter(s)
- IDG: Inconsistent data grids
- see subroutines as well

Here is the call graph for this function:



4.2.4.14 `cpl_error_code sm_comp_getmolspec (smspec * spec, const cpl_parameterlist * libfilepar, const double lim[2])`

Extracts and rebins library LBLRTM/RFM molecular radiance or transmission spectra. The file name and the decision between vacuum and air wavelengths have to be provided by a CPL parameter list. Each library file has to consist of the four columns wavelength, radiance/transmission for average atmospheric profiles, and radiance/transmission for minus and plus sigma atmospheric profiles. This structure is converted to the `smspec` structure, i.e. "lam", "flux", "dflux1" (negative error), and "dflux2" (positive error). The extraction-related wavelength range can be delimited regardless of the desired grid in order to reduce computing time by avoiding the extraction of wavelengths that do not contribute significantly to the composite sky spectrum.

INPUT:

Parameters

<code>spec</code>	desired wavelength grid
<code>libfilepar</code>	CPL parameter list containing file name (+ path) and vacuum/air parameter
<code>lim</code>	minimum and maximum wavelength extraction limits

OUTPUT:

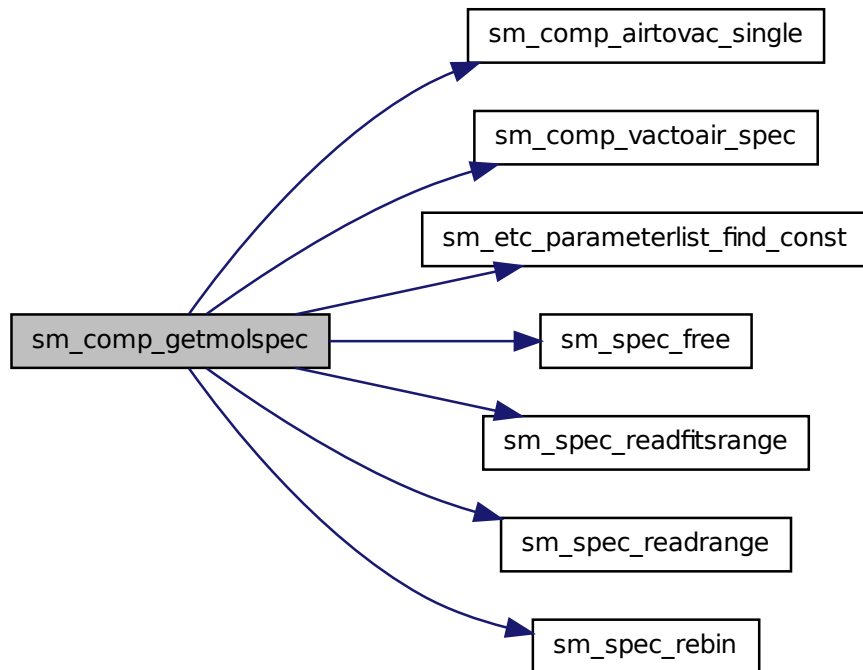
Parameters

<code>spec</code>	extracted and rebinned LBLRTM/RFM spectrum (radiance or transmission)
-------------------	---

ERRORS:

- NDA: No data
- IOD: Invalid order of data points
- IIP: Invalid input parameter(s)
- IOS: Invalid object structure
- FOF: File opening failed
- IFE: Invalid file name extension
- CPL: Access out of range [warning]

Here is the call graph for this function:



4.2.4.15 `cpl_error_code sm_comp_getmoonalbedo (smspec * albedo, const smparmodel modelpar)`

Calculates the wavelength-dependent disc-equivalent lunar albedo from the Moon illuminance model of Kieffer & Stone (2005) based on ROLO data. Except for the libration-dependent terms, the full model is implemented. The required model coefficients are read from a file. As user parameter, the angle between Sun and Moon as seen from Earth is provided by a [smparmodel](#) structure. For the calculation of the Moon albedo, it is converted into absolute Moon phase angle and selenographic longitude of the Sun. The latter is required for considering differences in the brightness of the waxing and the waning Moon.

INPUT:

Parameters

<i>albedo</i>	desired wavelength grid
<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)

OUTPUT:**Parameters**

<i>albedo</i>	spectrum of lunar albedo
---------------	--------------------------

ERRORS:

- IIP: Invalid input parameters
- FOF: File opening failed

4.2.4.16 `cpl_error_code sm_comp_lunskybright (smspec * spec, const smparmodel modelpar, const smspec * solspec, const smspec * molabstrans, const smspec * rscattrans, const smspec * mscattrans)`

Evaluates predicted lunar part of sky brightness following Krisciunas & Schaefer (1991, PASP 103, 1033).

Original implementation by J. Thorstensen. Modified by F. Patat to generalise to UBVRI filters. Spectroscopic version by S. Noll.

This routine uses the solar spectrum as a proxy for the unattenuated lunar spectrum and two different transmission curves: one with and one without molecular absorption. The second one is important for the estimate of the amount of scattered moonlight at the object position and includes Rayleigh scattering and aerosol extinction. The model depends on object and moon altitude, their angular separation, moon phase (expressed by angular distance of moon and sun), and moon distance.

Note

- The model becomes unreliable for moon-object distances around and below 30°. There, the main contribution comes from aerosol scattering, which is highly variable and, therefore, difficult to predict.
- The routine uses the library transmission curve selected depending on the object altitude and not the moon altitude (see [sm_comp_getmolspec](#)). This reduces the computing time significantly, but especially the flux in centres of telluric absorption features can be overestimated if the moon is much lower than the object. However, for the composite sky emission spectrum the flux differences should be lower than 10% even in the A-band.

INPUT:**Parameters**

<i>spec</i>	desired wavelength grid
<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)
<i>solspec</i>	solar spectrum
<i>molabstrans</i>	transmission curve from radiative transfer code
<i>rscattrans</i>	transmission curve for Rayleigh scattering
<i>mscattrans</i>	transmission curve for Mie scattering

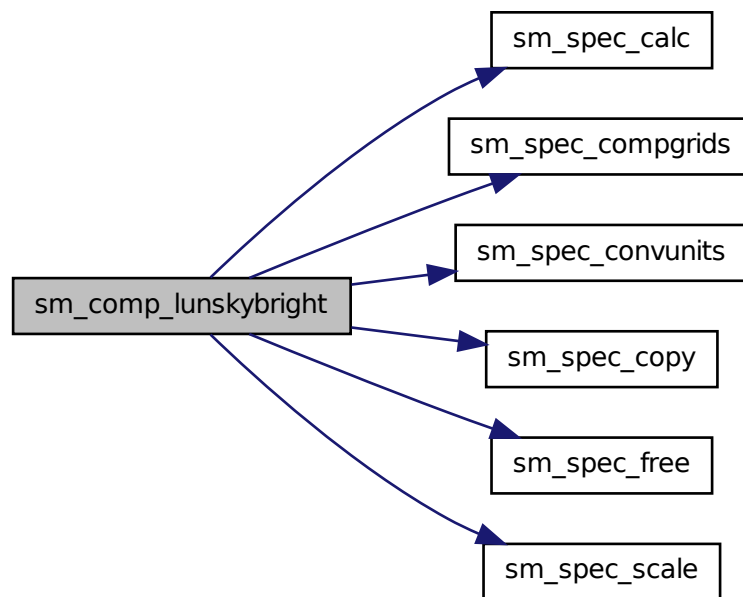
OUTPUT:**Parameters**

<i>spec</i>	spectrum of lunar sky brightness
-------------	----------------------------------

ERRORS:

- NDA: No data
- IIP: Invalid input parameter(s)
- IDG: Inconsistent data grids
- see subroutines as well

Here is the call graph for this function:



4.2.4.17 `cpl_error_code sm_comp_lunskybright_v (smspec * spec, const smparmodel
modelpar, const smspec * solspec, const smspec * molabstrans, const smspec
* rscattrans, const smspec * mscattrans)`

Calculates scattered moonlight. The Moon illuminance is approximated from Kieffer & Stone (2005) based on ROLO data. The scattering in the atmosphere is obtained from

3D single scattering calculations with multiple scattering correction (see [sm_scatter_moon](#) and subroutines) and the Cerro Paranal extinction curve of Patat et al. (2011). A simple estimate which depends on the optical depth dependent effective airmass is taken for the molecular absorption of the moonlight. The absorption by the stratospheric ozone layer is considered separately. The model depends on object and Moon altitude, their angular separation, Moon phase (expressed by angular distance of Moon and Sun), and Moon distance.

INPUT:

Parameters

<i>spec</i>	desired wavelength grid
<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)
<i>solspec</i>	solar spectrum
<i>molabstrans</i>	transmission curve from radiative transfer code
<i>rscattrans</i>	transmission curve for Rayleigh scattering
<i>mscattrans</i>	transmission curve for Mie scattering

OUTPUT:

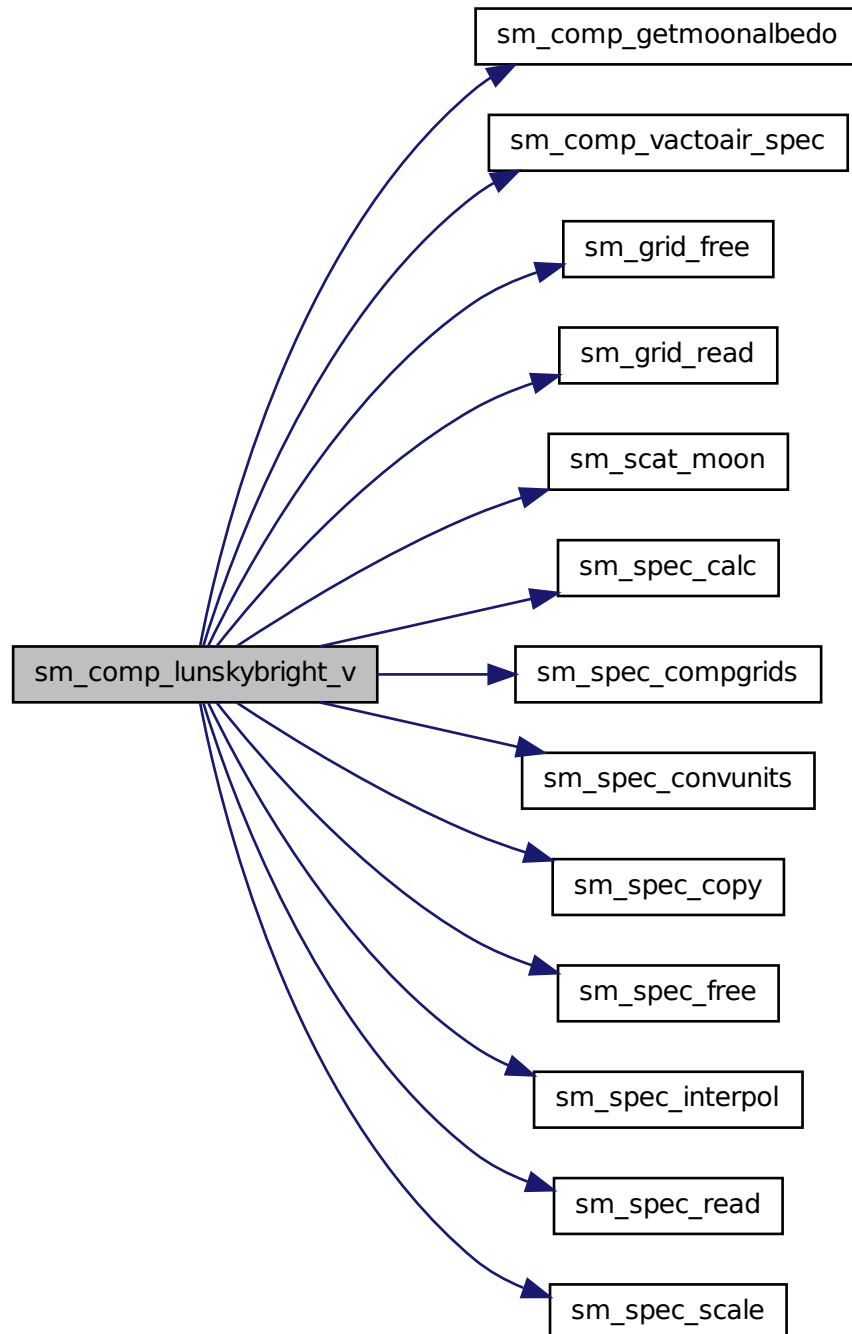
Parameters

<i>spec</i>	spectrum of lunar sky brightness
-------------	----------------------------------

ERRORS:

- NDA: No data
- IIP: Invalid input parameter(s)
- IDG: Inconsistent data grids
- see subroutines as well

Here is the call graph for this function:



4.2.4.18 `cpl_error_code sm_comp_readlibstruct (cpl_parameterlist * libfilepar, const smparmodel modelpar, const char * spectype)`

Builds file name of library spectrum or line list depending on input parameters such as library path, radiative transfer code, target altitude or airmass, period of the year, period of the night, precipitable water vapour, resolution, or file type. The required parameters are provided by a [smparmodel](#) structure except for the file type which has to be given directly and can be either "R" for radiance spectrum, "T" for transmission spectrum, or "L" for list of line transmissions. The file name format is given by a library structure file. The routine returns a CPL parameter list with the file name and the values of all parameters building the file name plus a parameter for vacuum/air wavelengths. The latter is added to indicate whether the library spectrum (always in vacuum wavelengths) has to be converted to air wavelengths.

