

Features of dual-purpose structure for heavy ion and light particles

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The different charge-to-mass ratio is essential in magneto-optics design. To achieve high luminosity sufficient beam lifetime must be guaranteed. Transition energy crossing must also be solved. The NICA collider will be used for both collider experiments with heavy ions and light polarized nuclei.

Keywords: Transition energy, Intrabeam scattering, Stochastic cooling, Resonant structure, Dual-purpose structure

I. INTRODUCTION

Regardless of the purpose of the synchrotron, always in the case of two modes, when multiply charged heavy particles and one or two charged light particles are accelerated, the problem arises of what the magneto-optical structure should look like to satisfy all the conditions of stable motion for both types of particles. Obviously, compared to light particles, multiply charged particles have a prevailing heating effect due to intra-beam scattering, and light particles have a greater chance of crossing through transition energy. All these effects are of great importance for colliders, where luminosity plays a decisive role. When developing a structure that meets all the requirements for differently charged particles, it is fundamentally important to have a retunable structure without introducing design differences. We called such a structure – dual-purpose or simple, dual.

In NICA complex the dual magneto-optical structure opens up the prospect of accelerating both heavy ions, such as gold, and light particles like protons and deuterons. The design of this structure requires a different approach due to the varying charge-to-mass ratios involved.

II. LIGHT PARTICLES

In a classical regular structure, the transition energy is approximately equal to the betatron tune $\gamma_{tr} \simeq \nu_s$ [?]. For the same magnetic rigidity $B\rho$, the maximum energy for light particles is greater than for heavy ions due to their charge-to-mass ratio. This means that a heavy-ion structure optimized for operating up to a certain transition energy would require overcoming that energy in order to operate with light particles. For this reason, a structure with varying transition energy can be considered.

A. Transition energy

In general, the transition energy is determined by the momentum compaction factor

$$\alpha = \frac{1}{\gamma_{tr}^2} = \frac{1}{C} \int_0^C \frac{D(s)}{\rho(s)} ds \quad (1)$$

where C – orbit length, $D(s)$ – dispersion function, $\rho(s)$ – radius of orbit curvature. It is a characteristic of the structure and remains constant regardless of the particle type. In the first order the slip-factor $\eta = \eta_0 = 1/\gamma_{tr}^2 - 1/\gamma^2$, and thus the frequency of synchrotron oscillations $\omega_s \sim \eta$ tends to zero when the beam energy approaches the transition value. In this case, the adiabaticity of the longitudinal phase motion is violated, which leads to instabilities, as well as the influence of nonlinear effect higher orders of momentum spread δ . The introduction of modulation into the $D(s)$ or $\rho(s)$, function leads to variations in the momentum compaction factor and, consequently, the transition energy.

B. Superperiodic modulation

The equation for the dispersion function with biperiodic variable focusing [?]

$$\frac{d^2 D}{ds^2} + [K(s) + \varepsilon k(s)] D = \frac{1}{\rho(s)}, \quad (2)$$

where $K(s) = \frac{e}{p} G(s)$, $\varepsilon k(s) = \frac{e}{p} \Delta G(s)$, $G(s)$ – gradient of magneto-optical lenses, $\Delta G(s)$ – superperiodic gradient modulation. Here is considered an additional perturbation to regular one $\varepsilon k(s) = \sum_{k=0}^{\infty} g_k \cos(k\phi)$, where g_k – k -th harmonic of the gradient modulation in the Fourier series expansion of the function. The solution for momentum compaction factor as follows for only gradient modulation

$$\alpha_s = \frac{1}{\nu^2} \left\{ 1 + \frac{1}{4(1 - kS/\nu)} \left(\frac{\bar{R}}{\nu} \right)^4 \frac{g_k^2}{[1 - (1 - kS/\nu)^2]^2} \right\}. \quad (3)$$

where \bar{R}_{arc} – the average value of the curvature, ν_x – betatron tune in horizontal plane on arc, S – number of superperiods per arc length. Eq. (??) considered without introducing curvature modulation due to the possibility of introducing a variation of transition energy into a stationary lattice. To raise the transition energy, it is necessary to reduce $\alpha_s = 1/\gamma_{tr}^{arc}$, this means that the expression under the sum sign must be negative, this is realizable under the condition $kS/\nu_{x,arc} > 1$.

