

*Mass, Scalar Curvature,
Kähler Geometry, and All That*

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Rutgers-Newark Colloquium
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Joint work with

Joint work with

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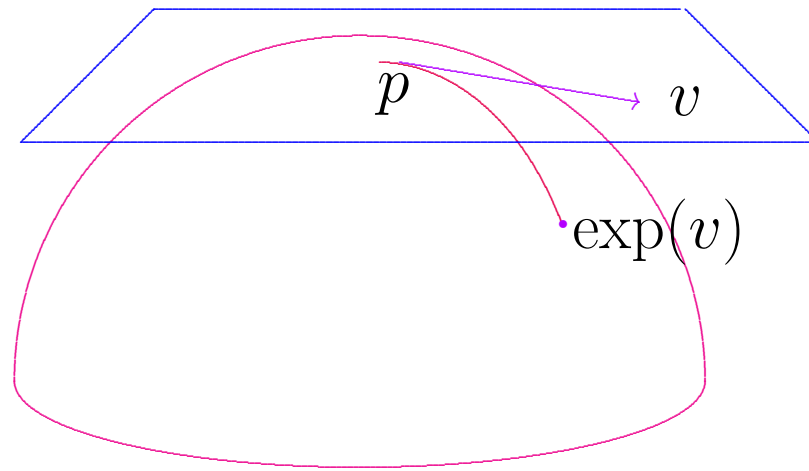
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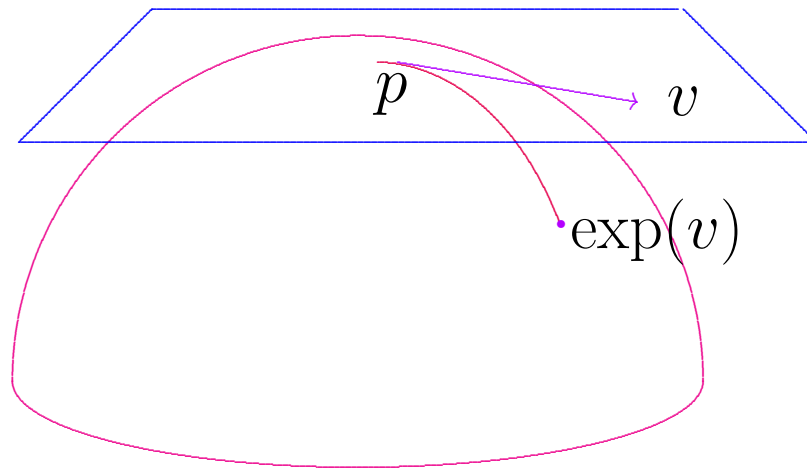
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which is a diffeomorphism on a neighborhood of 0:



Small ball B_ε maps to the ε distance ball in M :
points reachable from p by paths of length $< \varepsilon$.

The *scalar curvature*

$$s : M \rightarrow \mathbb{R}$$

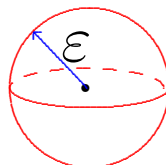
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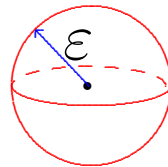


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A metric g is called *scalar-flat* if it satisfies $s \equiv 0$.

Similarly, the *Ricci curvature*

$$r : UTM \rightarrow \mathbb{R}$$

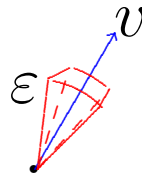
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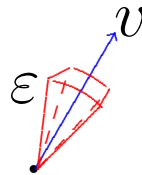


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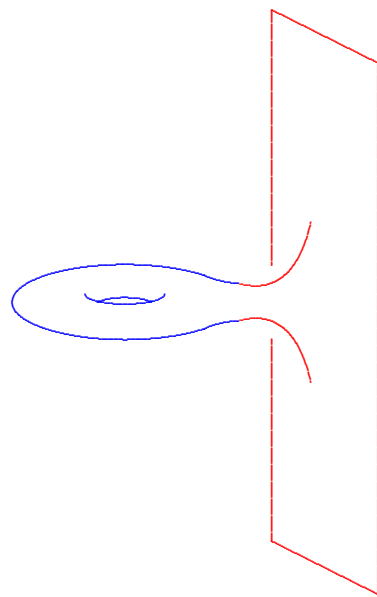


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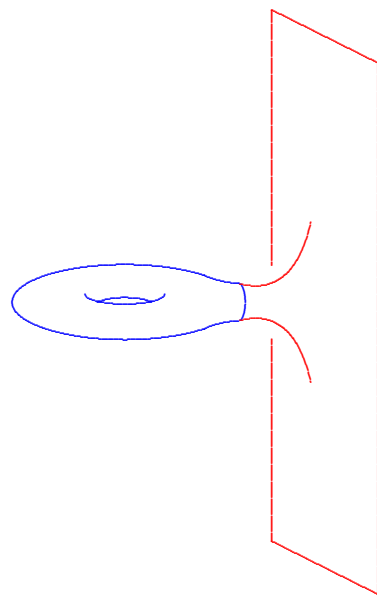
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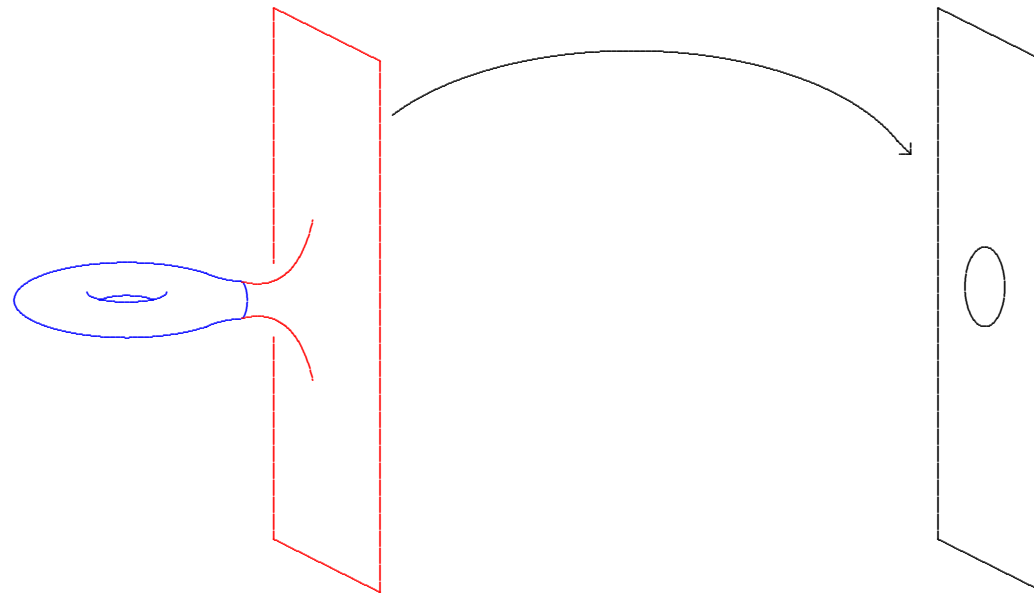
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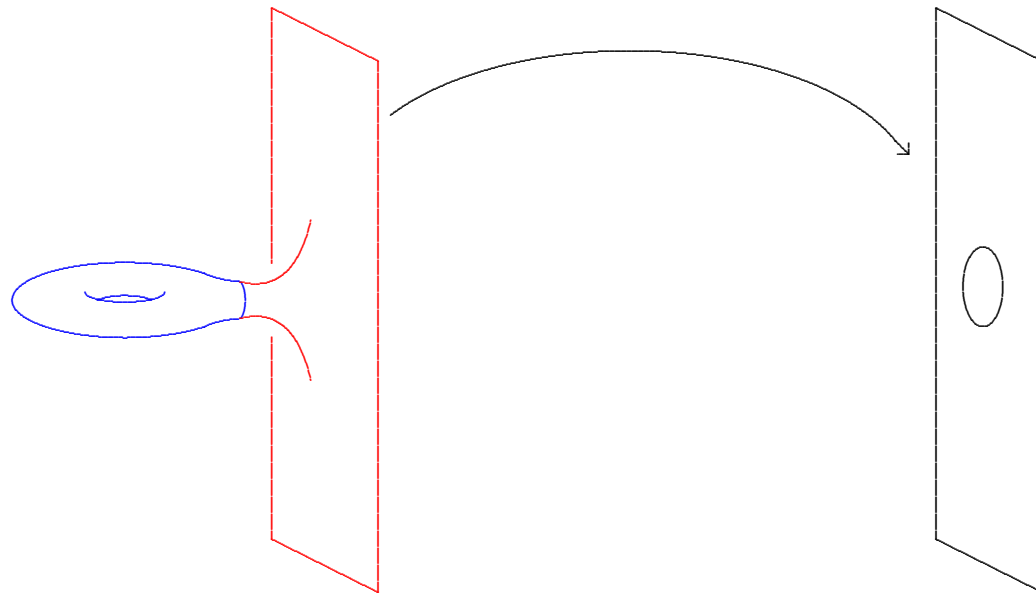


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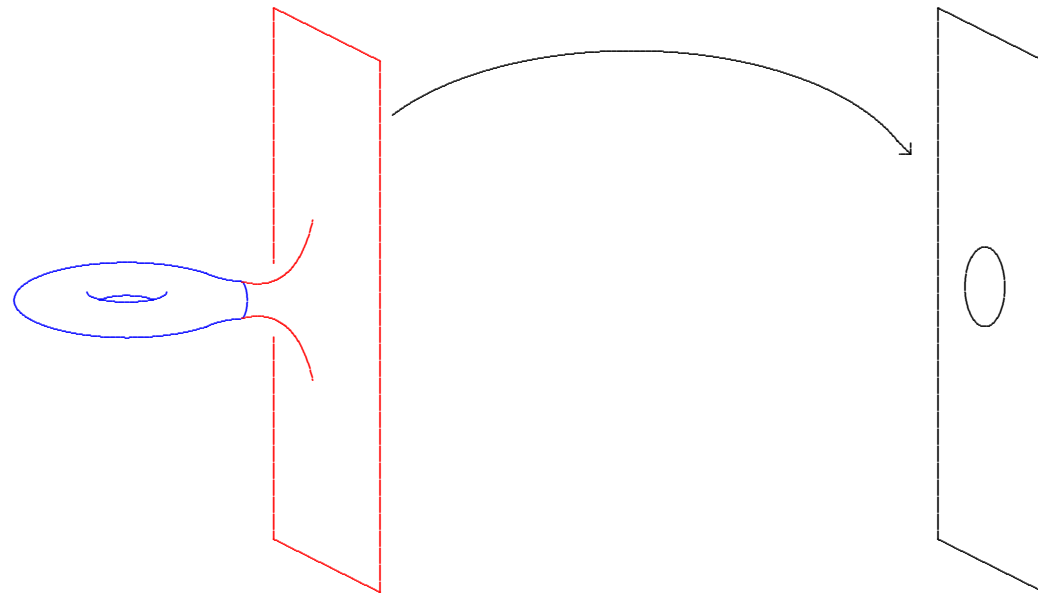


and an isometry $M - K \rightarrow \mathbb{R}^n - D^n$. (Euclidean)

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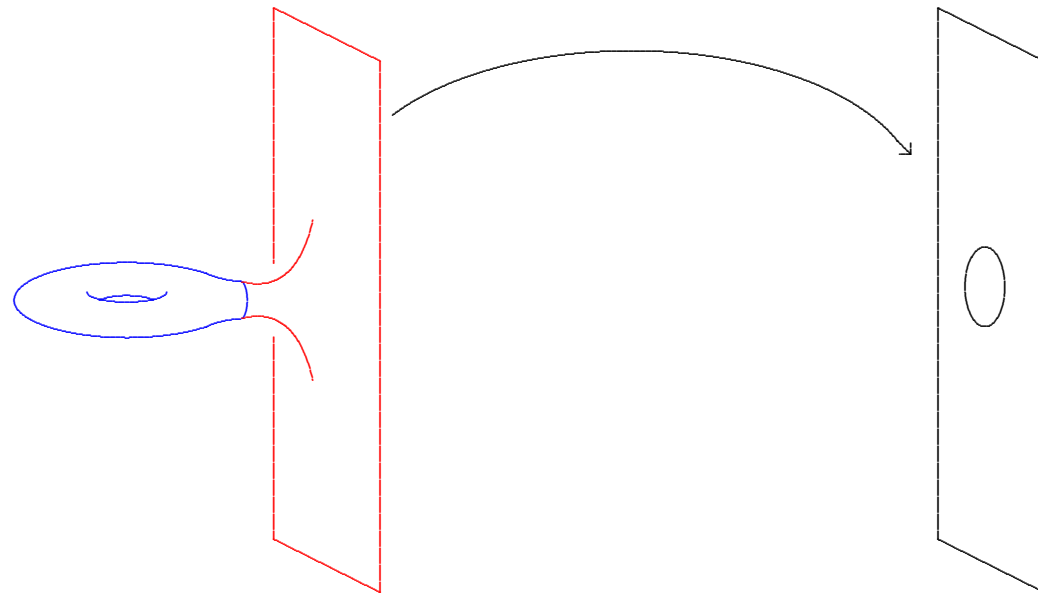


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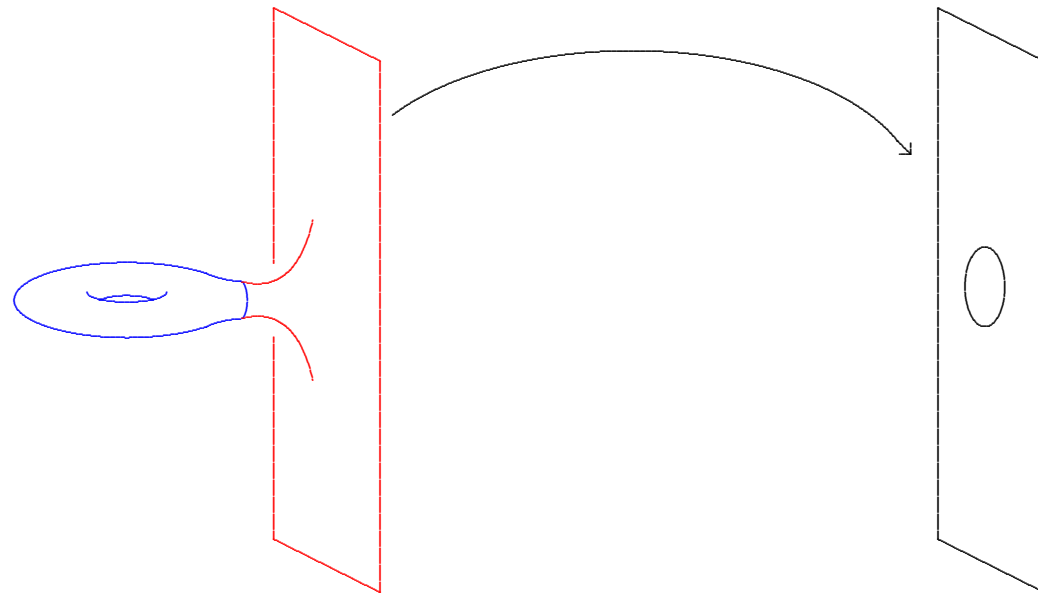
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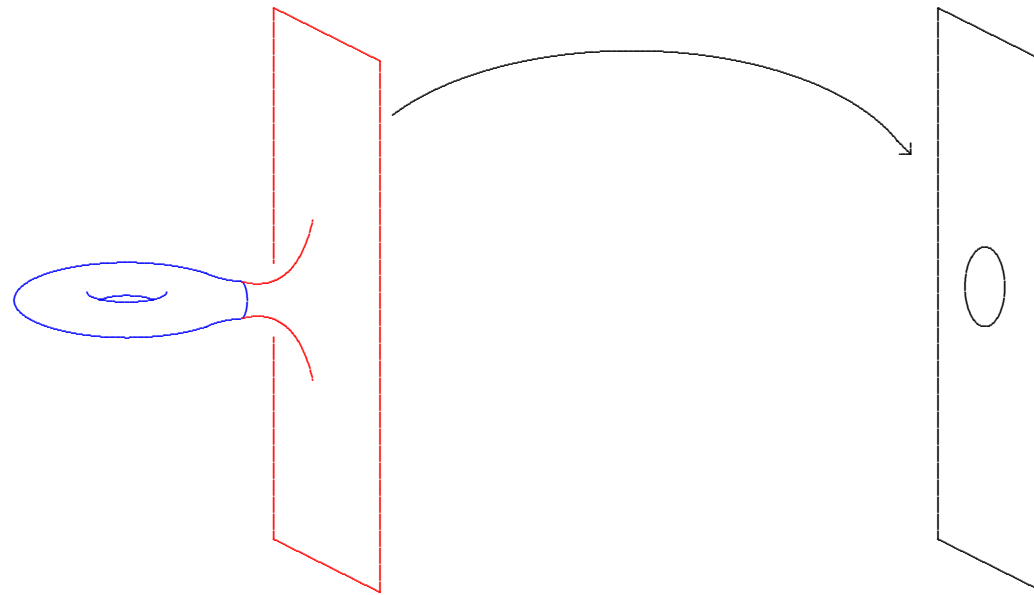
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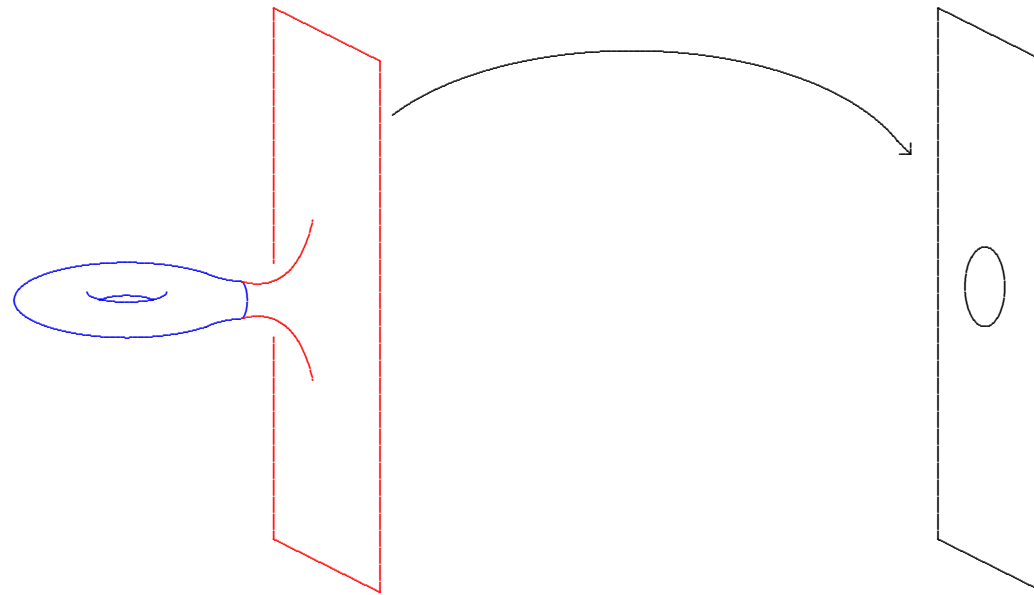
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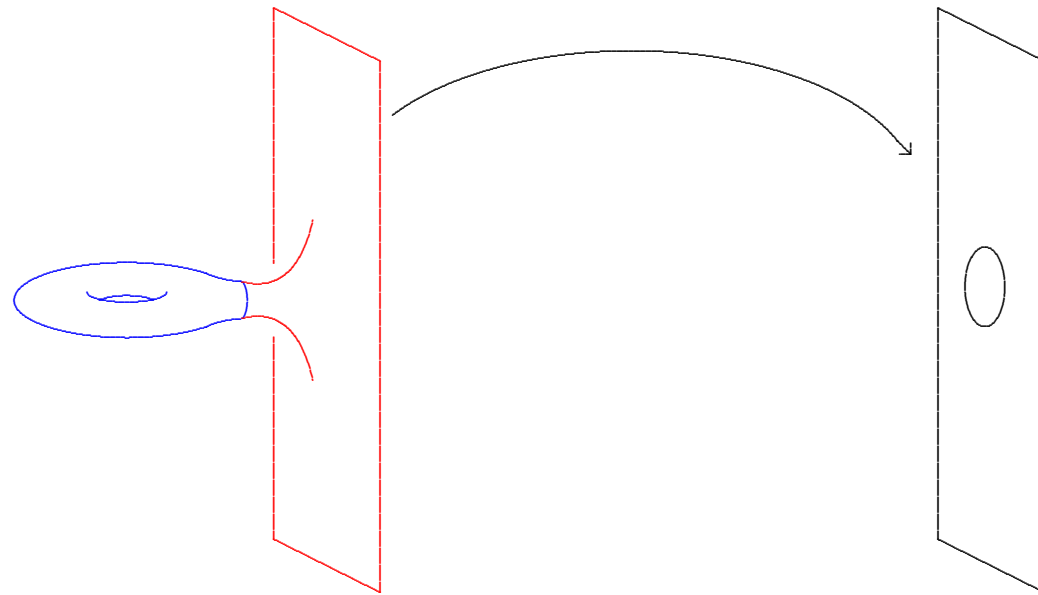
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Distance comparison...

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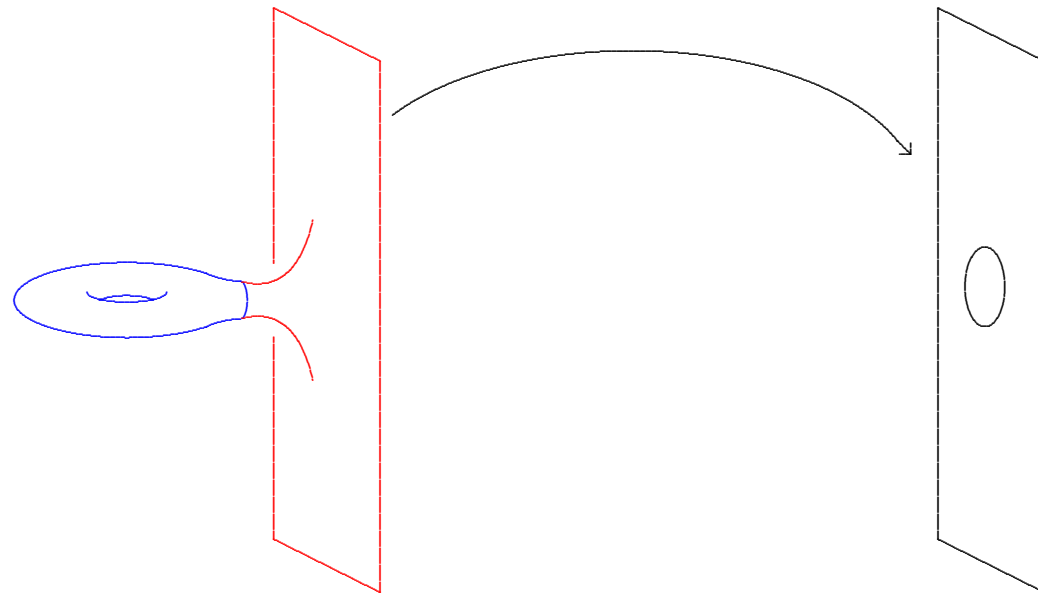
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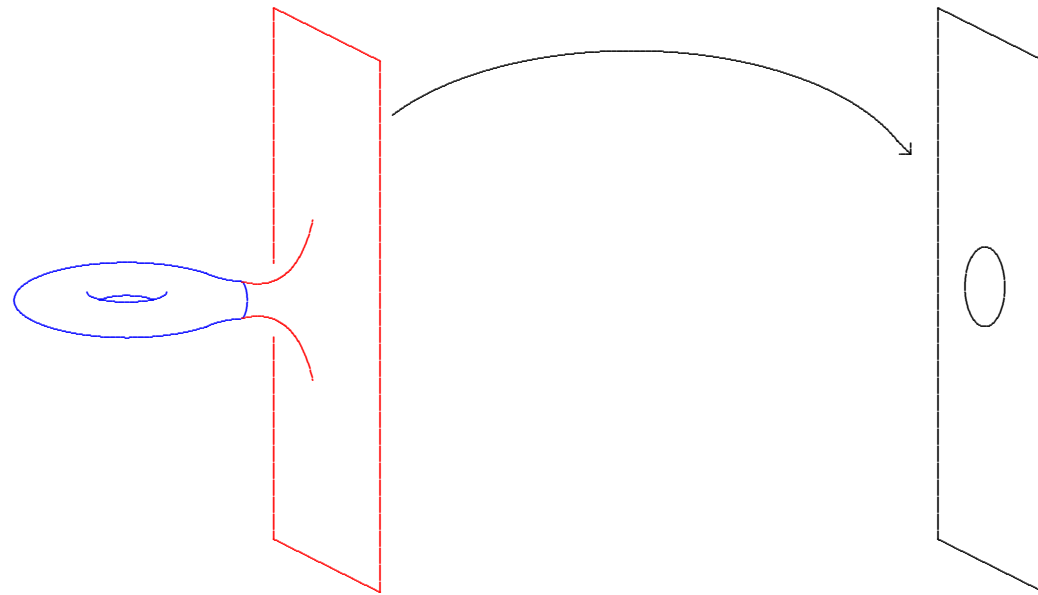
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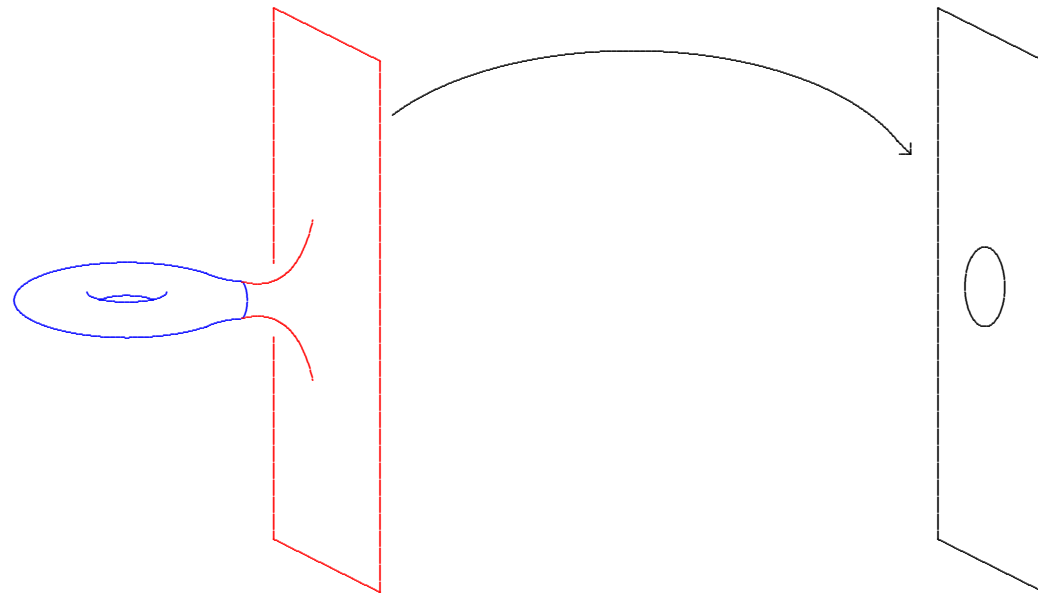
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Volume comparison...