INPUT:

Parameters

<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)
<i>spectype</i>	radiance ("R"), transmission ("T"), or line transmission ("L")?

OUTPUT:

Parameters

<i>libfilepar</i>	CPL parameter list containing file name (+ path) of library spectrum and its constituents
-------------------	---

ERRORS:

- FOF: File opening failed
- UFS: Unexpected file structure
- IIP: Invalid input parameter(s)

4.2.4.19 `cpl_error_code sm_comp_readvarpar (cpl_table * varpar, const smparmodel modelpar)`

Reads data related to airglow scaling from file and fills a CPL table with feature-related correction factors and uncertainties. The resulting factors are related to the standard strengths of airglow emission features of the upper atmosphere given by Hanuschik (2003) and Rousselot et al. (2000). The scaling factors depend on the emission layer width (related to airmass), the monthly-averaged solar flux in sfu, the season of the year, and the time of the day. In addition, the relative Doppler widths σ_D of the different features are provided by the output table. They are derived from the read mol masses and temperatures of the emission layers.

INPUT:

Parameters

<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)
-----------------	--

OUTPUT:**Parameters**

<i>varpar</i>	airglow variability data as CPL table with columns <ul style="list-style-type: none"> • <i>N</i>: feature number • <i>fac</i>: scaling factor for emission lines depending on feature • <i>dfac</i>: uncertainties of factors • <i>relwidth</i>: relative Doppler width
---------------	---

ERRORS:

- FOF: File opening failed
- UFS: Unexpected file structure
- IIP: Invalid input parameter(s)

4.2.4.20 `cpl_error_code sm_comp_scalelinetab (cpl_table * linedat, const smspec * linetab, const cpl_table * varpar)`

Writes line data in CPL table and corrects line fluxes by factors provided by *varpar* for the different features.

INPUT:**Parameters**

<i>linetab</i>	table of emission lines (provided as smspec structure; dflux1 contains feature number)
<i>varpar</i>	line parameters (scaling factors and Doppler widths) provided as CPL table

OUTPUT:**Parameters**

<i>linedat</i>	CPL table with columns "lam", "flux", "dflux", "feat", and "width"
----------------	--

ERRORS:

- NDA: No data
- ISD: Insufficient data points
- IOV: Invalid object value(s)

4.2.4.21 `cpl_error_code sm_comp_scaletranscurv (smspec * trans, const smparmodel modelpar)`

Scales transmission curve depending on airmass as calculated by the formula of Rozenberg (1966)

INPUT:

Parameters

<i>trans</i>	transmission curve at airmass = 1
<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)

OUTPUT:

Parameters

<i>trans</i>	transmission curve for given altitude
--------------	---------------------------------------

ERRORS:

- NDA: No data
- IIP: Invalid input parameter(s)

4.2.4.22 `cpl_error_code sm_comp_scatterstarlight (smspec * spec, const smparmodel modelpar, const smspec * molabstrans)`

Reads representative spectrum of scattered starlight, rebins it to given wavelength grid by means of linear interpolation, and considers telluric absorption for an effective air-mass of 1.25. The spectrum was calculated by means of integrated starlight photometric data based on Pioneer 10/11 observations (see Toller et al. 1987; Leinert et al. 1998; Melchior et al. 2007), spectroscopic data by Mattila (1980), and 3D atmospheric scattering calculations (see Wolstencroft & van Breda 1967; Staude 1975; Bernstein et al. 2002).

INPUT:

Parameters

<i>spec</i>	desired wavelength grid
<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)
<i>molabstrans</i>	transmission curve from radiative transfer code (scaled to airmass = 1)

OUTPUT:

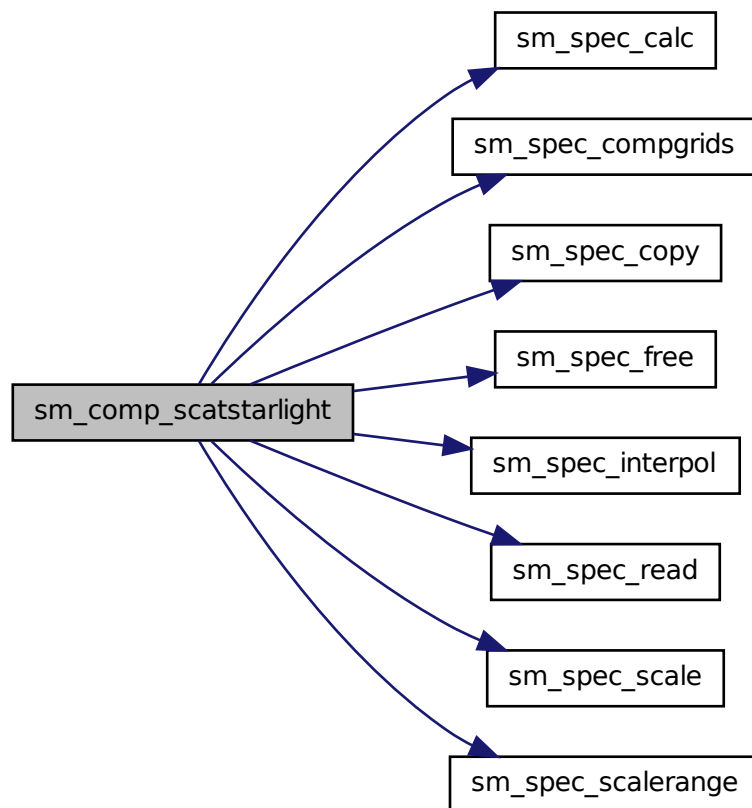
Parameters

<i>spec</i>	spectrum of scattered starlight
-------------	---------------------------------

ERRORS:

- IDG: Inconsistent data grids
- see [sm_spec_rebin](#)

Here is the call graph for this function:



4.2.4.23 `cpl_error_code sm_comp_skyemcomp (smspec * radiance, smspec * transmission, const smparmodel modelpar)`

Creates sky emission spectrum and transmission curve at object position for given wavelength grid and input parameters. Imports solar spectrum, LBLRTM or RFM radiance and transmission spectra (ASCII or FITS files), and airglow emission line and continuum table, estimates lunar and zodiacal sky brightness, considers scattered starlight, and calculates thermal telescope/instrument emission. Using the "incl" parameter of the

[smparmodel](#) structure, the selection of the seven different components for the computation of the output radiance spectrum can be modified.

INPUT:

Parameters

<i>radiance</i>	wavelength grid of sky emission spectrum
<i>transmission</i>	wavelength grid of transmission curve
<i>modelpar</i>	input parameters related to sky emission (see typedef of smparmodel)

OUTPUT:

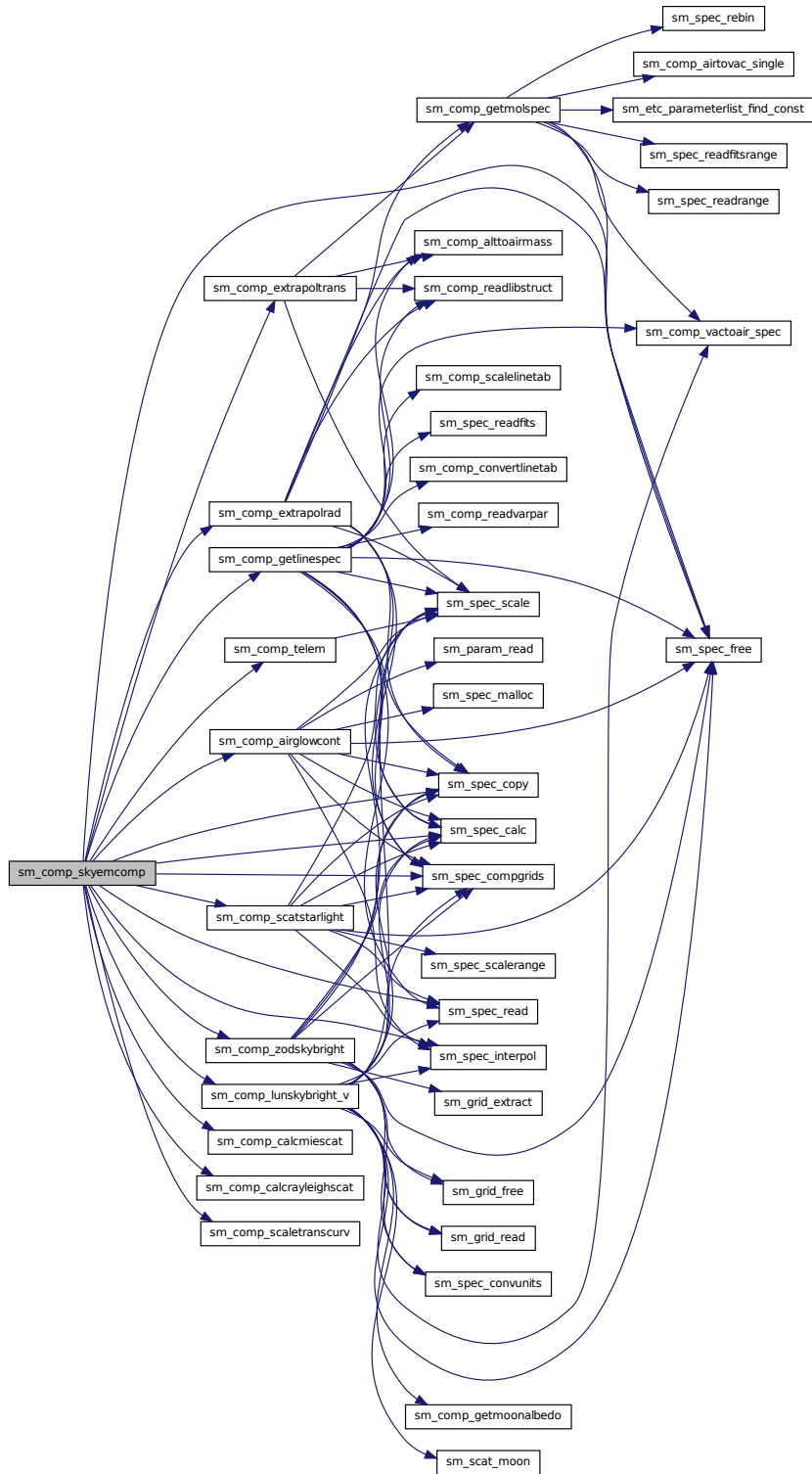
Parameters

<i>radiance</i>	sky emission spectrum (smspec structure, errors considered), units: μm , $\text{phot s}^{-1}\text{m}^{-2}\mu\text{m}^{-1}\text{arcsec}^{-2}$
<i>transmission</i>	transmission curve (smspec structure, errors considered), units: μm , [0,1]

ERRORS:

- NDA: No data
- IDG: Inconsistent data grids
- see subroutines as well

Here is the call graph for this function:



4.2.4.24 `cpl_error_code sm_comp_telem (smspec * spec, const smparmodel modelpar)`

Computes thermal emission by telescope/instrument. Several components are possible. The routine assumes grey bodies depending on emissivity and temperature for each component. The order of the components matters. The first emissivity and temperature have to be for the main mirror. Any emission or absorption which is not related to the listed optical components is not considered.

The routine assumes that the absorption inside the instrument is handled by a response curve, which is applied somewhere else. For this reason, the emission of the different components is raised depending on the position along the light path. For example, the emission of the main mirror is increased by the reciprocal of the mirror transmission. This factor makes sure that the telescope emission can be added to the other sky model components, for which the flux is reduced by absorption of the main mirror and all other optical components. The routine simulates the apparent telescope/instrument emission that is not removed in the course of typical astronomical flux calibration.

INPUT:

Parameters

<i>spec</i>	desired wavelength grid
<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)

OUTPUT:

Parameters

<i>spec</i>	apparent telescope/instrument emission
-------------	--

ERRORS:

- NDA: No data
- IIP: Invalid input parameter(s)
- IDR: Invalid data range

Here is the call graph for this function:



4.2.4.25 `double sm_comp_vactoir_single (const double lam)`

Converts vacuum wavelength [μm] to air wavelength by using the formula of Edlen (1966)

INPUT:**Parameters**

<i>lam</i>	vacuum wavelength in μm
------------	------------------------------------

RETURN:

- air wavelength in μm

ERRORS:

- none

4.2.4.26 `cpl_error_code sm_comp_vactoir_spec (smspec * spec)`

Converts spectrum of vacuum wavelengths to spectrum of air wavelengths by using the formula of Edlen (1966)

INPUT:**Parameters**

<i>spec</i>	spectrum with vacuum wavelengths
-------------	----------------------------------

OUTPUT:**Parameters**

<i>spec</i>	spectrum with air wavelengths
-------------	-------------------------------

ERRORS:

- none

4.2.4.27 `cpl_error_code sm_comp_zodskybright (smspec * spec, const smparmodel modelpar, const smspec * solspec, const smspec * molabstrans, const smspec * rscattrans, const smspec * mscattrans)`

Derives sky brightness caused by the zodiacal light. The procedure is based on recipes provided in Leinert et al. (1998) and uses a table containing the brightness of the zodiacal light at $0.5 \mu\text{m}$ [$\text{phot s}^{-1} \text{m}^{-2} \mu\text{m}^{-1} \text{arcsec}^{-2}$] dependent on the ecliptic coordinates of the object. For the zodiacal light spectrum a slightly reddened solar spectrum is taken.

The extinction of zodiacal light by molecular absorption, Rayleigh scattering, and aerosol extinction is considered by input transmission curves provided for airmass = 1. For the scattering curves the effective optical depth is reduced by recipes depending on the top-of-atmosphere line-of-sight zodiacal light intensity. The parametrisations used were obtained by 3D atmospheric scattering calculations for extended emission (see, e.g., Bernstein et al. 2002).

