

Quantum Mechanics

Lecture Notes

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1. sheet-1 :Dirac Delta Function

Consider the function $D_\varepsilon(x)$ given by

$$D_\varepsilon(x) = \begin{cases} \frac{1}{\varepsilon} & \text{for } -\frac{\varepsilon}{2} \leq x \leq \frac{\varepsilon}{2} \\ 0 & \text{for } |x| > \frac{\varepsilon}{2} \end{cases} \quad (1.1)$$

where ε is a positive parameter. The plot of the function is shown below

The integral of the function with respect to x is 1, i.e.,

$$\int_{-\infty}^{\infty} D_\varepsilon(x) dx = 1 \quad (1.2)$$

Now imagine making ε smaller. As we decrease ε , the function gets narrower and taller, but the integral of the function(i.e., the area under graph remains constant at the value 1). In the limit $\varepsilon \rightarrow 0$, the function $D_\varepsilon(x)$ collapses to a single point $x = 0$ and gets infinitely tall. So $\lim_{\varepsilon \rightarrow 0} D_\varepsilon(x)$ is not a function at all and the procedure of taking the limit is not justified.

However, we can make the limiting procedure meaningful if multiply $D_\varepsilon(x)$ by some well defined function $f(x)$, integrate over x and then take the limit $\varepsilon \rightarrow 0$. consider the integral

$$\int_{-\infty}^{\infty} D_\varepsilon(x) f(x) dx \quad (1.3)$$

where $f(x)$ is a well-defined function. If ε is significantly small, the variation of $f(x)$ over the effective integration interval $[-\varepsilon/2, \varepsilon/2]$ is negligible and $f(x)$ remains practically equal to $f(0)$, therefore,

$$\int_{-\infty}^{\infty} D_\varepsilon(x) f(x) dx \simeq f(0) \int_{-\infty}^{\infty} D_\varepsilon(x) dx = f(0) \quad (1.4)$$

The smaller the value of ε , the better the approximation. In the limit $\varepsilon \rightarrow 0$, the above equation is exact

$$\lim_{\varepsilon \rightarrow 0} \int_{-\infty}^{\infty} D_\varepsilon(x) f(x) dx = f(0) \quad (1.5)$$

Now, we define the delta function by the relation

$$\int_{-\infty}^{\infty} \delta(x)f(x)dx \stackrel{def}{=} \lim_{\epsilon \rightarrow 0} \int_{-\infty}^{\infty} D_{\epsilon}(x)f(x)dx = f(0) \quad (1.6)$$

This equation is valid for any function $f(x)$ defined at the origin. More generally, $\delta(x-x_0)$ is defined as,

$$\int_{-\infty}^{\infty} \delta(x-x_0)f(x)dx = f(x_0) \quad (1.7)$$

Actually, the integral notation $\int_{-\infty}^{\infty} \delta(x)f(x)dx$ is not justified because $\delta(x)$ is not really a function. Physically, there is no problem since it becomes impossible to distinguish between $D_{\epsilon}(x)$ and $\delta(x)$ as soon as ϵ becomes negligible compared to all distances involved in a physical problem. Whenever a mathematical difficulty might arise, all we need to do is to assume that $\delta(x)$ is actually $D_{\epsilon}(x)$ with ϵ extremely small but not strictly zero.

Formally, we can express $\delta(x)$ as a limit of a square of proper functions :

$$\lim_{\epsilon \rightarrow 0} D_{\epsilon}(x) \equiv \delta(x) \quad (1.8)$$

Here $D_{\epsilon}(x)$, which is a proper function of x is called the representation of the delta function. One representation is the "square function" given at the beginning. The representation is not unique. There are other functions which approach the delta function when appropriate limits are taken.

1.1 Other representation of delta function

1. Consider the function

$$D_{\epsilon}(x) = \frac{1}{\epsilon\sqrt{\pi}} e^{-x^2/\epsilon^2} \quad (\epsilon > 0) \quad (1.9)$$

For each value of the parameter ϵ , this function satisfies

$$\int_{-\infty}^{\infty} D_{\epsilon}(x)dx = 1 \quad (1.10)$$

When plotted against x , the function has a peak at the origin. The peak has a height of $\frac{1}{\epsilon\sqrt{\pi}}$ and a width of order ϵ (exactly how the width is defined doesn't matter). So if ϵ is allowed to become very small, the peak becomes very tall and very narrow. Outside the peak the function becomes extremely small. Thus we have

$$\delta(x) = \lim_{\epsilon \rightarrow 0} \frac{1}{\epsilon\sqrt{\pi}} e^{-x^2/\epsilon^2} \quad (1.11)$$

2. Consider another function

$$D_{\epsilon}(x) = \frac{1}{\pi} \frac{\sin(x/\epsilon)}{x} \quad (\epsilon > 0) \quad (1.12)$$

For any value of the parameter ϵ we have

$$\int_{-\infty}^{\infty} D_{\epsilon}(x)dx = 1 \quad (1.13)$$

A plot of the function $D_\varepsilon(x)$ shows that it has the value $\frac{1}{\varepsilon\pi}$ at $x = 0$ and it oscillates with decreasing amplitude as $|x|$ increases. The width of the central maxima is of the order of ε and the period of oscillation with respect to x is $2\pi\varepsilon$.

Thus the limit of this function as $\varepsilon \rightarrow 0$ has all the properties of the delta function : it becomes infinitely large at $x = 0$, it has unit integral, and infinitely rapid oscillations as $|x|$ increases means that the entire contribution to an integral containing this function comes from an infinitesimal neighborhood of $x = 0$.

We can therefore write,

$$\delta(x) = \lim_{\varepsilon \rightarrow 0} \frac{1}{\pi} \frac{\sin(x/\varepsilon)}{x} \quad (1.14)$$

3. We can also show that

$$\delta(x) = \lim_{\varepsilon \rightarrow 0} \frac{1}{2\varepsilon} e^{-|x|/\varepsilon} \quad (1.15)$$

$$\delta(x) = \lim_{\varepsilon \rightarrow 0} \frac{1}{\pi} \frac{\varepsilon}{x^2 + \varepsilon^2} \quad (1.16)$$

$$\delta(x) = \lim_{\varepsilon \rightarrow 0} \frac{\varepsilon}{\pi} \frac{\sin^2(x/\varepsilon)}{x^2} \quad (1.17)$$

1.2 Properties of the delta function

It is important to note that, because of its singular @@@@, the δ function cannot be the end result of a calculation, and has meaning only so long as a subsequent integral over its argument is carried out. With this understanding we can write down some relations between delta functions.