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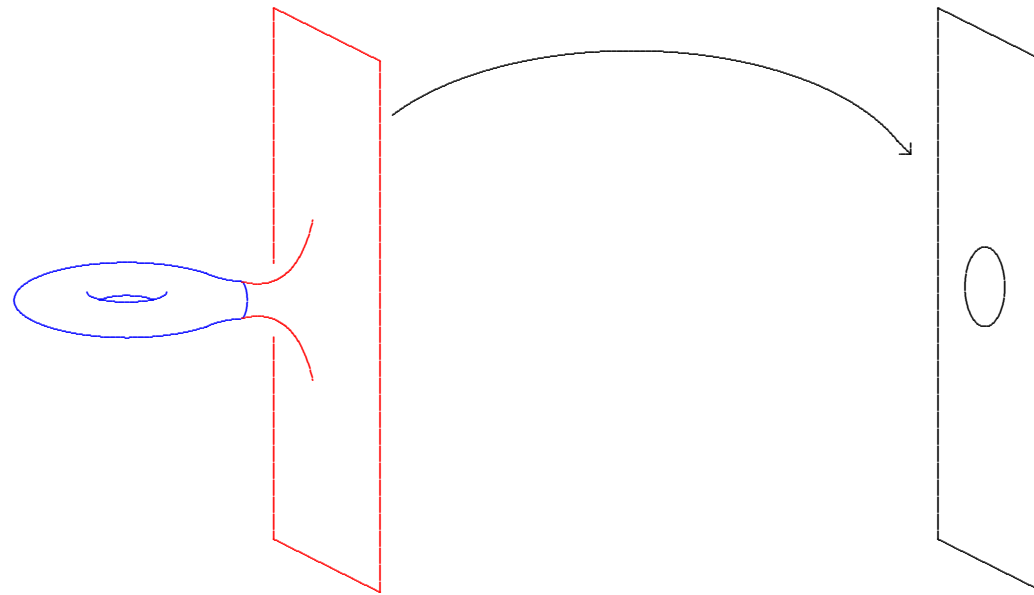
and an isometry $M \setminus K \rightarrow \mathbb{R}^n \setminus D^n$.

If M has scalar curvature ≥ 0 , is it flat?

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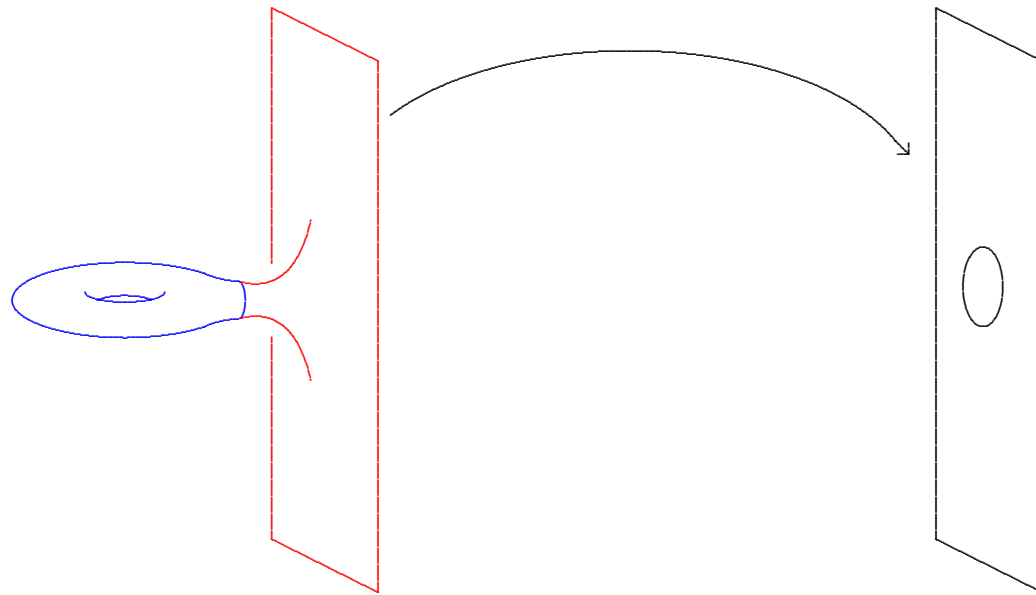
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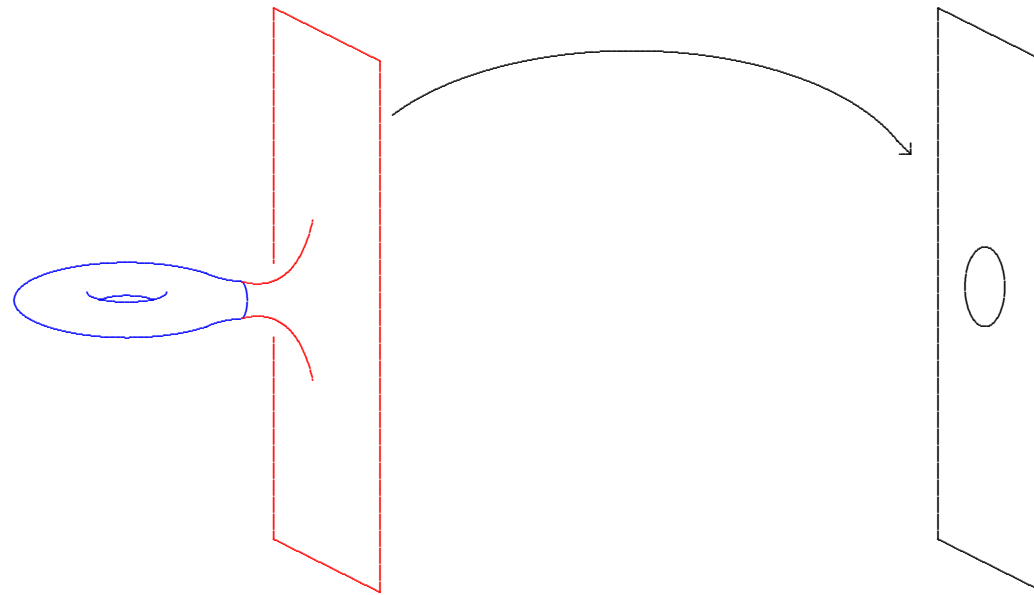
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at least if M spin, or if $n \leq 7$...

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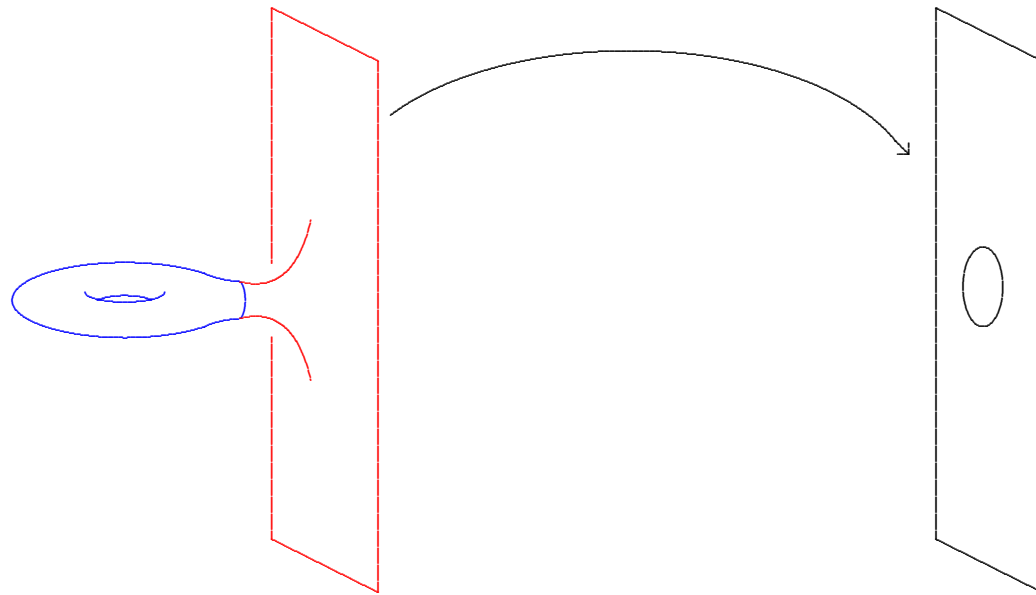
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Involves new idea from physics!

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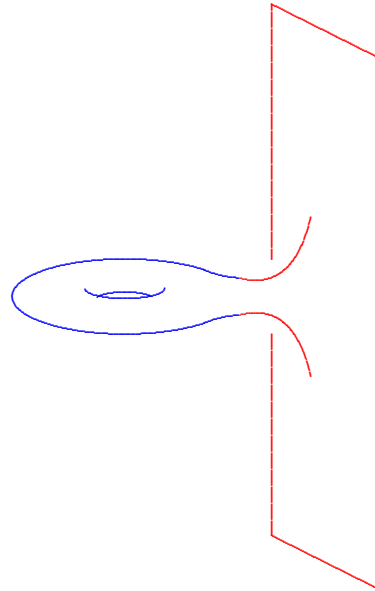


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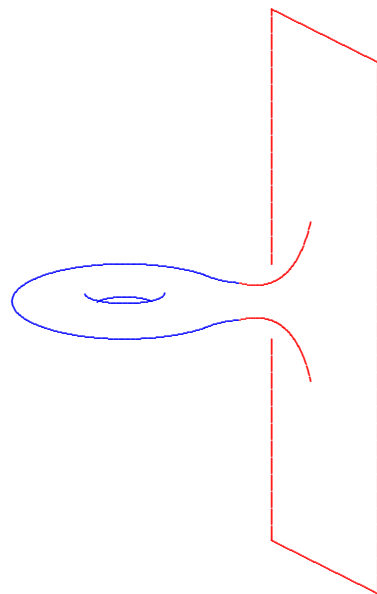
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Get result even with appropriate fall-off to Euclidean...

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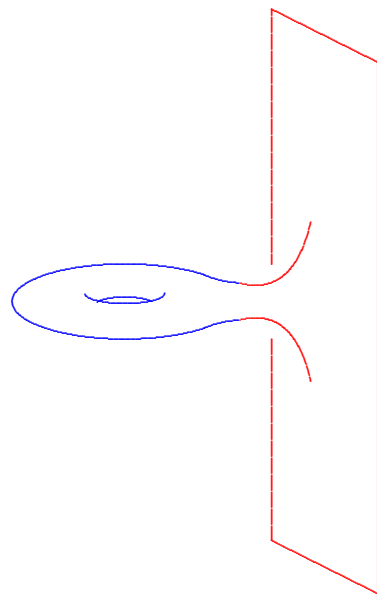


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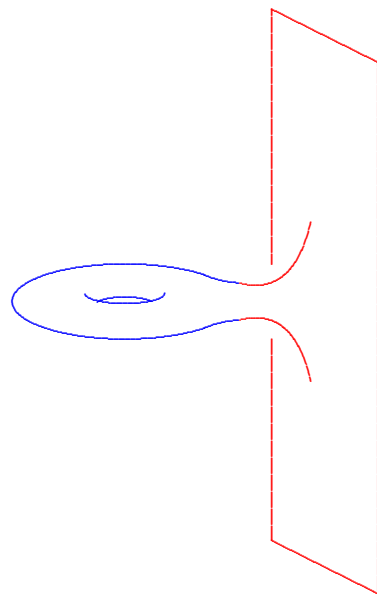
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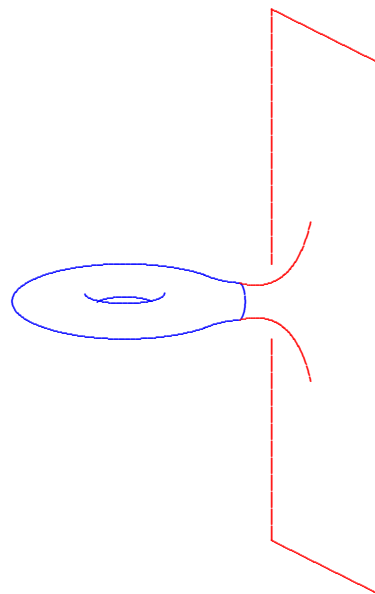
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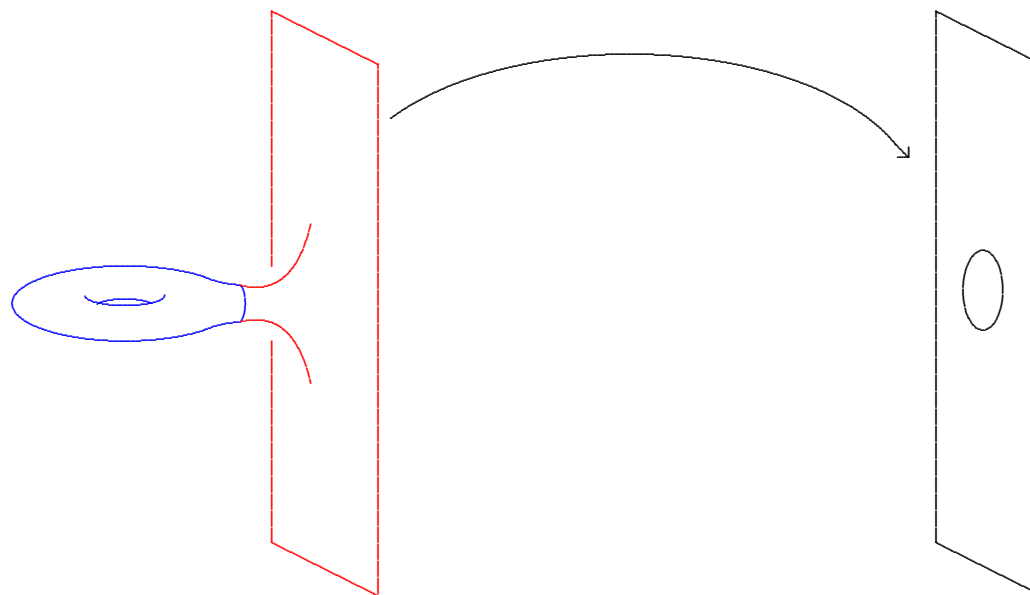


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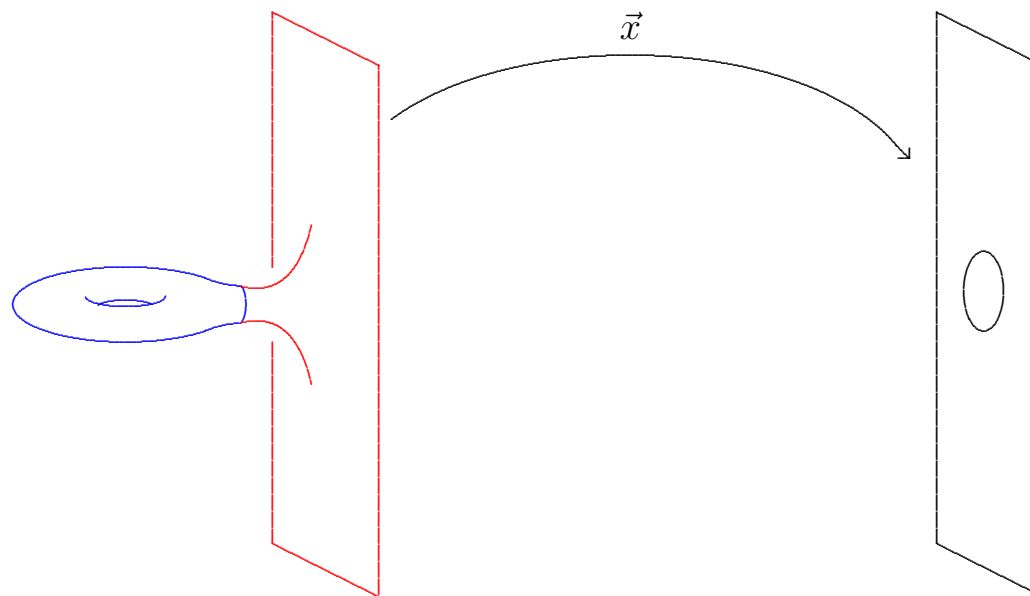
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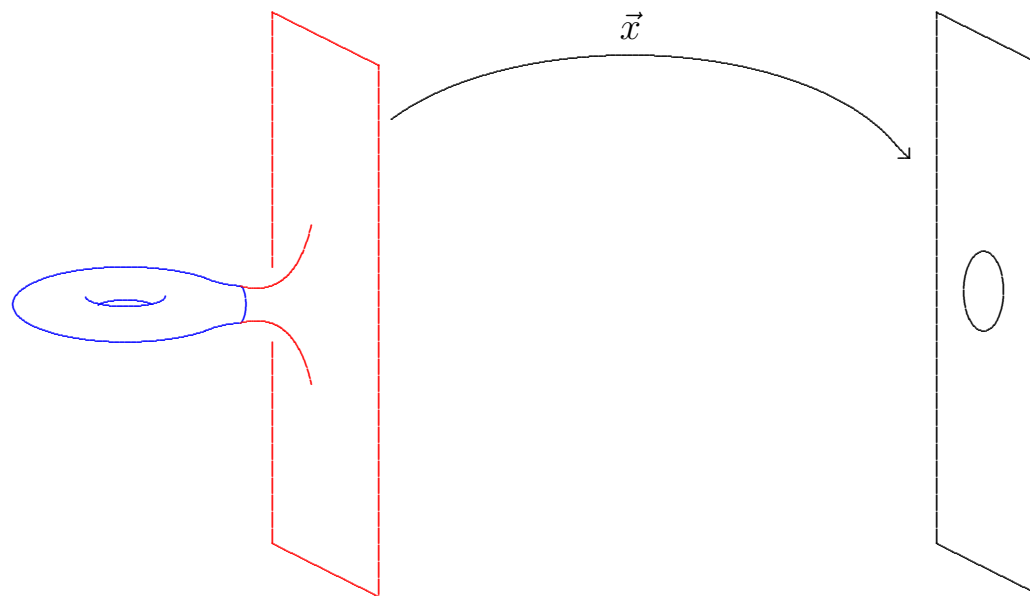