First harmonic $k = 1$ has a dominant influence the condition is implemented for $S = 4$, $\nu_{x,arc} = 3$. Fig. ?? shows 12

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70 FODO cells per arc, 3 FODO cells are combined into one su-
 71 perperiod [?]. Thus, due to the tune of betatron tune a multi-
 72 ple of 2π , the arc has the properties of a first-order achromat.
 73 For structure with the missing magnet technique introduced
 74 for one reason or another it can be also implemented (Fig.
 75 ??), but need to suppress dispersion at the edge of arc [?].

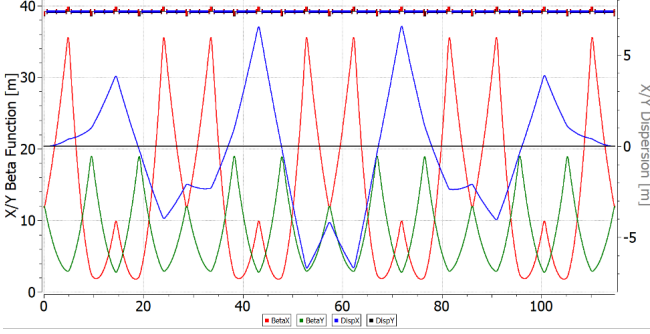


Fig. 1. (Color online) "Resonant" magneto-optic structure with dispersion modulation and increased transition energy.

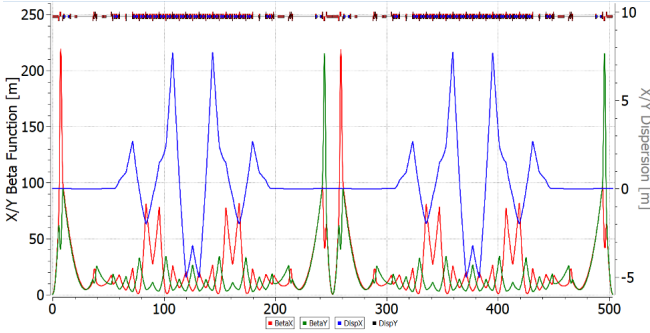


Fig. 2. (Color online) "Resonant" NICA magneto-optic adapted structure with increased transition energy and missing magnet.

III. HEAVY ION MODE

81 The lifetime of the beam luminosity in a collider exper-
 82 iment is achieved through the reduction of intra-beam scat-
 83 tering effects, coupled with the application of stochastic and
 84 electron beam cooling techniques. This approach assumes
 85 particular significance when dealing with high-intensity ion
 86 beams. The temporal evolution of emittance and momentum
 87 spread in the presence of cooling processes is governed by a
 88 set of equations

$$\begin{aligned} \frac{d\varepsilon}{dt} &= \underbrace{-\frac{1}{\tau_{tr}} \cdot \varepsilon}_{\text{cooling}} + \underbrace{\left(\frac{d\varepsilon}{dt}\right)_{\text{IBS}}}_{\text{heating}} \\ \frac{d\delta^2}{dt} &= \underbrace{-\frac{1}{\tau_{long}} \cdot \delta^2}_{\text{cooling}} + \underbrace{\left(\frac{d\delta^2}{dt}\right)_{\text{IBS}}}_{\text{heating}} \end{aligned} \quad (4)$$

90 where ε – transverse emittance, τ_{tr} – transverse cooling time,
 91 $\delta = \frac{\Delta p}{p}$ – momentum spread, τ_{long} – longitudinal cool-
 92 ing time. For time-independent, stationary values, the time
 93 derivatives become zero, then

$$\begin{aligned} \varepsilon_{st} &= \tau_{tr} \cdot \left(\frac{d\varepsilon}{dt}\right)_{\text{IBS}} \Big|_{\varepsilon=\varepsilon_{st}} \\ \delta_{st}^2 &= \tau_{long} \cdot \left(\frac{d\delta^2}{dt}\right)_{\text{IBS}} \Big|_{\delta^2=\delta_{st}^2} \end{aligned} \quad (5)$$

95 The benchmark for evaluating the effectiveness of a cooling
 96 technique can be determined by comparing the timescales of
 97 stochastic or electron cooling processes to the beam lifetime
 98 due to IBS over the entire energy spectrum.