Property 1 The delta function is a even function

$$\delta(-x) = \delta(x) \quad (1.18)$$

Property 2

$$x\delta(x) = 0 \quad (1.19)$$

Property 3

$$\delta(ax) = \frac{1}{|a|} \delta(x) \quad (1.20)$$

proof: Consider the integral

$$I = \int_{-\infty}^{\infty} \delta(ax)f(x)dx \quad (1.21)$$

Since the delta function is even in its argument, it doesn't matter if we replace a by $|a|$ in the argument. Thus

$$I = \int_{-\infty}^{\infty} \delta(|a|x)f(x)dx \quad (1.22)$$

Making the change in variable $y = |a|x$ we have,

$$I = \frac{1}{|a|} \int_{-\infty}^{\infty} \delta(y)f(y/|a|)dy$$

$$= \frac{1}{|a|} f(0)$$

or

$$\int_{-\infty}^{\infty} \delta(ax)f(x)dx = \frac{1}{|a|} \int_{-\infty}^{\infty} \delta(x)f(x)dx \quad (1.23)$$

or

$$\delta(ax) = \frac{1}{|a|} \delta(x) \quad (1.24)$$

Property 4 More generally

$$\delta(\phi(x)) = \sum_i \frac{\delta(x - x_i)}{\left| \frac{\partial \phi}{\partial x} \right|_{x=x_i}} \quad (1.25)$$

where the sum sums over the x_i 's which are simple roots of $\phi(x)$.

proof : let x_1, x_2, \dots, x_N be the simple roots of $\phi(x)$,

In the neighborhood of any one of the simple roots x_i , we can write

$$\phi(x) = (x - x_i)\psi(x) \quad (1.26)$$

where $\psi(x_i) \neq 0$. We have

$$\psi(x_i) = \left| \frac{\partial \phi(x)}{\partial x} \right|_{x=x_i} \quad (1.27)$$

Now, consider the integral

$$\begin{aligned} I &= \int_{-\infty}^{\infty} \delta(\phi(x))f(x)dx \\ &= \sum_{i=1}^N \int_{x_i-\epsilon}^{x_i+\epsilon} \delta[(x - x_i)\psi(x_i)]f(x)dx \\ &= \sum_{i=1}^N \frac{1}{|\psi(x_i)|} \int_{x_i-\epsilon}^{x_i+\epsilon} \delta(x - x_i)f(x)dx \\ &= \sum_{i=1}^N \frac{1}{\left| \frac{\partial \phi(x)}{\partial x} \right|_{x=x_i}} \int_{-\infty}^{\infty} \delta(x - x_i)f(x)dx \\ &= \sum_{i=1}^N \frac{1}{\left| \frac{\partial \phi}{\partial x} \right|_{x=x_i}} f(x_i) \end{aligned}$$

The above result is obtained if we write

$$\delta(\phi(x)) = \sum_{i=1}^N \frac{\delta(x - x_i)}{\left| \frac{\partial \phi}{\partial x} \right|_{x=x_i}} \quad (1.28)$$

Property 5 A frequently used example of the above result is

$$\delta(x^2 - a^2) = \frac{1}{2a} \delta(x - a) + \frac{1}{2a} \delta(x + a) \quad (a > 0) \quad (1.29)$$

Here

$$\phi(x) = x^2 - a^2 = (x - a)(x + a) \quad (1.30)$$

The two simple roots of $\phi(x)$ are at $x = a$ and $x = -a$. Now

$$\left| \frac{\partial \phi}{\partial x} \right|_{x=a} = |2x|_{x=a} = 2a \quad (1.31)$$

$$\left| \frac{\partial \phi}{\partial x} \right|_{x=-a} = |-2x|_{x=-a} = 2a \quad (1.32)$$

\therefore The above result follows.

Property 6

$$f(x)\delta(x-a) = f(a)\delta(x-a) \quad (1.33)$$

Property 7

$$\int \delta(x-y)\delta(y-a)dy = \delta(x-a) \quad (1.34)$$

1.2.1 Notes

Note : 1 We have the identity

$$x\delta(x) = 0 \quad (1.35)$$

The converse is also true and it can be shown that the equation

$$xu(x) = 0 \quad (1.36)$$

has the general solution

$$u(x) = c\delta(x) \quad (1.37)$$

Note : 2 We will now prove an identity which is particularly useful in Quantum Mechanics.

$$\lim_{\epsilon \rightarrow 0^+} \int_{-\infty}^{\infty} \frac{1}{x \pm i\epsilon} f(x) dx = \rho \int_{-\infty}^{\infty} \frac{dx}{x} f(x) \mp i\pi f(0) \quad (1.38)$$

or in short

$$\lim_{\epsilon \rightarrow 0^+} \frac{1}{x \pm i\epsilon} = \rho \left(\frac{1}{x} \right) \mp i\pi \delta(x) \quad (1.39)$$

Where it is understood that the second of these two equations have meaning only within an integral.

The symbol ρ means principle part of an integral where the integral has a single pole. The principle part is defined as

$$\rho \int_{-A}^B \frac{dx}{x} f(x) dx = \lim_{\eta \rightarrow 0^+} \left[\int_{-A}^{-\eta} + \int_{\eta}^B \right] \frac{dx}{x} f(x) \quad (1.40)$$

proof :

$$\frac{1}{x \pm i\epsilon} = \frac{x \mp i\epsilon}{x^2 + \epsilon^2} = \frac{x}{x^2 + \epsilon^2} \mp \frac{i\epsilon}{x^2 + \epsilon^2} \quad (1.41)$$

Now we have

$$\begin{aligned} \lim_{\epsilon \rightarrow 0^+} \frac{1}{\pi} \frac{\epsilon}{x^2 + \epsilon^2} &= \delta(x) \\ \lim_{\epsilon \rightarrow 0^+} (\mp) i \frac{\epsilon}{x^2 + \epsilon^2} &= \mp i\pi \delta(x) \end{aligned} \quad (1.42)$$