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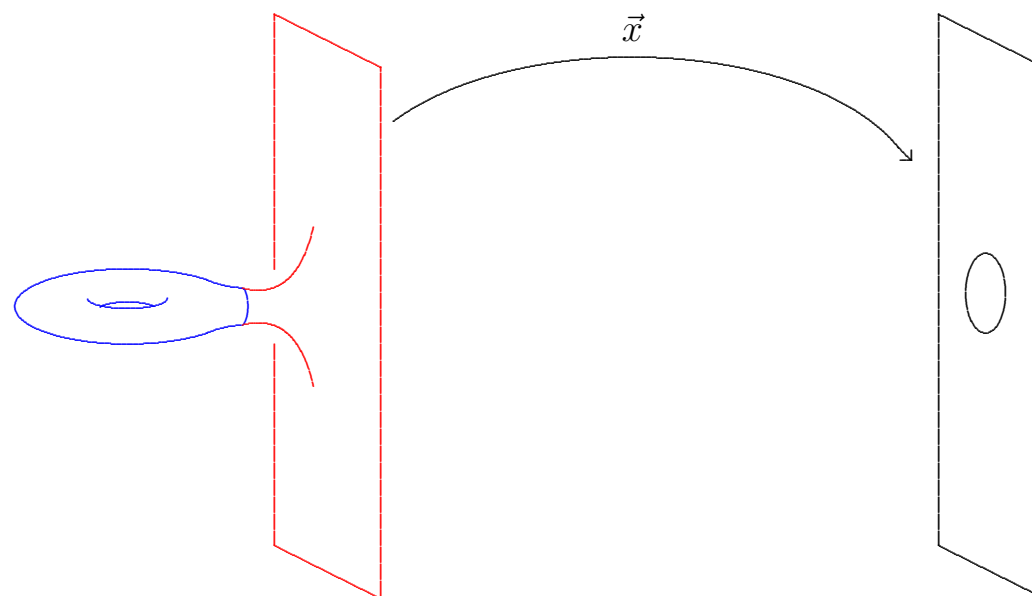
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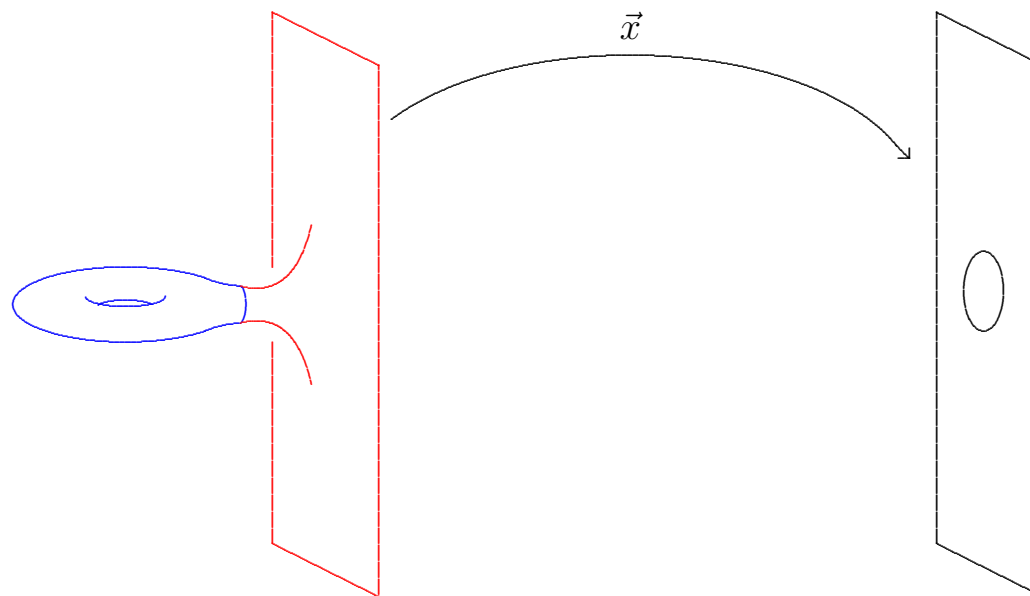
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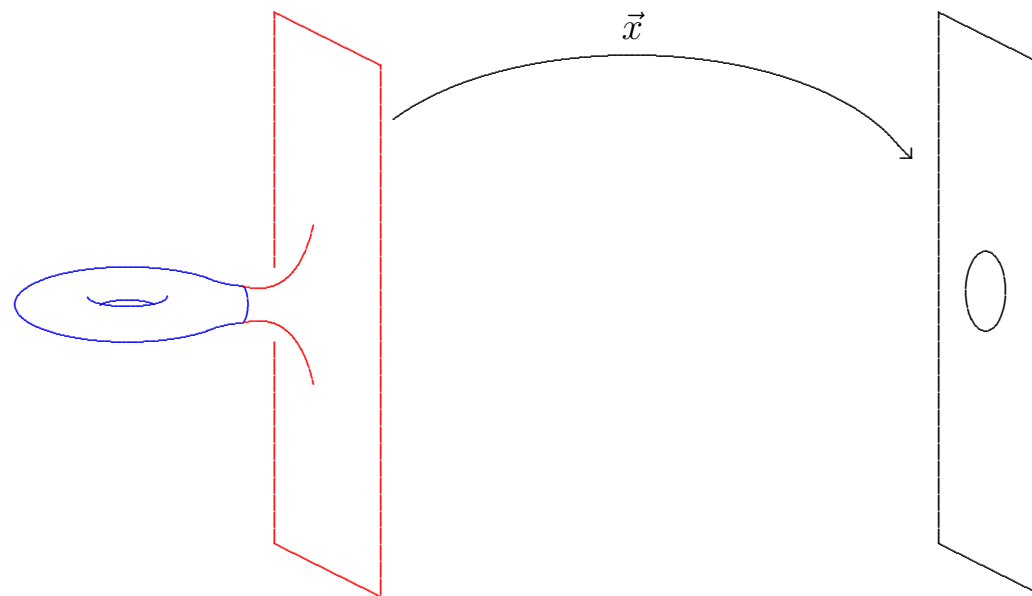
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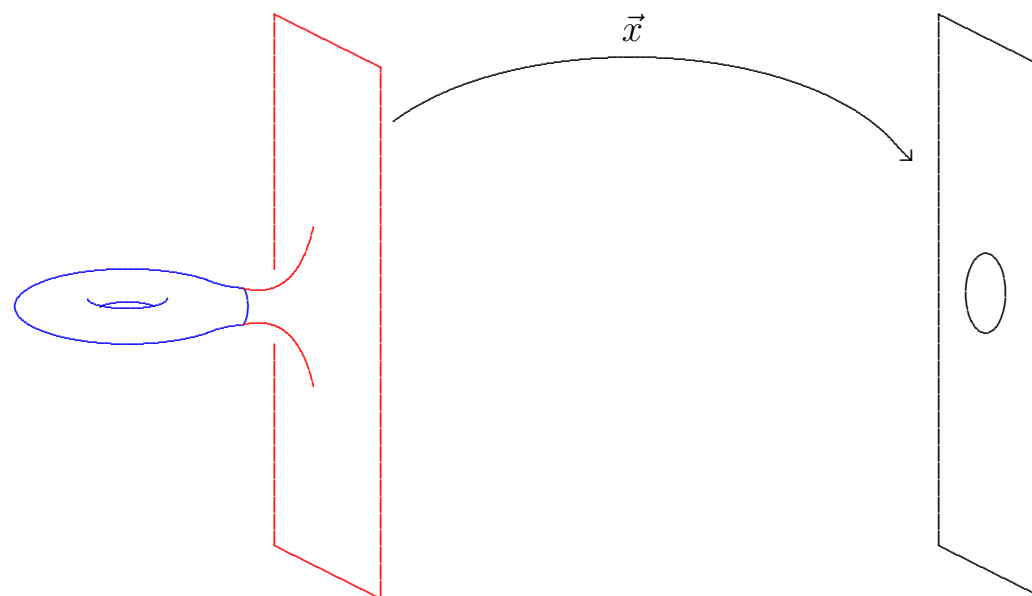
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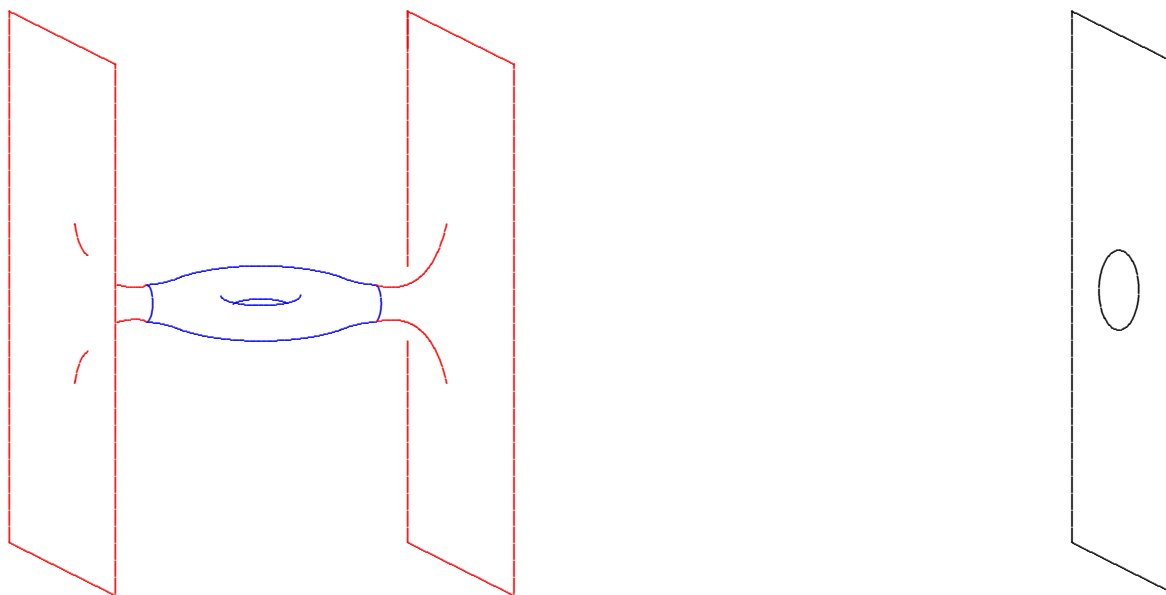
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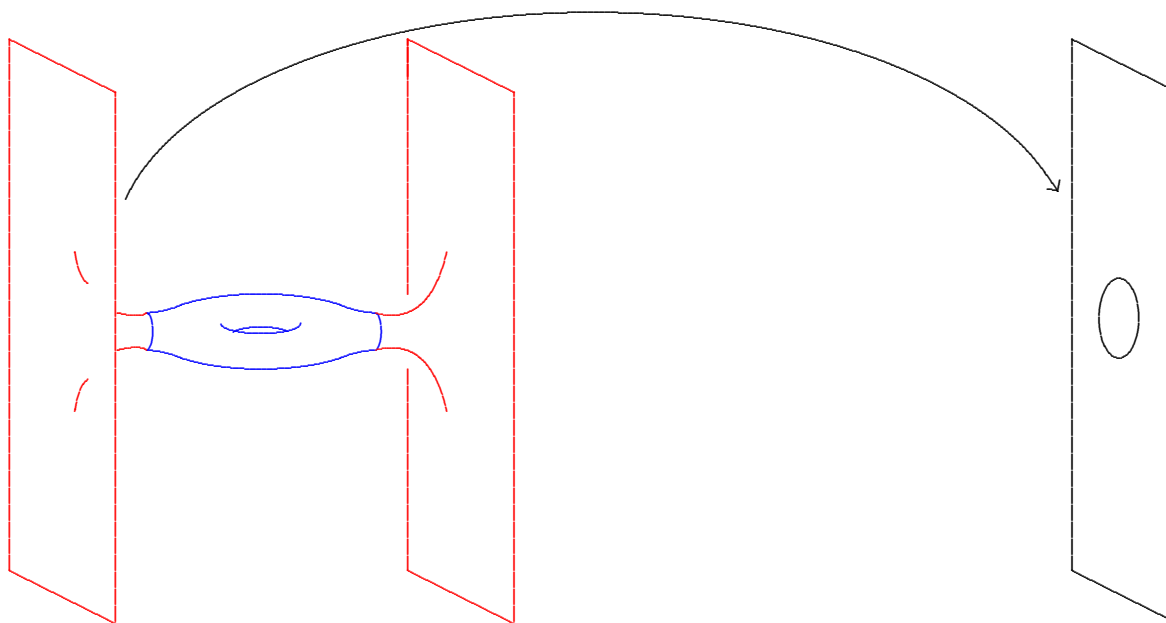
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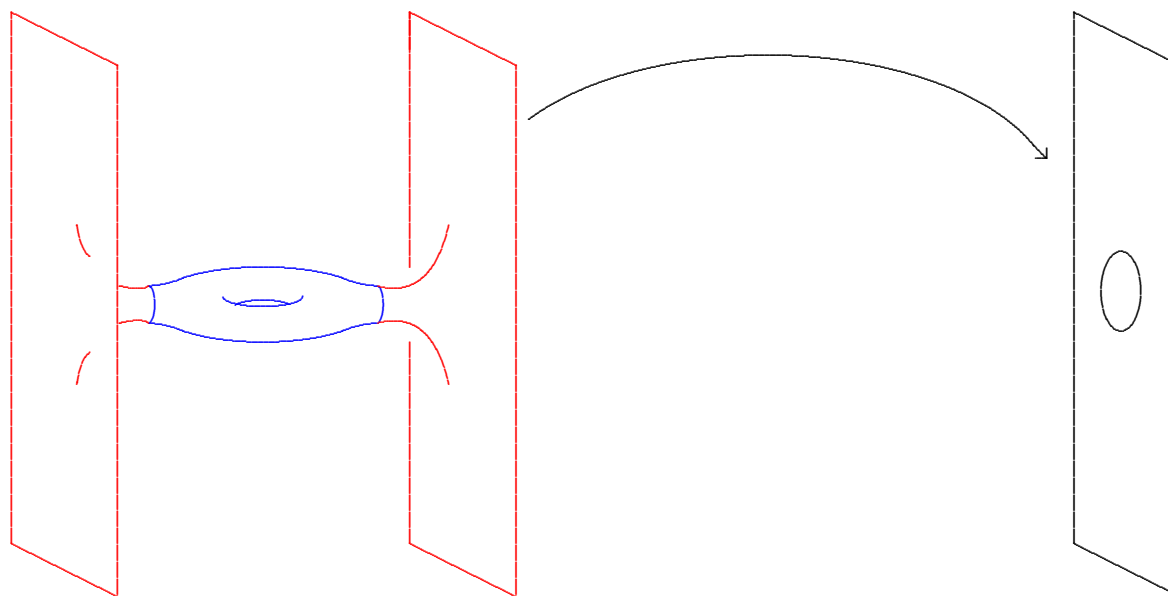
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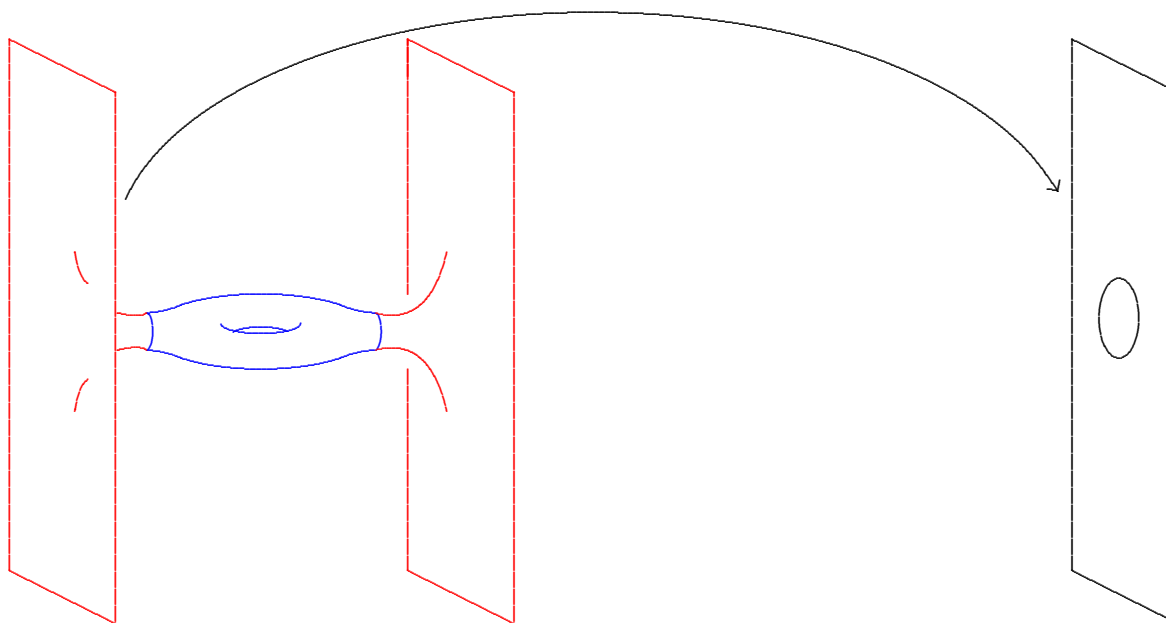
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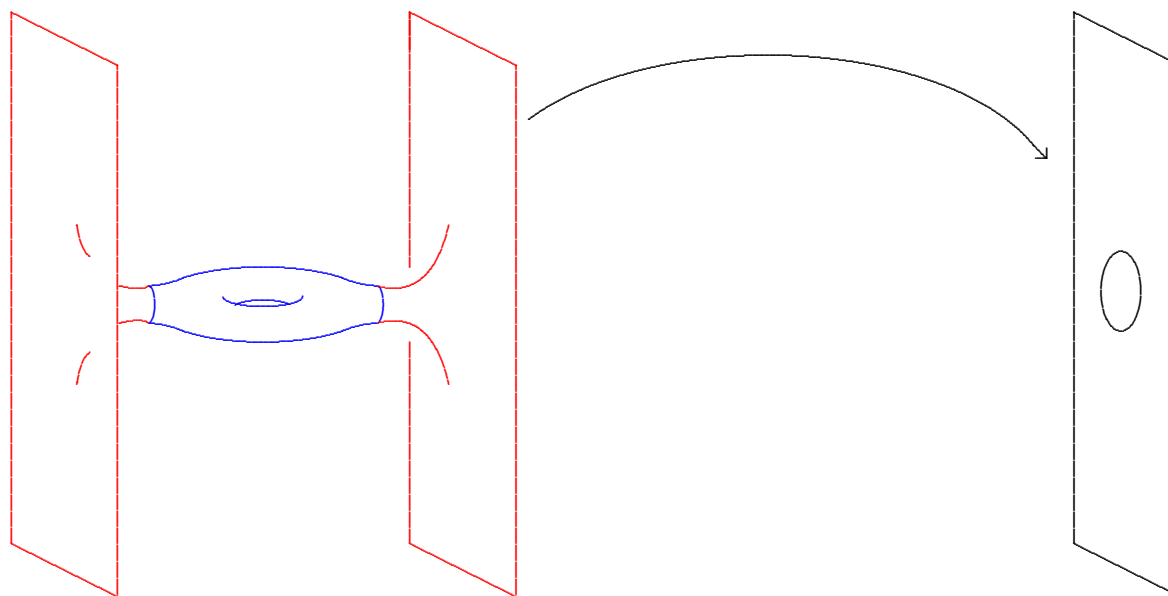
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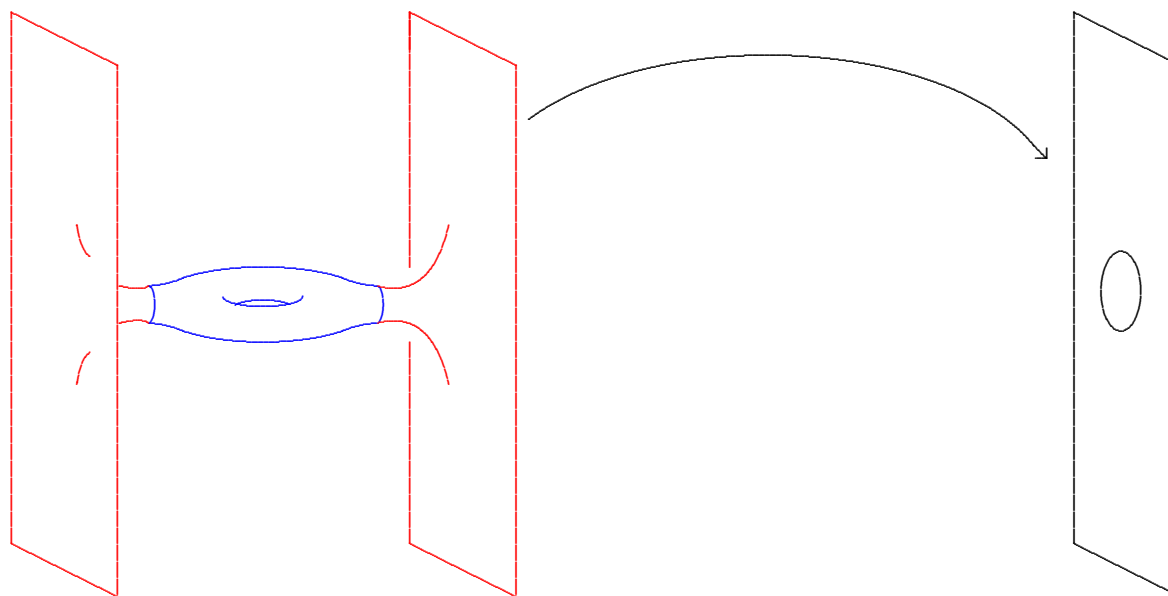
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Seems to depend on choice of coordinates!

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When $n = 3$, ADM mass in general relativity.

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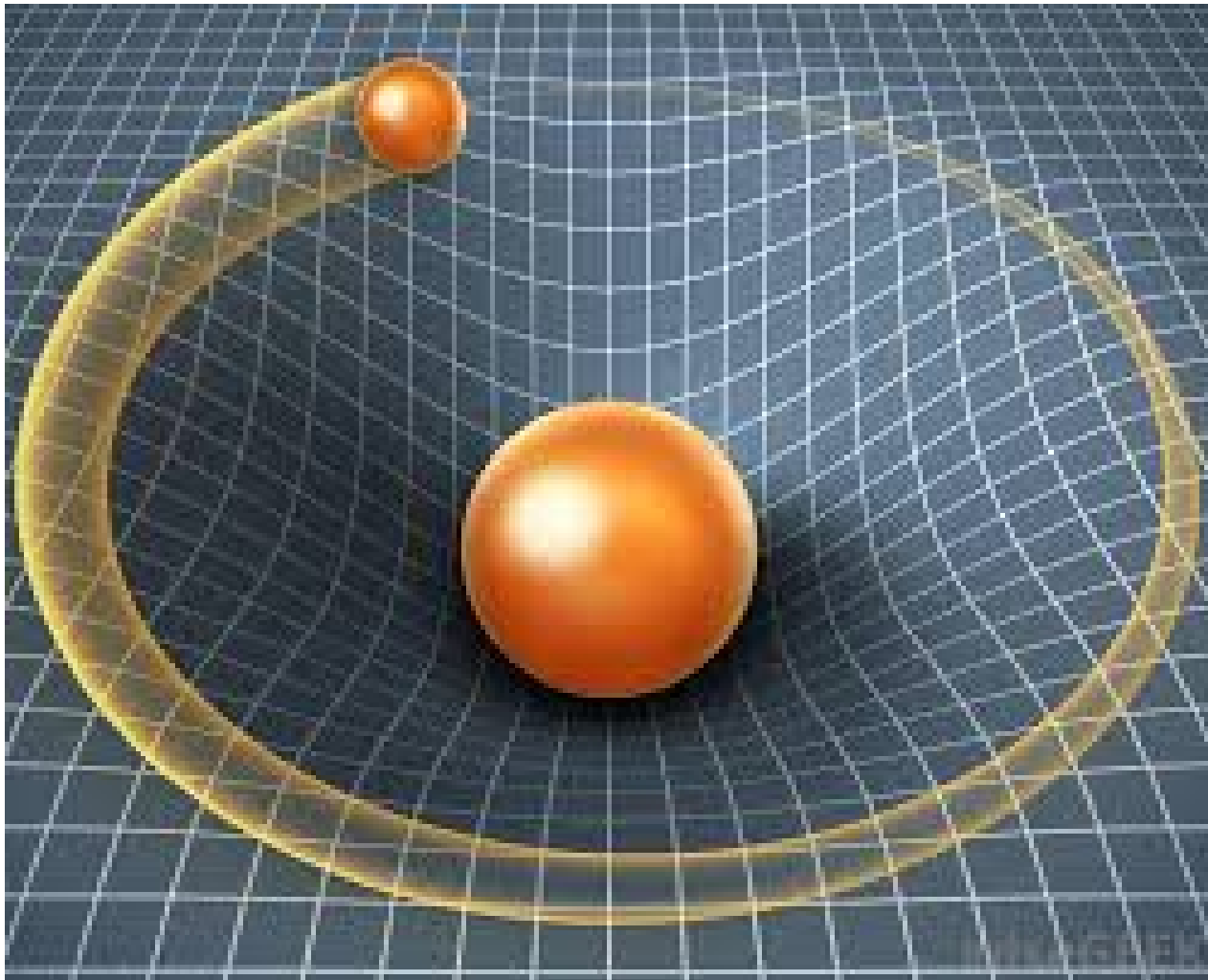
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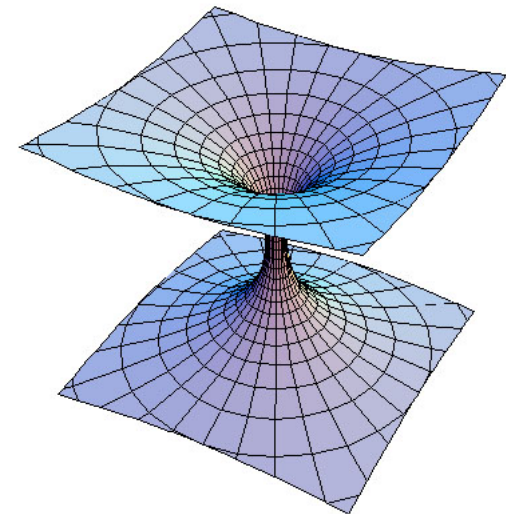
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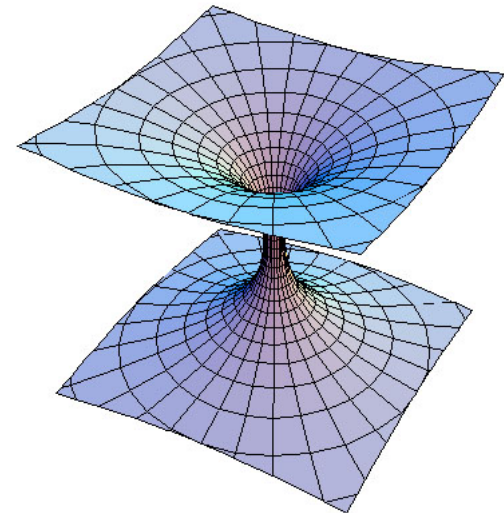
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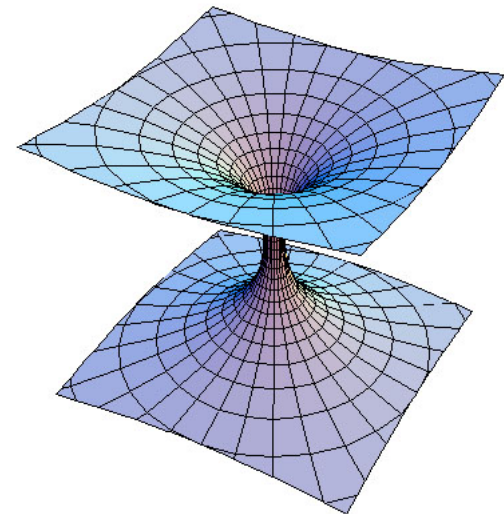
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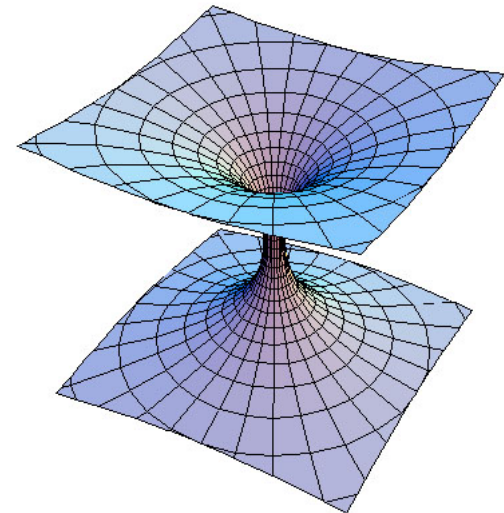
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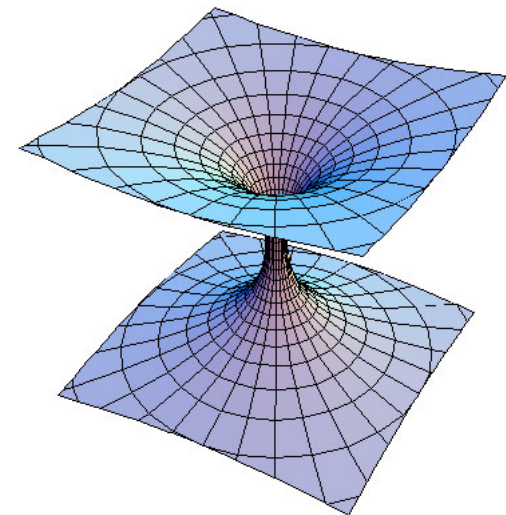
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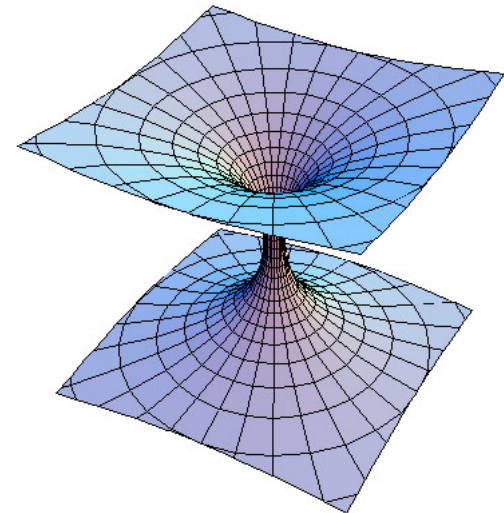
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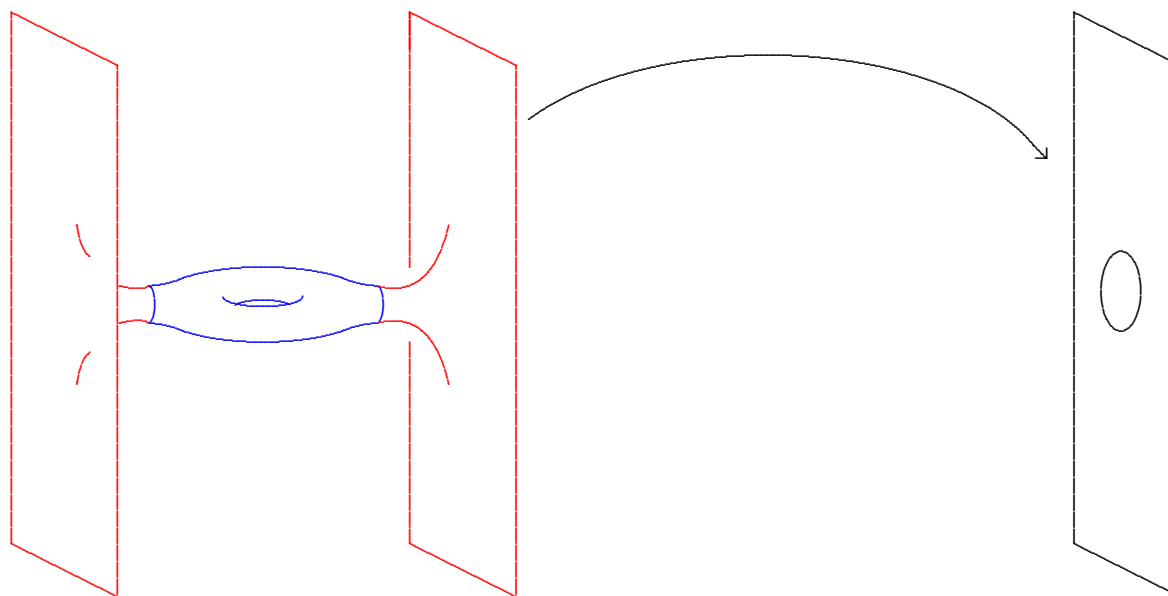
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Definition. A complete, non-compact Riemannian n -manifold (M^n, g) is called **asymptotically Euclidean (AE)** if there is a compact set $K \subset M$ such that each component of $M - K$ is diffeomorphic to $\mathbb{R}^n - D^n$ in such a manner that

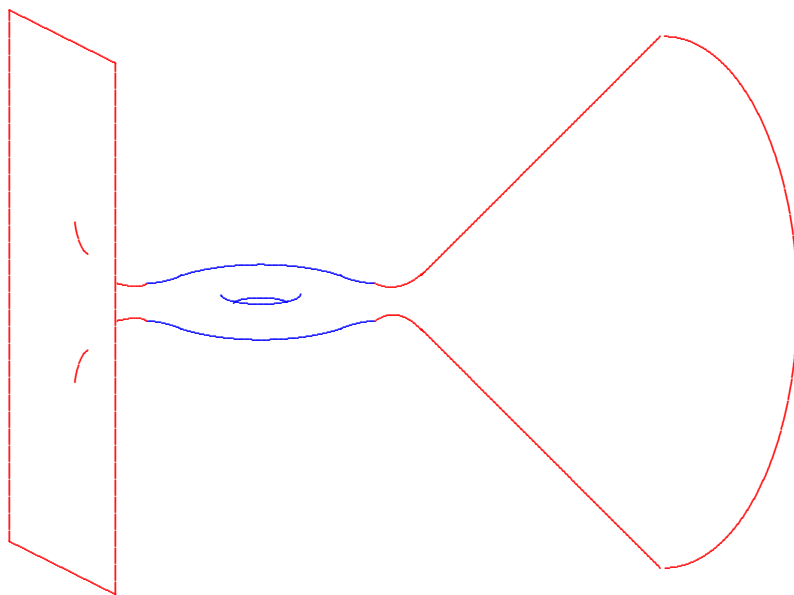


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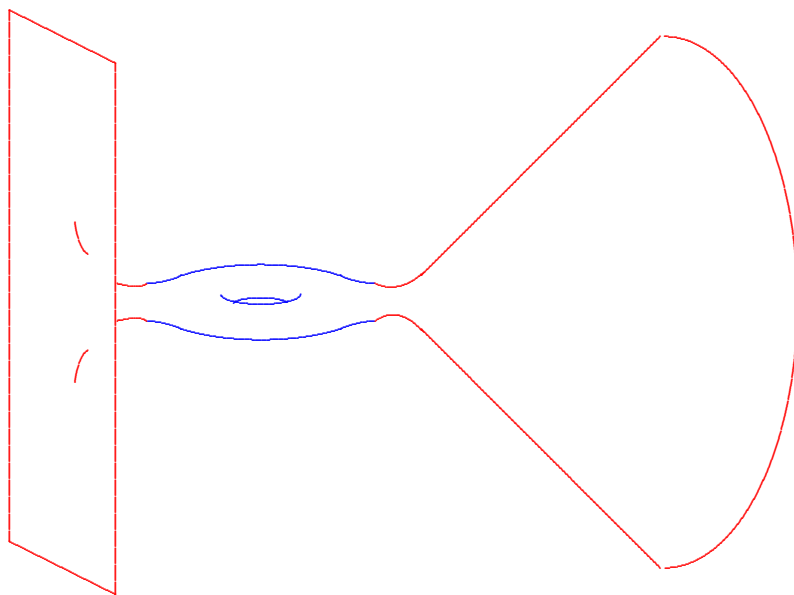
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A Generalization...