INPUT:

Parameters

<i>spec</i>	desired wavelength grid
<i>modelpar</i>	sky emission parameters (see typedef of smparmodel)
<i>solspec</i>	solar spectrum (note: normalisation for Colina et al. 1996)
<i>molabstrans</i>	transmission curve from radiative transfer code (scaled to airmass = 1)
<i>rscattrans</i>	transmission curve for Rayleigh scattering
<i>mscattrans</i>	transmission curve for Mie scattering

OUTPUT:

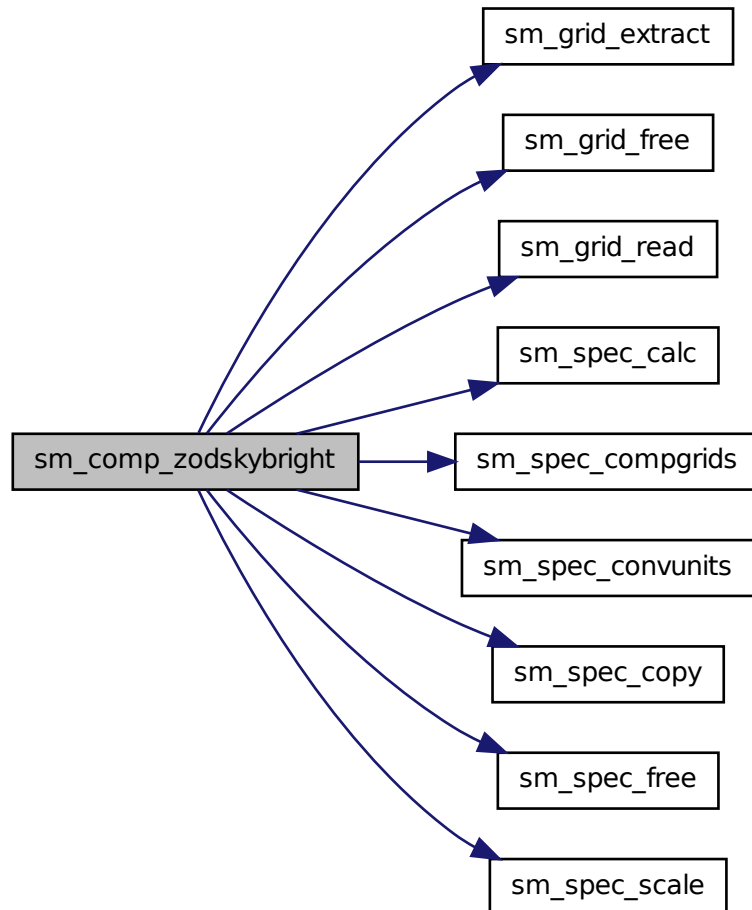
Parameters

<i>spec</i>	spectrum of zodiacal light
-------------	----------------------------

ERRORS:

- NDA: No data
- IIP: Invalid input parameter(s)
- IDG: Inconsistent data grids
- see subroutines as well

Here is the call graph for this function:



4.2.4.28 `cpl_error_code sm_etc_calcmode (cpl_table * skytable, const cpl_parameterlist * params)`

Reads wavelength grid as CPL table and sky model parameters as CPL parameter list and uses this information to compute a corresponding radiance spectrum and transmission curve which are committed to the ETC as a CPL table including the input wavelength grid and the derived flux and flux error data. The output columns are "lam", "flux", "dflux1", "dflux2", "trans", "dtrans1", and "dtrans2".

INPUT:**Parameters**

<i>skytable</i>	input wavelength grid as CPL table
<i>params</i>	CPL sky model parameter list

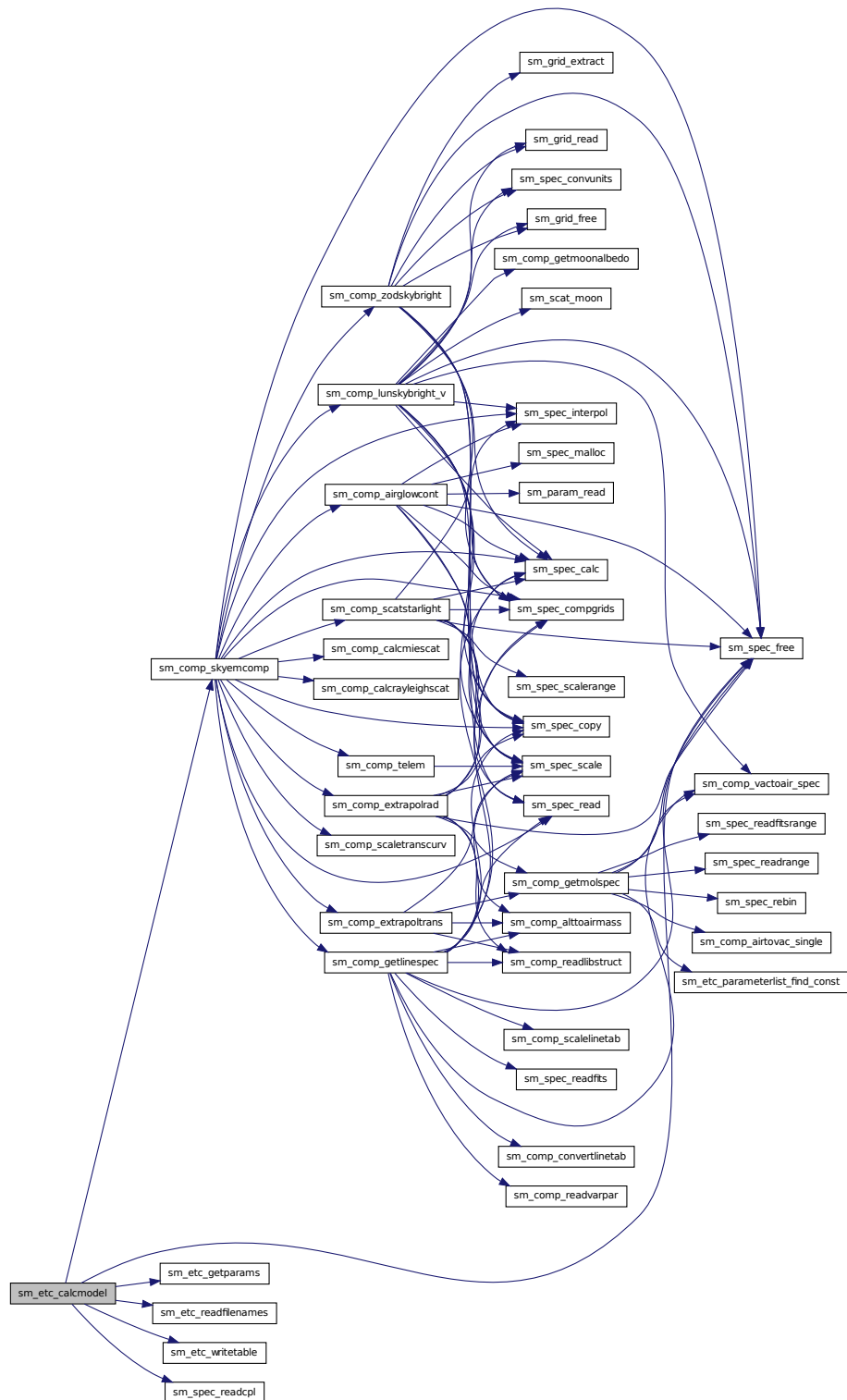
OUTPUT:**Parameters**

<i>skytable</i>	CPL table with full radiance and transmission data
-----------------	--

ERRORS:

- see subroutines

Here is the call graph for this function:



4.2.4.29 `cpl_error_code sm_etc_getparams (smparmodel * modelpar, const cpl_parameterlist * params)`

Reads sky model parameters from a CPL parameter list and put them in a [smparmodel](#) structure. The strings of comma-separated emissivity and temperature values are converted into double arrays. As only non-input parameter the airmass is added to the [smparmodel](#) structure. It is computed from altitude angle by using the formula of Rozenberg (1966).

INPUT:

Parameters

<i>params</i>	CPL sky model parameter list
---------------	------------------------------

OUTPUT:

Parameters

<i>modelpar</i>	smparmodel structure containing the read parameters
-----------------	---

ERRORS:

- see subroutines

4.2.4.30 `const cpl_parameter* sm_etc_parameterlist_find_const (const cpl_parameterlist * self, const char * name, cpl_type type)`

Error-handling wrapper around `cpl_parameterlist_find_const()`, catching errors of mismatching parameter name and/or type.

INPUT:

Parameters

<i>self</i>	CPL parameter list
<i>name</i>	parameter name
<i>type</i>	CPL type of parameter

RETURN:

- CPL parameter

ERRORS:

- CPL_ERROR_DATA_NOT_FOUND
- CPL_ERROR_TYPE_MISMATCH

4.2.4.31 `cpl_error_code sm_etc_readfilenames (smparmodel * modelpar, const cpl_parameterlist * params)`

Reads data paths and file names from a file (path from CPL parameter list and name from header file) and put them in a [smparmodel](#) structure.

INPUT:

Parameters

<i>params</i>	CPL sky model parameter list
---------------	------------------------------

OUTPUT:

Parameters

<i>modelpar</i>	smparmodel structure containing the read parameters
-----------------	---

ERRORS:

- FOF: File opening failed
- see [sm_param_readcheck](#) as well

4.2.4.32 `cpl_error_code sm_etc_splitstring (cpl_array * outval, char * instring)`

Splits a string of comma-separated values and converts the substrings into double values.

INPUT:

Parameters

<i>instring</i>	comma-separated list of values
-----------------	--------------------------------

OUTPUT:

Parameters

<i>outval</i>	CPL array of double values
---------------	----------------------------

ERRORS:

- ISM: Insufficient memory
- NDA: No data

4.2.4.33 `cpl_error_code sm_etc_writetable (cpl_table * skytable, const smspec * radiance, const smspec * transmission)`

Writes radiance spectrum and transmission curve which are provided as [smspec](#) structures to the CPL table of the ETC interface.

INPUT:**Parameters**

<i>skytable</i>	input wavelength grid as CPL table
<i>radiance</i>	radiance spectrum as smspec structure
<i>transmission</i>	transmission curve as smspec structure

OUTPUT:**Parameters**

<i>skytable</i>	CPL table with full radiance and transmission data
-----------------	--

ERRORS:

- NDA: No wavelength data
- IDG: Inconsistent data grids

Chapter 5

Data Structure Documentation

5.1 `_smdat_` Struct Reference

```
#include <sm_general.h>
```

Data Fields

- double `lam`
- double `flux`
- double `dflux1`
- double `dflux2`

5.1.1 Detailed Description

Structure for data points

Parameters

<i>lam</i>	wavelength
<i>flux</i>	flux
<i>dflux1</i>	(lower) flux error
<i>dflux2</i>	upper flux error

5.1.2 Field Documentation

5.1.2.1 `double _smdat_::dflux1`

5.1.2.2 `double _smdat_::dflux2`

5.1.2.3 `double _smdat_::flux`

5.1.2.4 double _smdat_::lam

The documentation for this struct was generated from the following file:

- [src/sm_general.h](#)

5.2 _smgrid_ Struct Reference

```
#include <sm_general.h>
```

Data Fields

- int [nx](#)
- int [ny](#)
- double * [xpos](#)
- double * [ypos](#)
- double ** [val](#)

5.2.1 Detailed Description

Structure for coordinate grids

Parameters

<i>nx</i>	number of x coordinate values
<i>ny</i>	number of y coordinate values
<i>*xpos</i>	vector of x coordinate values
<i>*ypos</i>	vector of y coordinate values
<i>**val</i>	matrix of data values for the coordinate grid

5.2.2 Field Documentation

5.2.2.1 int _smgrid_::nx

5.2.2.2 int _smgrid_::ny

5.2.2.3 double** _smgrid_::val

5.2.2.4 double* _smgrid_::xpos

5.2.2.5 double* _smgrid_::ypos

The documentation for this struct was generated from the following file:

- [src/sm_general.h](#)

5.3 `_smparam_` Struct Reference

```
#include <sm_general.h>
```

Data Fields

- char `c` [`SM_LENLINE+2`]
- int `i`
- double `d`
- int `n`

5.3.1 Detailed Description

Structure for a parameter value in different data types

Parameters

<code>c</code>	string (maximum length <code>SM_LENLINE</code>)
<code>i</code>	integer number
<code>d</code>	float number of double precision
<code>n</code>	number of value (for arrays; counted backwards)

5.3.2 Field Documentation

5.3.2.1 char `_smparam_::c`[`SM_LENLINE+2`]

5.3.2.2 double `_smparam_::d`

5.3.2.3 int `_smparam_::i`

5.3.2.4 int `_smparam_::n`

The documentation for this struct was generated from the following file:

- src/[sm_general.h](#)