A. Stochastic cooling

100 Let's consider stochastic cooling using the approximate
 101 theory developed by D.Mohl [? ?]. Based on his main find-
 102 ings, the cooling rate can be determined using the following
 103 expression

$$\frac{1}{\tau_{tr,1}} = \frac{W}{N} \left[\underbrace{2g \cos \theta \left(1 - 1/M_{pk}^2\right)}_{\text{coherent effect(cooling)}} - \underbrace{g^2 (M_{kp} + U)}_{\text{incoherent effect(heating)}} \right] \quad (6)$$

105 where $W = f_{\max} - f_{\min}$ – system bandwidth, N – effec-
 106 tive number of particles, recalculated based on the ratio of
 107 orbit length to the beam length, considering its distribution,
 108 g – fraction of observed sample error corrected per turn, U
 109 – the ratio of noise to signal, M_{pk} , M_{kp} – mixing factors be-
 110 tween the pickup-kicker and the kicker-pickup, respectively.
 111 Eq. (??) in the absence of noise at $g = g_0 = \frac{1-M_{pk}^2}{M_{kp}}$ reaches
 112 the maximum

$$\begin{aligned} \frac{1}{\tau_{tr}} &= \frac{W}{N} \frac{\left(1 - 1/M_{pk}^2\right)^2}{M_{kp}} \\ \frac{1}{\tau_l} &= 2 \frac{W}{N} \frac{\left(1 - 1/M_{pk}^2\right)^2}{M_{kp}} \end{aligned} \quad (7)$$

114 The mixing coefficients are defined as

$$\begin{aligned} M_{pk} &= \frac{1}{2(f_{\max} + f_{\min}) \eta_{pk} T_{pk} \frac{\Delta p}{p}}, \\ M_{kp} &= \frac{1}{2(f_{\max} - f_{\min}) \eta_{kp} T_{kp} \frac{\Delta p}{p}} \end{aligned} \quad (8)$$

116 where $\eta_{pk} T_{pk} \delta$, $\eta_{kp} T_{kp} \delta$ – relative particle displacement times
 117 (mixing), η_{pk} , η_{kp} – slip-factor, as a first approximation $\eta_{pk} =$
 118 $\alpha_{pk} - 1/\gamma^2$, $\eta_{kp} = \alpha_{kp} - 1/\gamma^2$, α_{pk} , α_{kp} – first-order of lo-
 119 cal momentum compaction factors, T_{pk} , T_{kp} – the absolute
 120 times between the pickup-kicker and kicker-pickup, respec-
 121 tively. The stochastic cooling times of Eq. (??) depend on

the ratio of the effective particle density to the cooling system bandwidth and the properties of magneto-optics, local momentum compaction factors α_{pk} , α_{kp} .

The maximum value of the frequency band is determined by the requirement that the "Schottky" beam bands do not overlap. In the simplest case, this can be expressed:

$$f_{\max} < \frac{1}{\eta_{pk} T_{pk} \frac{\Delta p}{p}} \quad (9)$$

thus, a mixing factor $M_{pk} > 1$. Otherwise, the cooling efficiency becomes zero. Thus, for a given number of particles, it is desirable to achieve the highest possible frequency band. From an electron perspective, modern technologies allow for the implementation of a 10 GHz frequency band [?], however, its use is not always feasible due to the large magnitude of the slip-factor η_{pk} and momentum spread δ .