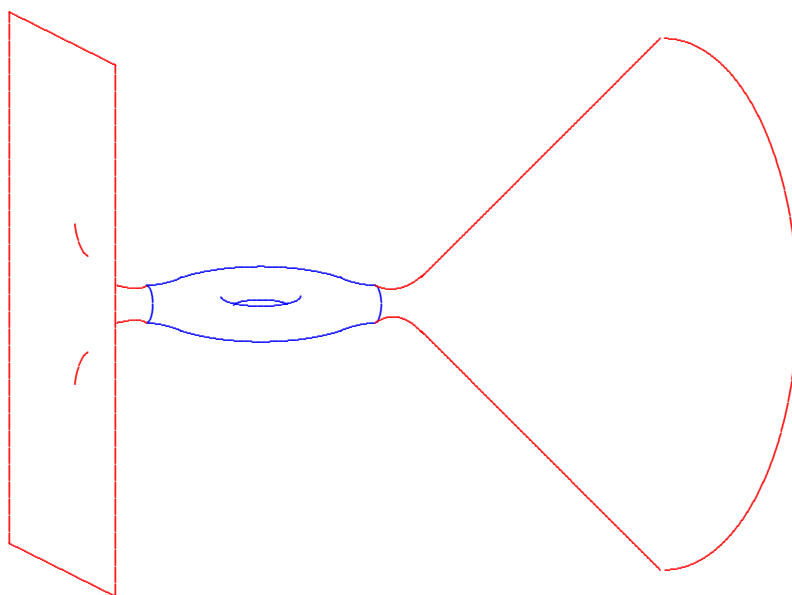
Definition. *Complete, non-compact n -manifold (M^n, g) is asymptotically locally Euclidean*



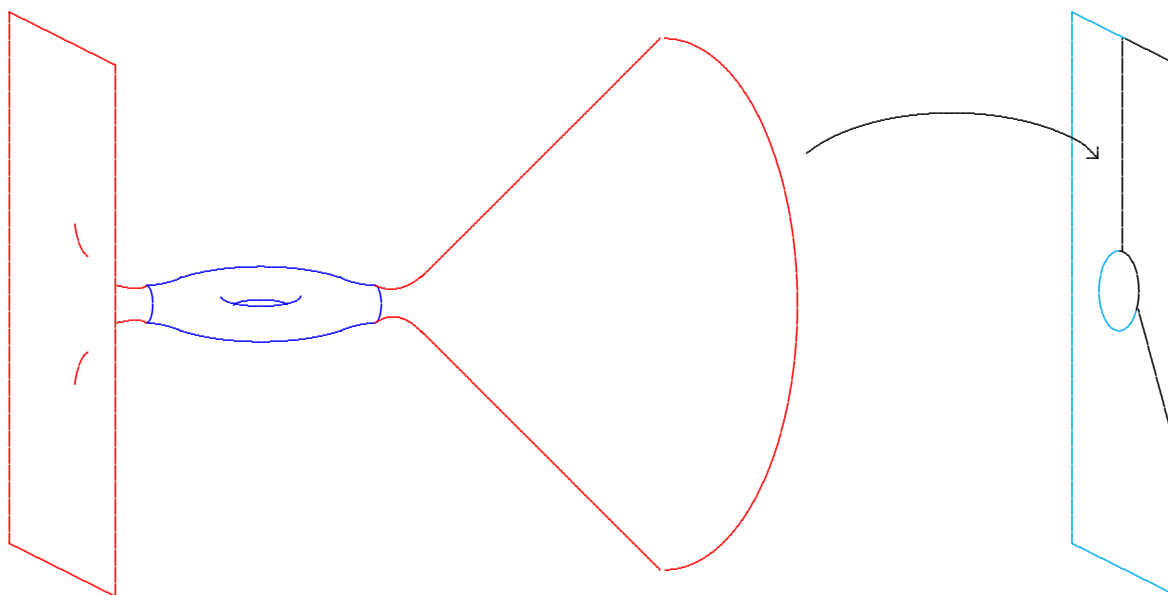
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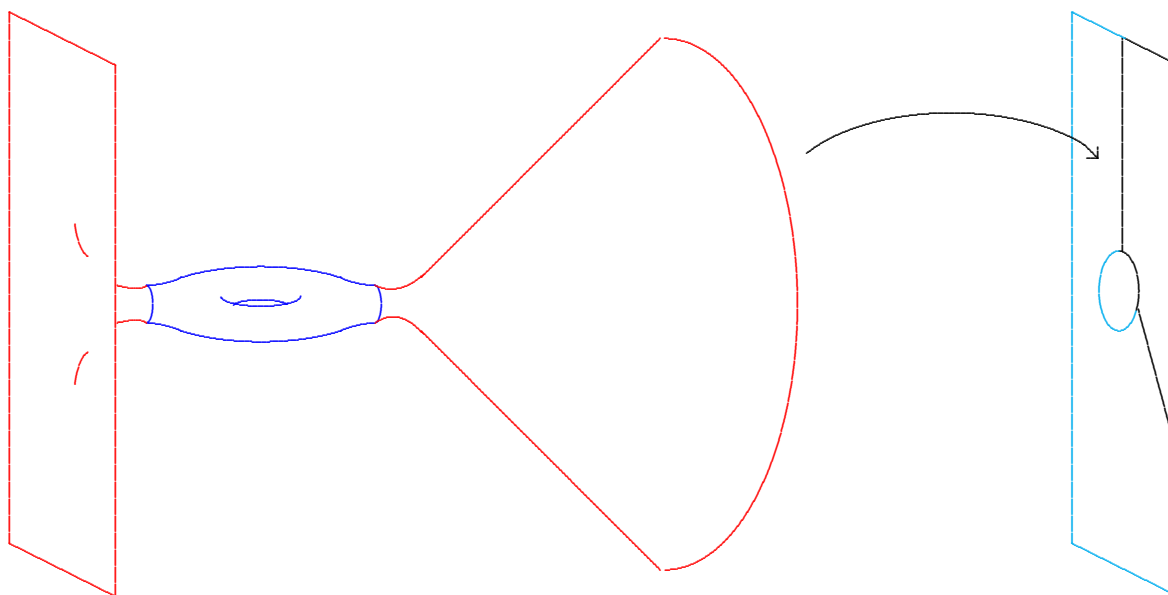
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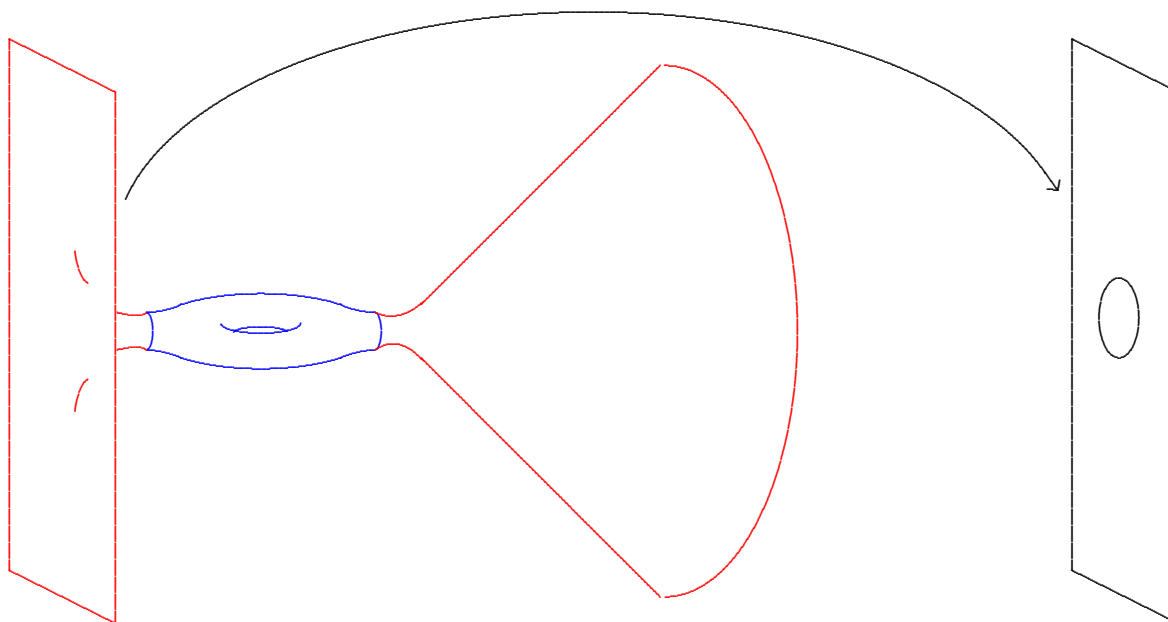
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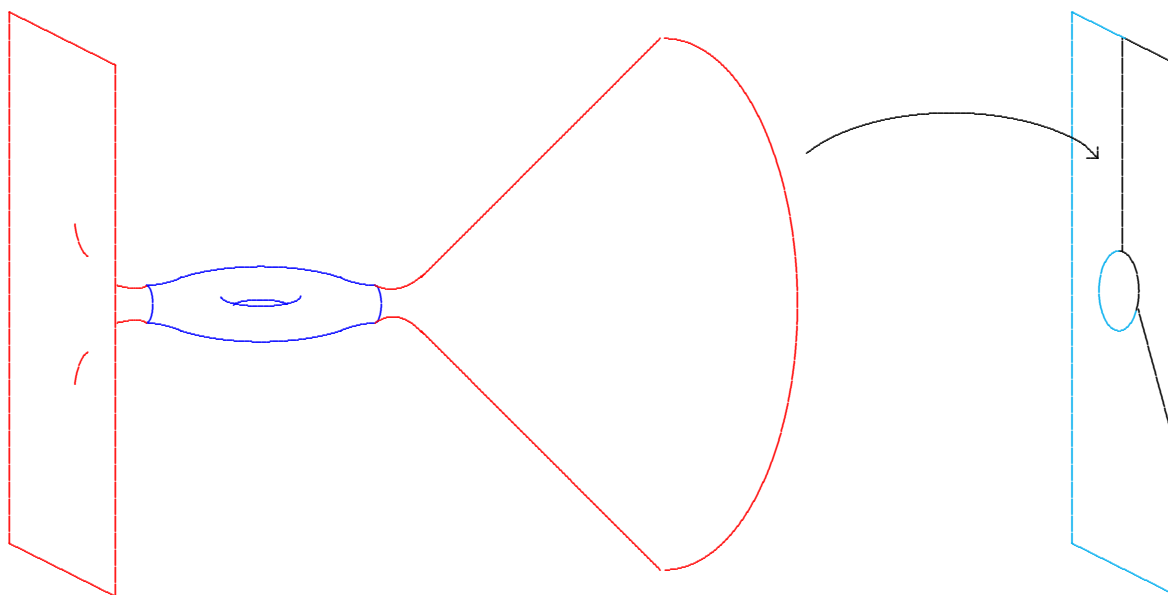
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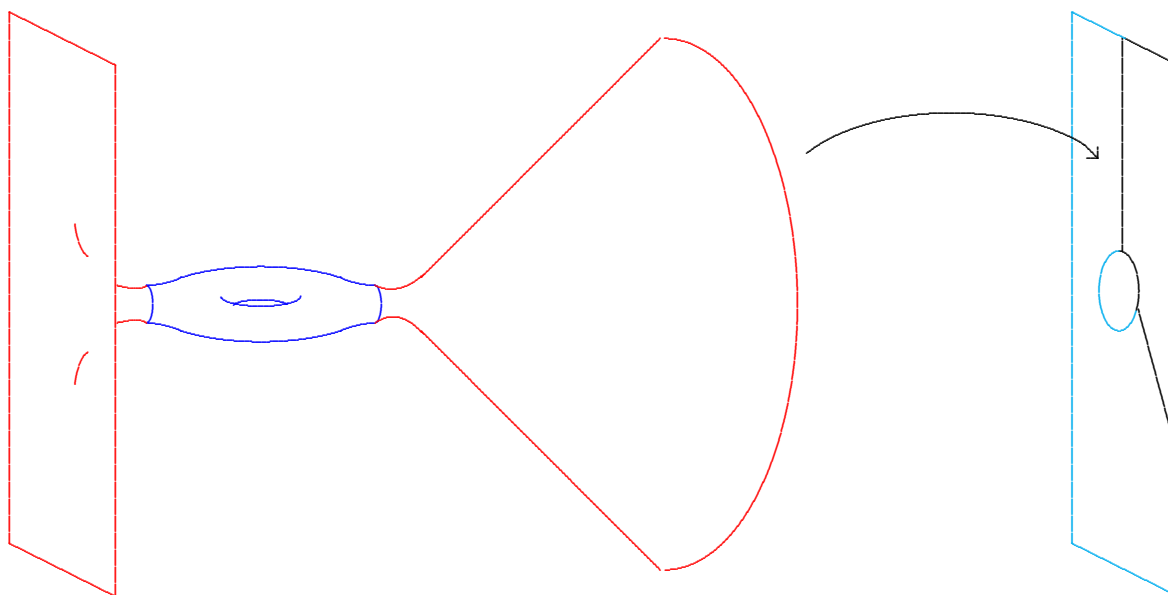
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Why consider **ALE** spaces?

Key examples:

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By contrast, any **Ricci-flat AE** manifold must be flat, by the Bishop-Gromov inequality...

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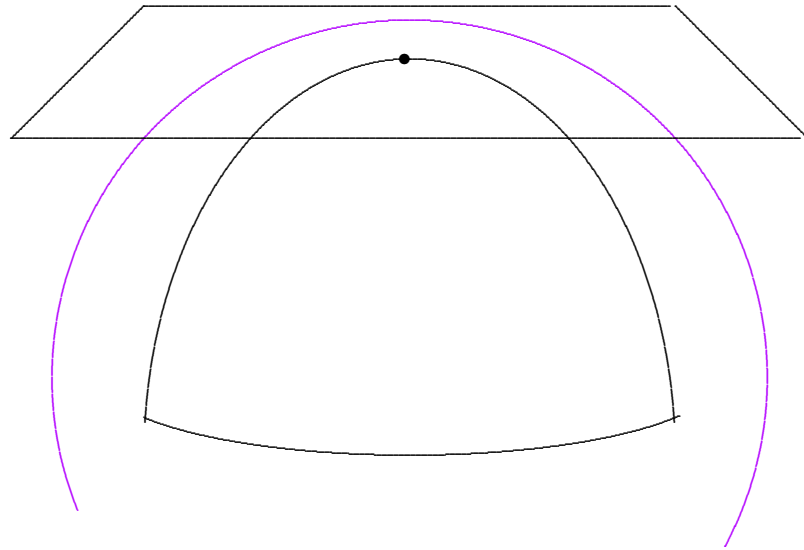
The G-H metrics are **hyper-Kähler**, and were soon rediscovered independently by Hitchin.

$(M^n, g):$

holonomy

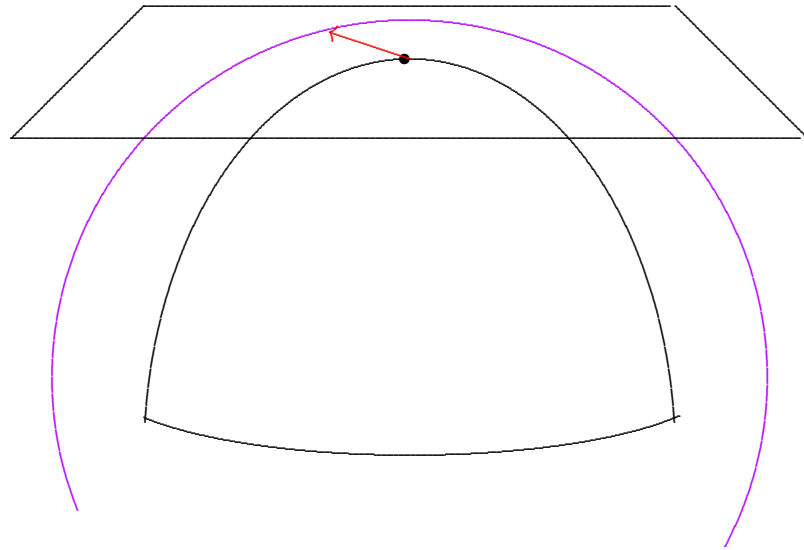
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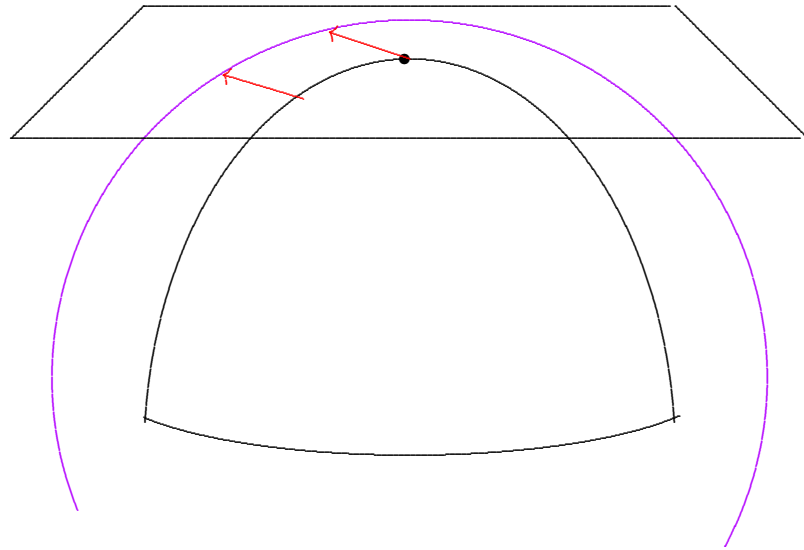
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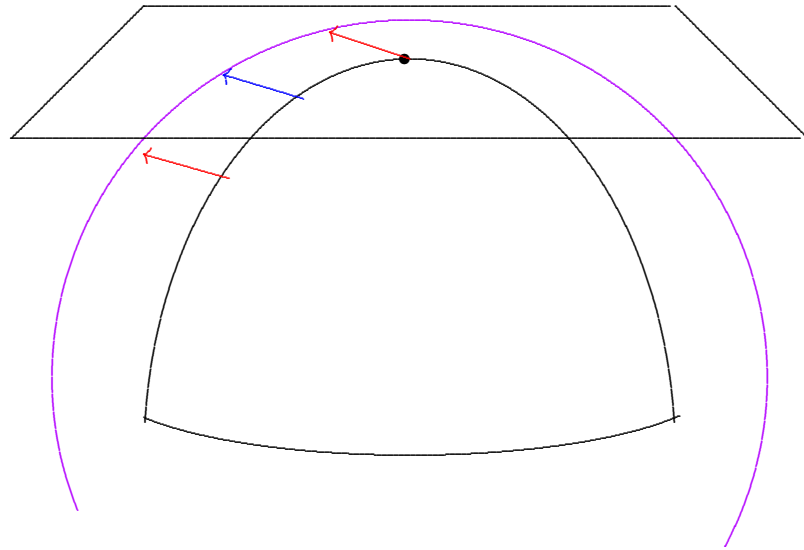
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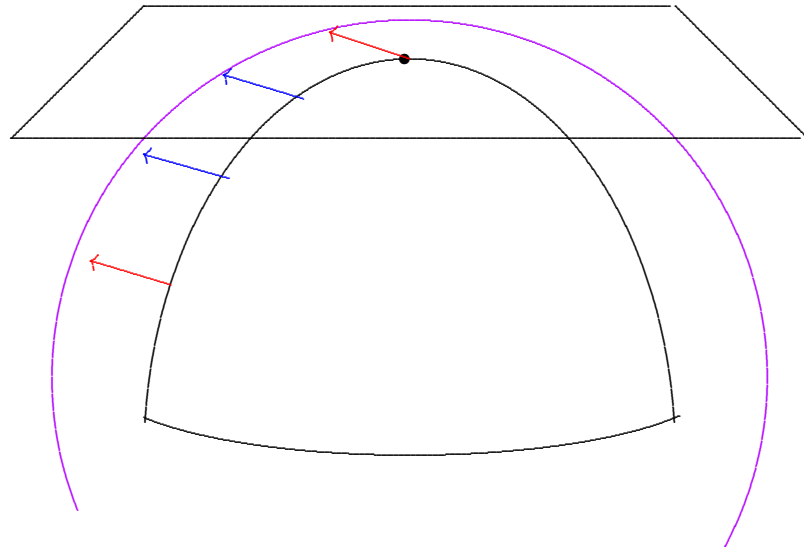
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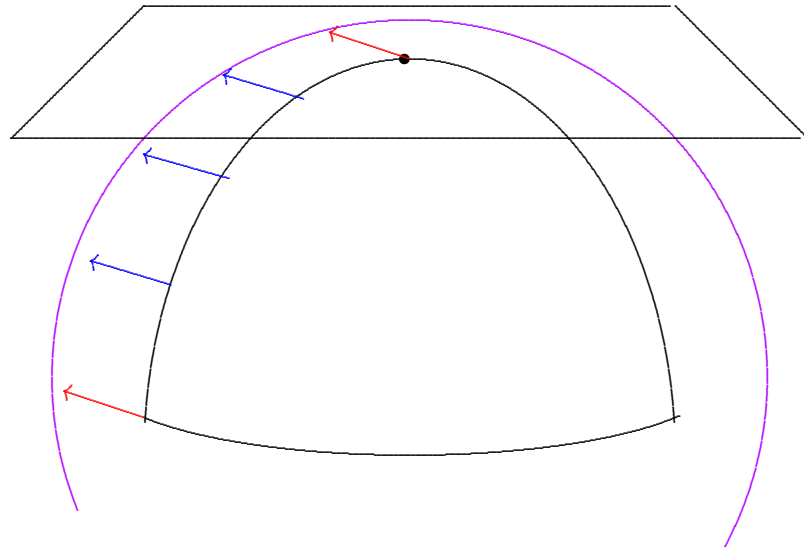
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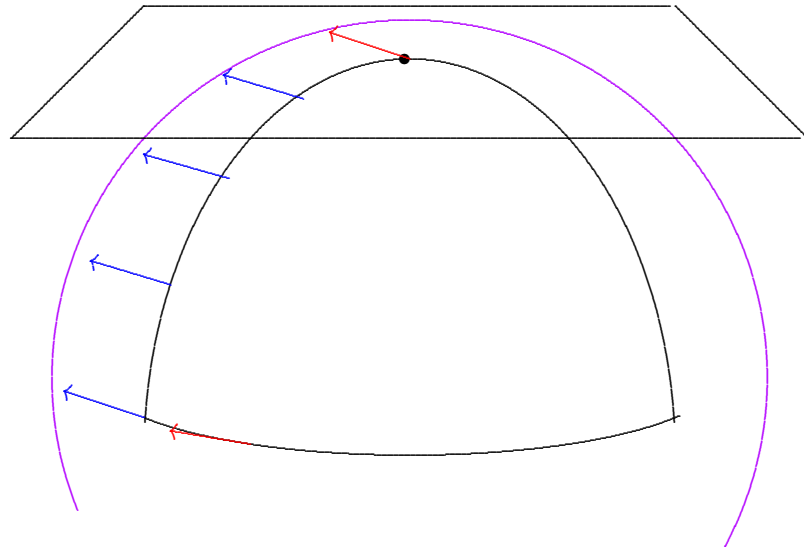
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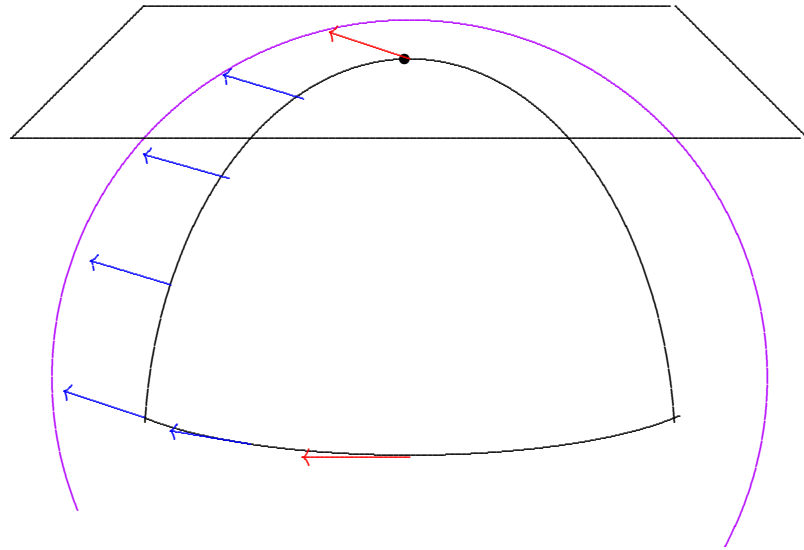
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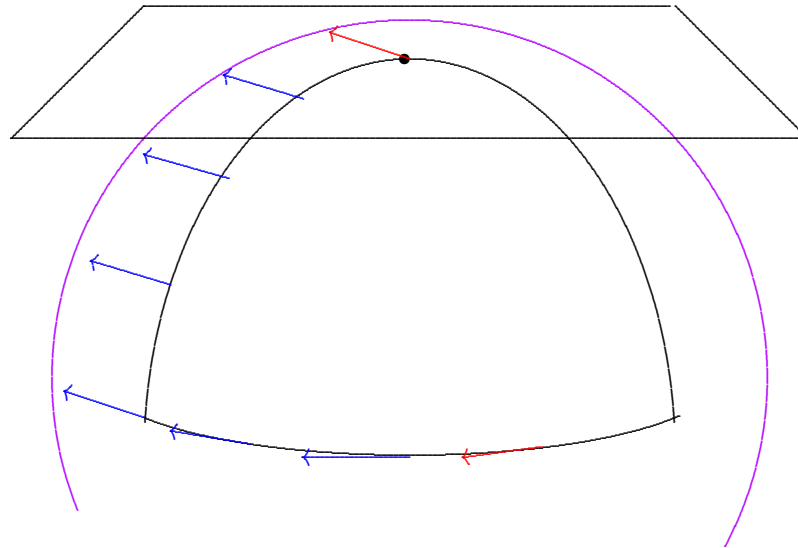
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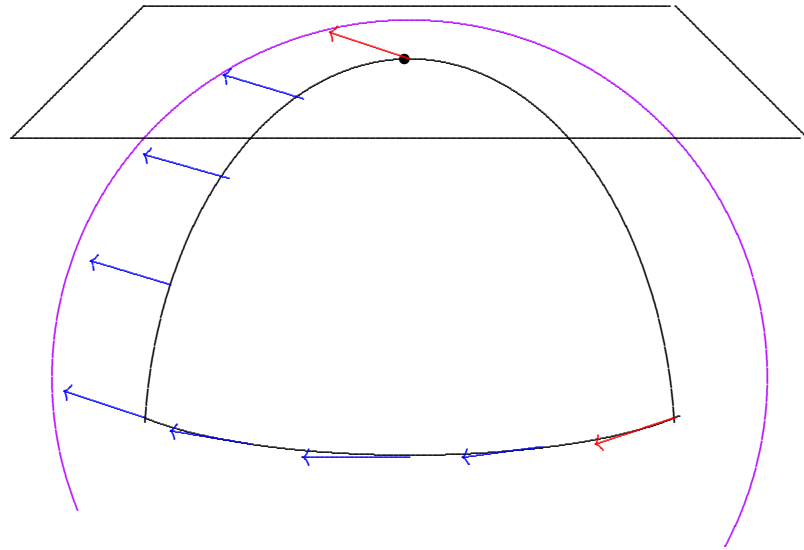
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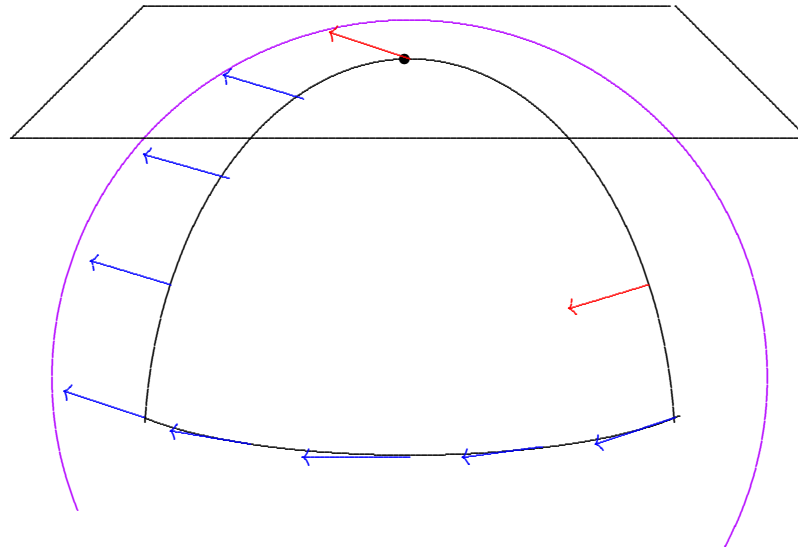
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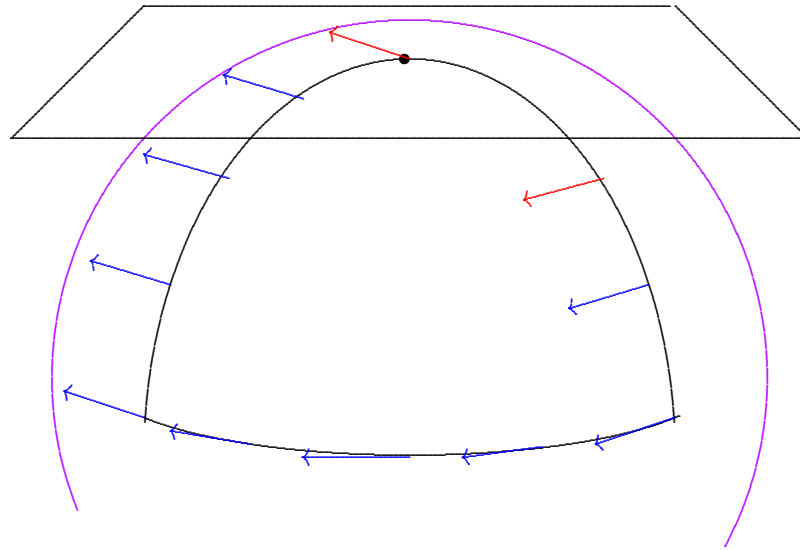
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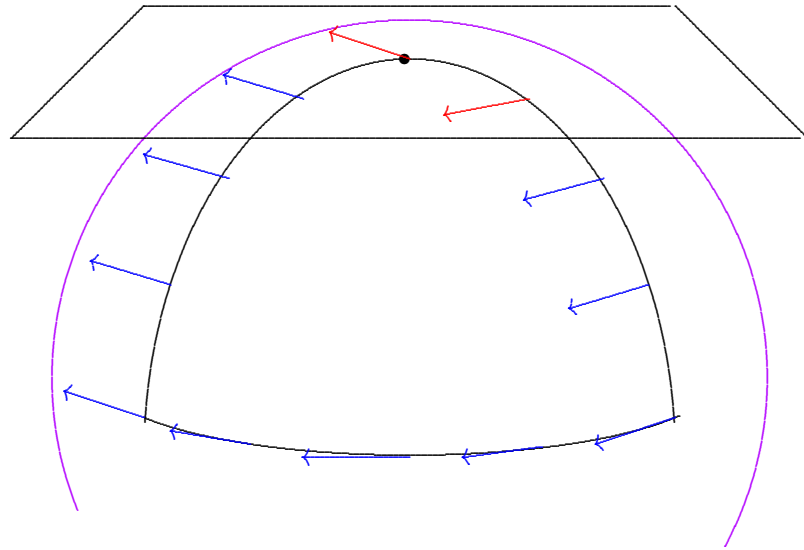
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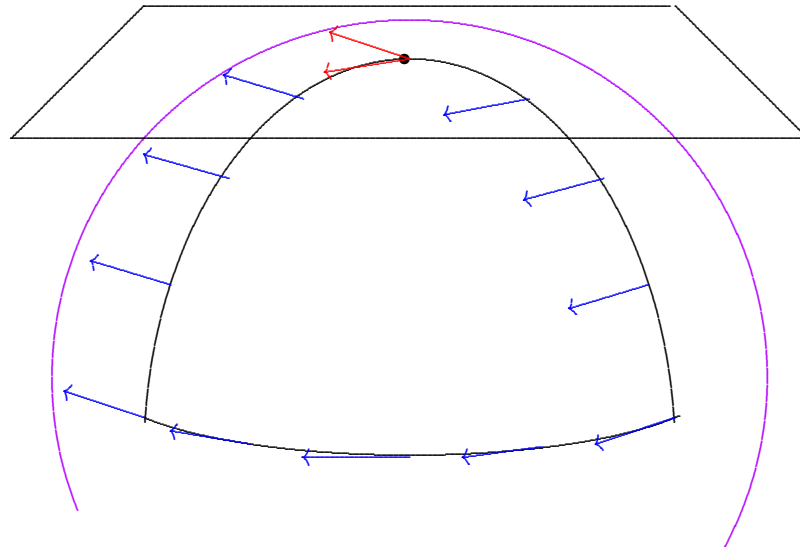
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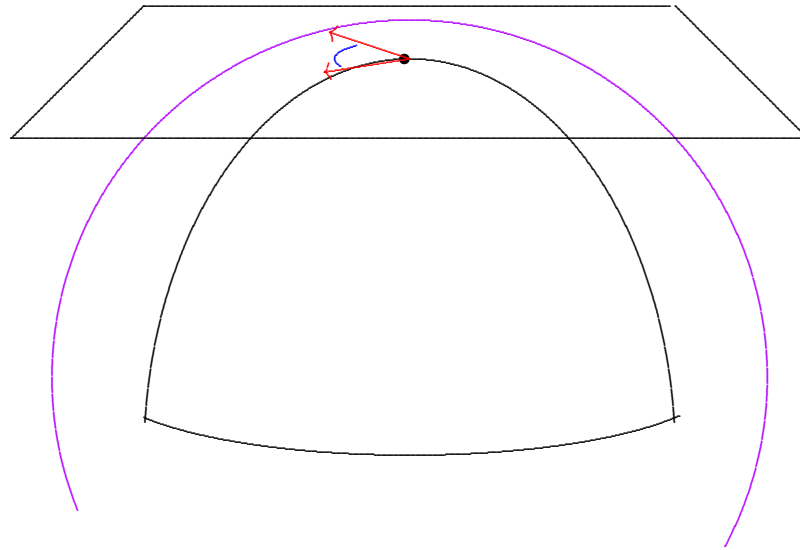
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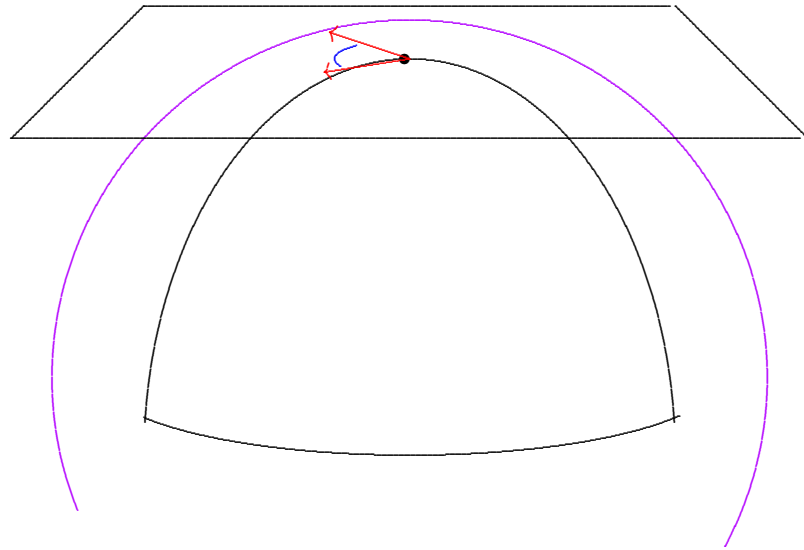
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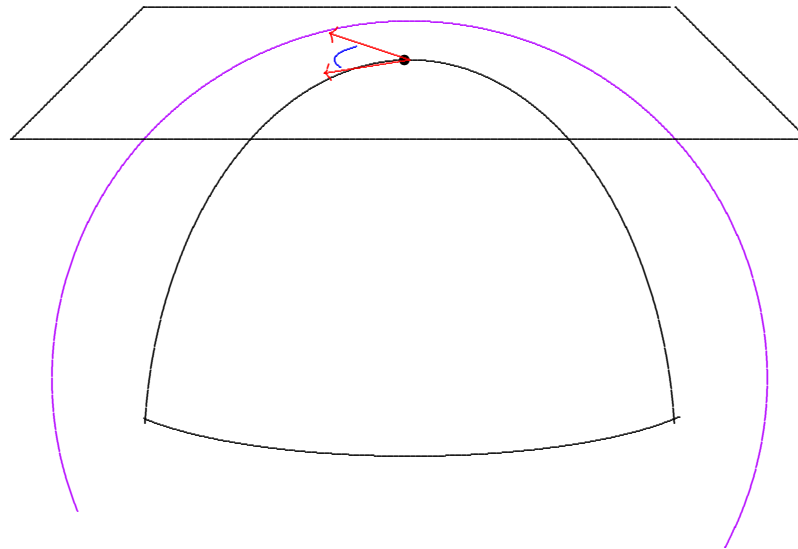
holonomy $\subset \mathbf{O}(n)$