5.4 `_smparmodel_` Struct Reference

```
#include <sm_skyemcomp.h>
```

Data Fields

- double `alt`
- double `airmass`

- double [alpha](#)
- double [rho](#)
- double [altmoon](#)
- double [moondist](#)
- double [lon_ecl](#)
- double [lat_ecl](#)
- int [ncomp](#)
- double [eps](#) [SM_MAXPAR]
- double [temp](#) [SM_MAXPAR]
- double [msolflux](#)
- int [season](#)
- int [time](#)
- char [vac_air](#) [SM_LENLINE+1]
- double [pwv](#)
- char [rtcode](#) [SM_LENLINE+1]
- double [resol](#)
- char [libpath](#) [SM_MAXLEN+1]
- char [libstruct](#) [SM_LENLINE+1]
- char [libstruct1](#) [SM_LENLINE+1]
- char [libstruct2](#) [SM_LENLINE+1]
- char [datapath](#) [SM_MAXLEN+1]
- char [solspecname](#) [SM_LENLINE+1]
- char [lunirrname](#) [SM_LENLINE+1]
- char [miephasename](#) [SM_LENLINE+1]
- char [mscatname](#) [SM_LENLINE+1]
- char [o3transname](#) [SM_LENLINE+1]
- char [starspecname](#) [SM_LENLINE+1]
- char [zodtabname](#) [SM_LENLINE+1]
- char [linetabname](#) [SM_LENLINE+1]
- char [vardatname](#) [SM_LENLINE+1]
- char [acontname](#) [SM_LENLINE+1]
- char [incl](#) [SM_LENLINE+1]

5.4.1 Detailed Description

Structure for sky emission parameters

Parameters

<i>alt</i>	altitude of object above horizon [0,90]
<i>airmass</i>	line-of-sight airmass
<i>alpha</i>	separation of sun and moon as seen from earth [0,180]
<i>rho</i>	separation of moon and object [0,180]
<i>altmoon</i>	altitude of moon above horizon [-90,90]
<i>moondist</i>	distance to moon (mean distance = 1; [0.945,1.055])
<i>lon_ecl</i>	heliocentric ecliptic longitude of object [-180,180]
<i>lat_ecl</i>	ecliptic latitude of object [-90,90]

<i>ncomp</i>	number of emissivity/temperature pairs (\leq SM_MAXPAR)
<i>eps[]</i>	emissivity (factor for conversion black body \rightarrow grey body)
<i>temp[]</i>	temperature in K
<i>msolflux</i>	monthly-averaged solar radio flux [s.f.u.]
<i>season</i>	period of the year (1: Dec/Jan, ..., 6: Oct/Nov.; 0: entire year)
<i>time</i>	period of the night (x/3 of night, x = 1,2,3; 0: entire night)
<i>vac_air</i>	vac[uum] or air wavelengths
<i>pwv</i>	precipitable water vapour in mm (-1: bimonthly mean)
<i>rtcode</i>	radiative transfer code L(BLRTM) or R(FM) for molecular spectra
<i>resol</i>	resolution of molecular spectra in library (crucial for run time)
<i>libpath</i>	directory path for library of molecular spectra
<i>libstruct</i>	name of file containing the structure of the selected LBLRTM/RFM library
<i>libstruct1</i>	name of structure file for time-dependent library
<i>libstruct2</i>	name of structure file for PWV-dependent library
<i>datapath</i>	directory path for input data files
<i>solspec-name</i>	solar spectrum
<i>lunirrname</i>	file for lunar irradiance model parameters
<i>miephase-name</i>	file for Mie scattering phase functions
<i>mscatname</i>	file for multiple scattering correction factors
<i>o3transname</i>	file for UV/optical ozone transmission
<i>starspec-name</i>	mean spectrum of scattered starlight
<i>zodtabname</i>	V-brightness of zodiacal light [$10^{-8} \text{ W m}^{-2} \mu\text{m}^{-1} \text{ sr}^{-1}$]
<i>linetabname</i>	sky line table
<i>vardatname</i>	file for airglow lines scaling parameters
<i>acontname</i>	file for airglow continuum (scaling parameters included)
<i>incl</i>	rules inclusion of sky emission components <ul style="list-style-type: none"> • format: "xxxxxxx" where x = "Y" (yes) or x = "N" (no) • position: <ul style="list-style-type: none"> – 1: scattered moonlight – 2: scattered starlight – 3: zodiacal light – 4: thermal emission by telescope/instrument – 5: molecular emission of lower atmosphere – 6: sky emission lines of upper atmosphere – 7: airglow continuum (residual continuum)

5.4.2 Field Documentation

5.4.2.1 char _smparmodel_::acontname[SM.LENLINE+1]

- 5.4.2.2 double _smparmodel_::airmass
- 5.4.2.3 double _smparmodel_::alpha
- 5.4.2.4 double _smparmodel_::alt
- 5.4.2.5 double _smparmodel_::altmoon
- 5.4.2.6 char _smparmodel_::datapath[SM_MAXLEN+1]
- 5.4.2.7 double _smparmodel_::eps[SM_MAXPAR]
- 5.4.2.8 char _smparmodel_::incl[SM_LENLINE+1]
- 5.4.2.9 double _smparmodel_::lat_ecl
- 5.4.2.10 char _smparmodel_::libpath[SM_MAXLEN+1]
- 5.4.2.11 char _smparmodel_::libstruct[SM_LENLINE+1]
- 5.4.2.12 char _smparmodel_::libstruct1[SM_LENLINE+1]
- 5.4.2.13 char _smparmodel_::libstruct2[SM_LENLINE+1]
- 5.4.2.14 char _smparmodel_::linetabname[SM_LENLINE+1]
- 5.4.2.15 double _smparmodel_::lon_ecl
- 5.4.2.16 char _smparmodel_::lunirrname[SM_LENLINE+1]
- 5.4.2.17 char _smparmodel_::miephasename[SM_LENLINE+1]
- 5.4.2.18 double _smparmodel_::moondist
- 5.4.2.19 char _smparmodel_::mscatname[SM_LENLINE+1]
- 5.4.2.20 double _smparmodel_::msolflux
- 5.4.2.21 int _smparmodel_::ncomp
- 5.4.2.22 char _smparmodel_::o3transname[SM_LENLINE+1]
- 5.4.2.23 double _smparmodel_::pwv
- 5.4.2.24 double _smparmodel_::resol
- 5.4.2.25 double _smparmodel_::rho

- 5.4.2.26 char _smparmodel_::rtcode[SM_LENLINE+1]
- 5.4.2.27 int _smparmodel_::season
- 5.4.2.28 char _smparmodel_::solspecname[SM_LENLINE+1]
- 5.4.2.29 char _smparmodel_::starspecname[SM_LENLINE+1]
- 5.4.2.30 double _smparmodel_::temp[SM_MAXPAR]
- 5.4.2.31 int _smparmodel_::time
- 5.4.2.32 char _smparmodel_::vac_air[SM_LENLINE+1]
- 5.4.2.33 char _smparmodel_::vardatname[SM_LENLINE+1]
- 5.4.2.34 char _smparmodel_::zodtabname[SM_LENLINE+1]

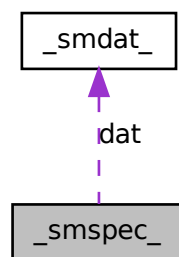
The documentation for this struct was generated from the following file:

- src/[sm_skyemcomp.h](#)

5.5 _smspec_ Struct Reference

```
#include <sm_general.h>
```

Collaboration diagram for _smspec_:



Data Fields

- int [type](#)

- `int n`
- `smdat * dat`

5.5.1 Detailed Description

Structure for spectra

Parameters

<i>type</i>	number of columns (1 = no flux, 2 = no errors, 3 = symmetric error, 4 = lower and upper error)
<i>n</i>	number of wavelengths
<i>*dat</i>	vector of data points defined by <code>smdat</code> structure

5.5.2 Field Documentation

5.5.2.1 `smdat*_smspec_::dat`

5.5.2.2 `int _smspec_::n`

5.5.2.3 `int _smspec_::type`

The documentation for this struct was generated from the following file:

- `src/sm_general.h`

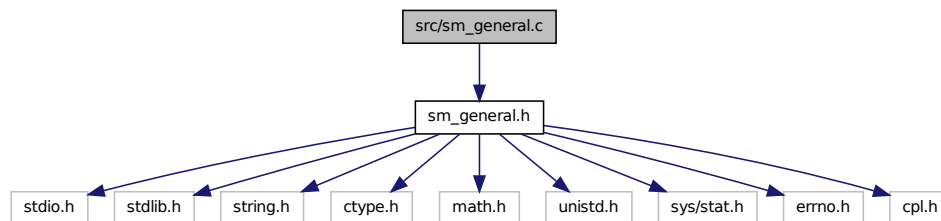
Chapter 6

File Documentation

6.1 src/sm_general.c File Reference

```
#include <sm_general.h>
```

Include dependency graph for sm_general.c:



Functions

- `cpl_error_code` [sm_spec_malloc](#) (`smspec *spec`, `const int size`)
- `cpl_error_code` [sm_spec_create](#) (`smspec *outspec`, `const double limlam[2]`, `const double dlam`)
- `cpl_error_code` [sm_spec_read](#) (`smspec *spec`, `const char *filename`)
- `cpl_error_code` [sm_spec_readrange](#) (`smspec *spec`, `const char *filename`, `const double limlam[2]`, `const int step`)
- `cpl_error_code` [sm_spec_readcpl](#) (`smspec *spec`, `const cpl_table *cpltab`)
- `cpl_error_code` [sm_spec_readfits](#) (`smspec *spec`, `const char *filename`)
- `cpl_error_code` [sm_spec_readfitsrange](#) (`smspec *spec`, `const char *filename`, `const double limlam[2]`, `const int step`)

- cpl_error_code [sm_spec_copy](#) (smspec *outspec, const smspec *inspec)
- cpl_error_code [sm_spec_compgrids](#) (const smspec *spec1, const smspec *spec2)
- cpl_error_code [sm_spec_join](#) (smspec *spec, const smspec *errfunc, const int errflag)
- cpl_error_code [sm_spec_split](#) (smspec *spec, smspec *errfunc, const int errflag)
- cpl_error_code [sm_spec_changetype](#) (smspec *spec, const int type)
- cpl_error_code [sm_spec_scalerange](#) (smspec *spec, const double limlam[2], const char op, const double c)
- cpl_error_code [sm_spec_scale](#) (smspec *spec, const char op, const double c)
- cpl_error_code [sm_spec_modval](#) (smspec *spec, const double lam, const char op, const double c)
- cpl_error_code [sm_spec_calc](#) (smspec *spec, const char op, const smspec *opspec)
- cpl_error_code [sm_spec_funct](#) (smspec *spec, const char *funct, const char baselab)
- cpl_error_code [sm_spec_functnoerr](#) (smspec *spec, const char *funct)
- cpl_error_code [sm_spec_convunits](#) (smspec *spec, const double factor, const int lamexp)
- cpl_error_code [sm_spec_changegrid](#) (smspec *spec, const double factor, const char *scale)
- cpl_error_code [sm_spec_average](#) (const smspec *spec, const double limlam[2], double mean[3])
- cpl_error_code [sm_spec_write](#) (const smspec *spec, const char *filename)
- cpl_error_code [sm_spec_print](#) (const smspec *spec)
- cpl_error_code [sm_spec_writcpl](#) (cpl_table *cpltab, const smspec *spec)
- cpl_error_code [sm_spec_writfits](#) (const smspec *spec, const char *filename)
- cpl_error_code [sm_spec_free](#) (smspec *spec)
- cpl_error_code [sm_spec_extract](#) (smspec *outspec, const smspec *inspec, const double limlam[2])
- cpl_error_code [sm_spec_rebin](#) (smspec *outspec, const smspec *inspec)
- cpl_error_code [sm_spec_interpol](#) (smspec *outspec, const smspec *inspec)
- cpl_error_code [sm_spec_convolve](#) (smspec *spec, const int nkpix, const double *kernel)
- cpl_error_code [sm_param_read](#) (FILE *stream, smparam par[])
- cpl_error_code [sm_param_readcheck](#) (FILE *stream, smparam par[], const char *parname, const int npar)
- cpl_error_code [sm_param_check](#) (smparam par[], const char *parname, const int npar)
- cpl_error_code [sm_grid_malloc](#) (smgrid *xy, const int nx, const int ny)
- cpl_error_code [sm_grid_read](#) (smgrid *xy, const char *filename)
- cpl_error_code [sm_grid_extract](#) (const smgrid *xy, const double x0, const double y0, double *outval)
- cpl_error_code [sm_grid_write](#) (const smgrid *xy, const char *filename)
- cpl_error_code [sm_grid_print](#) (const smgrid *xy)
- cpl_error_code [sm_grid_free](#) (smgrid *xy)
- cpl_error_code [sm_basic_chdir](#) (const char *dir)
- cpl_error_code [sm_basic_access](#) (const char *pathname, const int mode)
- cpl_error_code [sm_basic_mkdir](#) (const char *dir, const mode_t mode)

- cpl_error_code [sm_basic_createdir](#) (const char *dir, const mode_t mode)
- void [sm_basic_initstring](#) (char *str, const long n)
- cpl_boolean [sm_basic_isnumber](#) (char *str)
- char * [sm_basic_replacestring](#) (char *instring, char *oldsubstr, char *newsubstr)
- char * [sm_basic_rmcntrl](#) (char *str)
- void [sm_basic_rmcntrl_inplace](#) (char *str)
- char * [sm_basic_strtrim](#) (char *str)
- void [sm_basic_strtrim_inplace](#) (char *str)
- void [sm_basic_terminatestring](#) (char *str)
- cpl_error_code [sm_basic_interpollin](#) (const double *x_out, double *y_out, const long n_out, const double *x_ref, const double *y_ref, const long n_ref)