Now consider the first term on the right hand side of equation 1.42. We multiply this term by a function $f(x)$ which is regular at the origin and then integrate over x . We get

$$\lim_{\epsilon \rightarrow 0^+} \int_{-\infty}^{\infty} \frac{xf(x)}{x^2 + \epsilon^2} dx = \lim_{\epsilon \rightarrow 0^+} \left[\lim_{\eta \rightarrow 0^+} \left[\int_{-\infty}^{-\eta} \frac{xf(x)}{x^2 + \epsilon^2} + \int_{-\eta}^{\eta} \frac{xf(x)}{x^2 + \epsilon^2} + \int_{\eta}^{\infty} \frac{xf(x)}{x^2 + \epsilon^2} dx \right] \right] \quad (1.43)$$

Note that we take the limit over η first and then we take the limit over ε . Consider now the second integral above

$$\lim_{\eta \rightarrow 0^+} \int_{-\eta}^{+\eta} \frac{xf(x)}{x^2 + \varepsilon^2} = f(0) \lim_{\eta \rightarrow 0^+} \frac{1}{2} [\ln(x^2 + \varepsilon^2)]_{x=-\eta}^{x=\eta} = 0 \quad (1.44)$$

If we now reverse the order of the evaluation of limits in equation 1.43, the $\varepsilon \rightarrow 0$ limit causes no difficulties in the other two integrals. Thus we have

$$\begin{aligned} & \lim_{\varepsilon \rightarrow 0^+} \int_{-\infty}^{\infty} \frac{x dx}{x^2 + \varepsilon^2} f(x) \\ &= \lim_{\eta \rightarrow 0^+} \lim_{\varepsilon \rightarrow 0^+} \left[\int_{-\infty}^{-\eta} + \int_{\eta}^{\infty} \right] \frac{x dx}{x^2 + \varepsilon^2} f(x) \\ &= \lim_{\eta \rightarrow 0^+} \left[\int_{-\infty}^{-\eta} + \int_{\eta}^{\infty} \right] \frac{dx}{x} f(x) \\ &= \rho \int_{-\infty}^{\infty} \frac{1}{x} f(x) dx \end{aligned}$$

This establishes the identity.

1.3 Derivatives of the delta function

One may define the derivative $\delta'(x)$ of the delta function. When ε is small, the derivative of $D_\varepsilon(x)$ has two peaks close to the origin, one peak is positive and the other is negative as drawn in the figure below

As $\varepsilon \rightarrow 0$, each of these peaks becomes very narrow and very tall, and two peaks each approach very close to the origin.

Now the integration by parts gives

$$\int_{-\infty}^{\infty} dx D'_\varepsilon(x) f(x) = [D_\varepsilon(x) f(x)]_{-\infty}^{\infty} - \int_{-\infty}^{\infty} dx D_\varepsilon(x) f'(x) \quad (1.45)$$

Because $D_\varepsilon(x)$ tens to zero as $x \rightarrow \pm\infty$, the first term on the right hand side vanishes unless $f(x)$ @@@@ @@@@ violently at infinity. So by letting $\varepsilon \rightarrow 0$, we arrive at the definition of $\delta'(x)$

$$\int_{-\infty}^{\infty} \delta'(x) f(x) dx = - \int_{-\infty}^{\infty} \delta(x) f'(x) dx = -f'(0) \quad (1.46)$$

From this we immediately get

$$x \delta'(x) = -\delta(x) \quad (1.47)$$

Conversely, it can be shown that the general solution of the equation

$$x u(x) = \delta(x) \quad (1.48)$$

can be written as

$$u(x) = -\delta'(x) + c \delta(x) \quad (1.49)$$

Where the second term arises from the homogeneous equation $x \delta(x) = 0$. From equation 1.46 it also follows that

$$\delta'(-x) = -\delta'(x) \quad (1.50)$$

The n^{th} order derivative of $\delta(x)$ can be defined in the same way. We find

$$\int_{-\infty}^{\infty} \delta^{(n)}(x) f(x) dx = (-1)^n f(0) \quad (1.51)$$

We can prove the following properties:

$$\delta^{(m)}(x) = (-1)^m \delta^{(m)}(-x) \quad (1.52)$$

$$x^{m+1} \delta^{(m)}(x) = 0 \quad (1.53)$$

$$x \delta^{(m)}(x) = -m \delta^{(m-1)}(x) \quad (1.54)$$

1.4 Integration of the delta function

Consider the indefinite integral

$$\Theta_\varepsilon(x) = \int_{-\infty}^x D_\varepsilon(y) dy \quad (1.55)$$

A graph of $\Theta_\varepsilon(x)$ vs x is shown below

As $\varepsilon \rightarrow 0$, the step in the function $\Theta_\varepsilon(x)$ gets progressively steeper, until, finally, the function changes abruptly from 0 to 1 at $x = 0$. Thus taking the limit $\varepsilon \rightarrow 0$ in equation 1.55 we have

$$\Theta(x) = \int_{-\infty}^x \delta(x) dx \quad (1.56)$$

Where

$$\Theta(x) = \begin{cases} 1 & \text{for } x > 0 \\ 0 & \text{for } x < 0 \end{cases} \quad (1.57)$$

If we differentiate equation 1.56 with respect to x , we get

$$\frac{d\Theta(x)}{dx} = \delta(x) \quad (1.58)$$

1.5 Three - dimensional delta function

$$\delta(\vec{r}) \stackrel{\text{def}}{=} \delta(x)\delta(y)\delta(z) \quad (1.59)$$

In other words, $\delta(\vec{r})$ is zero if any of the coordinates x, y, z is not equal to zero and $\delta(\vec{r})$ tends to infinity at the origin, i.e., when $x = 0, y = 0, z = 0$, such that

$$\int_{\text{volume}} \delta(\vec{r}) d^3 r = 1 \quad (1.60)$$

if the volume of the integration contains the origin. We also have

$$\int \delta(\vec{r}) f(\vec{r}) d^3 r = f(0) \quad (1.61)$$

where again the volume of the integration includes the origin. **Note:**

$$\delta(\vec{r} - \vec{r}') = \delta(x - x')\delta(y - y')\delta(z - z') \quad (1.62)$$

$$\int_V \delta(\vec{r} - \vec{r}') d^3 r = 1 \quad (1.63)$$

where the volume of integration includes the point \vec{r}' . Otherwise the integral is zero.

$$\int_V \delta(\vec{r} - \vec{r}') f(\vec{r}) d^3 r = f(\vec{r}') \quad (1.64)$$

if V includes the point \vec{r}' .