Kähler metrics:

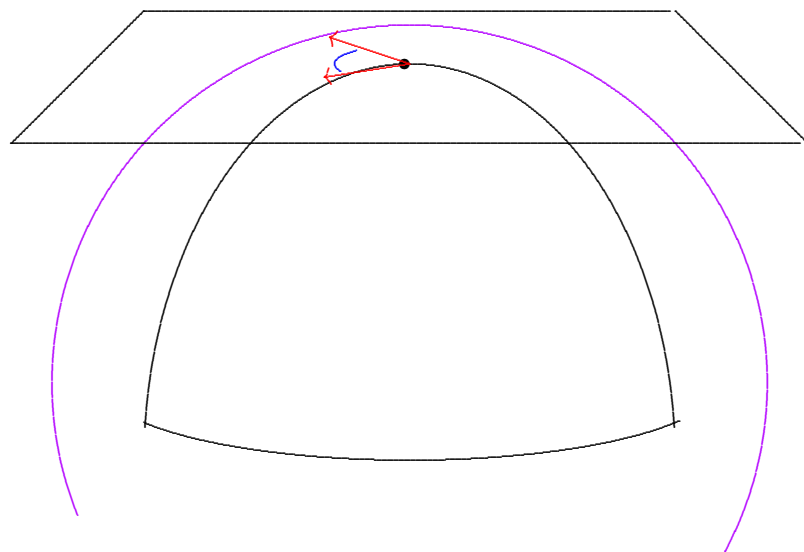
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holonomy



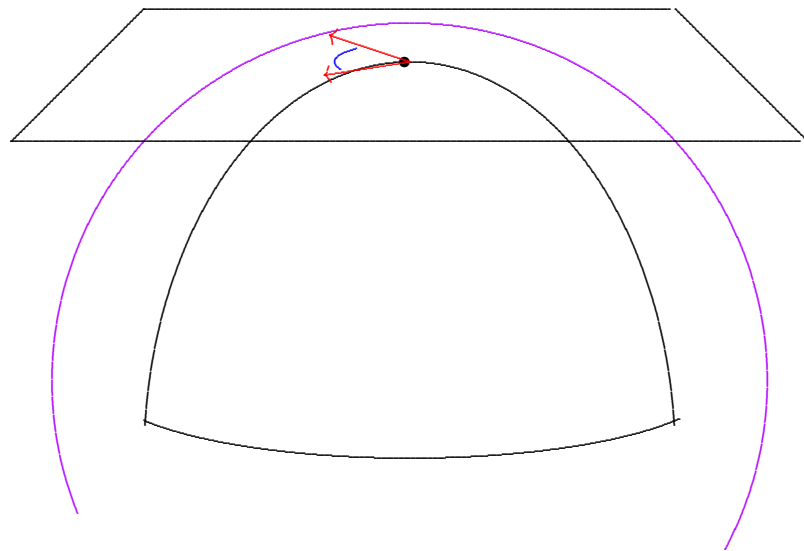
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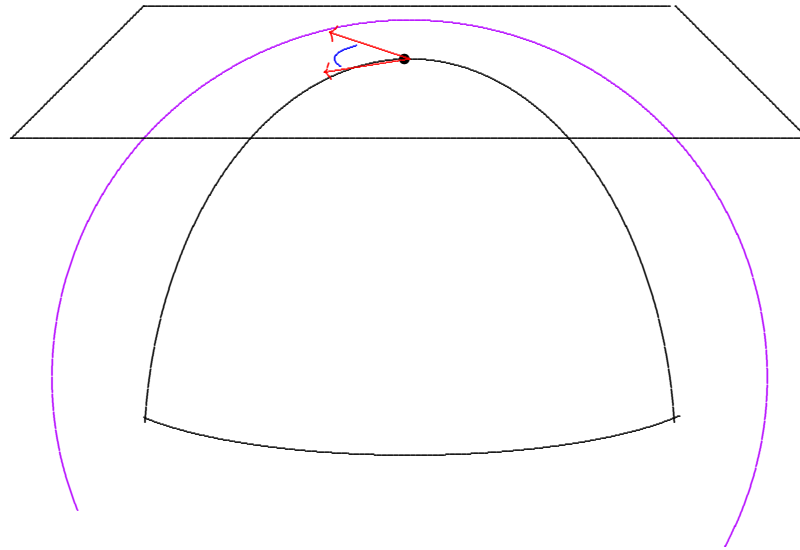
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$$\mathbf{U}(m) := \mathbf{O}(2m) \cap \mathbf{GL}(m, \mathbb{C})$$

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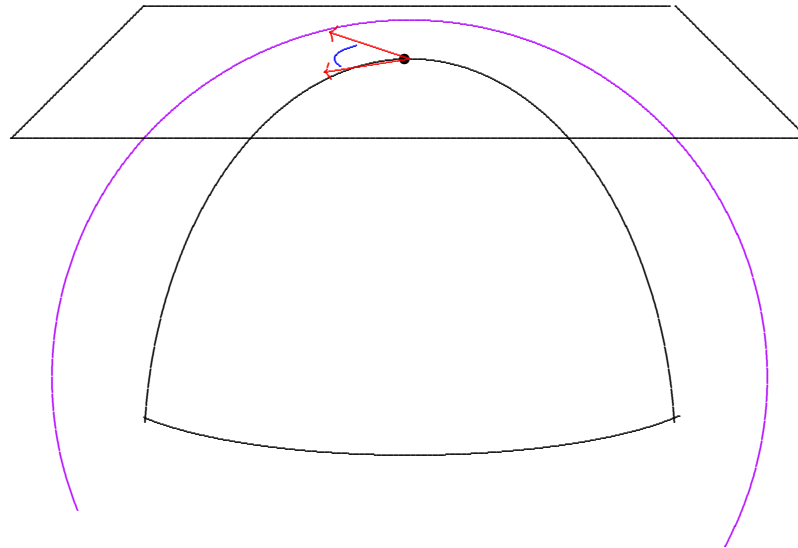
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Makes tangent space a complex vector space!

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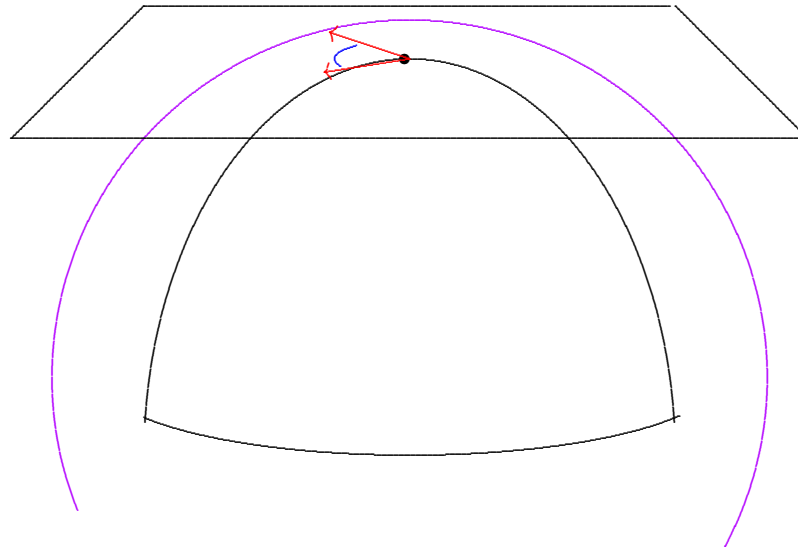
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$$J : TM \rightarrow TM, \quad J^2 = -\text{identity}$$

“almost-complex structure”

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Makes tangent space a complex vector space!

Invariant under parallel transport!

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$$d\omega = 0$$

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$$[\omega] \in H^2(M)$$

“Kähler class”

Kähler metrics:

(M^{2m}, g) Kähler \iff holonomy $\subset \mathbf{U}(m)$

$\iff \exists$ almost complex-structure J with $\nabla J = 0$
and $g(J\cdot, J\cdot) = g$.

$\iff (M^{2m}, g)$ is a complex manifold & $\exists J$ -invariant
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Kähler magic:

$$r = - \sum_{j,k=1}^m \frac{\partial^2}{\partial z^j \partial \bar{z}^k} \log \det[g_{p\bar{q}}] \left[dz^j \otimes d\bar{z}^k + d\bar{z}^k \otimes dz^j \right]$$

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Kähler magic:

If we define the Ricci form by

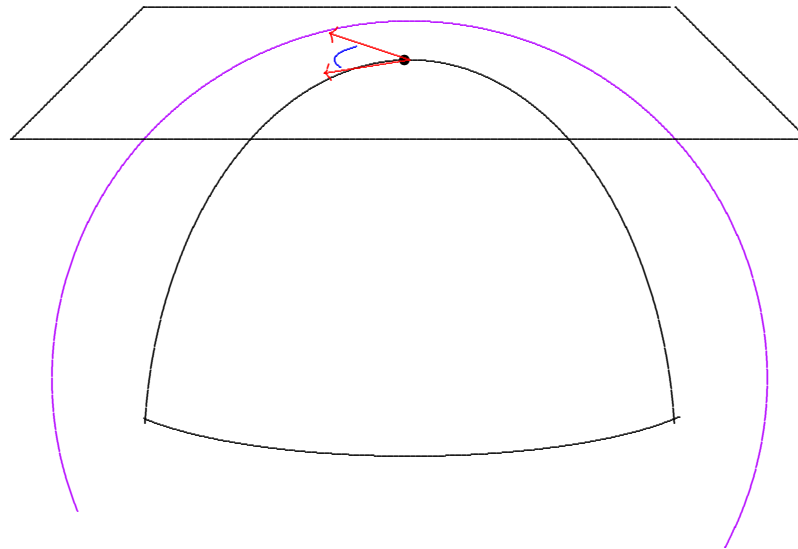
$$\rho = r(J\cdot, \cdot)$$

then $i\rho$ is curvature of canonical line bundle $\Lambda^{m,0}$.

Kähler metrics:

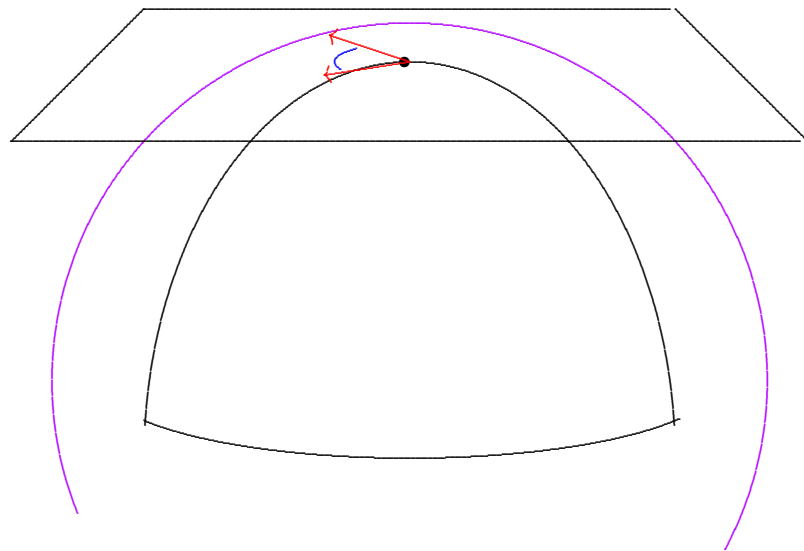
(M^{2m}, g) :

holonomy



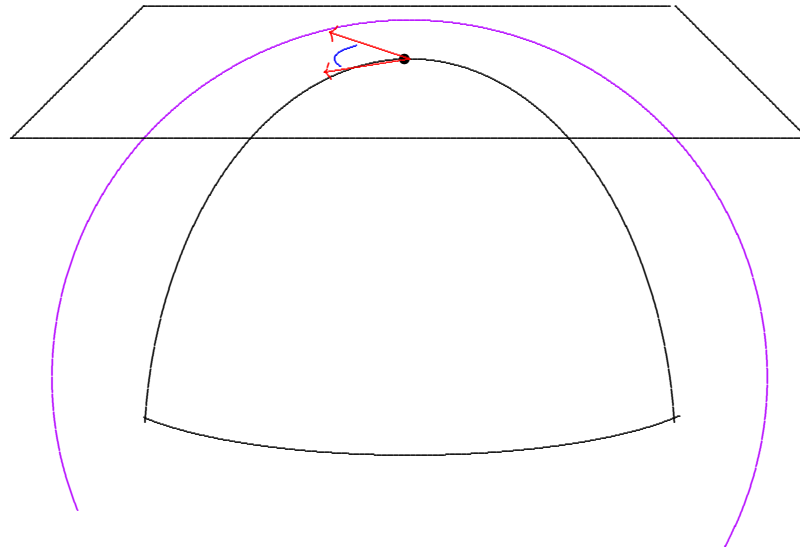
Kähler metrics:

(M^{2m}, g) : Ricci-flat Kähler \Longleftarrow holonomy $\subset \mathbf{SU}(m)$



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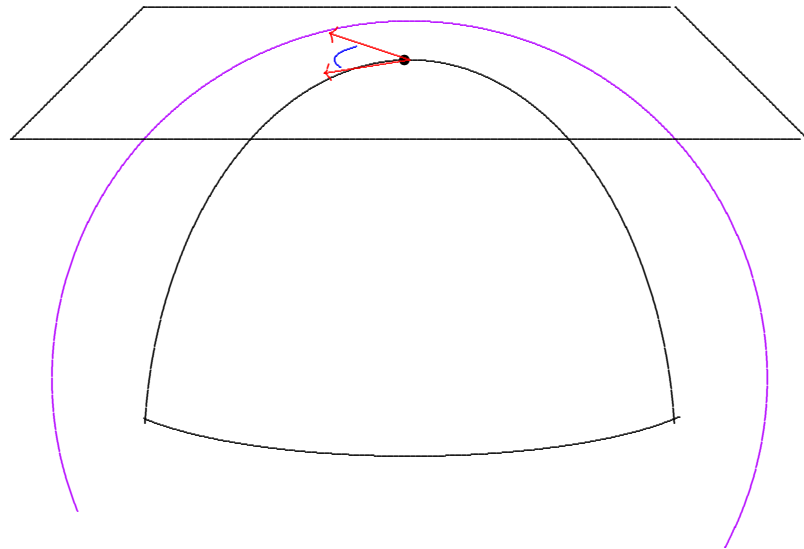
(M^{2m}, g) : Ricci-flat Kähler \Longleftarrow holonomy $\subset \mathbf{SU}(m)$



$$\mathbf{SU}(m) \subset \mathbf{U}(m) : \quad \{A \mid \det A = 1\}$$

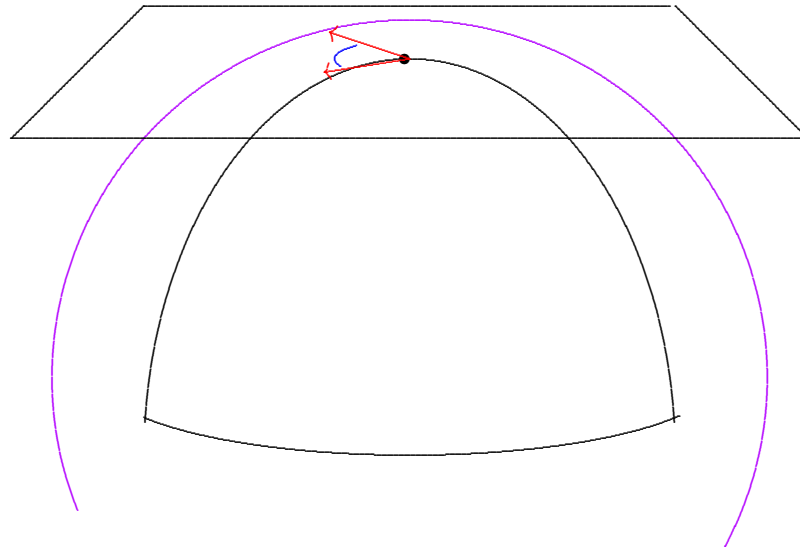
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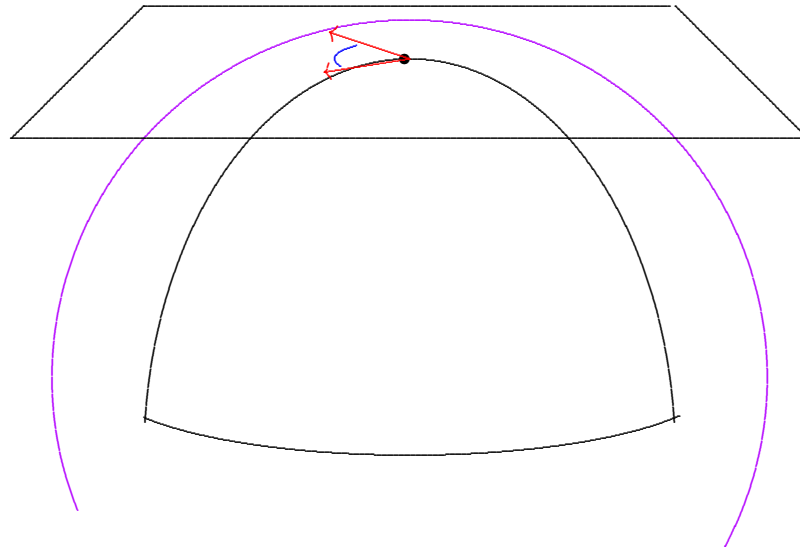


if M is simply connected.