Eq. (??) has been derived for coasting beam. Particle density for a single harmonic RF resonator is described by a Gaussian distribution

$$\rho(s) = \frac{N_{\text{bunch}}}{\sigma_{\text{bunch}} \sqrt{2\pi}} \cdot e^{-\frac{s^2}{2\sigma_{\text{bunch}}^2}} \quad (10)$$

where s – the distance from the beam center, σ_{bunch} – the dispersion of the particle distribution, and N_{bunch} – the number of particles. Assuming that cooling is at its minimum at the center ($s = 0$), the effective particle number at orbit length C_{orb} can be calculated as follows:

$$N = \frac{N_{\text{bunch}}}{4\sigma_{\text{bunch}}} \cdot C_{\text{orb}} \quad (11)$$

For a beam generated by a multiharmonic barrier-type RF system, so-called "Barrier Bucket", the particles distribution in the beam can be considered approximately uniform along its entire length. The effective particles number is determined by a simple ratio of the beam length to the total orbit length: To summarize, the effective value of particles depends on their distribution and is determined by their form-factor $F_{\text{bunch}} = \sqrt{2\pi} \div 4$

$$N = N_{\text{bunch}} \cdot \frac{C_{\text{orb}}}{F_{\text{bunch}} \cdot \sigma_{\text{bunch}}} \quad (12)$$

As example let us consider the case of NICA with maximal form-factor $F_{\text{bunch}} = 4$ with $C_{\text{orb}} = 503.04$ m, $\sigma_{\text{bunch}} = 0.6$ m, $N_{\text{bunch}} = 2.2 \cdot 10^9$. Considering the accumulated FNAL [?] experience, quite realistic values for the frequency band are $f_{\max} = 8$ GHz and $f_{\min} = 2$ GHz. For NICA $f_{\max} = 4$ GHz and $f_{\min} = 2$ GHz. With these parameters, the maximum achievable cooling rate is $1/\tau_{\text{tr}} = 1/230 \text{ s}^{-1}$.

Based on Eq. (??), it is evident that asymptotic growth may occur in two scenarios:

1. slip-factor approaches the value $\eta \rightarrow \frac{1}{2(f_{\max} + f_{\min})T_{pk}\delta}$, the beam Schottky spectrum becomes continuous and $M_{pk} \rightarrow 1$;

2. slip-factor approaches zero, mixing between the kicker to the pickup does not occur and $M_{kp} \rightarrow \infty$.

The efficiency of stochastic cooling depends on the properties of magneto-optical structure. In classical "regular" structures, transition energy is acquired through the horizontal frequency $\gamma_{\text{tr}} \approx \nu_x$ and slip-factor $\eta = 1/\gamma_{\text{tr}}^2 - 1/\gamma^2$ can achieve zero. To avoid asymptotic growth, it is necessary to vary the slip-factor which means γ_{tr} . This is possible in "resonant" structure, where transition energy can be increased or even reach complex value. In more exotic case, can be used "combined" structure then η_{pk} (pickup-kicker) with real transition energy at one arc

$$\eta_{pk} = 1/\gamma_{\text{tr}}^2 - 1/\gamma^2 \quad (13)$$

compensated by η_{kp} (kicker-pickup) with complex transition energy at another

$$\eta_{kp} = -1/\gamma_{\text{tr}}^2 - 1/\gamma^2 \quad (14)$$

for the whole ring. Such structure achieves the required ratio of mixing factors for a maximum cooling rate close to ideal [?]. Let us delve deeper declared structures in more detail.

The behaviour of the β -functions and D the dispersion across the entire "regular" ring illustrates at Fig. ???. The straight sections, which remain constant in all structures, are essential for analyzing of the resonant characteristics of the entire structure. Their arrangement does not affect the intra-beam scattering and transition energy. To suppress dispersion in the "regular" structure, 'missing magnets' technic implemented on both sides of the arc. The "resonant" structure is based on the principle of resonant modulation of the dispersion function and can be obtained from a "regular" one by introducing additional family of focusing quadrupoles. To suppress dispersion can be used either two edge focusing quadrupoles on both sides of the arc or only two families of focusing quadrupoles on the arc, when an integer number of betatron oscillations is reached.