A useful formula Consider the integral

$$\begin{aligned} \int_{-\infty}^{\infty} e^{ikx} dx &= \lim_{L \rightarrow \infty} \int_{-L}^L e^{ikx} dx \\ &= \lim_{L \rightarrow \infty} \frac{1}{ik} (e^{ikL} - e^{-ikL}) \\ &= \lim_{L \rightarrow \infty} \frac{2}{k} \left(\frac{e^{ikL} - e^{-ikL}}{2i} \right) \\ &= \lim_{L \rightarrow \infty} \frac{2}{k} \sin(kL) \\ &= 2\pi \lim_{L \rightarrow \infty} \frac{\sin(kL)}{\pi k} \\ &= 2\pi \delta(k) \end{aligned} \quad (1.65)$$

using

$$\lim_{\epsilon \rightarrow 0} \frac{\sin(x/\epsilon)}{\pi x} = \delta(x) \quad (1.66)$$

Thus

$$\int_{-\infty}^{\infty} e^{ikx} dx = 2\pi \delta(k) \quad (1.67)$$

In equation 1.67 if we integrate with respect to k , we would have *delta(x)* on the right hand side,

$$\int_{-\infty}^{\infty} e^{ikx} dk = 2\pi \delta(x) \quad (1.68)$$

Also note that in equation 1.67 we are integrating over its full range of values. Making a change of variable $x \rightarrow -x$ does not change the value of the integral. Hence we also have

$$\int_{-\infty}^{\infty} e^{-ikx} dx = 2\pi \delta(k) \quad (1.69)$$

Similarly, in equation 1.68 making the change $k \rightarrow -k$, doesn't change the value of the integral. So we could also write

$$\int_{-\infty}^{\infty} e^{-ikx} dk = 2\pi \delta(x) \quad (1.70)$$

Thus in summary

$$\int_{\pm\infty}^{\infty} e^{-ikx} dx = 2\pi \delta(k) \int_{\pm\infty}^{\infty} e^{-ikx} dk = 2\pi \delta(x) \quad (1.71)$$

In three dimensions

$$\int_{all\ space} e^{\pm i\vec{k}\cdot\vec{r}} d^3\vec{r} = (2\pi)^3 \delta(\vec{k}) \quad (1.72)$$

$$\int_{all\ space} e^{\pm i(\vec{k}-\vec{k}')\cdot\vec{r}} d^3\vec{r} = (2\pi)^3 \delta(\vec{k}-\vec{k}') \quad (1.73)$$

$$\int_{all\ space} e^{\pm i\vec{k}\cdot(\vec{r}-\vec{r}')} d^3\vec{r} = (2\pi)^3 \delta(\vec{r}-\vec{r}') \quad (1.74)$$

(1.75)

1.6 Fourier Transformation

We can always express a function $f(x)$ in the form

$$f(x) = \int_{-\infty}^{\infty} e^{ikx} \tilde{f}(k) dk \quad (1.76)$$

where $\tilde{f}(k)$ is a function of k , called the fourier transform of $f(x)$. From eqnarray 1.76 we can write

$$\begin{aligned} \int_{-\infty}^{\infty} e^{-ik'x} f(x) dx &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} e^{i(k-k')x} \tilde{f}(k) dk dk' \\ &= 2\pi \int_{-\infty}^{\infty} \delta(k-k') \tilde{f}(k) dk \\ &= 2\pi \tilde{f}(k') \end{aligned} \quad (1.77)$$

Thus

$$\tilde{f}(k) = \frac{1}{2\pi} \int_{-\infty}^{\infty} e^{-ikx} f(x) dx \quad (1.78)$$

Thus functions $f(x)$ and $\tilde{f}(k)$ are Fourier transform of each other. We can write eqnarray 1.76 and 1.78 in a more symmetrical fashion as follows:

$$f(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} e^{ikx} \tilde{f}(k) dk \quad (1.79)$$

$$\tilde{f}(k) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{\infty} e^{-ikx} f(x) dx \quad (1.80)$$

In three dimension, we can write

$$f(\vec{r}) = \frac{1}{(2\pi)^{3/2}} \int_{all\ k-space} e^{i\vec{k}\cdot\vec{r}} \tilde{f}(\vec{k}) d^3k \quad (1.81)$$

multiplying eqnarray 1.81 by $e^{-i\vec{k}'\cdot\vec{r}}$ and integrating over \vec{r} , we have

$$\begin{aligned} \int_{all\ space} e^{-i\vec{k}'\cdot\vec{r}} f(\vec{r}) d^3r &= \frac{1}{(2\pi)^{3/2}} \int d^3r \int d^3k e^{i(\vec{k}-\vec{k}')\cdot\vec{r}} \tilde{f}(\vec{k}) \\ &= \frac{1}{(2\pi)^{3/2}} \int d^3k (2\pi)^3 \delta(\vec{k}-\vec{k}') \tilde{f}(\vec{k}) \\ &= \frac{1}{(2\pi)^{-3/2}} \tilde{f}(\vec{k}') \end{aligned} \quad (1.82)$$

Therefore

$$\tilde{f}(\vec{k}) = \frac{1}{(2\pi)^{3/2}} \int e^{-i\vec{k} \cdot \vec{r}} f(\vec{r}) d^3 r \quad (1.83)$$

Thus we can write

$$f(\vec{r}) = \frac{1}{(2\pi)^{3/2}} \int e^{i\vec{k} \cdot \vec{r}} \tilde{f}(\vec{k}) d^3 k \quad (1.84)$$

$$\tilde{f}(\vec{k}) = \frac{1}{(2\pi)^{3/2}} \int e^{-i\vec{k} \cdot \vec{r}} f(\vec{r}) d^3 r \quad (1.85)$$

Parseval's Identity:

We can now prove the important identity

$$\int |f(\vec{r})|^2 d^3 r = \int |\tilde{f}(\vec{k})|^2 d^3 k \quad (1.86)$$

proof:

$$\begin{aligned} \int |f(\vec{r})|^2 d^3 r &= \int f(\vec{r}) f^*(\vec{r}) d^3 r \\ &= \int d^3 r \frac{1}{(2\pi)^{3/2}} \int e^{i\vec{k} \cdot \vec{r}} \tilde{f}(\vec{k}) d^3 k \frac{1}{(2\pi)^{3/2}} \int e^{-i\vec{k}' \cdot \vec{r}} \tilde{f}^*(\vec{k}') d^3 k' \\ &= \frac{1}{(2\pi)^3} \int d^3 k d^3 k' \tilde{f}(\vec{k}) \tilde{f}^*(\vec{k}') \int d^3 r e^{i(\vec{k}-\vec{k}') \cdot \vec{r}} \\ &= \frac{1}{(2\pi)^3} \int d^3 k d^3 k' \tilde{f}(\vec{k}) \tilde{f}^*(\vec{k}') (2\pi)^3 \delta(\vec{k} - \vec{k}') \\ &= \int d^3 k \tilde{f}(\vec{k}) \tilde{f}^*(\vec{k}) \\ &= \int |\tilde{f}(\vec{k})|^2 d^3 k \end{aligned} \quad (1.87)$$

2. sheet-2 : Linear Vector Space

References

1. Quantum Mechanics - Shanker
2. Quantum Mechanics - Sakurai

2.1 Definition

A linear vector space V is a collection of objects ψ_a, ψ_b, \dots , called vectors, which satisfy the following postulates:

1. If ψ_a and ψ_b are vectors in V , there is a unique vector $\psi_a + \psi_b$ in V , called the sum of ψ_a and ψ_b . In other words, an operation called addition is defined in the vector space such that the space is closed under addition.
2. The vector addition is commutative and associative, i.e.,

$$\psi_a + \psi_b = \psi_b + \psi_a \quad (2.1)$$

$$\psi_a + (\psi_b + \psi_c) = (\psi_a + \psi_b) + \psi_c \quad (2.2)$$

3. There is a vector in V called the null vector and denoted by ϕ satisfying

$$\psi_a + \phi = \phi + \psi_a \quad (2.3)$$

for every ψ_a in V .

4. For every vector ψ_a in V there is another vector ψ'_a in V such that

$$\psi_a + \psi'_a = \phi \quad (2.4)$$

we denote ψ'_a as $-\psi_a$.

(Note : we use the notation $\psi_a - \psi_b$ to mean $\psi_a + (-\psi_b)$).

5. If ψ_a is a vector and λ is an arbitrary number (real or complex), called a scalar, there is a uniquely defined vector $\lambda \psi_a$ in V satisfying.

(a)

$$\lambda(\psi_a + \psi_b) = \lambda \psi_a + \lambda \psi_b \quad (2.5)$$

i.e., multiplication is distributive with respect to vector addition.

(b)

$$(\lambda\mu)\psi_a = \lambda(\mu\psi_a) \quad (2.6)$$

i.e., multiplication by a scalar is associative.

(c)

$$(\lambda + \mu)\psi_a = \lambda\psi_a + \mu\psi_b \quad (2.7)$$

i.e., multiplication is distributive with respect to addition scalars.

(d) Multiplication by scalars 0 and 1 are defined by

$$0\psi_a = \phi \quad (2.8)$$

$$1\psi_a = \psi_a \quad (2.9)$$

for any ψ_a in V .

2.2 Example of linear vector space

1. Consider all real numbers x in the range $-\infty$ to ∞ , i.e.,

$$-\infty < x < \infty \quad (2.10)$$

$$x \in \text{Re} \quad (\text{Re is the set of all real numbers}) \quad (2.11)$$

Take any two real numbers x_1 and x_2 . If we add two real numbers we get another real number in Re . Thus

$$x_1 + x_2 = \text{Re} \quad (2.12)$$

Next take any real number x . If we multiply x by another real number λ , we get a real number in Re , i.e.,

$$\lambda x \in \text{Re} \quad (2.13)$$

If we take a real number x , then there exists another real number $-x$ such that

$$x + (-x) = 0 \quad (2.14)$$

So the real numbers form a vector space with the real numbers themselves as vectors in the space. The number 0 is the null vector ϕ of the space. The scalars λ by which the vectors are multiplied are also real numbers.

Thus the real numbers form a real linear vector space over a field which are also real numbers. The addition and multiplication are just the normal addition and multiplication of the real numbers.

2. The set of $n-tuples$ of numbers (x_1, x_2, \dots, x_n) . When the addition of vectors and multiplication by a scalar are defined by

$$(x_1, x_2, \dots, x_n) + (y_1, y_2, \dots, y_n) = (x_1 + y_1, x_2 + y_2, \dots, x_n + y_n) \quad (2.15)$$

and

$$\lambda(x_1, x_2, \dots, x_n) = (\lambda x_1, \lambda x_2, \dots, \lambda x_n) \quad (2.16)$$

3. The collection of all square-integrable complex valued functions of a real variable form a vector space. consider all functions

$$f : \mathbb{R} \rightarrow \mathbb{C} \quad (2.17)$$

here, \mathbb{R} = set of real numbers

\mathbb{C} = set of complex numbers

such that

$$\int_{-\infty}^{\infty} f^*(x)f(x)dx \equiv \int_{-\infty}^{\infty} |f(x)|^2 dx < \infty \text{ (i.e., finite)} \quad (2.18)$$

The sum of two functions and the product of a function by a complex scalar are defined in the usual way.

The reason the square-integrable functions form a (complex) vector space is that the space is closed under addition. The vectors of the space are the square-integrable functions. In other words, it can be shown that if f and g are two vectors, in this case two functions $f(x)$ and $g(x)$ both of which are square integrable, then $f(x) + g(x)$ is also square integrable and have the sum belong to the vector space.

proof:

Let $f(x)$ and $g(x)$ be two square integrable function, i.e.,

$$\int_{-\infty}^{\infty} |f(x)|^2 dx < \infty \quad (2.19)$$

and

$$\int_{-\infty}^{\infty} |g(x)|^2 dx < \infty \quad (2.20)$$

then using the inequality

$$\int_{-\infty}^{\infty} |f+g|^2 dx \leq \left[\sqrt{\int_{-\infty}^{\infty} |f|^2 dx} \sqrt{\int_{-\infty}^{\infty} |g|^2 dx} \right]^2 \quad (2.21)$$

it is obvious that

$$\int_{-\infty}^{\infty} |f+g|^2 dx < \infty \quad (2.22)$$

(i.e., finite).