Hyper-Kähler metrics:

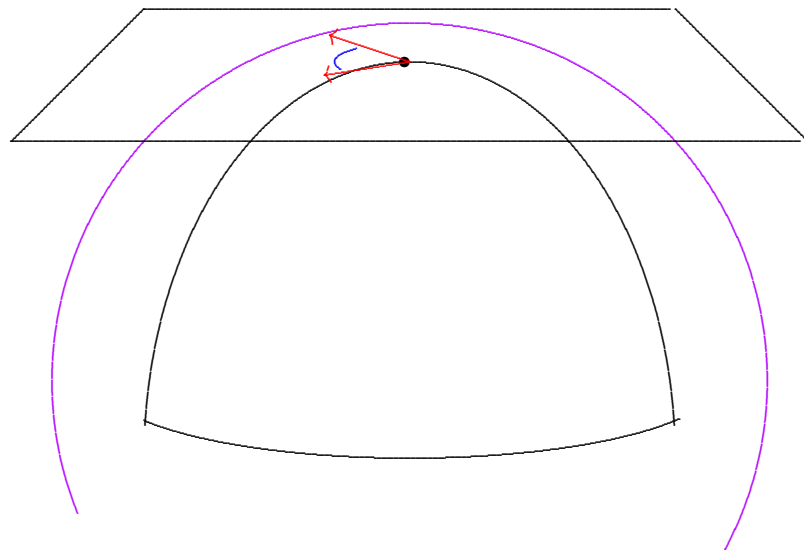
$(M^{4\ell}, g)$

holonomy



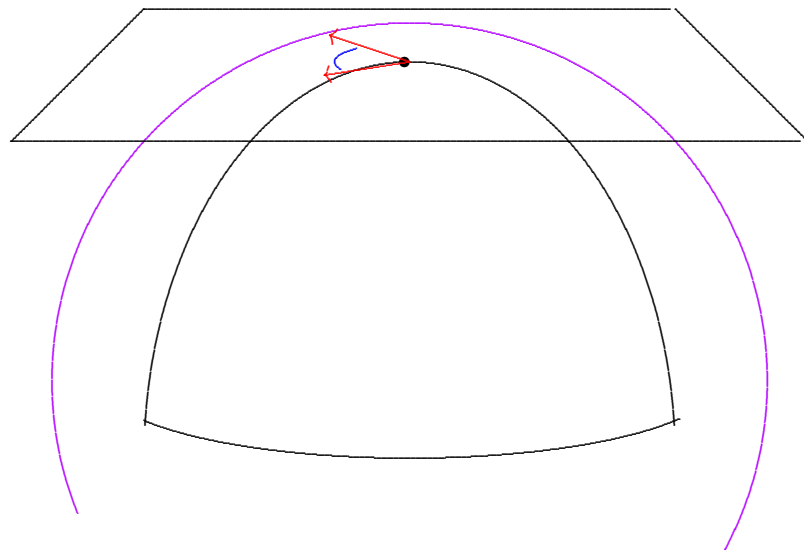
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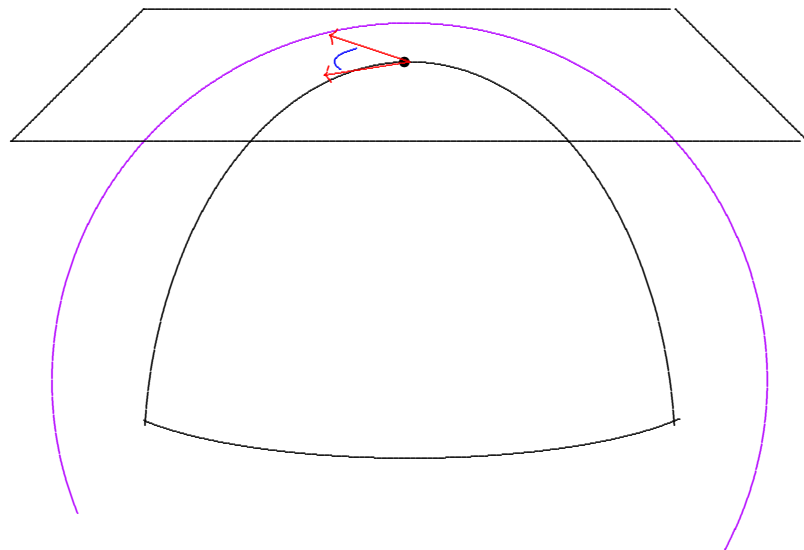
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$$\mathbf{Sp}(\ell) := \mathbf{O}(4\ell) \cap \mathbf{GL}(\ell, \mathbb{H})$$

Hyper-Kähler metrics:

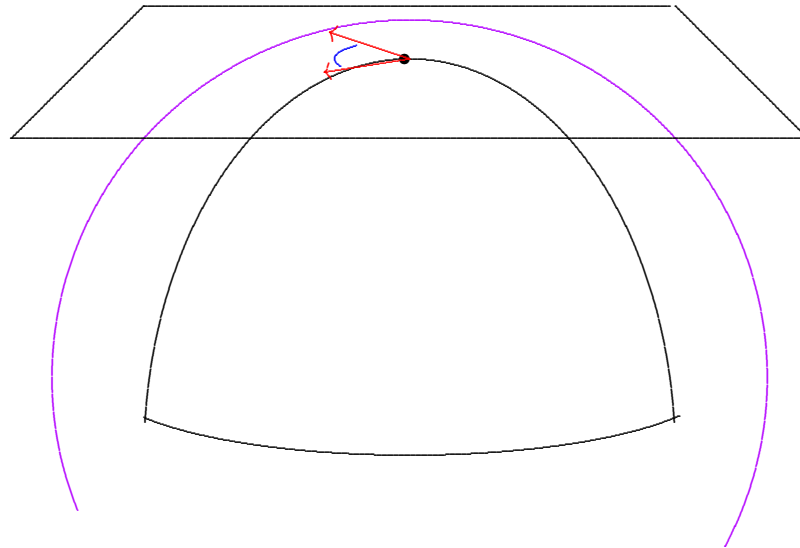
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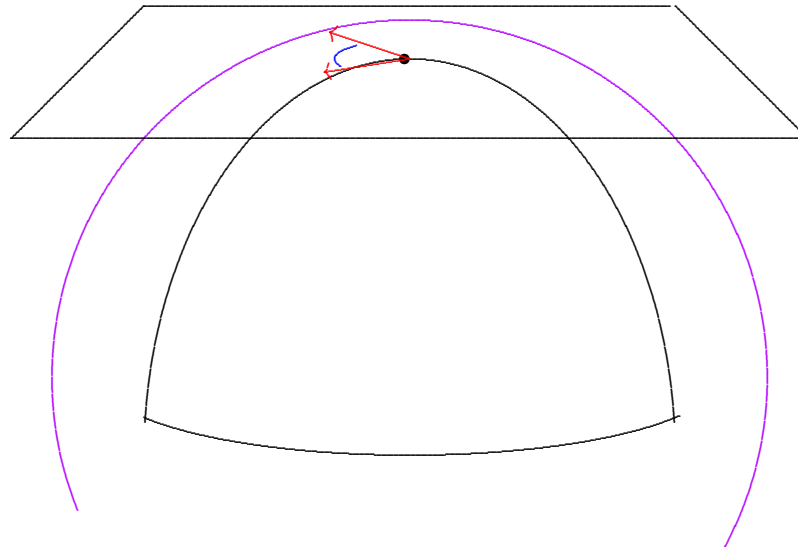


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in many ways!

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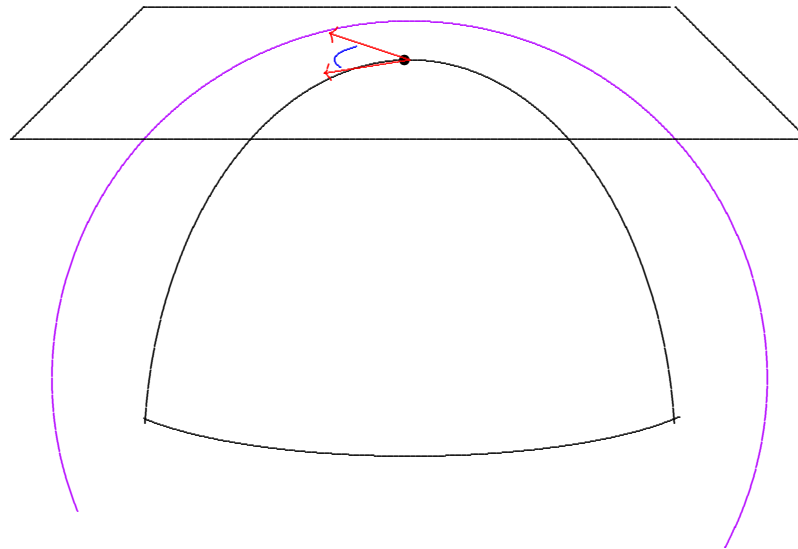


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in many ways! (For example, permute $i, j, k \dots$)

Hyper-Kähler metrics:

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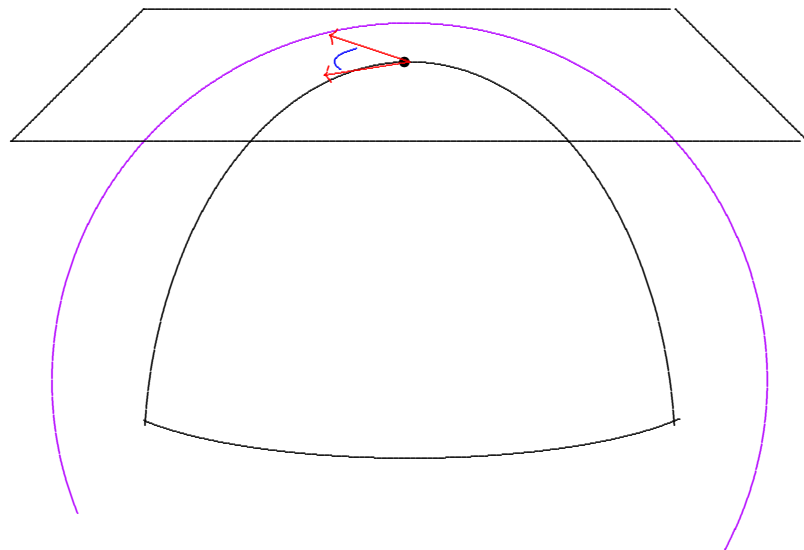
$$\mathbf{Sp}(\ell) \subset \mathbf{SU}(2\ell)$$

Ricci-flat and Kähler,

for many different complex structures!

Hyper-Kähler metrics:

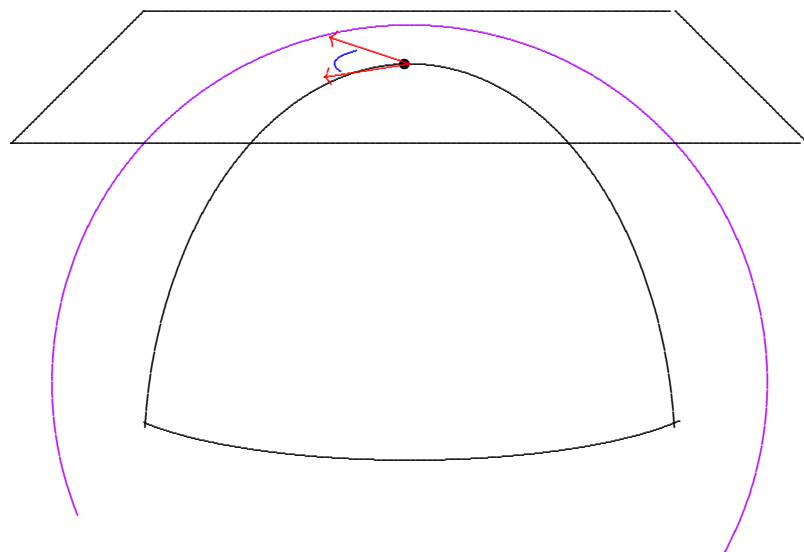
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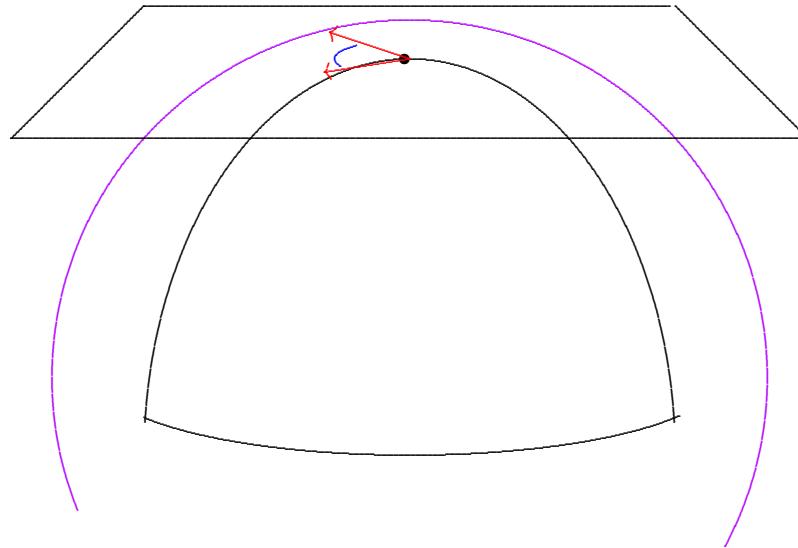
(M^4, g) hyper-Kähler \iff holonomy $\subset \mathbf{Sp}(1)$



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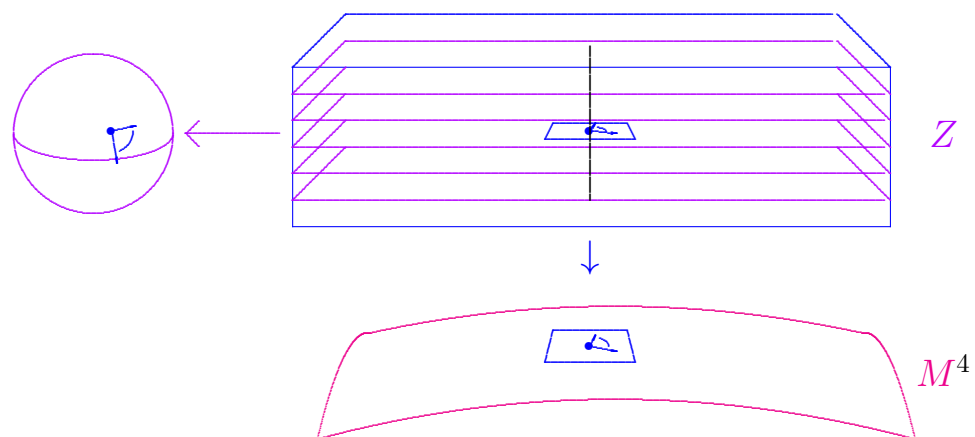


$$\mathbf{Sp}(1) = \mathbf{SU}(2)$$

When (M^4, g) simply connected:

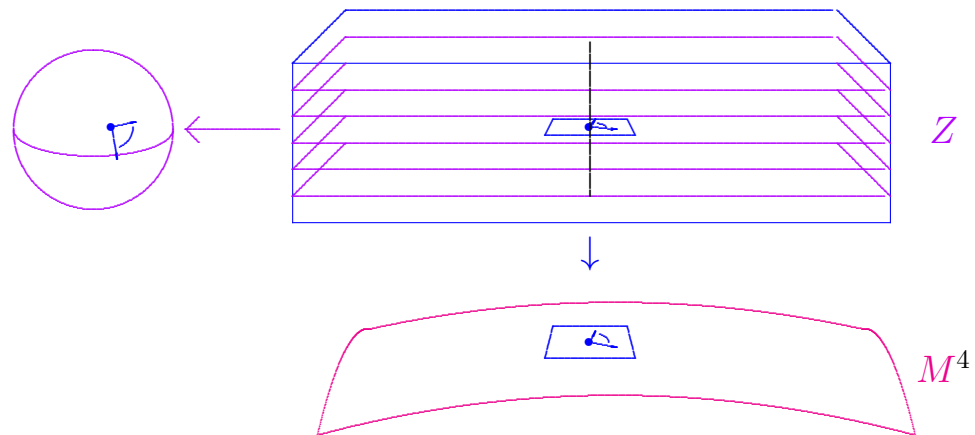
hyper-Kähler \iff Ricci-flat Kähler.

All these complex structures can be repackaged



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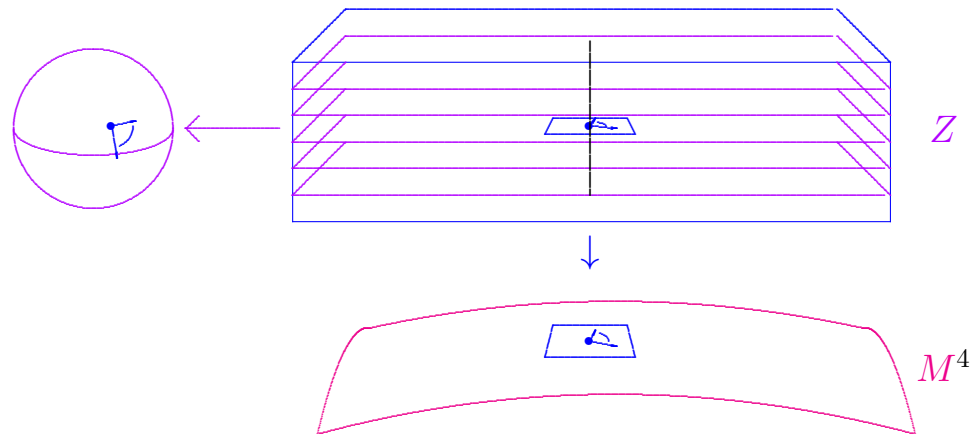
Penrose Twistor Space (Z, J) ,



All these complex structures can be repackaged as

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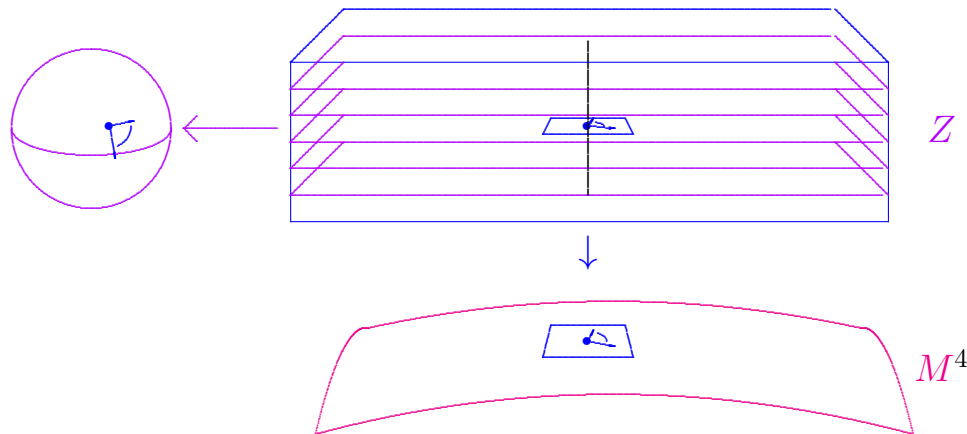
which is a complex 3-manifold.



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Penrose Twistor Space (Z, J) ,

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Riemannian non-linear graviton construction.

Key examples:

Term **ALE** coined by Gibbons & Hawking, 1979.

They wrote down various explicit **Ricci-flat ALE** 4-manifolds they called **gravitational instantons**.

Their examples have just one end, with

$$\Gamma \cong \mathbb{Z}_\ell \subset \mathbf{SU}(2) \subset \mathbf{O}(4).$$

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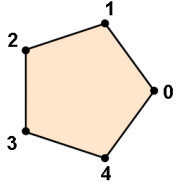
Hitchin conjectured that similar metrics would exist for each finite $\Gamma \subset \mathbf{SU}(2)$.

This conjecture was proved by Kronheimer, 1986.

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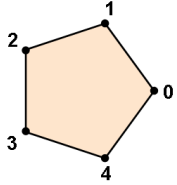
$$\mathbb{Z}_{k+1}$$



$$xy + z^{k+1} = 0$$

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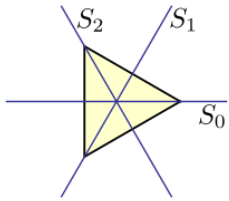
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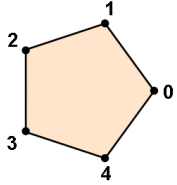
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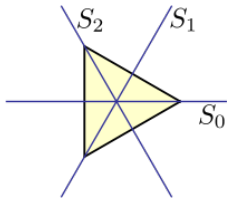
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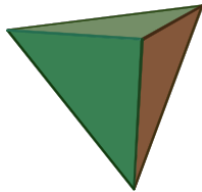
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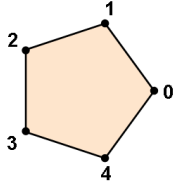
$$T^*$$



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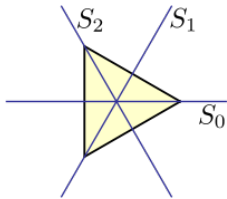
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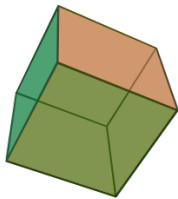
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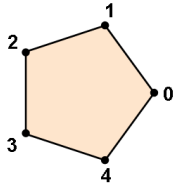
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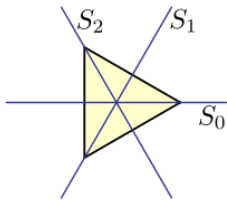
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$$\mathbb{Z}_{k+1}$$



$$xy + z^{k+1} = 0$$



$$\text{Dih}_{k-2}^*$$



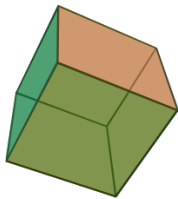
$$x^2 + z(y^2 + z^{k-2}) = 0$$



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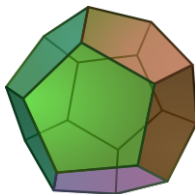
$$x^2 + y^3 + z^4 = 0$$



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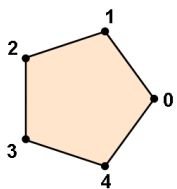
$$x^2 + y^3 + yz^3 = 0$$



$$I^*$$

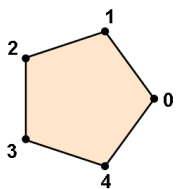


$$x^2 + y^3 + z^5 = 0$$

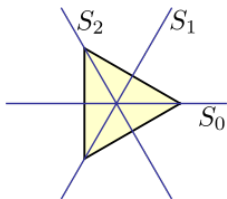


$$\mathbb{Z}_{k+1} \longleftrightarrow A_k$$

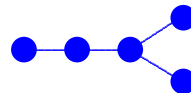


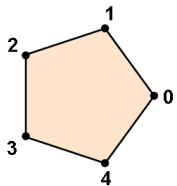


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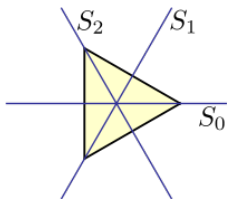


$$\text{Dih}_{k-2}^* \longleftrightarrow D_k$$

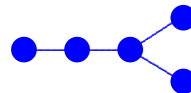




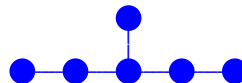
$$\mathbb{Z}_{k+1} \longleftrightarrow A_k$$

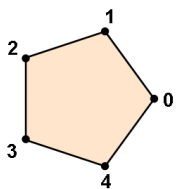


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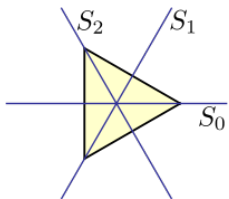


$$T^* \longleftrightarrow E_6$$

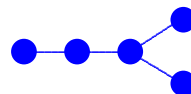




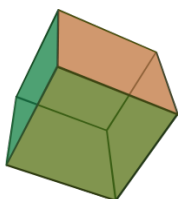
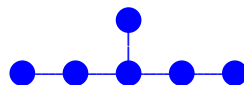
$$\mathbb{Z}_{k+1} \longleftrightarrow A_k$$



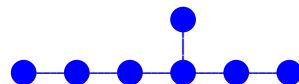
$$\text{Dih}_{k-2}^* \longleftrightarrow D_k$$

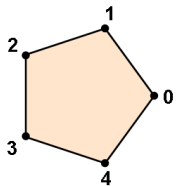


$$T^* \longleftrightarrow E_6$$

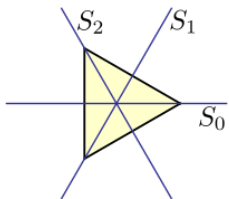


$$O^* \longleftrightarrow E_7$$

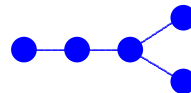




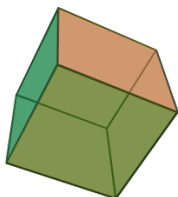
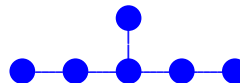
$$\mathbb{Z}_{k+1} \longleftrightarrow A_k$$



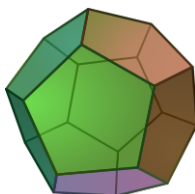
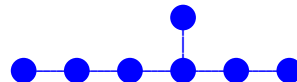
$$\text{Dih}_{k-2}^* \longleftrightarrow D_k$$



$$T^* \longleftrightarrow E_6$$



$$O^* \longleftrightarrow E_7$$

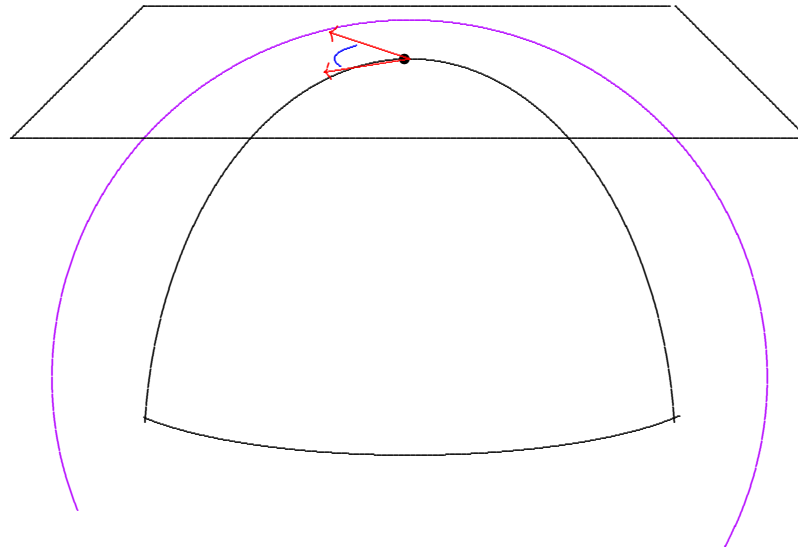


$$I^* \longleftrightarrow E_8$$



Hyper-Kähler metrics:

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$$\mathbf{Sp}(1) = \mathbf{SU}(2)$$

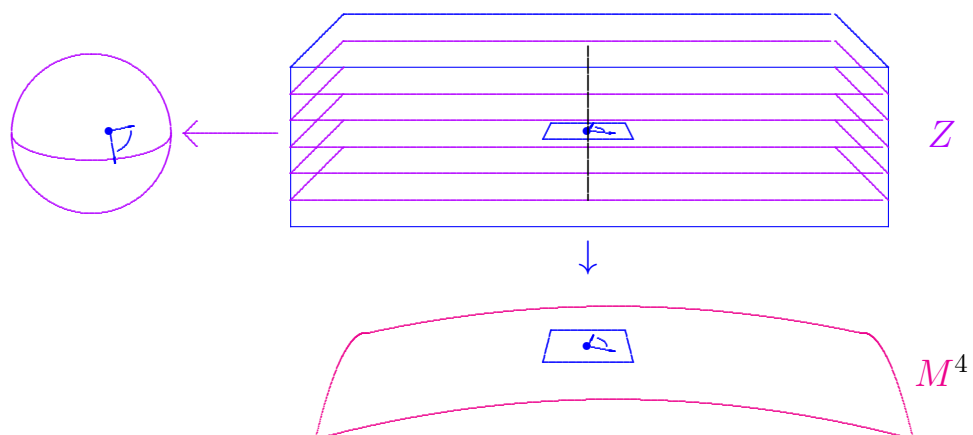
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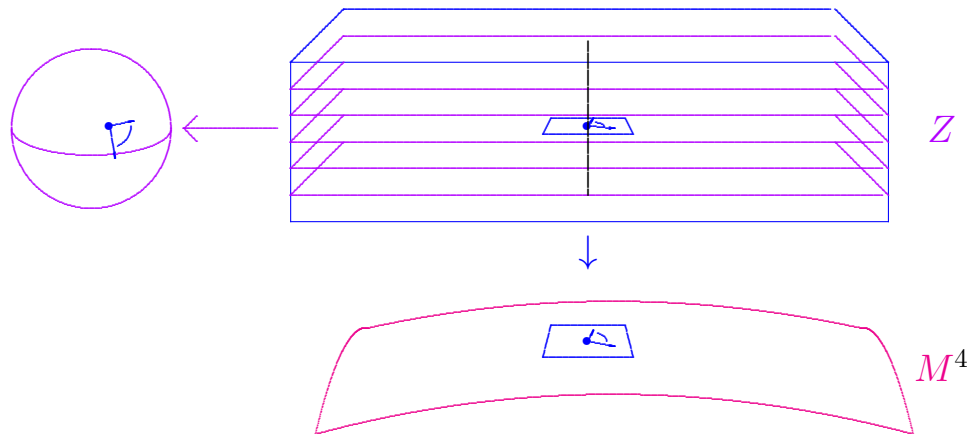
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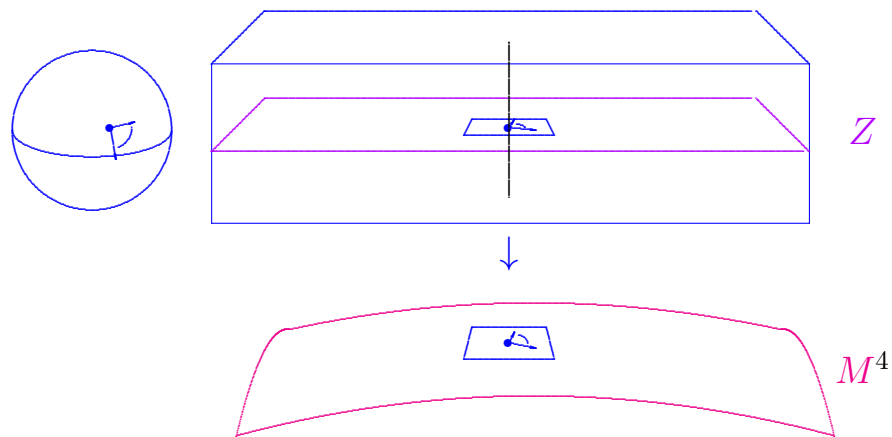


But similar for scalar-flat Kähler surfaces (M^4, g, J) !