6.1.1 Detailed Description

Basic routines used for the sky model

Author

Stefan Noll & ESO In-Kind Team Innsbruck

Since

22 Sep 2009

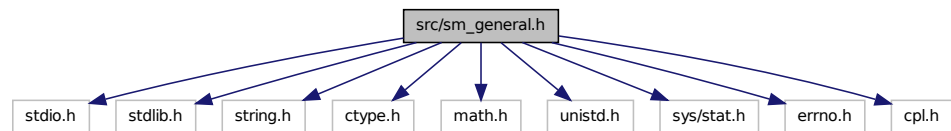
Date

18 Apr 2013

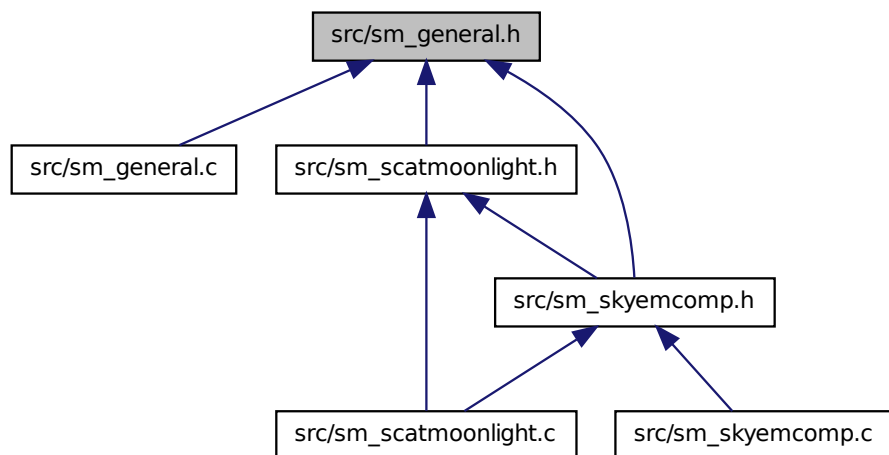
6.2 src/sm_general.h File Reference

```
#include <stdio.h>
#include <stdlib.h>
#include <string.h>
#include <ctype.h>
#include <math.h>
#include <unistd.h>
#include <sys/stat.h>
#include <errno.h>
#include <cpl.h>
```

Include dependency graph for `sm_general.h`:



This graph shows which files directly or indirectly include this file:



Data Structures

- struct `_smdat_`
- struct `_smspec_`
- struct `_smparam_`
- struct `_smgrid_`

Defines

- #define `SM_GENERAL_H`
- #define `SM_BOOL`

- #define [F](#) SM_F
- #define [T](#) SM_T
- #define [SM_NMAXERR](#) 100
- #define [SM_LENLINE](#) 160
- #define [SM_MAXLEN](#) 4000
- #define [SM_MAXPAR](#) 21
- #define [SM_TOL](#) 1e-7
- #define [SM_SR_IN_ARCSEC2](#) 4.254517e+10
- #define [SM_GAS_CONST](#) 8.31447
- #define [SM_MOL_AIR_DRY](#) 0.0289644
- #define [SM_GRAV_ACC](#) 9.80665
- #define [SM_ERROR_FOF](#) SM_ERROR_FOPEN
- #define [SM_ERROR_ISM](#) SM_ERROR_NOMEM
- #define [SM_ERROR_EIS](#) SM_ERROR_SUBROUTINE
- #define [SM_ERROR_ISD](#) SM_ERROR_INSUFF_DATA
- #define [SM_ERROR_FOPEN_TXT](#) "File opening failed"
- #define [SM_ERROR_UFS_TXT](#) "Unexpected file structure"
- #define [SM_ERROR_IFE_TXT](#) "Invalid file name extension"
- #define [SM_ERROR_BDR_TXT](#) "Bad directory"
- #define [SM_ERROR_NDA_TXT](#) "No data"
- #define [SM_ERROR_ISD_TXT](#) "Insufficient data points"
- #define [SM_ERROR_IDG_TXT](#) "Inconsistent data grids"
- #define [SM_ERROR_IDR_TXT](#) "Invalid data range"
- #define [SM_ERROR_IOD_TXT](#) "Invalid order of data points"
- #define [SM_ERROR_IIP_TXT](#) "Invalid input parameter(s)"
- #define [SM_ERROR_IOV_TXT](#) "Invalid object value(s)"
- #define [SM_ERROR_IOS_TXT](#) "Invalid object structure"
- #define [SM_ERROR_SUBROUTINE_TXT](#) "Error in subroutine"
- #define [SM_ERROR_ACCES_TXT](#) "Permission denied"
- #define [SM_ERROR_LOOP_TXT](#) "Too many symbolic links"
- #define [SM_ERROR_NAMETOOLONG_TXT](#) "Pathname too long"
- #define [SM_ERROR_NOENT_TXT](#) "File/dir does not exist"
- #define [SM_ERROR_NOTDIR_TXT](#)
- #define [SM_ERROR_ROFS_TXT](#)
- #define [SM_ERROR_FAULT_TXT](#)
- #define [SM_ERROR_INVAL_TXT](#) "Mode was incorrectly specified"
- #define [SM_ERROR_IO_TXT](#) "I/O error occurred"
- #define [SM_ERROR_NOMEM_TXT](#) "Insufficient memory"
- #define [SM_ERROR_TXTBSY_TXT](#)
- #define [SM_ERROR_EXIST_TXT](#) "File/dir already exists"
- #define [SM_ERROR_NOSPC_TXT](#) "No space left on device"
- #define [SM_ERROR_PERM_TXT](#)
- #define [SM_ERROR_LINK_TXT](#) "Could not create symbolic link"
- #define [SM_ERROR_UNDEF_TXT](#) "Undefined error"
- #define [SM_ERROR_FOF_TXT](#) SM_ERROR_FOPEN_TXT
- #define [SM_ERROR_ISM_TXT](#) SM_ERROR_NOMEM_TXT
- #define [SM_ERROR_EIS_TXT](#) SM_ERROR_SUBROUTINE_TXT

Typedefs

- typedef enum [SM_BOOL](#) [smbool](#)
- typedef enum [_sm_error_code_](#) [sm_error_code](#)
- typedef struct [_smdat_](#) [smdat](#)
- typedef struct [_smspec_](#) [smspec](#)
- typedef struct [_smparam_](#) [smparam](#)
- typedef struct [_smgrid_](#) [smgrid](#)

Enumerations

- enum [SM_BOOL](#) { [SM_F](#), [SM_T](#) }
- enum [_sm_error_code_](#) {
[SM_ERROR_FOPEN](#) = [CPL_ERROR_EOL](#) + 11, [SM_ERROR_UFS](#) = [CPL_ERROR_EOL](#) + 12, [SM_ERROR_IFE](#) = [CPL_ERROR_EOL](#) + 13, [SM_ERROR_NDA](#) = [CPL_ERROR_EOL](#) + 20,
[SM_ERROR_INSUFF_DATA](#) = [CPL_ERROR_EOL](#) + 21, [SM_ERROR_IDG](#) = [CPL_ERROR_EOL](#) + 22, [SM_ERROR_IDR](#) = [CPL_ERROR_EOL](#) + 23, [SM_ERROR_IOD](#) = [CPL_ERROR_EOL](#) + 24,
[SM_ERROR_IIP](#) = [CPL_ERROR_EOL](#) + 30, [SM_ERROR_IOV](#) = [CPL_ERROR_EOL](#) + 31, [SM_ERROR_IOS](#) = [CPL_ERROR_EOL](#) + 32, [SM_ERROR_SUBROUTINE](#) = [CPL_ERROR_EOL](#) + 40,
[SM_ERROR_ACCES](#) = [CPL_ERROR_EOL](#) + 50, [SM_ERROR_LOOP](#) = [CPL_ERROR_EOL](#) + 51, [SM_ERROR_NAMETOOLONG](#) = [CPL_ERROR_EOL](#) + 52, [SM_ERROR_NOENT](#) = [CPL_ERROR_EOL](#) + 53,
[SM_ERROR_NOTDIR](#) = [CPL_ERROR_EOL](#) + 54, [SM_ERROR_ROFS](#) = [CPL_ERROR_EOL](#) + 55, [SM_ERROR_FAULT](#) = [CPL_ERROR_EOL](#) + 56, [SM_ERROR_INVALID](#) = [CPL_ERROR_EOL](#) + 57,
[SM_ERROR_IO](#) = [CPL_ERROR_EOL](#) + 58, [SM_ERROR_NOMEM](#) = [CPL_ERROR_EOL](#) + 59, [SM_ERROR_TXTBSY](#) = [CPL_ERROR_EOL](#) + 60, [SM_ERROR_EXIST](#) = [CPL_ERROR_EOL](#) + 61,
[SM_ERROR_NOSPC](#) = [CPL_ERROR_EOL](#) + 62, [SM_ERROR_PERM](#) = [CPL_ERROR_EOL](#) + 63, [SM_ERROR_BADUSERINPUT](#) = [CPL_ERROR_EOL](#) + 70, [SM_ERROR_LINK](#) = [CPL_ERROR_EOL](#) + 71,
[SM_ERROR_RFM](#) = [CPL_ERROR_EOL](#) + 81, [SM_ERROR_UNDEF](#) = [CPL_ERROR_EOL](#) + 80 }

Functions

- [cpl_error_code](#) [sm_spec_malloc](#) ([smspec](#) *spec, const int size)
- [cpl_error_code](#) [sm_spec_create](#) ([smspec](#) *outspec, const double limlam[2], const double dlam)
- [cpl_error_code](#) [sm_spec_read](#) ([smspec](#) *spec, const char *filename)
- [cpl_error_code](#) [sm_spec_readrange](#) ([smspec](#) *spec, const char *filename, const double limlam[2], const int step)

- cpl_error_code [sm_spec_readcpl](#) (smspec *spec, const cpl_table *cpltab)
- cpl_error_code [sm_spec_readfits](#) (smspec *spec, const char *filename)
- cpl_error_code [sm_spec_readfitsrange](#) (smspec *spec, const char *filename, const double limlam[2], const int step)
- cpl_error_code [sm_spec_copy](#) (smspec *outspec, const smspec *inspec)
- cpl_error_code [sm_spec_compgrids](#) (const smspec *spec1, const smspec *spec2)
- cpl_error_code [sm_spec_join](#) (smspec *spec, const smspec *errfunc, const int errflag)
- cpl_error_code [sm_spec_split](#) (smspec *spec, smspec *errfunc, const int errflag)
- cpl_error_code [sm_spec_changetype](#) (smspec *spec, const int type)
- cpl_error_code [sm_spec_scalerange](#) (smspec *spec, const double limlam[2], const char op, const double c)
- cpl_error_code [sm_spec_scale](#) (smspec *spec, const char op, const double c)
- cpl_error_code [sm_spec_modval](#) (smspec *spec, const double lam, const char op, const double c)
- cpl_error_code [sm_spec_calc](#) (smspec *spec, const char op, const smspec *opspec)
- cpl_error_code [sm_spec_funct](#) (smspec *spec, const char *funct, const char baselab)
- cpl_error_code [sm_spec_functnoerr](#) (smspec *spec, const char *funct)
- cpl_error_code [sm_spec_convunits](#) (smspec *spec, const double factor, const int lamexp)
- cpl_error_code [sm_spec_changegrid](#) (smspec *spec, const double factor, const char *scale)
- cpl_error_code [sm_spec_average](#) (const smspec *spec, const double limlam[2], double mean[3])
- cpl_error_code [sm_spec_write](#) (const smspec *spec, const char *filename)
- cpl_error_code [sm_spec_print](#) (const smspec *spec)
- cpl_error_code [sm_spec_writecpl](#) (cpl_table *cpltab, const smspec *spec)
- cpl_error_code [sm_spec_writelfits](#) (const smspec *spec, const char *filename)
- cpl_error_code [sm_spec_free](#) (smspec *spec)
- cpl_error_code [sm_spec_extract](#) (smspec *outspec, const smspec *inspec, const double limlam[2])
- cpl_error_code [sm_spec_rebin](#) (smspec *outspec, const smspec *inspec)
- cpl_error_code [sm_spec_interpol](#) (smspec *outspec, const smspec *inspec)
- cpl_error_code [sm_spec_convolve](#) (smspec *spec, const int nkpix, const double *kernel)
- cpl_error_code [sm_param_read](#) (FILE *stream, smparam par[])
- cpl_error_code [sm_param_readcheck](#) (FILE *stream, smparam par[], const char *parname, const int npar)
- cpl_error_code [sm_param_check](#) (smparam par[], const char *parname, const int npar)
- cpl_error_code [sm_grid_malloc](#) (smgrid *xy, const int nx, const int ny)
- cpl_error_code [sm_grid_read](#) (smgrid *xy, const char *filename)
- cpl_error_code [sm_grid_extract](#) (const smgrid *xy, const double x0, const double y0, double *outval)
- cpl_error_code [sm_grid_write](#) (const smgrid *xy, const char *filename)
- cpl_error_code [sm_grid_print](#) (const smgrid *xy)

- cpl_error_code [sm_grid_free](#) (smgrid *xy)
- cpl_error_code [sm_basic_chdir](#) (const char *dir)
- cpl_error_code [sm_basic_access](#) (const char *pathname, const int mode)
- cpl_error_code [sm_basic_mkdir](#) (const char *dir, const mode_t mode)
- cpl_error_code [sm_basic_createdir](#) (const char *dir, const mode_t mode)
- void [sm_basic_initstring](#) (char *str, const long n)
- cpl_boolean [sm_basic_isnumber](#) (char *str)
- char * [sm_basic_replacestring](#) (char *instr, char *oldsubstr, char *newsubstr)
- char * [sm_basic_rmcntrl](#) (char *str)
- void [sm_basic_rmcntrl_inplace](#) (char *str)
- char * [sm_basic_strtrim](#) (char *str)
- void [sm_basic_strtrim_inplace](#) (char *str)
- void [sm_basic_terminatestring](#) (char *str)
- cpl_error_code [sm_basic_interpollin](#) (const double *x_out, double *y_out, const long n_out, const double *x_ref, const double *y_ref, const long n_ref)