4. The set of all $n \times n$ matrices with complex elements form a complex linear vector space. For illustration let us take 2×2 matrix A

$$A = \begin{pmatrix} a & b \\ c & d \end{pmatrix} \quad (2.23)$$

where the elements a, b, c and d can, in general, be complex. Then A belongs to a vector (complex) space. We have

$$\phi = \begin{pmatrix} 0 & 0 \\ 0 & 0 \end{pmatrix} \quad (2.24)$$

$$\mathbb{I} = \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} \quad (2.25)$$

$$-A = \begin{pmatrix} -a & -b \\ -c & -d \end{pmatrix} \quad (2.26)$$

$$\lambda A = \begin{pmatrix} \lambda a & \lambda b \\ \lambda c & \lambda d \end{pmatrix} \quad (2.27)$$

The set of all these matrices form a vector space.

2.3 Inner product space or a unitary vector space

For a general linear vector space, product of vectors (i.e. multiplication of two vectors) need not be defined. However, we will restrict ourselves to spaces in which a scalar product or an inner product is defined.

A linear vector space is called unitary if a scalar product is defined in it. To every pair of vectors ψ_a and ψ_b in V there corresponds a unique scalar (in general complex), called the scalar product. The scalar product is defined to have the following properties :

$$(\psi_a, \psi_b) = (\psi_b, \psi_a)^* \quad (2.28)$$

$$(\psi_a, \lambda \psi_b) = \lambda (\psi_a, \psi_b) \quad (2.29)$$

$$(\lambda \psi_a, \psi_b) = \lambda^* (\psi_a, \psi_b) \quad (2.30)$$

$$(\psi_a, \psi_b + \psi_c) = (\psi_a, \psi_b) + (\psi_a, \psi_c) \quad (2.31)$$

$$(\psi_a, \psi_a) \geq 0 ; \text{the equality holds only if } \psi_a \text{ is the null vector} \quad (2.32)$$

It follows from the above postulated properties of the scalar product, that the scalar product is linear with respect to post factors, i.e.,

$$(\psi_a, \lambda \psi_b + \mu \psi_c) = \lambda (\psi_a, \psi_b) + \mu (\psi_a, \psi_c) \quad (2.33)$$

and anti-linear with respect to the pre factors, i.e.,

$$(\lambda \psi_a + \mu \psi_b, \psi_c) = \lambda^* (\psi_a, \psi_c) + \mu^* (\psi_b, \psi_c) \quad (2.34)$$

2.4 Examples of scalar product

Example 1 Consider the vector space consisting of all square integrable functions of a real variable in the domain $[a, b]$. This space is denoted by $L^2[a, b]$.

Suppose

$$f \in L^2[a, b] \quad (2.35)$$

i.e.,

$$\int_a^b f^*(x) f(x) dx \equiv \int_a^b |f(x)|^2 dx < \infty \quad (2.36)$$

We can define the scalar product of two vectors f and g as

$$(f, g) \stackrel{def}{=} \int_a^b f^*(x)g(x)dx = \text{complex number} \quad (2.37)$$

We can show

$$|(f, g)| = \left[\sqrt{\int_a^b |f(x)|^2 dx} \right] \left[\sqrt{\int_a^b |g(x)|^2 dx} \right] \quad (2.38)$$

Since both f and g are square integrable, $|(f, g)|$ is finite, i.e., the scalar product of f and g exists.

The scalar product defined above satisfies all the properties that a scalar product is postulated to have.

Example 2 Now consider the vector space consisting of n -tuples of complex numbers. Such a vector space is denoted as \mathbb{C}^n .

A vector $\psi_a \in \mathbb{C}^n$ may be expressed as

$$\psi_a = (a_1 \ a_2 \ \dots \ a_n)^T \quad (2.39)$$

The scalar product may then be defined as

$$(\psi_a, \psi_b) \stackrel{def}{=} (a_1^* \ a_2^* \ \dots \ a_n^*) \begin{pmatrix} b_1 \\ b_2 \\ \vdots \\ b_n \end{pmatrix} \quad (2.40)$$

$$= \sum_{i=1}^n a_i^* b_i \quad (2.41)$$

This scalar product also satisfies all the properties of a scalar product.

Example 3 Euclidean 3-space \mathbb{R}^3 . The vectors of \mathbb{R}^3 are 3-tuples of real numbers which could be represented as column vectors. Thus if ψ_a and ψ_b are in \mathbb{R}^3 ,

$$\psi_a = \begin{pmatrix} a_1 \\ a_2 \\ a_3 \end{pmatrix} \quad (2.42)$$

$$\psi_b = \begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix} \quad (2.43)$$

where a_i and b_i are real.

We could define the scalar product of ψ_a and ψ_b as

$$(\psi_a, \psi_b) \stackrel{def}{=} (a_1 \ a_2 \ a_3) \begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix} \quad (2.44)$$

$$= \psi_a^T \psi_b \quad (2.45)$$

$$= \sum_{i=1}^3 a_i b_i \quad (2.46)$$

This scalar product also has all the postulated properties of a scalar product.

In case of the vector space \mathbb{R}^3 , the vectors ψ_a and ψ_b could be represented as directed lines \vec{a} and \vec{b} in a three dimensional coordinate system.

figure

The Scalar product (ψ_a, ψ_b) is the usual dot product

$$\vec{a} \cdot \vec{b} = a_1 b_1 + a_2 b_2 + a_3 b_3 \quad (2.47)$$

$$= |\vec{a}| |\vec{b}| \cos(\theta) \quad (2.48)$$

Where $|\vec{a}|$ and $|\vec{b}|$ are the magnitudes of the vectors \vec{a} and \vec{b} defined as

$$|\vec{a}| \equiv \sqrt{(\psi_a, \psi_a)} = \sqrt{a_1^2 + a_2^2 + a_3^2} \quad (2.49)$$

$$|\vec{b}| \equiv \sqrt{(\psi_b, \psi_b)} = \sqrt{b_1^2 + b_2^2 + b_3^2} \quad (2.50)$$

2.5 Norm of a vector

If a vector space is enclosed with a scalar product, then the scalar product gives us the concept of the "magnitude" or "length" of a vector. In a general vector space the "magnitude" or "length" of a vector is called the norm of the vector. We simply define the norm of a vector ψ_a as

$$\|\psi_a\| \stackrel{\text{def}}{=} \sqrt{(\psi_a, \psi_a)} \quad (2.51)$$

The norm has the following properties:

1.

$$\|\psi_a\| \geq 0 \quad (2.52)$$

the equality holds only if the vector is null.