Any scalar-flat Kähler surface (M^4, g, J) has a

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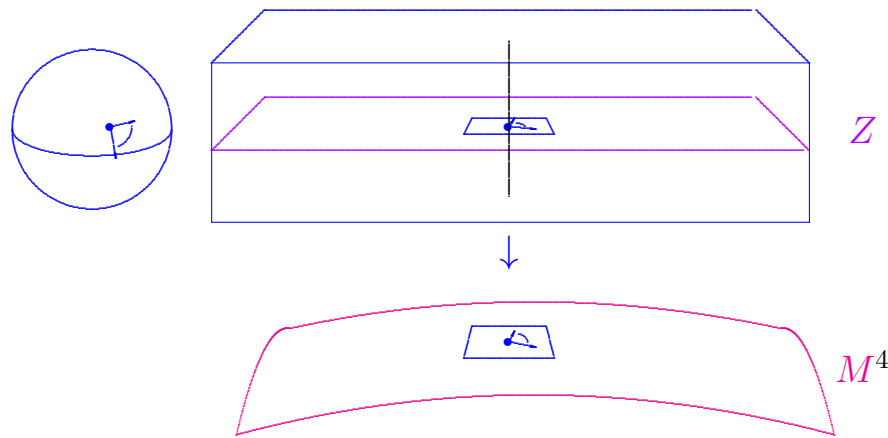
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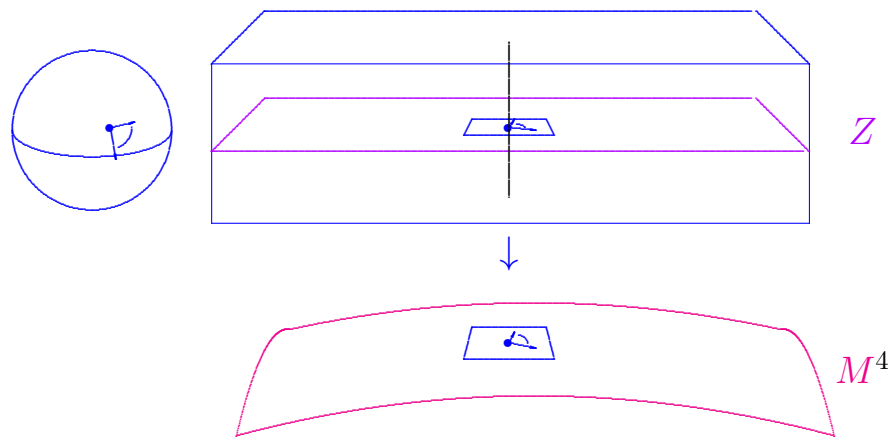


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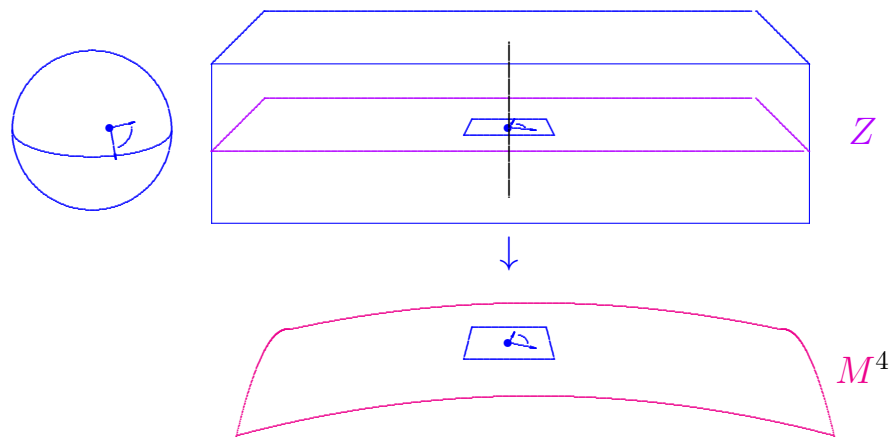
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Constructed ALE examples on line bundles $L \rightarrow \mathbb{CP}_1$ with $c_1 < 0$, and on their blow-ups.

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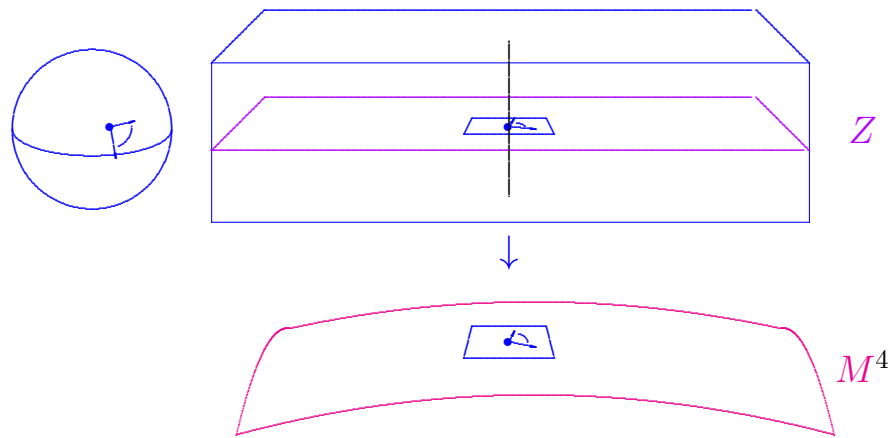
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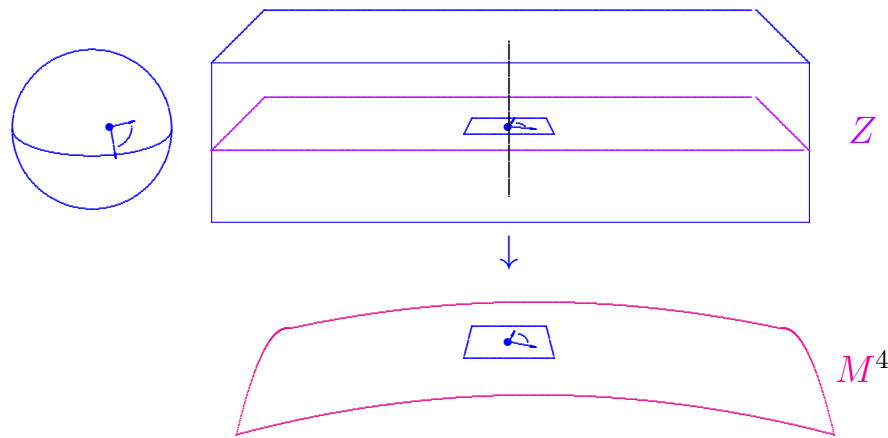


These **ALE** spaces arise naturally in the study of compact Einstein or Bach-flat 4-manifolds as bubbling modes for sequences of metrics.

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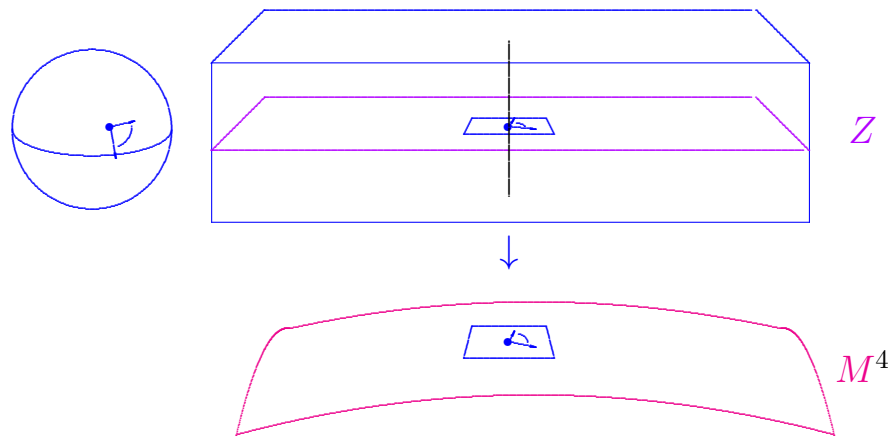


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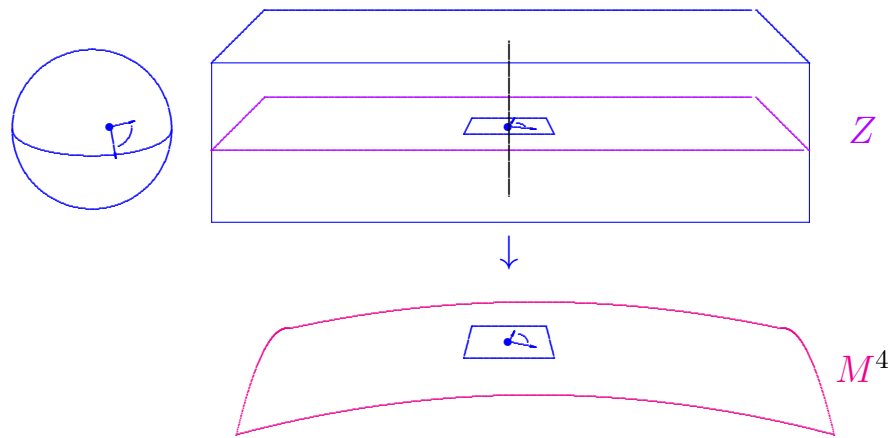
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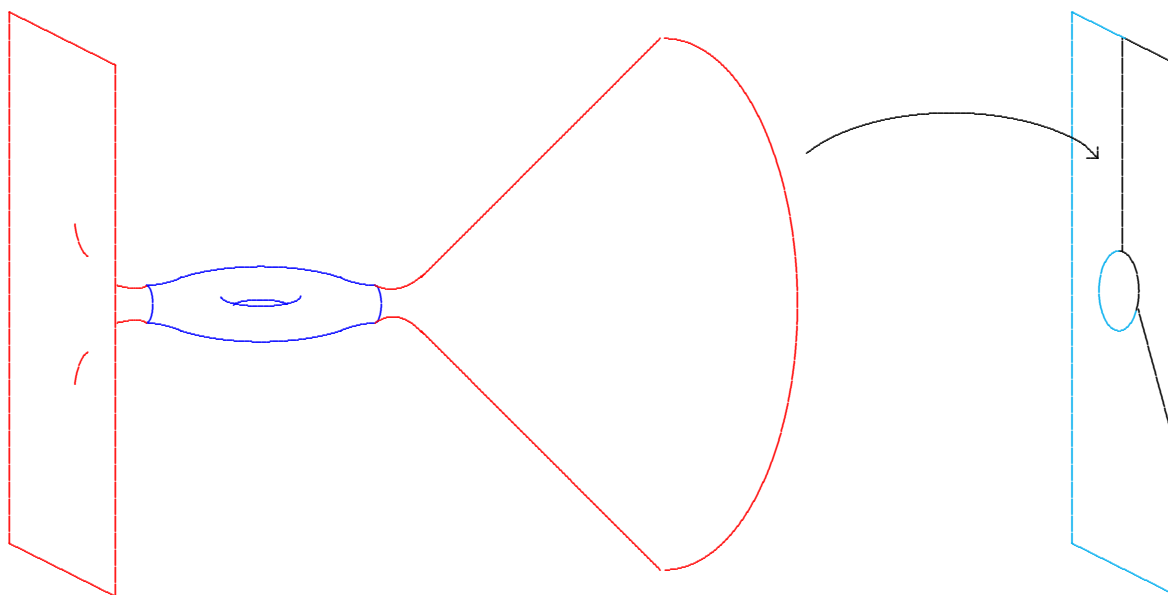


Lots more ALE scalar-flat Kähler surfaces now known:

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But full classification remains an open problem.

Definition. Complete, non-compact n -manifold (M^n, g) is asymptotically locally Euclidean (ALE) if \exists compact set $K \subset M$ such that $M - K \approx \coprod_i (\mathbb{R}^n - D^n) / \Gamma_i$, where $\Gamma_i \subset \mathbf{O}(\mathbf{n})$, such that



$$g_{jk} = \delta_{jk} + O(|x|^{1-\frac{n}{2}-\varepsilon})$$

$$g_{jk,\ell} = O(|x|^{-\frac{n}{2}-\varepsilon}), \quad s \in L^1$$

Mass still meaningful in this context...

Definition. *The mass (at a given end) of an ALE n -manifold is defined to be*

$$m(M, g) := \lim_{\varrho \rightarrow \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} [g_{ij,i} - g_{ii,j}] \nu^j \alpha_E$$

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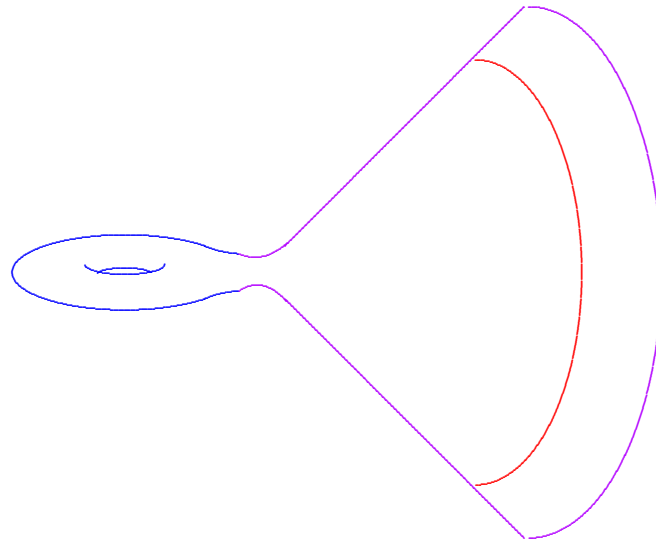
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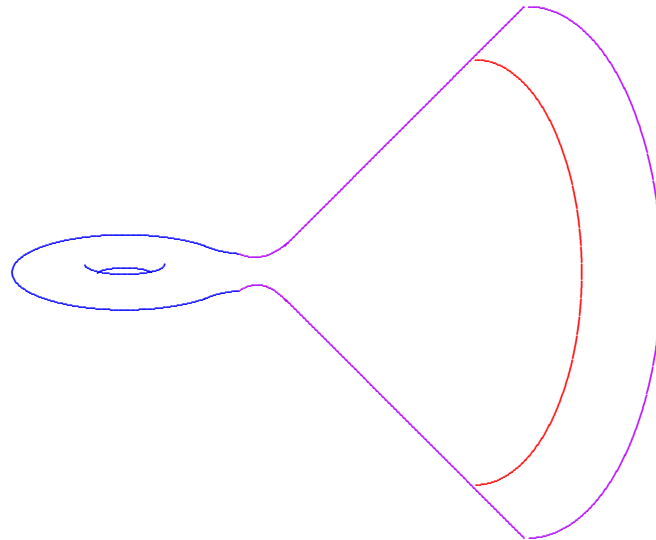


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Bartnik/Chruściel (1986): With weak fall-off conditions, the mass is well-defined & coordinate independent.

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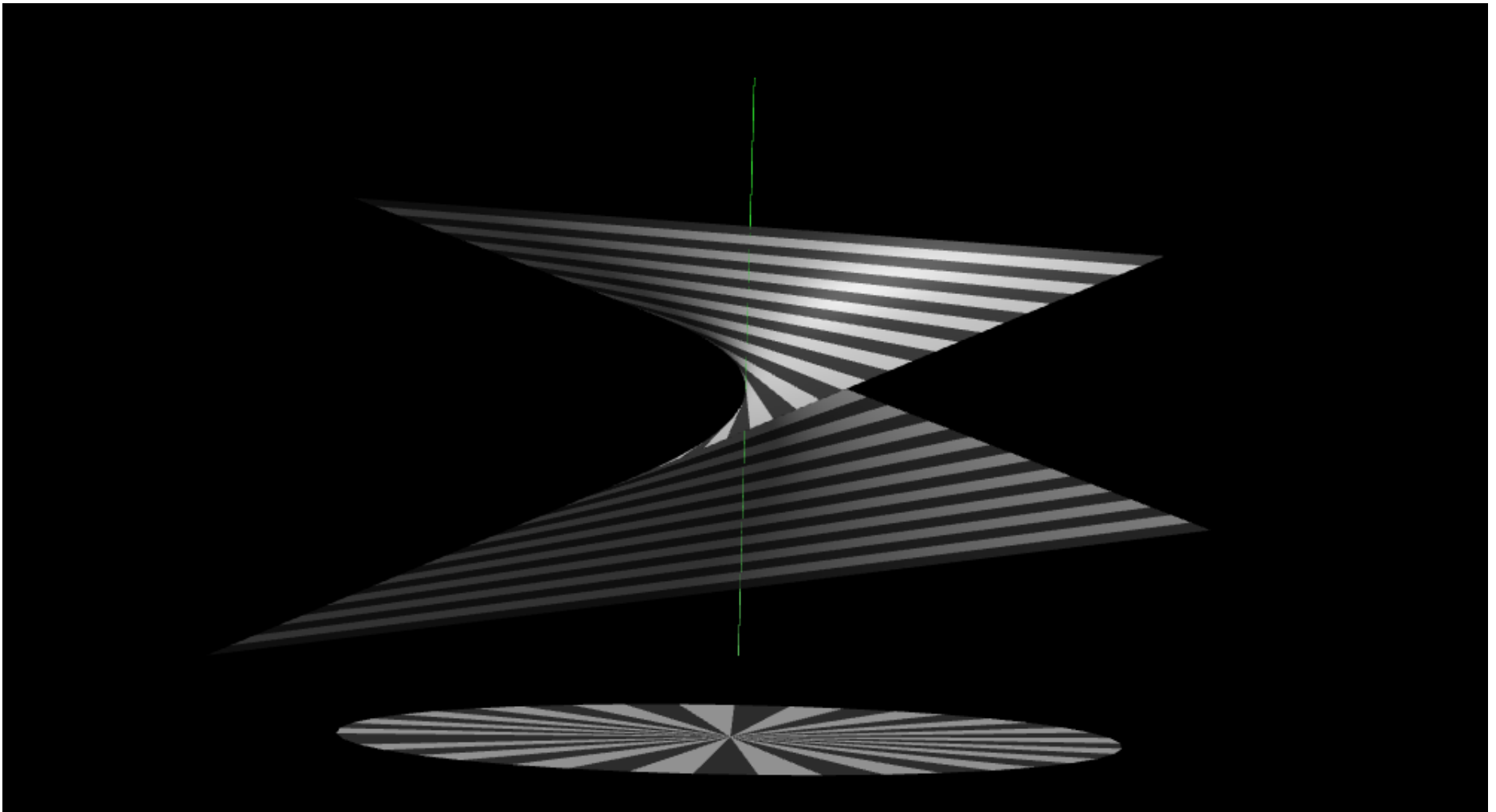
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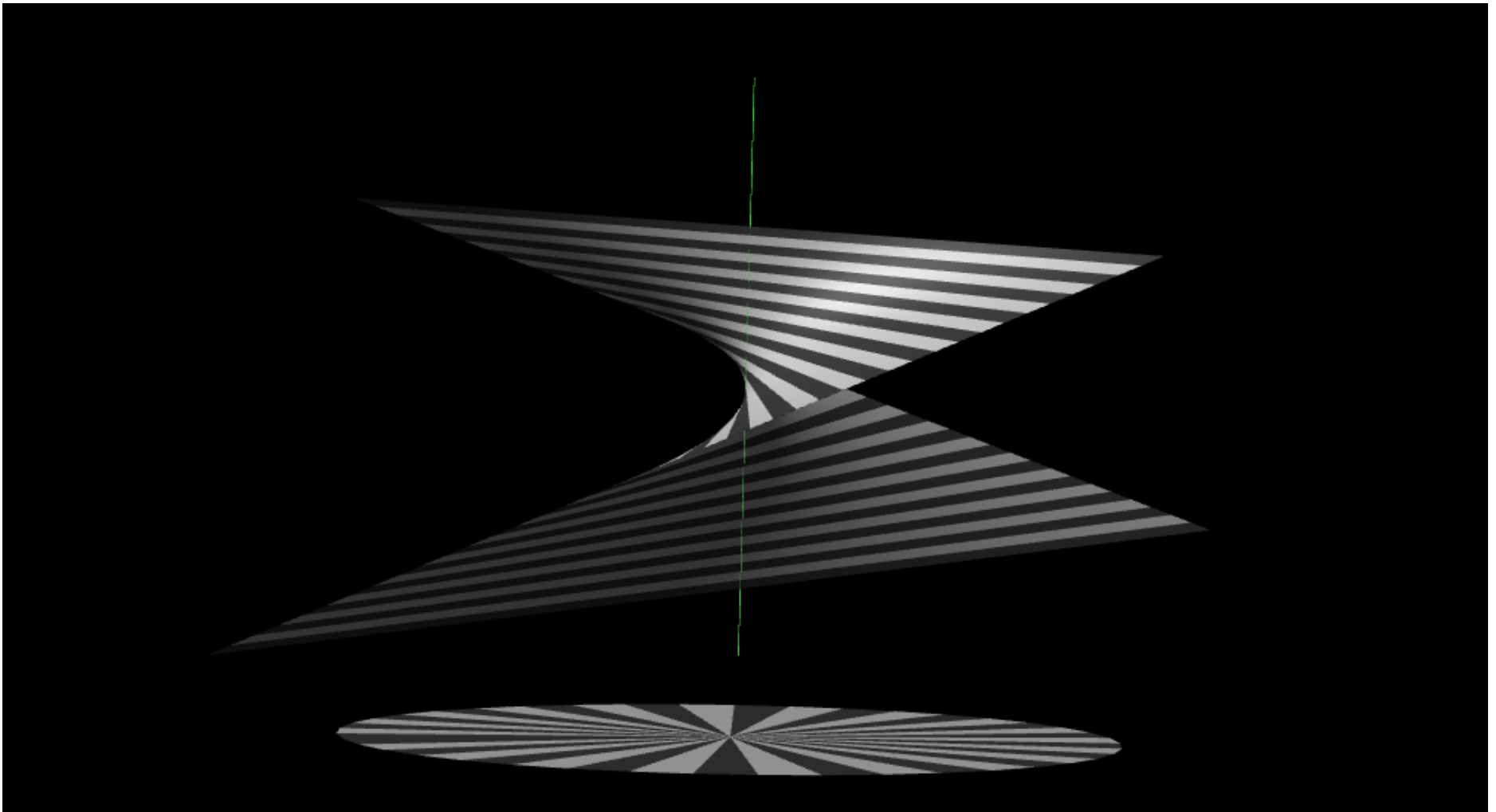
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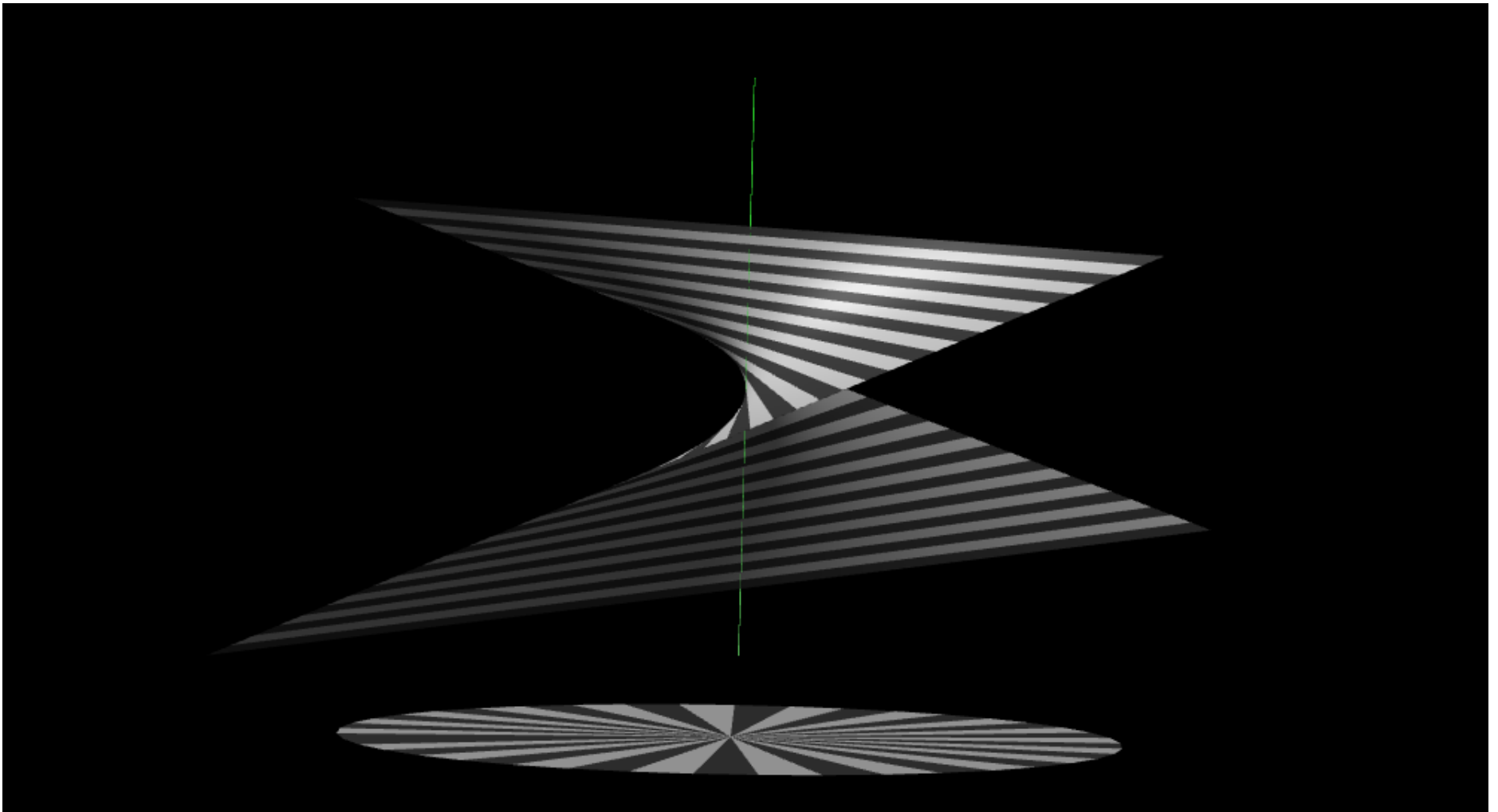
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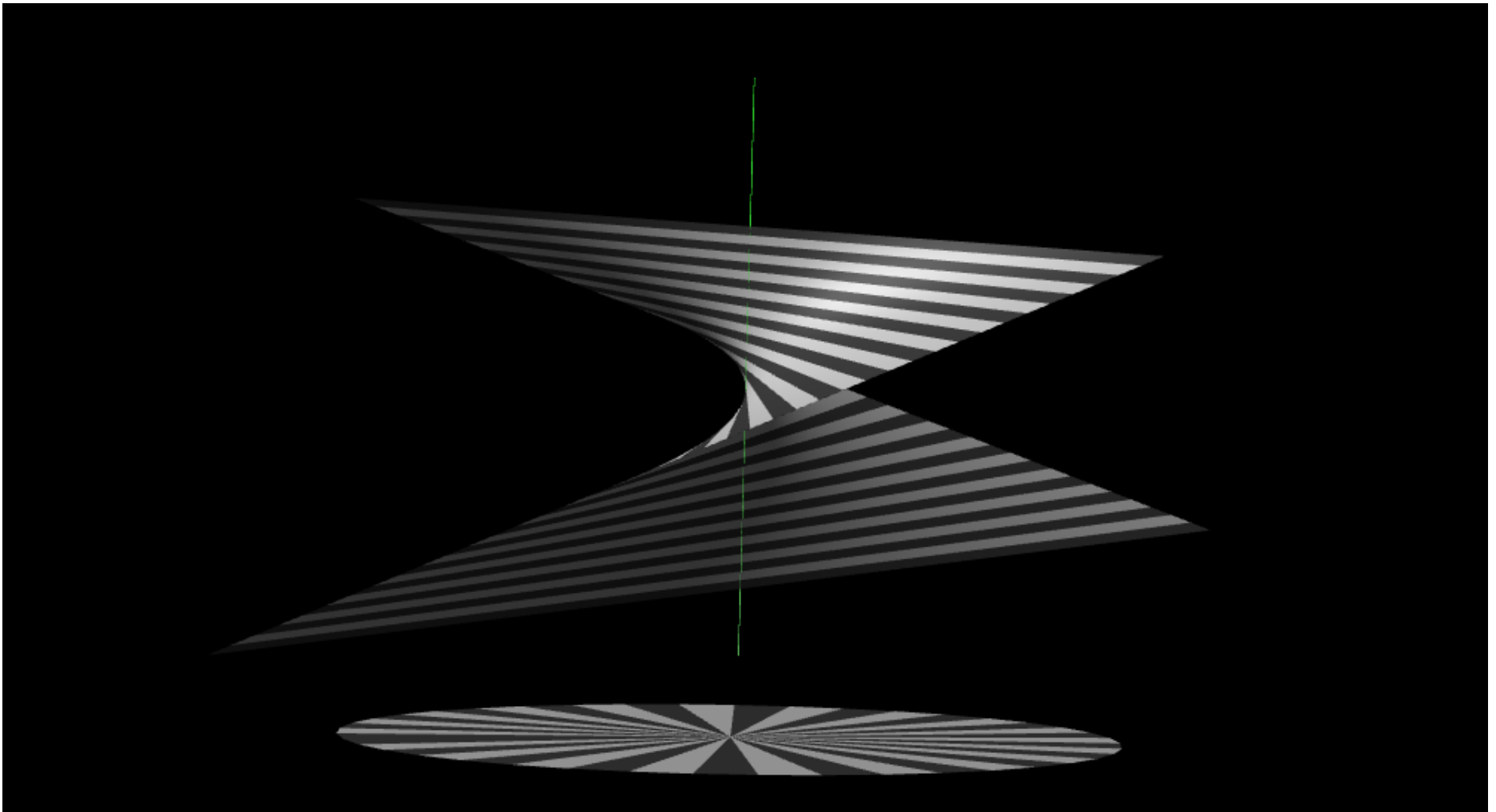
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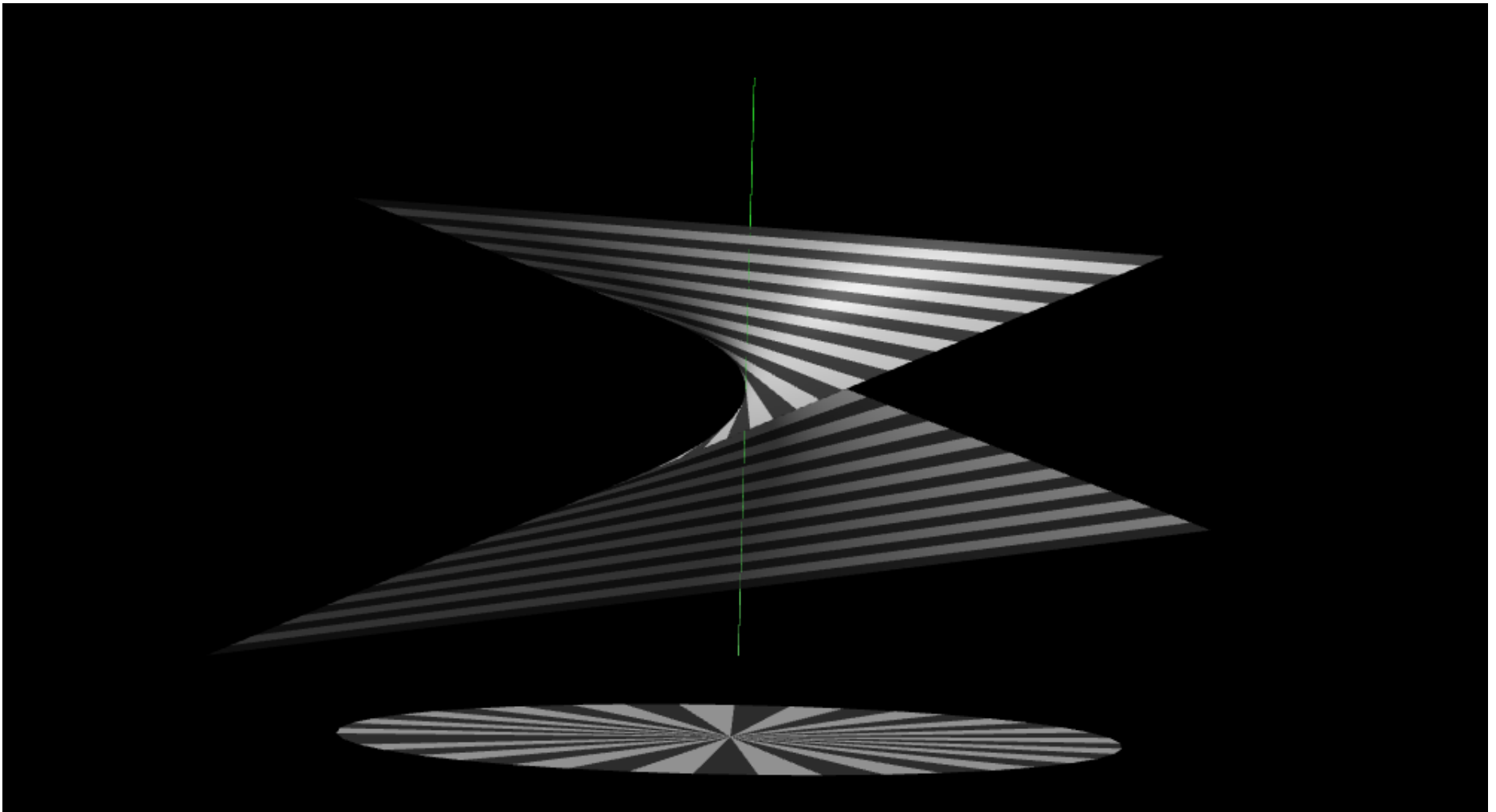
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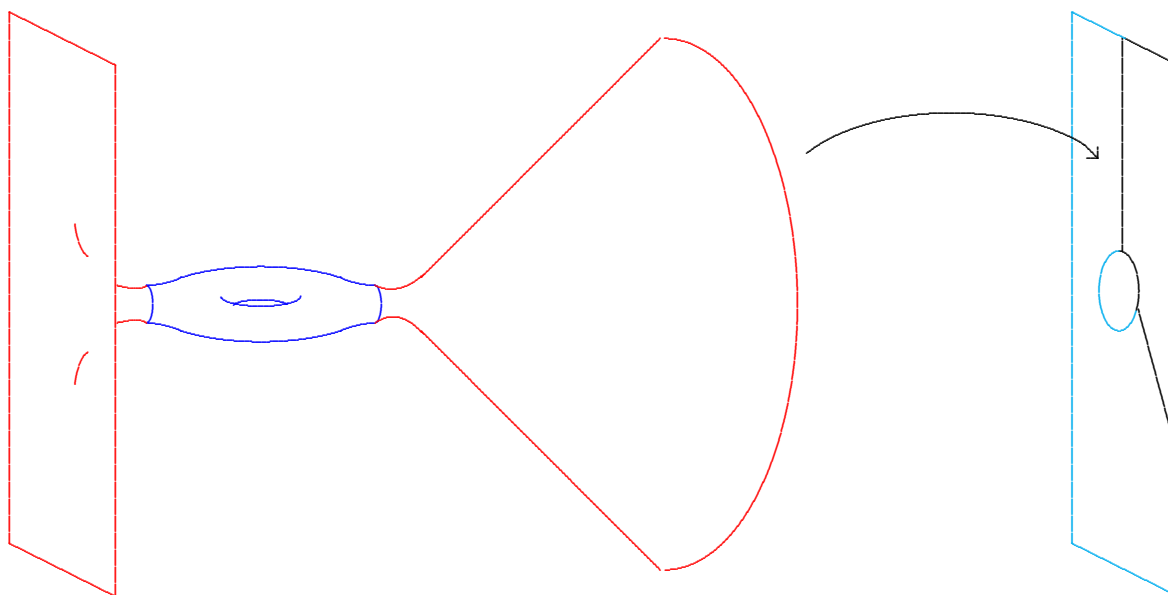
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Lemma. *Any ALE Kähler manifold*