6.2.1 Detailed Description

Header for basic routines used for the sky model

Author

Stefan Noll & ESO In-Kind Team Innsbruck

Since

22 Sep 2009

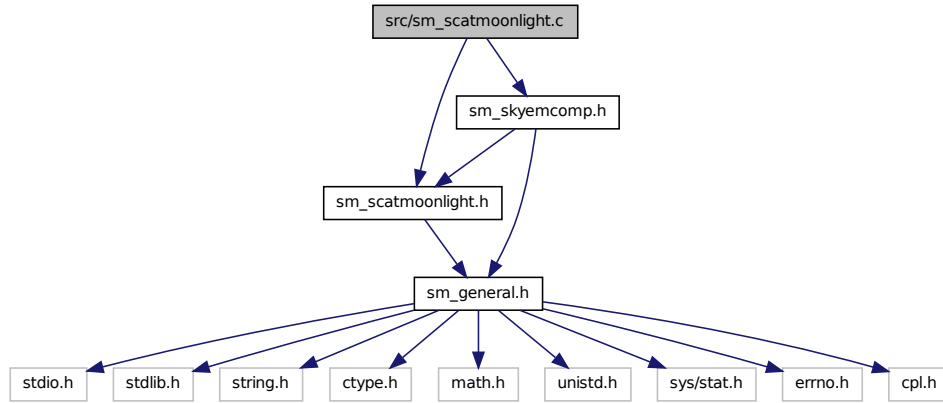
Date

11 Dec 2012

6.3 src/sm_scatmoonlight.c File Reference

```
#include <sm_scatmoonlight.h>
#include <sm_skyemcomp.h>
```


Include dependency graph for sm_scatmoonlight.c:



Functions

- `cpl_error_code sm_scat_moon (smspec *spec, const double z, const double zmoon, const double rho, const smspec *molabstrans, const smspec *rscattrans, const smspec *mscattrans, const smgrid *miephase, const smgrid *multiscat)`
- `cpl_error_code sm_scat_createdatatab (cpl_table *tab)`
- `cpl_error_code sm_scat_calcextcurv (cpl_table *tautab)`
- `cpl_error_code sm_scat_calctau (double *lam0, double *tau0r, double *tau0m, const double tau0, const cpl_table *tautab)`
- `double sm_scat_tau2lam (const double tau0, const smspec *scattrans)`
- `cpl_error_code sm_scat_calcscatmoon (double *iscat, double *xeff, const int ns, const double z0, const double zeta, const double theta, const double flux0, const double tau0, const double cextr, const double cexm, const double cexta, const smgrid *miephase, const double lam0)`
- `double sm_scat_geteffcoldens (const double z, const double sig, const int ns, const int phasefunc)`
- `double sm_scat_getmolecdens (const double h)`
- `double sm_scat_getaerosoldens (const double h)`
- `double sm_scat_getmiephasefunc (const double theta)`
- `cpl_error_code sm_scat_calcscatspec (smspec *spec, const cpl_table *tab, const smspec *molabstrans, const smspec *rscattrans, const smspec *mscattrans)`

6.3.1 Detailed Description

Routines for calculation of scattered moonlight

Author

Stefan Noll & ESO In-Kind Team Innsbruck

Since

05 Jun 2012

Date

21 Jan 2013

6.3.2 Function Documentation**6.3.2.1** `cpl_error_code sm_scatt_calcextcurv (cpl_table * tautab)`

Calculates Rayleigh and Mie extinction curves and the total optical depths depending on wavelength. The parametrisations are optimised for Cerro Paranal. The covered wavelengths range from 0.3 to 30.0 μm .

INPUT:**Parameters**

<i>tautab</i>	empty CPL table
---------------	-----------------

OUTPUT:**Parameters**

<i>tautab</i>	table with optical depths by Rayleigh scattering and aerosol extinction
---------------	---

ERRORS:

- none

6.3.2.2 `cpl_error_code sm_scatt_calcscatmoon (double * iscat, double * xeff, const int ns, const double z0, const double zeta, const double theta, const double flux0, const double tau0, const double cextr, const double cextm, const double cexta, const smgrid * miephase, const double lam0)`

Calculation of intensity of scattered moonlight, which is scattered into the given line of sight through the Earth's atmosphere, and effective airmass relative to a direct path from zenith to observer. Rayleigh scattering, aerosol extinction, and molecular absorption are considered. For the latter, the same particle distribution as for Rayleigh scattering is assumed (optimal for molecular oxygen). The wavelength-dependent phase functions for Mie scattering were calculated based on optimised log-normal aerosol distributions (Warneck & Williams 2012) and the Bohren-Huffman Mie scattering algorithm (IDL code BHMIE). The single scattering formalism in an spherical atmosphere used is described by Wolstencroft & van Breda (1967), Staude (1975), and Bernstein et al. (2002). As in the latter reference polarisation is neglected. For estimating the contribution of multiple scattering, 3D double scattering calculations were performed, which consider mixed

scattering at air molecules, aerosols, and the surface. The results of the multiple scattering calculations are provided as data table to this routine.

INPUT:**Parameters**

<i>ns</i>	maximum number of data points (- 1) along a path in the atmosphere
<i>z0</i>	line-of-sight zenith distance in rad
<i>zeta</i>	zenith distance of radiation source in rad
<i>theta</i>	angle between line of sight and radiation source
<i>flux0</i>	flux of radiation source
<i>tau0</i>	optical depth at zenith
<i>cextr</i>	Rayleigh extinction cross section in cm^2
<i>cextm</i>	Mie extinction cross section in cm^2
<i>cexta</i>	molecular absorption cross section in cm^2
<i>miephase</i>	grid structure for Mie scattering phase functions
<i>lam0</i>	wavelength for Mie scattering

OUTPUT:**Parameters**

<i>iscat</i>	intensity scattered into line-of-sight
<i>xeff</i>	effective airmass

ERRORS:

- IIP: Invalid input parameter(s)

6.3.2.3 `cpl_error_code sm_scat_calcspec (smspec * spec, const cpl_table * tab, const smspec * molabstrans, const smspec * rscattrans, const smspec * mscattrans)`

Calculates spectrum of scattering intensities. The conversion of optical depths into wavelengths is performed by means of the transmission curves for Rayleigh and Mie scattering. Scattering intensities are interpolated. The scattering intensities are roughly corrected for molecular absorption by means of effective airmasses corresponding to the extinction optical depths based on scattering and absorption.

INPUT:**Parameters**

<i>spec</i>	desired wavelength grid
<i>tab</i>	list of scattering intensities and effective airmasses for different optical depths
<i>molabstrans</i>	transmission curve from radiative transfer code
<i>rscattrans</i>	transmission curve for Rayleigh scattering
<i>mscattrans</i>	transmission curve for aerosol extinction

OUTPUT:

Parameters

<i>spec</i>	spectrum of scattering intensities
-------------	------------------------------------

ERRORS:

- IDG: Inconsistent data grids

6.3.2.4 `cpl_error_code sm_scatt_calctau (double * lam0, double * tau0r, double * tau0m, const double tau0, const cpl_table * tautab)`

Retrieves wavelength, Rayleigh optical depth, and Mie optical depth for an input total optical depth. The extinction curves used are optimised for Cerro Paranal. For optical depths beyond the covered wavelength range, the limiting wavelength and optical depths consistent with the Rayleigh and Mie fractions for this wavelength are returned.

INPUT:**Parameters**

<i>tau0</i>	optical depth at zenith
<i>tautab</i>	CPL table with extinction curves

OUTPUT:**Parameters**

<i>lam0</i>	wavelength related to input optical depth
<i>tau0r</i>	optical depth by Rayleigh scattering
<i>tau0m</i>	optical depth by Mie extinction

ERRORS:

- none

6.3.2.5 `cpl_error_code sm_scatt_createdatatab (cpl_table * tab)`

Writes CPL table with grid of zenithal optical depths for atmospheric scattering calculations.

INPUT:**Parameters**

<i>tab</i>	empty CPL table
------------	-----------------

OUTPUT:**Parameters**

<i>tab</i>	grid of zenithal optical depths
------------	---------------------------------

ERRORS:

- none

6.3.2.6 double sm_scat_getaerosoldens (const double *h*)

Provides aerosol density in cm^{-3} for given height in km. Uses exponential formula. The constants are from Staude (1975).

INPUT:**Parameters**

<i>h</i>	height above ground in km
----------	---------------------------

RETURN:

- aerosol density in cm^{-3}

ERRORS:

- none

6.3.2.7 double sm_scat_geteffcoldens (const double *z*, const double *sig*, const int *ns*, const int *phasefunc*)

Calculates effective column density in cm^{-2} for a path from top of atmosphere to point of reference by considering the given zenith distance. The size of the volume elements for the integration increases with height in order to consider the decrease of density with height.

INPUT:**Parameters**

<i>z</i>	zenith distance in rad at point of reference
<i>sig</i>	height in km of point of reference relative to centre of earth
<i>ns</i>	maximum number of data points (- 1) along a path in the atmosphere
<i>phasefunc</i>	0 = Rayleigh scattering, 1 = Mie (aerosol) scattering

RETURN:

- effective column density in cm^{-2} for given path

ERRORS:

- IIP: Invalid input parameter(s)

6.3.2.8 double sm_scat_getmiephasefunc (const double *theta*)

Provides phase function for Mie scattering. For scattering angles up to 20 deg the empirical Green et al. (1971) phase function for aerosols is used. The phase function for larger scattering angles is extrapolated by a linear fit of the phase function in a $\log P - \log \theta$ diagram. This simple approach neglects the typical increase of phase functions for large scattering angles. However, in view of the strong peak for forward scattering (P covers about two orders of magnitude), the details of the phase function beyond the empirical part are not crucial for scattering calculations.

INPUT:

Parameters

<i>theta</i>	scattering angle
--------------	------------------

RETURN:

- value of phase function for given scattering angle

ERRORS:

- none

6.3.2.9 double sm_scat_getmolecdens (const double *h*)

Provides molecular density in cm^{-3} for given height in km. Uses barometric formula. The constants are from Staude (1975).

INPUT:

Parameters

<i>h</i>	height above ground in km
----------	---------------------------

RETURN:

- molecular density in cm^{-3}

ERRORS:

- none

6.3.2.10 cpl_error_code sm_scat_moon (smspec * *spec*, const double *z*, const double *zmoon*, const double *rho*, const smspec * *molabstrans*, const smspec * *rscattrans*, const smspec * *mscattrans*, const smgrid * *miephase*, const smgrid * *multiscat*)

Atmospheric scattering calculations for moonlight. Required input parameters are line-of-sight zenith distance, zenith distance of Moon, and angle between line of sight and

Moon. All parameters have to be provided in deg. Be aware that there are impossible combinations of these three parameters, which cause a programme termination without scattering results. In the case of a successful completion, the procedure returns a spectrum of scattering intensities per arcsec² for an input total Moon flux of 1. The required conversion from τ to wavelength is carried out by means of the input zenithal transmission curves for Rayleigh and Mie scattering. Before multiplication, the transmission curve for molecular absorption is adapted to the effective airmasses related to the scattering paths of both scattering processes. This procedure works best for Rayleigh scattering and absorption by O₂.

INPUT:**Parameters**

<i>spec</i>	desired wavelength grid
<i>z</i>	line-of-sight zenith distance in deg
<i>zmoon</i>	zenith distance of Moon in deg
<i>rho</i>	angle between line of sight and Moon in deg
<i>molabstrans</i>	transmission curve from radiative transfer code
<i>rscattrans</i>	transmission curve for Rayleigh scattering
<i>mscattrans</i>	transmission curve for Mie scattering
<i>miephase</i>	grid structure for Mie scattering phase functions
<i>multiscat</i>	grid structure for multiple scattering correction factors

OUTPUT:**Parameters**

<i>spec</i>	spectrum of scattered moonlight (input flux = 1)
-------------	--

ERRORS:

- IIP: Invalid input parameter(s)
- IDG: Inconsistent data grids

6.3.2.11 double sm_scat_tau2lam (const double *tau0*, const smspec * *scattrans*)

Conversion of an optical depth into a wavelength by means of a given transmission curve.