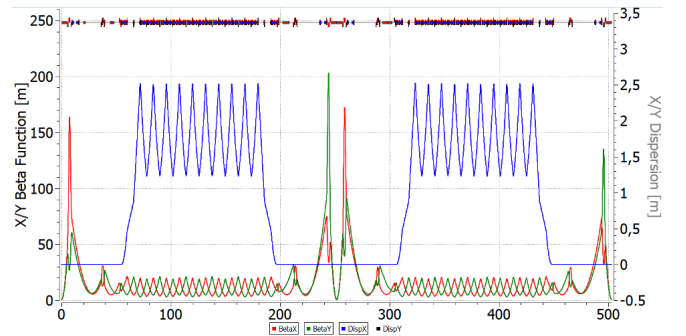


Fig. 3. (Color online) "Regular" FODO NICA magneto-optic structure with missing magnets.

The case of a "combined" structure, one arc operates in a regular mode, while the other employs resonant modulation (Fig. ??). Such choice is based on the principle of compensation,

as described by Eqs. ??, ??, which requires a greater modulation depth of the quadrupoles than in purely "resonant" structure with increased transition energy. As illustrated in

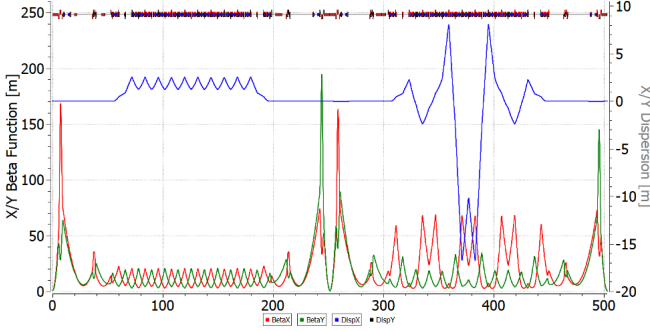


Fig. 4. (Color online) "Combined" NICA magneto-optic structure with real and complex transition energies in arcs.

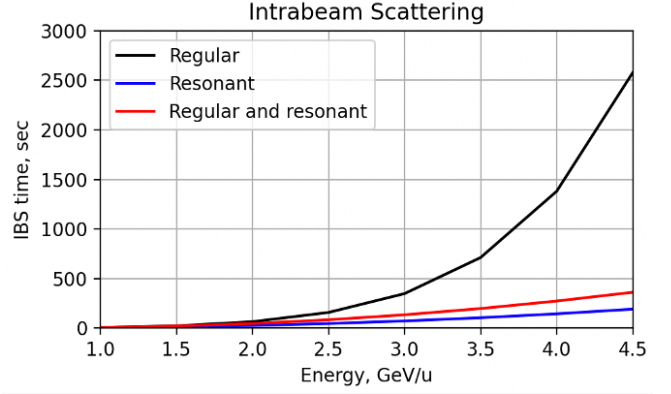


Fig. 6. (Color online) The dependence of the beam lifetime due to intra-beam scattering in "regular", "resonant" and "combined" structures on the beam energy for heavy ion beam.

Fig. ??, "resonant" optics with increased transition energy, the second asymptotic is at higher energy compared to the "regular" structure. In "combined" magneto-optics, the cooling efficiency is closer to the ideal value in a large energy range from 2.5 to 4.5 GeV, while in "regular" optics the cooling rate is almost two times lower at the most optimal point ~ 3 GeV. This behaviour is explained by absence of the second point of asymptotic growth.