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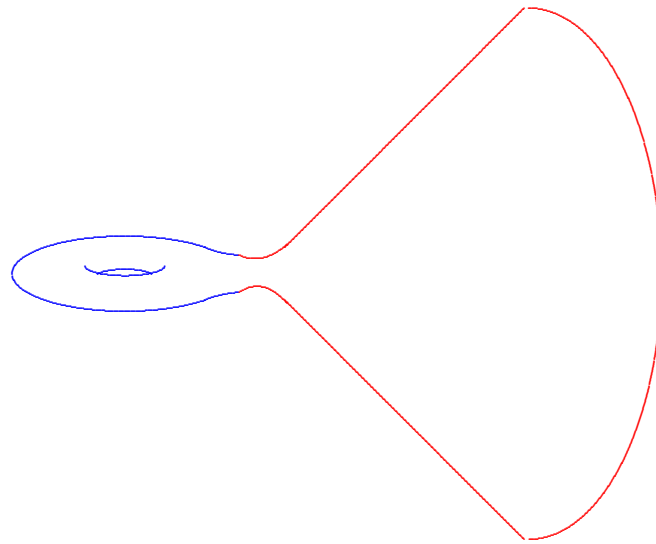
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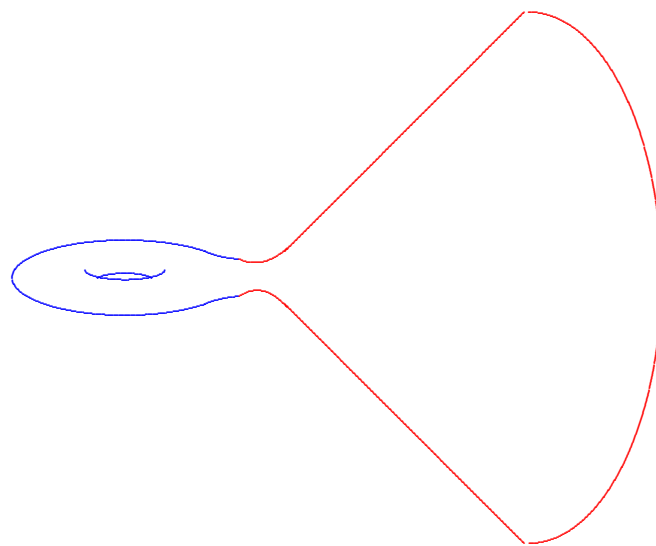
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Does not depend on the choice of an end!

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Theorem A.

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In fact, we will see that there is an explicit formula for the mass in terms of these data!

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Non-minimal resolutions typically admit families of such metrics for which the mass can be continuously deformed from negative to positive.

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- $\langle \cdot, \cdot \rangle$ is pairing between $H_c^2(M)$ and $H^{2m-2}(M)$.

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$$\frac{4\pi^m(2m-1)}{(m-1)!}m(M,g) = -\frac{4\pi}{(m-1)!}\langle \clubsuit(c_1), [\omega]^{m-1}\rangle + \int_M s_g d\mu_g$$

For a compact Kähler manifold (M^{2m}, g, J) ,

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So the mass is a “boundary correction” to the topological formula for the total scalar curvature.

Theorem C. Any *ALE Kähler manifold* (M, g, J) of complex dimension m has mass given by

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Corollary. Any *ALE* scalar-flat Kähler manifold (M, g, J) of complex dimension m has mass given by

$$m(M, g) = -\frac{\langle \clubsuit(c_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}}.$$

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So **Theorem A** is an immediate consequence!

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$$g^{jk} (g_{j\ell,k} - g_{jk,\ell}) \nu^\ell \alpha_E = -\star d \log \left(\sqrt{\det g} \right) + O(\varrho^{-3-\varepsilon}).$$

$$m(M,g)=-\lim_{\varrho\rightarrow\infty}\frac{1}{12\pi^2}\int_{S_{\varrho}/\Gamma}\star d\left(\log\sqrt{\det g}\right)$$

$$m(M, g) = - \lim_{\varrho \rightarrow \infty} \frac{1}{12\pi^2} \int_{S_{\varrho}/\Gamma} \star d \left(\log \sqrt{\det g} \right)$$

Now set $\theta = \frac{i}{2}(\partial - \bar{\partial}) \left(\log \sqrt{\det g} \right)$, so that

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However, since $s = 0$,

$$d(\theta \wedge \omega) = \rho \wedge \omega = \frac{s}{4} \omega^2 = 0.$$

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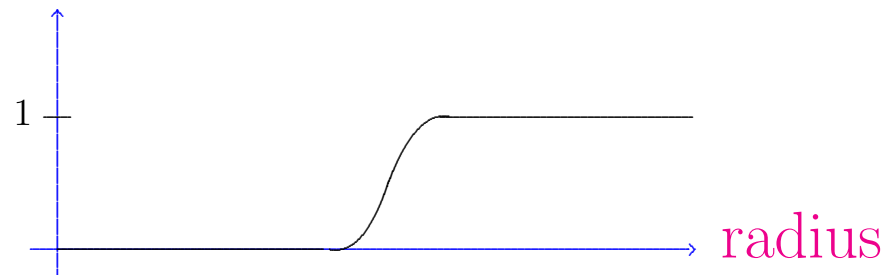
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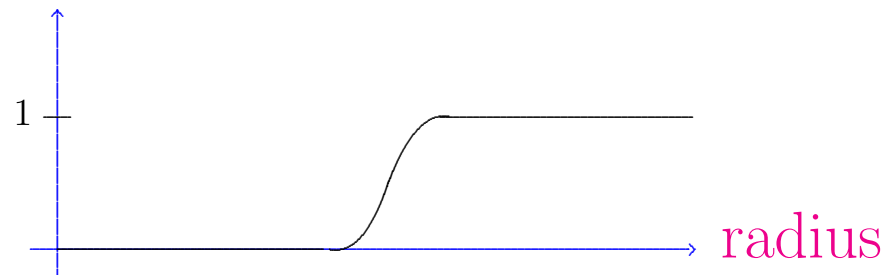
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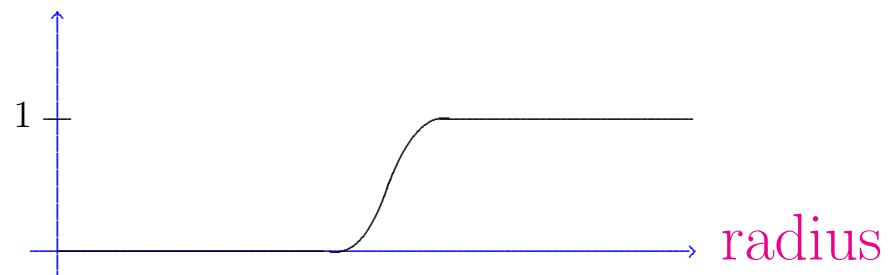
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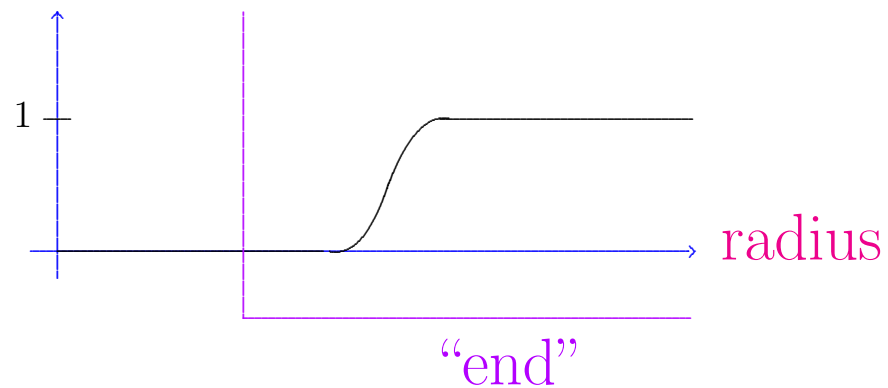
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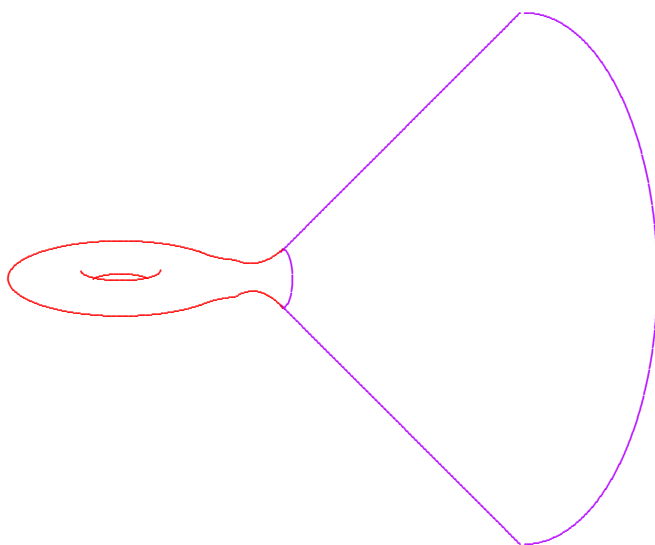
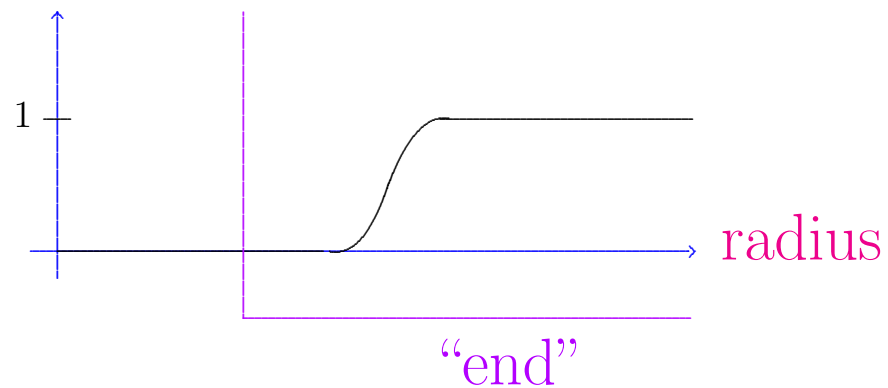
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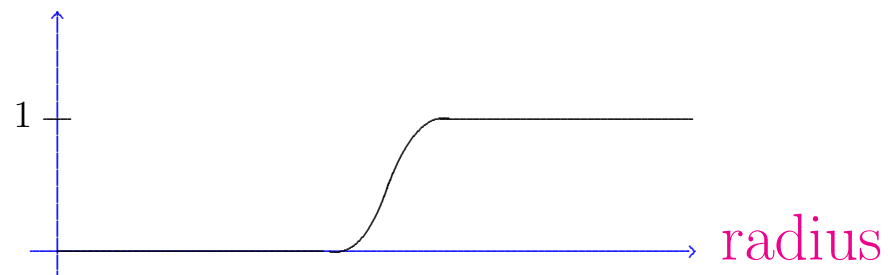
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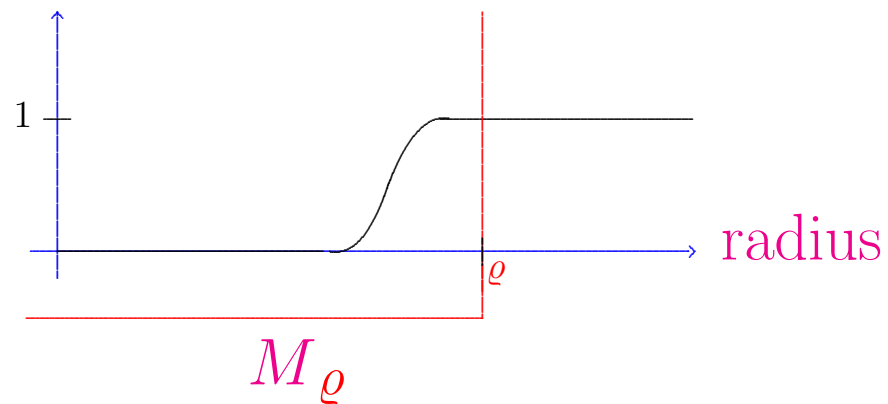
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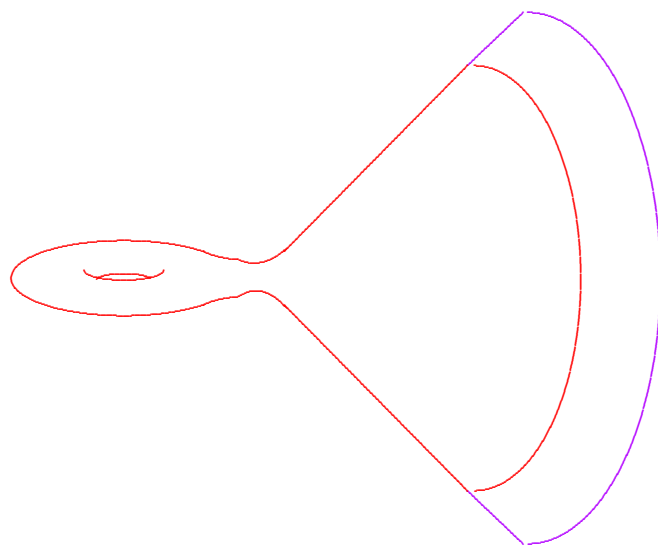
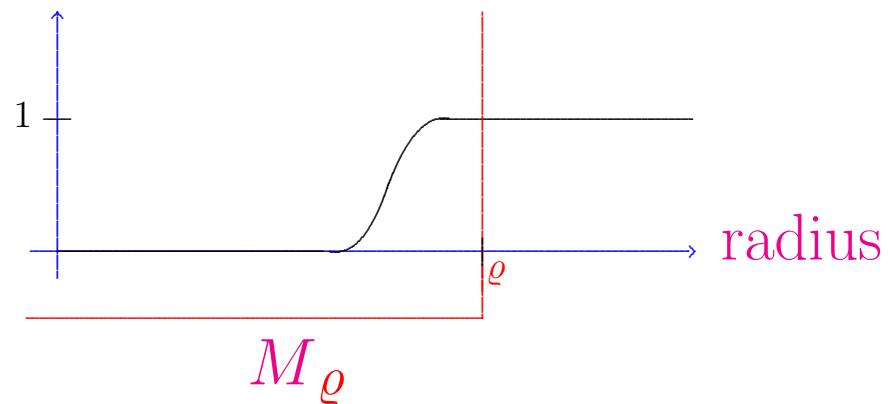
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Compactly supported, because $d\theta = \rho$ near infinity.

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where M_ϱ defined by radius $\leq \varrho$.

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by Stokes' theorem.

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- $s \equiv 0$; and
- Complex structure J standard at infinity.

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Seen in “gravitational instantons”

and other explicit examples.

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$$J = J_0 + O(\varrho^{-3}), \quad \nabla J = O(\varrho^{-4})$$

in suitable asymptotic coordinates adapted to g .

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Complete analytic family encodes info about J .

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This has some interesting consequences...

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Proof actually shows something stronger!

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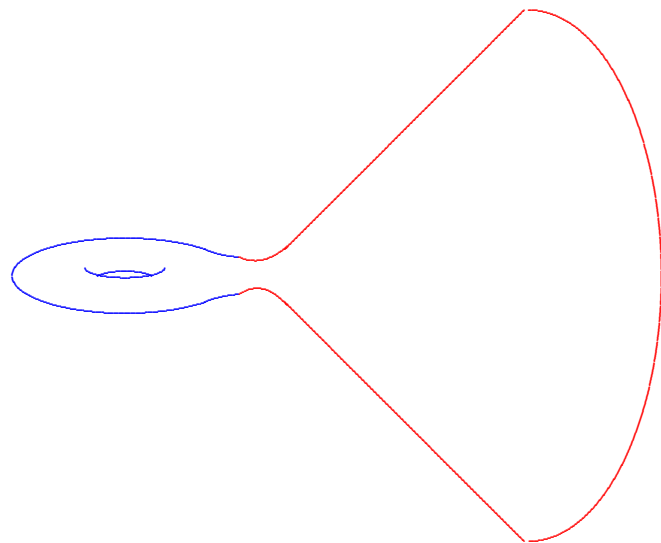
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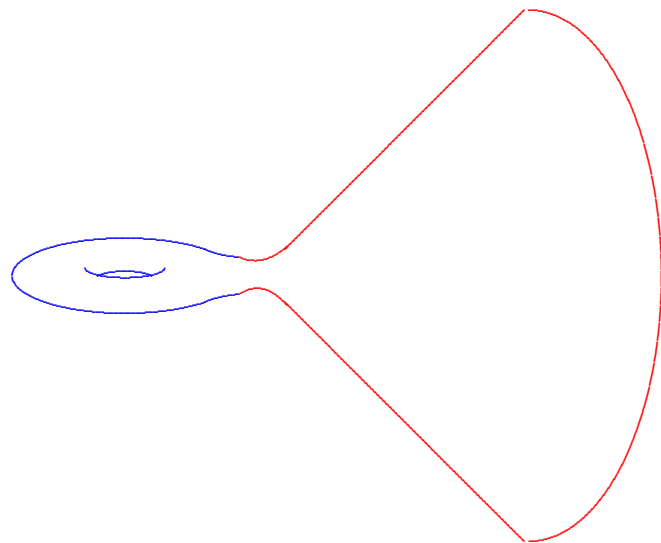
so the mass formula implies the claim.

$$m(M, g) = -\frac{\langle \clubsuit(c_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M s_g d\mu_g$$



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