INPUT:**Parameters**

<i>tau0</i>	optical depth at zenith
<i>scattrans</i>	transmission curve

RETURN:

- wavelength related to input optical depth

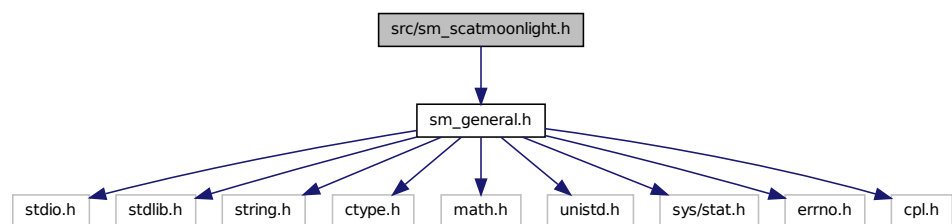
ERRORS:

- none

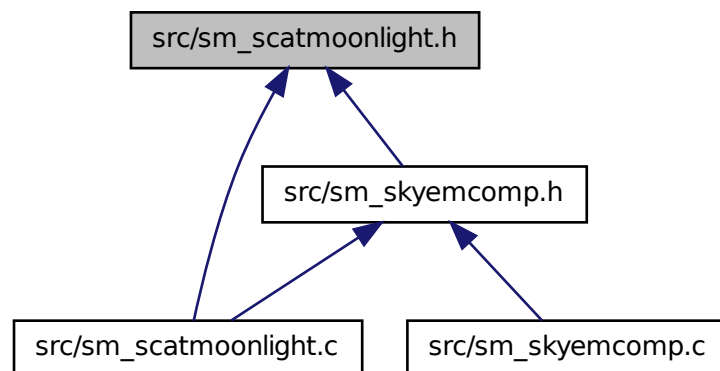
6.4 src/sm_scatmoonlight.h File Reference

```
#include <sm_general.h>
```

Include dependency graph for sm_scatmoonlight.h:



This graph shows which files directly or indirectly include this file:



- `#define SM_SCATMOONLIGHT_H`
- `#define SM_NS 20`
- `#define SM_LNSCALE 0.5`

- #define [SM_R](#) 6371.
- #define [SM_HMAX](#) 200.
- #define [SM_H](#) 2.64
- #define [SM_NO_R](#) 2.67e19
- #define [SM_H0_R](#) 7.99
- #define [SM_NO_M](#) 1.11e4
- #define [SM_H0_M](#) 1.2
- #define [SM_SSA_M](#) 0.97
- #define [SM_NTAU1](#) 7
- #define [SM_DTAU1](#) 0.01
- #define [SM_NTAU2](#) 31
- #define [SM_DTAU2](#) 0.03
- #define [SM_NTAU3](#) 20
- #define [SM_DTAU3](#) 0.2
- #define [SM_REFTAU0](#) 0.02
- cpl_error_code [sm_scat_moon](#) (smspec *spec, const double z, const double zmoon, const double rho, const smspec *molabstrans, const smspec *rscattrans, const smspec *mscattrans, const smgrid *miephase, const smgrid *multiscat)
- cpl_error_code [sm_scat_createdatatab](#) (cpl_table *tab)
- cpl_error_code [sm_scat_calcextcurv](#) (cpl_table *tautab)
- cpl_error_code [sm_scat_calctau](#) (double *lam0, double *tau0r, double *tau0m, const double tau0, const cpl_table *tautab)
- double [sm_scat_tau2lam](#) (const double tau0, const smspec *scattrans)
- cpl_error_code [sm_scat_calcscatmoon](#) (double *iscat, double *xeff, const int ns, const double z0, const double zeta, const double theta, const double flux0, const double tau0, const double cextr, const double cextm, const double cexta, const smgrid *miephase, const double lam0)
- double [sm_scat_geteffcoldens](#) (const double z, const double sig, const int ns, const int phasefunc)
- double [sm_scat_getmolecdens](#) (const double h)
- double [sm_scat_getaerosoldens](#) (const double h)
- double [sm_scat_getmiephasefunc](#) (const double theta)
- cpl_error_code [sm_scat_calcscatspec](#) (smspec *spec, const cpl_table *tab, const smspec *molabstrans, const smspec *rscattrans, const smspec *mscattrans)

6.4.1 Detailed Description

Header for routines for calculation of scattered moonlight

Author

Stefan Noll & ESO In-Kind Team Innsbruck

Since

05 Jun 2012

Date

25 Jan 2013

6.4.2 Define Documentation

6.4.2.1 #define SM_DTAU1 0.01

First τ step size

6.4.2.2 #define SM_DTAU2 0.03

Second τ step size

6.4.2.3 #define SM_DTAU3 0.2

Second τ step size

6.4.2.4 #define SM_H 2.64

Altitude of observer in km

6.4.2.5 #define SM_H0_M 1.2

Scale height of aerosol number density in km

6.4.2.6 #define SM_H0_R 7.99

Scale height of atmosphere in km

6.4.2.7 #define SM_HMAX 200.

Top of atmosphere in km

6.4.2.8 #define SM_LNSCALE 0.5

Scaling factor in km for unitless logarithmic bins

6.4.2.9 #define SM_N0_M 1.11e4

Density of aerosols at surface in cm^{-3}

6.4.2.10 #define SM_N0_R 2.67e19

Density of atmosphere at surface in cm^{-3}

6.4.2.11 #define SM_NS 20

Number of points for integration of particle number density

6.4.2.12 #define SM_NTAU1 7

Number of τ values for first step size

6.4.2.13 #define SM_NTAU2 31

Number of τ values for second step size

6.4.2.14 #define SM_NTAU3 20

Number of τ values for third step size

6.4.2.15 #define SM_R 6371.

Radius of Earth in km

6.4.2.16 #define SM_REFTAU0 0.02

Reference scattering optical depth for effective absorption airmass calculation

6.4.2.17 #define SM_SCATMOONLIGHT_H

Number of points for integration of particle number density

6.4.2.18 #define SM_SSA_M 0.97

Single scattering albedo for aerosols

6.4.3 Function Documentation**6.4.3.1 cpl_error_code sm_scat_calcextcurv (cpl_table * *tautab*)**

Calculates Rayleigh and Mie extinction curves and the total optical depths depending on wavelength. The parametrisations are optimised for Cerro Paranal. The covered wavelengths range from 0.3 to 30.0 μm .

INPUT:**Parameters**

<i>tautab</i>	empty CPL table
---------------	-----------------

OUTPUT:**Parameters**

<i>tautab</i>	table with optical depths by Rayleigh scattering and aerosol extinction
---------------	---

ERRORS:

- none

6.4.3.2 `cpl_error_code sm_scattermoon (double * iscat, double * xeff, const int ns, const double z0, const double zeta, const double theta, const double flux0, const double tau0, const double cextr, const double cextm, const double cexta, const smgrid * miephase, const double lam0)`

Calculation of intensity of scattered moonlight, which is scattered into the given line of sight through the Earth's atmosphere, and effective airmass relative to a direct path from zenith to observer. Rayleigh scattering, aerosol extinction, and molecular absorption are considered. For the latter, the same particle distribution as for Rayleigh scattering is assumed (optimal for molecular oxygen). The wavelength-dependent phase functions for Mie scattering were calculated based on optimised log-normal aerosol distributions (Warneck & Williams 2012) and the Bohren-Huffman Mie scattering algorithm (IDL code BHMIE). The single scattering formalism in a spherical atmosphere used is described by Wolstencroft & van Breda (1967), Staude (1975), and Bernstein et al. (2002). As in the latter reference polarisation is neglected. For estimating the contribution of multiple scattering, 3D double scattering calculations were performed, which consider mixed scattering at air molecules, aerosols, and the surface. The results of the multiple scattering calculations are provided as data table to this routine.

INPUT:**Parameters**

<i>ns</i>	maximum number of data points (- 1) along a path in the atmosphere
<i>z0</i>	line-of-sight zenith distance in rad
<i>zeta</i>	zenith distance of radiation source in rad
<i>theta</i>	angle between line of sight and radiation source
<i>flux0</i>	flux of radiation source
<i>tau0</i>	optical depth at zenith
<i>cextr</i>	Rayleigh extinction cross section in cm ²
<i>cextm</i>	Mie extinction cross section in cm ²
<i>cexta</i>	molecular absorption cross section in cm ²
<i>miephase</i>	grid structure for Mie scattering phase functions
<i>lam0</i>	wavelength for Mie scattering

OUTPUT:**Parameters**

<i>iscat</i>	intensity scattered into line-of-sight
<i>xeff</i>	effective airmass

ERRORS:

- IIP: Invalid input parameter(s)

6.4.3.3 `cpl_error_code sm_scat_calcspec (smspec * spec, const cpl_table * tab, const smspec * molabstrans, const smspec * rscattrans, const smspec * mscattrans)`

Calculates spectrum of scattering intensities. The conversion of optical depths into wavelengths is performed by means of the transmission curves for Rayleigh and Mie scattering. Scattering intensities are interpolated. The scattering intensities are roughly corrected for molecular absorption by means of effective airmasses corresponding to the extinction optical depths based on scattering and absorption.

INPUT:**Parameters**

<i>spec</i>	desired wavelength grid
<i>tab</i>	list of scattering intensities and effective airmasses for different optical depths
<i>molabstrans</i>	transmission curve from radiative transfer code
<i>rscattrans</i>	transmission curve for Rayleigh scattering
<i>mscattrans</i>	transmission curve for aerosol extinction

OUTPUT:**Parameters**

<i>spec</i>	spectrum of scattering intensities
-------------	------------------------------------

ERRORS:

- IDG: Inconsistent data grids

6.4.3.4 `cpl_error_code sm_scat_calctau (double * lam0, double * tau0r, double * tau0m, const double tau0, const cpl_table * tautab)`

Retrieves wavelength, Rayleigh optical depth, and Mie optical depth for an input total optical depth. The extinction curves used are optimised for Cerro Paranal. For optical depths beyond the covered wavelength range, the limiting wavelength and optical depths consistent with the Rayleigh and Mie fractions for this wavelength are returned.

INPUT:**Parameters**

<i>tau0</i>	optical depth at zenith
<i>tautab</i>	CPL table with extinction curves

OUTPUT:**Parameters**

<i>lam0</i>	wavelength related to input optical depth
<i>tau0r</i>	optical depth by Rayleigh scattering
<i>tau0m</i>	optical depth by Mie extinction

ERRORS:

- none

6.4.3.5 `cpl_error_code sm_scatter_createdatab (cpl_table * tab)`

Writes CPL table with grid of zenithal optical depths for atmospheric scattering calculations.

INPUT:**Parameters**

<i>tab</i>	empty CPL table
------------	-----------------

OUTPUT:**Parameters**

<i>tab</i>	grid of zenithal optical depths
------------	---------------------------------

ERRORS:

- none

6.4.3.6 `double sm_scatter_getaerosoldens (const double h)`

Provides aerosol density in cm^{-3} for given height in km. Uses exponential formula. The constants are from Staude (1975).

INPUT:**Parameters**

<i>h</i>	height above ground in km
----------	---------------------------

RETURN:

- aerosol density in cm^{-3}

ERRORS:

- none

6.4.3.7 `double sm_scat_geteffcoldens (const double z, const double sig, const int ns, const int phasefunc)`

Calculates effective column density in cm^{-2} for a path from top of atmosphere to point of reference by considering the given zenith distance. The size of the volume elements for the integration increases with height in order to consider the decrease of density with height.

INPUT:

Parameters

<i>z</i>	zenith distance in rad at point of reference
<i>sig</i>	height in km of point of reference relative to centre of earth
<i>ns</i>	maximum number of data points (- 1) along a path in the atmosphere
<i>phasefunc</i>	0 = Rayleigh scattering, 1 = Mie (aerosol) scattering

RETURN:

- effective column density in cm^{-2} for given path

ERRORS:

- IIP: Invalid input parameter(s)

6.4.3.8 `double sm_scat_getmiephasefunc (const double theta)`

Provides phase function for Mie scattering. For scattering angles up to 20 deg the empirical Green et al. (1971) phase function for aerosols is used. The phase function for larger scattering angles is extrapolated by a linear fit of the phase function in a $\log P - \log \theta$ diagram. This simple approach neglects the typical increase of phase functions for large scattering angles. However, in view of the strong peak for forward scattering (P covers about two orders of magnitude), the details of the phase function beyond the empirical part are not crucial for scattering calculations.

INPUT:

Parameters

<i>theta</i>	scattering angle
--------------	------------------

RETURN:

- value of phase function for given scattering angle

ERRORS:

- none

6.4.3.9 double sm_scat_getmolecdens (const double *h*)

Provides molecular density in cm^{-3} for given height in km. Uses barometric formula. The constants are from Staude (1975).

INPUT:

Parameters

<i>h</i>	height above ground in km
----------	---------------------------

RETURN:

- molecular density in cm^{-3}

ERRORS:

- none

6.4.3.10 cpl_error_code sm_scat_moon (smspec * *spec*, const double *z*, const double *zmoon*, const double *rho*, const smspec * *molabstrans*, const smspec * *rscattrans*, const smspec * *mscattrans*, const smgrid * *miephase*, const smgrid * *multiscat*)

Atmospheric scattering calculations for moonlight. Required input parameters are line-of-sight zenith distance, zenith distance of Moon, and angle between line of sight and Moon. All parameters have to be provided in deg. Be aware that there are impossible combinations of these three parameters, which cause a programme termination without scattering results. In the case of a successful completion, the procedure returns a spectrum of scattering intensities per arcsec^2 for an input total Moon flux of 1. The required conversion from τ to wavelength is carried out by means of the input zenithal transmission curves for Rayleigh and Mie scattering. Before multiplication, the transmission curve for molecular absorption is adapted to the effective airmasses related to the scattering paths of both scattering processes. This procedure works best for Rayleigh scattering and absorption by O_2 .

INPUT:

Parameters

<i>spec</i>	desired wavelength grid
<i>z</i>	line-of-sight zenith distance in deg
<i>zmoon</i>	zenith distance of Moon in deg
<i>rho</i>	angle between line of sight and Moon in deg
<i>molabstrans</i>	transmission curve from radiative transfer code
<i>rscattrans</i>	transmission curve for Rayleigh scattering
<i>mscattrans</i>	transmission curve for Mie scattering
<i>miephase</i>	grid structure for Mie scattering phase functions
<i>multiscat</i>	grid structure for multiple scattering correction factors

OUTPUT:

Parameters

<i>spec</i>	spectrum of scattered moonlight (input flux = 1)
-------------	--

ERRORS:

- IIP: Invalid input parameter(s)
- IDG: Inconsistent data grids

6.4.3.11 double sm_scat_tau2lam (const double *tau0*, const smspec * *scattrans*)

Conversion of an optical depth into a wavelength by means of a given transmission curve.

