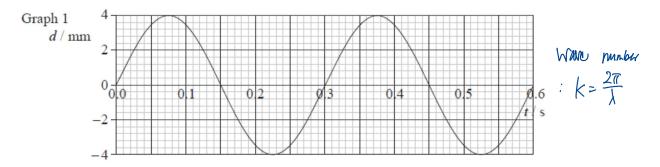
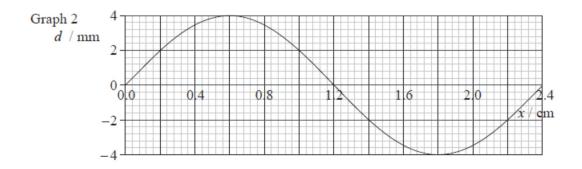
Wave Motion Tutorial

Question 1

In the graphs shown below, Graph 1 shows the variation with time t of the displacement d of a travelling (progressive) wave. Graph 2 shows the variation with distance x along the **same** wave of its displacement d.





- (i) State what is meant by a travelling wave.
- (ii) Determine the amplitude, wavelength, frequency and speed of this wave from the two graphs (Graph 1 & Graph 2) provided. [4]

[1]

Amplitude = 4mm
frequency =
$$0.3$$
 Hz
 $\lambda = 2.4$ Cm
 $\lambda = 0.024$ x 0.3
= 0.08 m/s

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Ouestion 2

A radio telescope is an instrument consisting of a large parabolic dish aerial to collect and focus radio waves to a radio receiver that is capable of very high amplification. Such a system looks similar to those used at any satellite ground station.

The radio telescopes are used to detect radio-frequency radiation that are emitted by extraterrestrial sources such as distant galaxies. This is because such sources emit enormous amount of radio waves.

One such galaxy is Cygnus A, which is a very distant galaxy that is too faint to be seen by its visible emission. Its radio emission is enormous, making it the second brightest object in the $I = \frac{F_{\pm}}{4\pi \rho^2}$ sky to the radio eye.

Total power emitted by Cygnus A,

$$P_{\text{total}} = 10^{37} \text{ W}$$

Total power received by the Earth,

$$P_{\text{Earth}} = 10 \text{ W}.$$

Diameter of the dish aerial at Jodrell Bank, d = 80 m

Diameter of the Earth,

$$D = 1.3 \text{ x} 10^7 \text{ m}$$

(i) What do you understand from the statement 'the second brightest object to the radio eye'?

(ii) Estimate the intensity of the wave received from Cygnus A on Earth.

I = \frac{10}{Cross \text{ sections} \text{ Max}} = \frac{10}{\pi(0.65\text{k})0^7\text{l}^2} = 7.5\text{ x } \text{ 10}^{-1/4}

(iii) Estimate the distance to Cygnus A from the Earth.

$$I = \frac{Power received by earth}{(ross sertimal area} = \frac{10}{T(0.65x|0^7)^2} = 7.54 \times 10^{-14}$$

$$I = \frac{P_6}{1 - R_0^2} \Rightarrow R = \frac{P_6}{1 - R_0^2} = 3.25 \times 10^{24} \text{ m}$$

I = $\frac{P_S}{4\pi R^2}$ $\Rightarrow R = \int \frac{R}{4\pi T} = 3.25 \times [0^{24} \text{ m}]$ (iv) Show that the magnitude of the power from Cygnus A collected by the radio telescope's

aerial is approximately of the order of 10⁻¹⁰ W.

$$P_{\text{tole}} = I \times \pi \left(\frac{J}{2}\right)^{2}$$

$$= (7.54 \times 10^{-14}) \times \pi \times \left(\frac{80}{2}\right)^{2}$$

$$= 3.8 \times 10^{-10} \text{ W}$$

Power Received by telescope

Question 3

State the characteristic features of

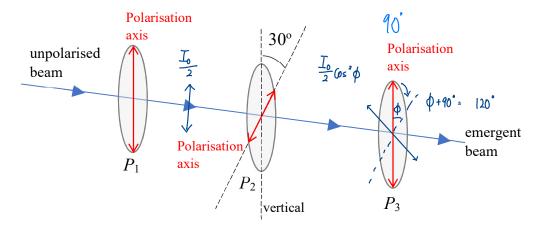
- longitudinal wave, and (i)
- (ii) transverse wave.

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Polarisation Can be any direction

Question 4

A beam of unpolarised light passes through Polaroids P_1 , P_2 and P_3 as shown below.

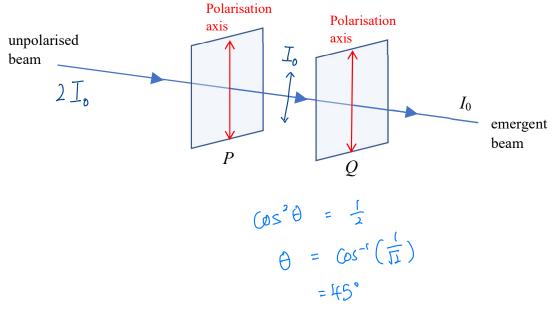


The polarisation axes of both Polaroid P_1 and P_3 are vertical but that of Polaroid P_2 is inclined at an angle of 30° to the vertical. In order to obtain the minimum intensity for the emergent beam, through what angle must Polaroid P_3 be rotated in the clockwise direction?

120'

Question 5

Two pieces of Polaroids, P and Q, are arranged such that their polarisation axes are parallel and vertical, as shown below. The intensity of the emergent light is I_0 . Through what angle must Q be rotated so that the intensity of the emergent light decreases to $\frac{1}{2}I_0$?



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Answers:

Question 1

(ii) amplitude =
$$4.0$$
mm
wavelength = 2.4 cm
frequency = $1/0.3 = 3.3$ Hz
speed = $3.3 \times 2.4 = 8.0$ cms⁻¹

Question 2

Question 4

Angle that P_3 must be rotated in clockwise direction = 120 °.

Question 5

Angle that Q must be rotated = 45 °.