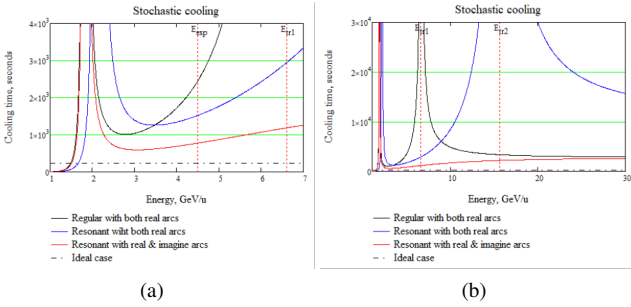


Fig. 5. (Color online) The dependence of stochastic cooling time on the energy for various structures. Energy range (a) 1 – 7, (b) 0-30 GeV per nuclon.

B. Intrabeam Scattering

Intra-beam scattering represents a fundamental limitation on the beam lifetime in the collider. Consequently, the selection of an appropriate cooling technique hinges on comparing its characteristic time scales with the rate at which the beam is heated due to intra-beam scattering. This is derived from the fundamental principles governing this process

$$\frac{1}{\tau_{\text{IBS}}} = \frac{\sqrt{\pi}}{4} \frac{cZ^2 r_p^2 L_C}{A} \cdot \frac{N}{C_{\text{orb}}} \cdot \frac{\langle \beta_x \rangle}{\beta^3 \gamma^3 \varepsilon_x^{5/2} \langle \sqrt{\beta_x} \rangle} \times \left(\left\langle \frac{D_x^2 + \dot{D}_x^2}{\beta_x^2} \right\rangle - \frac{1}{\gamma^2} \right) \quad (15)$$

Unlike stochastic cooling, the IBS rate increases as decreasing energy $1/\gamma^3$. In addition, the expression in parentheses is proportional to the slip-factor η . Therefore, it should be expected that in optics with a value η close to zero, the heating rate should decrease. Fig. ?? shows the dependences of the heating time constant in the three above-mentioned structures calculated using MADX programs [?] for the parameters of the heavy ion beam ^{197}Au of the NICA collider with maximum luminosity $10^{27} \text{ sm}^{-2} \text{ s}^{-1}$. In the context of light nuclei, such as protons and deuterons, the IBS time experiences a significant increase as the charge decreases. Consequently, the issue of intra-beam scattering becomes critical for heavy-ion beam. From the comparison of the IBS lifetime with the cooling time it can be concluded that in a regular structure, stochastic cooling is able to balance intra-beam scattering in the energy range $W \geq 4.5$ GeV. In order to apply stochastic cooling over the entire energy range, it is obvious that we must sacrifice the luminosity of the beam at low energies by increasing the emittance. In resonant structures, the IBS time is notably reduced. This is explained by the fact that the structure has a greater ratio $\left\langle \frac{D_x^2 + \dot{D}_x^2}{\beta_x^2} \right\rangle$ between the dispersion and the beam β -function than in the case of a regular. Thus, for the case of heavy ions, the configuration should be regular and minimally modulated. Electron cooling is used in the regular structure to cool the beam lower 4.5 GeV [? ?].

TABLE 1. Main parameters of structures.

Structure	Regular	Resonant	Combined
Energy, GeV per nuclon	4.5	12.6	12.6
Transition energy γ_{tr}	7	15	150
Modulation depth	–	25%	45%
Stochastic cooling at 4.5 GeV, s	2500	1500	800
IBS time at 4.5 GeV, s	2500	400	250

IV. SUMMARY

The dual magneto-optical structure is proposed for accelerating both heavy ion and light particle beams, exemplified by the NICA facility. For light particles, due to charge-to-mass ratio, experiment energy can rise more than transition energy of the lattice, which is optimal for heavy ions. Using dispersion modulation, transition energy increases or even reaches a complex value in a “resonant” magneto-optic structure. However, due to modulation of β -function and D dispersion, the

time of intra-beam scattering decreases, which is crucial for multiply charged heavy particles. For this reason, a “regular” magneto-optic structure with minimally modulated dispersion and β -function is optimal in the heavy-ion mode. Despite the fact that stochastic cooling in “regular” structures is significantly weaker than in “resonant” and “combined” ones, it can compensate IBS effect.

No special changes are required to convert the “regular” structure into a “resonant” one. It is enough only to introduce a separate family of quadrupoles

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