From Galaxies to Stars: Bridging the Expanse

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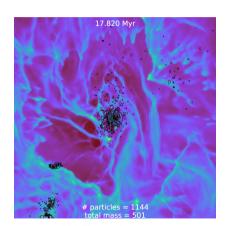
Clouds in Reality and... Not

What's the difference?



NASA; Carina Nebula

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Wall et al. 2020

The Big Problems

Resolution & Initial Conditions

Self consistent galactic scale simulations with resolution down to sub-tenth parsec scales and include Nbody individual stellar dynamics and individual stellar feedback all at once? A little tough.



The Big Problems

Resolution & Initial Conditions

- Self consistent galactic scale simulations with resolution down to sub-tenth parsec scales and include Nbody individual stellar dynamics and individual stellar feedback all at once?
 A little tough.
- Creating our own isolated clouds from scratch? "Creative liberties..."

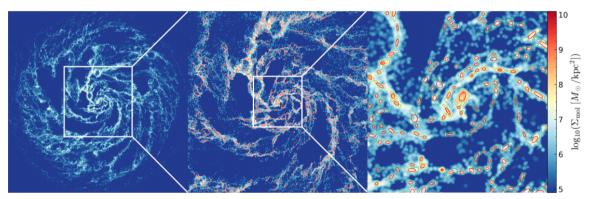


Stars From "Realistic" Clouds

- Clouds that formed under the influence of galactic dynamics.
- Track dynamics and feedback of individual stars.
- High resolution to quantify star-gas interactions.



Clouds from Galactic Simulations

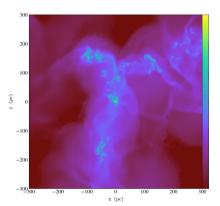


GMC identification [Li, H. et al. 2020]

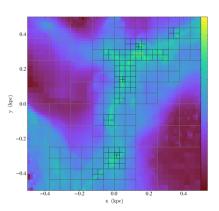


From AREPO to FLASH

(1st attempt)

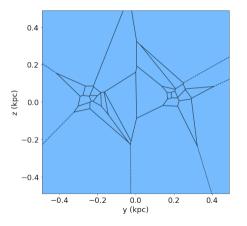


Cloud from raw AREPO data



Cloud-in-cell mapping onto AMR FLASH grid_

Voronoi Mesh to AMR Grid



0.4 0.2 z (kpc) 0.0 -0.2-0.4100 pc -0.4-0.20.0 0.2 0.4 y (kpc)

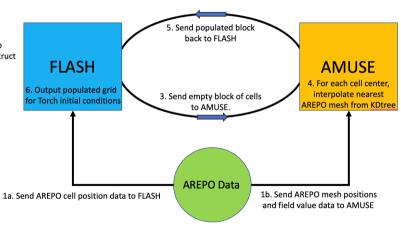
Figure: Voronoi mesh from 20 points

Figure: AMR grid from 20 points



VorAMR: Logic path

2a. Convert cells to particles and construct refined AMR grid



2b. Construct KDtree with field values assigned to leaf nodes