2.

$$\|\psi_a + \psi_b\| \leq \|\psi_a\| + \|\psi_b\| \quad (2.53)$$

This is called the triangle inequality

3.

$$\|\psi_a - \psi_b\| = \|\psi_b - \psi_a\| \quad (2.54)$$

2.6 Metric included by the scalar product

The norm included by the scalar product allows us to develop the concept of "distance" between vectors in a vector space. We say two vectors ψ_a and ψ_b are 'close' if $\|\psi_a - \psi_b\|$ is small. The metric in a vector space assigns a real number to the vector $\psi_a - \psi_b$. This real number is a measure of how close the two vectors are. We simply define the metric $\mathbf{d}(\psi_a, \psi_b)$ as

$$\mathbf{d}(\psi_a, \psi_b) \stackrel{\text{def}}{=} \|\psi_a - \psi_b\| \quad (2.55)$$

Thus, if there are three vectors ψ_a , ψ_b and ψ_c and if $\mathbf{d}(\psi_a, \psi_b) < \mathbf{d}(\psi_a, \psi_c)$ then we say ψ_a is closer to ψ_b than to ψ_c .

2.7 Schwarz's inequality

We will now prove a very important inequality called Schwarz inequality which states

$$|(\psi_a, \psi_b)| \leq \sqrt{(\psi_a, \psi_a)(\psi_b, \psi_b)} \quad (2.56)$$

or

$$|(\psi_a, \psi_b)| \leq \|\psi_a\| \|\psi_b\| \quad (2.57)$$

proof : Let

$$\psi = \psi_a + \lambda \psi_b \quad (2.58)$$

Then

$$(\psi, \psi) = (\psi_a + \lambda \psi_b, \psi_a + \lambda \psi_b) \quad (2.59)$$

$$= (\psi_a, \psi_a) + \lambda (\psi_a, \psi_b) + \lambda^* (\psi_b, \psi_a) + |\lambda|^2 (\psi_b, \psi_b) \geq 0 \quad (2.60)$$

The best inequality is obtained if λ is chosen so as to minimize the left hand side of the above equation. By differentiation, the value of λ which accomplishes this is found to be

$$\lambda = -\frac{(\psi_b, \psi_a)}{\psi_b, \psi_b} \quad (2.61)$$

Substituting this value of λ in the above equation yields the Schwarz inequality.

We note that the equality sign holds if and only if $(\psi, \psi) = 0$, i.e., ψ is the null vector, i.e., $\psi = \phi$, in other words

$$\psi_a + \lambda \psi_b = \phi(\text{null}) \quad (2.62)$$

i.e.,

$$\psi_a = -\lambda \psi_b + \phi \quad (2.63)$$

$$\psi_a = -\lambda \psi_b \quad (2.64)$$

Hence, the equality holds if ψ_a and ψ_b are multiple of each other, or if ψ_a and ψ_b are "parallel". It follows from the Schwarz inequality that the scalar product (ψ_a, ψ_b) is finite if the norms of ψ_a and ψ_b are finite.

2.8 Analogy of Schwarz inequality with vectors in a three-dimensional Euclidean space \mathbb{R}^3

In \mathbb{R}^3 , the vectors can be represented by directed lines (i.e. arrows). We have the scalar product ordinary vectors in the form

$$\vec{A} \cdot \vec{B} = |\vec{A}| |\vec{B}| \cos(\theta) \quad (2.65)$$

Since cosine of any angle lies between -1 and $+1$, we have

$$|\vec{A} \cdot \vec{B}| \leq |\vec{A}| |\vec{B}| \quad (2.66)$$

The analogue of this equation for a general vector space is the Schwarz inequality

$$|(\psi_a, \psi_b)| \leq \|\psi_a\| \|\psi_b\| \quad (2.67)$$

2.9 Orthogonality and linear independence

A vector whose norm is unity is called a unit vector. For any given non-null vector, a unit vector can be formed by defining the vector by its norm. Thus

$$u_a = \frac{\psi_a}{\|\psi_a\|} \quad (2.68)$$

is normalized.

Two vectors ψ_a and ψ_b are orthogonal if their inverse product is zero, i.e., if

$$(\psi_a, \psi_b) = 0 \quad (2.69)$$

The unit vectors u_1, u_2, \dots, u_N form an orthogonal set if they are mutually orthonormal , i.e.,

$$(u_i, u_j) = \delta_{ij} \quad (2.70)$$

2.9.1 Linear independence

The set of vectors $\psi_1, \psi_2, \dots, \psi_N$ are linearly independent if none of them can be expressed as a linear combination of the others. Mathematically this means that the equation

$$\sum_{j=1}^N c_j \psi_j = 0 \quad (2.71)$$

cannot be satisfied except by $c_j = 0$ for all j .

2.10 Orthonormality and linear independence

A dot of mutually orthogonal vectors (not necessarily normalized) are necessarily linearly independent. The converse is not true, however. That is, a set of linearly independent vectors may not be mutually orthogonal.

It is always possible to orthonormalize a set of linearly independent vectors. By this we mean that from a given set of N linearly independent vectors, it is possible to form a set of N orthonormal vectors. This procedure is called **Schmidt orthonormalization method**.