INPUT:

Parameters

<i>tau0</i>	optical depth at zenith
<i>scattrans</i>	transmission curve

RETURN:

- wavelength related to input optical depth

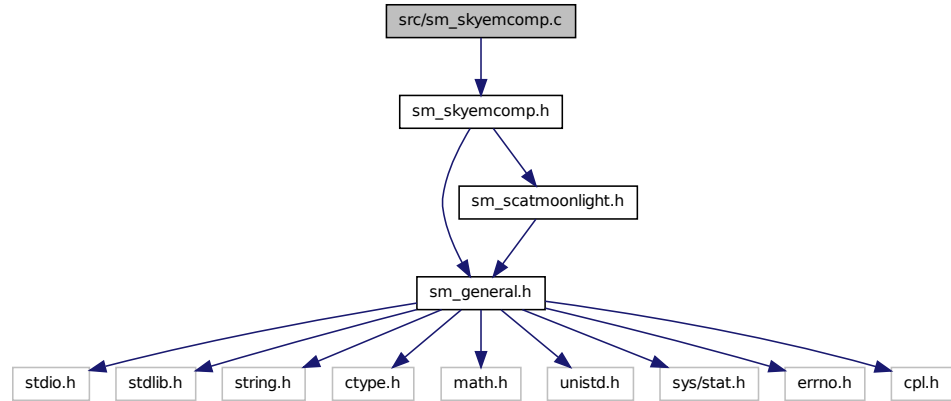
ERRORS:

- none

6.5 src/sm_skyemcomp.c File Reference

```
#include <sm_skyemcomp.h>
```

Include dependency graph for `sm_skyemcomp.c`:



Functions

- `cpl_error_code` [sm_etc_calcmoel](#) (`cpl_table *skytable`, `const cpl_parameterlist *params`)
- `cpl_error_code` [sm_etc_getparams](#) (`smparmodel *modelpar`, `const cpl_parameterlist *params`)
- `const cpl_parameter *` [sm_etc_parameterlist_find_const](#) (`const cpl_parameterlist *self`, `const char *name`, `cpl_type type`)
- `cpl_error_code` [sm_etc_splitstring](#) (`cpl_array *outval`, `char *instring`)
- `cpl_error_code` [sm_etc_readfilenames](#) (`smparmodel *modelpar`, `const cpl_parameterlist *params`)
- `cpl_error_code` [sm_etc_writetable](#) (`cpl_table *skytable`, `const smspec *radiance`, `const smspec *transmission`)
- `cpl_error_code` [sm_comp_skyemcomp](#) (`smspec *radiance`, `smspec *transmission`, `const smparmodel modelpar`)
- `cpl_error_code` [sm_comp_extrapoltrans](#) (`smspec *spec`, `const smparmodel modelpar`, `const double lim[2]`)
- `double` [sm_comp_altoairmass](#) (`const double alt`)
- `cpl_error_code` [sm_comp_getmolspec](#) (`smspec *spec`, `const cpl_parameterlist *libfilepar`, `const double lim[2]`)
- `cpl_error_code` [sm_comp_readlibstruct](#) (`cpl_parameterlist *libfilepar`, `const smparmodel modelpar`, `const char *spectype`)
- `double` [sm_comp_vactoir_single](#) (`const double lam`)
- `cpl_error_code` [sm_comp_vactoir_spec](#) (`smspec *spec`)
- `double` [sm_comp_airtovac_single](#) (`const double lam`)

- cpl_error_code [sm_comp_airtovac_spec](#) (smspec *spec)
- double [sm_comp_calcrayleighscat1](#) (const double lam)
- cpl_error_code [sm_comp_calcrayleighscat](#) (smspec *rscattrans)
- double [sm_comp_calcmiescat1](#) (const double lam)
- cpl_error_code [sm_comp_calcmiescat](#) (smspec *mscattrans)
- cpl_error_code [sm_comp_scaletranscurv](#) (smspec *trans, const [smparmodel](#) modelpar)
- cpl_error_code [sm_comp_lunskybright](#) (smspec *spec, const [smparmodel](#) modelpar, const [smspec](#) *solspec, const [smspec](#) *molabstrans, const [smspec](#) *rscattrans, const [smspec](#) *mscattrans)
- cpl_error_code [sm_comp_lunskybright_v](#) (smspec *spec, const [smparmodel](#) modelpar, const [smspec](#) *solspec, const [smspec](#) *molabstrans, const [smspec](#) *rscattrans, const [smspec](#) *mscattrans)
- cpl_error_code [sm_comp_getmoonbedo](#) (smspec *albedo, const [smparmodel](#) modelpar)
- cpl_error_code [sm_comp_scatstarlight](#) (smspec *spec, const [smparmodel](#) modelpar, const [smspec](#) *molabstrans)
- cpl_error_code [sm_comp_zodskybright](#) (smspec *spec, const [smparmodel](#) modelpar, const [smspec](#) *solspec, const [smspec](#) *molabstrans, const [smspec](#) *rscattrans, const [smspec](#) *mscattrans)
- cpl_error_code [sm_comp_telem](#) (smspec *spec, const [smparmodel](#) modelpar)
- cpl_error_code [sm_comp_extrapolrad](#) (smspec *spec, const [smspec](#) *trans, const [smparmodel](#) modelpar, const double lim[2])
- cpl_error_code [sm_comp_getlinespec](#) (smspec *spec, const [smparmodel](#) modelpar, const [smspec](#) *rscattrans, const [smspec](#) *mscattrans)
- cpl_error_code [sm_comp_readvarpar](#) (cpl_table *varpar, const [smparmodel](#) modelpar)
- cpl_error_code [sm_comp_scalelinetab](#) (cpl_table *linedat, const [smspec](#) *linetab, const cpl_table *varpar)
- cpl_error_code [sm_comp_convertlinetab](#) (smspec *spec, const cpl_table *linedat)
- cpl_error_code [sm_comp_convertlinetabo](#) (smspec *spec, const cpl_table *linedat)
- cpl_error_code [sm_comp_airglowcont](#) (smspec *spec, const [smparmodel](#) modelpar, const [smspec](#) *molabstrans, const [smspec](#) *rscattrans, const [smspec](#) *mscattrans)

6.5.1 Detailed Description

Routines for the ETC sky model

Author

Stefan Noll & ESO In-Kind Team Innsbruck

Since

02 Oct 2009

Date

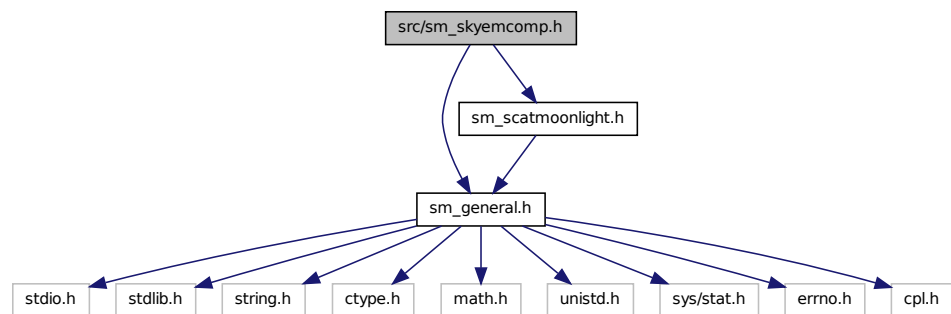
31 Oct 2013

6.6 src/sm_skyemcomp.h File Reference

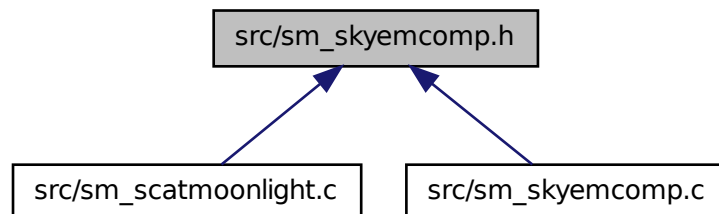
```
#include <sm_general.h>
```

```
#include <sm_scatmoonlight.h>
```

Include dependency graph for sm_skyemcomp.h:



This graph shows which files directly or indirectly include this file:



Data Structures

- struct [_smparmodel_](#)

Defines

- #define [SM_SKYEMCOMP_H](#)

- #define [SM_FILENAMESLIST](#) "sm_filenames.dat"
- #define [SM_RRSTEP](#) 1000
- #define [SM_LAM_UNIT](#) 1e-6
- #define [SM_RADMINLAM](#) 1.3 * 1e-6 / SM_LAM_UNIT
- #define [SM_ERAD](#) 6371.
- #define [SM_SIGMAX](#) 4.
- #define [SM_NSIGBIN](#) 20

Typedefs

- typedef struct [_smparmodel](#) [smparmodel](#)

Functions

- cpl_error_code [sm_etc_calcmmodel](#) (cpl_table *skytable, const cpl_parameterlist *params)
- cpl_error_code [sm_etc_getparams](#) (smparmodel *modelpar, const cpl_parameterlist *params)
- const cpl_parameter * [sm_etc_parameterlist_find_const](#) (const cpl_parameterlist *self, const char *name, cpl_type type)
- cpl_error_code [sm_etc_splitstring](#) (cpl_array *outval, char *instring)
- cpl_error_code [sm_etc_readfilenames](#) (smparmodel *modelpar, const cpl_parameterlist *params)
- cpl_error_code [sm_etc_writetable](#) (cpl_table *skytable, const [smspec](#) *radiance, const [smspec](#) *transmission)
- cpl_error_code [sm_comp_skyemcomp](#) ([smspec](#) *radiance, [smspec](#) *transmission, const [smparmodel](#) modelpar)
- cpl_error_code [sm_comp_extrapoltrans](#) ([smspec](#) *spec, const [smparmodel](#) modelpar, const double lim[2])
- double [sm_comp_altoairmass](#) (const double alt)
- cpl_error_code [sm_comp_getmolspec](#) ([smspec](#) *spec, const cpl_parameterlist *libfilepar, const double lim[2])
- cpl_error_code [sm_comp_readlibstruct](#) (cpl_parameterlist *libfilepar, const [smparmodel](#) modelpar, const char *spectype)
- double [sm_comp_vactoir_single](#) (const double lam)
- cpl_error_code [sm_comp_vactoir_spec](#) ([smspec](#) *spec)
- double [sm_comp_airtovac_single](#) (const double lam)
- cpl_error_code [sm_comp_airtovac_spec](#) ([smspec](#) *spec)
- double [sm_comp_calcrayleighscat1](#) (const double lam)
- cpl_error_code [sm_comp_calcrayleighscat](#) ([smspec](#) *rscattrans)
- double [sm_comp_calcmiescat1](#) (const double lam)
- cpl_error_code [sm_comp_calcmiescat](#) ([smspec](#) *mscattrans)
- cpl_error_code [sm_comp_scaletranscurv](#) ([smspec](#) *trans, const [smparmodel](#) modelpar)
- cpl_error_code [sm_comp_lunskybright](#) ([smspec](#) *spec, const [smparmodel](#) modelpar, const [smspec](#) *solspec, const [smspec](#) *molabstrans, const [smspec](#) *rscattrans, const [smspec](#) *mscattrans)

- `cpl_error_code sm_comp_lunskybright_v` (`smspec *spec`, `const smparmodel modelpar`, `const smspec *solspec`, `const smspec *molabstrans`, `const smspec *rscattrans`, `const smspec *mscattrans`)
- `cpl_error_code sm_comp_getmoonabedo` (`smspec *albedo`, `const smparmodel modelpar`)
- `cpl_error_code sm_comp_scstarlight` (`smspec *spec`, `const smparmodel modelpar`, `const smspec *molabstrans`)
- `cpl_error_code sm_comp_zodskybright` (`smspec *spec`, `const smparmodel modelpar`, `const smspec *solspec`, `const smspec *molabstrans`, `const smspec *rscattrans`, `const smspec *mscattrans`)
- `cpl_error_code sm_comp_telem` (`smspec *spec`, `const smparmodel modelpar`)
- `cpl_error_code sm_comp_extrapolrad` (`smspec *spec`, `const smspec *trans`, `const smparmodel modelpar`, `const double lim[2]`)
- `cpl_error_code sm_comp_getlinespec` (`smspec *spec`, `const smparmodel modelpar`, `const smspec *rscattrans`, `const smspec *mscattrans`)
- `cpl_error_code sm_comp_readvarpar` (`cpl_table *varpar`, `const smparmodel modelpar`)
- `cpl_error_code sm_comp_scalelinetab` (`cpl_table *linedat`, `const smspec *linetab`, `const cpl_table *varpar`)
- `cpl_error_code sm_comp_convertlinetab` (`smspec *spec`, `const cpl_table *linedat`)
- `cpl_error_code sm_comp_convertlinetabo` (`smspec *spec`, `const cpl_table *linedat`)
- `cpl_error_code sm_comp_airglowcont` (`smspec *spec`, `const smparmodel modelpar`, `const smspec *molabstrans`, `const smspec *rscattrans`, `const smspec *mscattrans`)

6.6.1 Detailed Description

Header for routines concerning the ETC sky model

Author

Stefan Noll & ESO In-Kind Team Innsbruck

Since

02 Oct 2009

Date

31 Oct 2013