2.11 Schmidt orthonormalization method

Suppose $\psi_1, \psi_2, \dots, \psi_N$ is a set of linearly independent vectors. Let

$$u_1 = \frac{\psi_1}{\|\psi_1\|} \quad (2.72)$$

Then $(u_1, u_1) = 1$, i.e., u_1 is normalized. Next construct the vector ψ'_2 as follows:

$$\psi'_2 = \psi_2 - u_1(u_1, \psi_2) \quad (2.73)$$

i.e., to obtain ψ'_2 we have subtracted the 'component' of ψ_2 along the u_1 "direction". Then it follows that

$$(u_1, \psi'_2) = (u_1, \psi_2) - (u_1, u_1)(u_1, \psi_2) \quad (2.74)$$

$$= (u_1, \psi_2) - (u_1, \psi_2) \quad (2.75)$$

$$= 0 \quad (2.76)$$

i.e., ψ'_2 is orthogonal to u_1 . We then normalize ψ'_2 , i.e.,

$$u_2 = \frac{\psi'_2}{\|\psi'_2\|} \quad (2.77)$$

We can continue the process until we exhaust all the vectors. For example, in the next step we can write

$$\psi'_3 = \psi_3 - u_1(u_1, \psi_3) - u_2(u_2, \psi_3) \quad (2.78)$$

We note immediately that ψ'_3 is orthogonal to both u_1 and u_2 , i.e.,

$$(u_1, \psi'_3) = (u_2, \psi'_3) = 0 \quad (2.79)$$

We normalize ψ'_3 to get u_3 , i.e.,

$$u_3 = \frac{\psi'_3}{\|\psi'_3\|} \quad (2.80)$$

Finally, in the $N-th$ step, we write

$$\psi'_N = \psi_N - u_1(u_1, \psi_N) - u_2(u_2, \psi_N) - \dots - u_{N-1}(u_{N-1}, \psi_N) \quad (2.81)$$

ψ'_N is orthogonal to u_1, u_2, \dots, u_{N-1} , i.e.,

$$(u_1, \psi'_N) = (u_2, \psi'_N) = \dots = (u_{N-1}, \psi'_N) = 0 \quad (2.82)$$

Normalizing ψ'_N we get

$$u_N = \frac{\psi'_N}{\|\psi'_N\|} \quad (2.83)$$

Thus, the set $\{u_1, u_2, \dots, u_N\}$ is an orthonormal set of vectors.

2.12 Dimension of a vector space

The vector space V is said to be n -dimensional if there exists n linearly independent vectors, but if $n+1$ vectors are linearly dependent the dimension may be finite or infinite.

2.13 Complete vector space

Before defining what a complete vector space is we will give some other definitions.

A sequence of vectors $\{\psi_n\}$ in the vector space V is called a **Cauchy sequence** if for every $\varepsilon > 0$ there exists an integer N such that

$$\|\psi_n - \psi_m\| < \varepsilon \quad (2.84)$$

if $n, m > N$. In other words, the vectors in the sequence come 'closer' if the index increases. In particular

$$\|\psi_n - \psi_m\| \rightarrow 0 \text{ as } n, m \rightarrow \infty \quad (2.85)$$

2.14 Convergence of a sequence of vectors in a vector space

A sequence $\{\psi_n\}$, $n = 1, 2, \dots$ in a vector space V converges to a vector ψ in V if for every $\varepsilon > 0$ there exists an integer N such that

$$\|\psi_n - \psi_m\| < \varepsilon \quad (2.86)$$

if $n > N$, that is if

$$\lim_{n \rightarrow \infty} \|\psi - \psi_n\| = 0 \quad (2.87)$$

then

$$\{\psi_n\} \rightarrow \psi \quad (2.88)$$

and the sequence is called a convergent sequence.

Now we can show every convergent sequence is a Cauchy sequence.

proof

Let $\{\psi_n\} \rightarrow \psi$. Then

$$\|\psi_n - \psi_m\| = \|\psi_n - \psi + \psi - \psi_m\| \quad (2.89)$$

$$\leq \|\psi_n - \psi\| + \|\psi - \psi_m\| \quad (2.90)$$

here the triangle inequality is used.

Since $\{\psi_n\}$ is a convergent sequence, each term on the right hand side tends to zero as n and m tends to infinity. Hence $\|\psi_n - \psi_m\| \rightarrow 0$ as $n, m = \infty$, i.e., the sequence $\{\psi_i\}$ with $i = 1, 2, \dots$ is a Cauchy sequence.

The converse of the above statement is not true in general. In other words, *a Cauchy sequence in a vector space may not converge to a vector in the space*. It can be shown that for a finite dimensional vector space the converse is true, i.e., in a finite dimensional vector space a Cauchy sequence is always a convergent sequence. Exceptions may arise in infinite-dimensional vector space.

2.14.1 An example of a vector space whose Cauchy sequence does not converge to a vector in the vector space

Consider the vector space consisting of all continuous function of a single real variable x in the range $[-1, 1]$. In this vector space consider a sequence $\{f_k(x)\}$, $k = 1, 2, \dots$ of the following form:

$$f_k(x) = \begin{cases} 1 & \text{for } \frac{1}{k} \leq x \leq 1 \\ \frac{kx+1}{2} & \text{for } -\frac{1}{k} < x < \frac{1}{k} \\ 0 & \text{for } -1 \leq x \leq -\frac{1}{k} \end{cases} \quad (2.91)$$

$k = 1, 2, 3, \dots$

The graph of the sequence of functions is shown below:

figure

Note that, in this example each $f_k(x)$ is continuous, but their first derivatives are discontinuous.

Let us define the scalar product in this space as

$$(f, g) = \int_{-1}^1 f^*(x)g(x)dx \quad (2.92)$$

So that the metric $\mathbf{d}(f, g)$, i.e., the "distance" between vectors f and g can be defined as

$$\mathbf{d}(f, g) \equiv ||f - g|| \quad (2.93)$$

$$= \sqrt{\int_{-1}^{+1} (f^*(x) - g^*(x))(f(x) - g(x)) dx} \quad (2.94)$$

With this metric we can show that the sequence f_k defined above is indeed a Cauchy sequence. However, looking at the graph above, we see that as k becomes large, f_k approaches the θ function

$$\theta(x) = \begin{cases} 0 & \text{for } x < 0 \\ 1 & \text{for } x > 0 \end{cases} \quad (2.95)$$

Which is a discontinuous function at $x = 0$.

We show that graph of $\theta(x)$ below:

figure

Thus the Cauchy sequence $\{f_k(x)\}$ of continuous function is converging to a discontinuous function which lies outside the vector space V .

If, instead of choosing all continuous function, we had chosen all square integrable function as defining the vector space, then any Cauchy sequence in the vector space would converge to a vector in the space.

Definition : A linear vector space is said to be complete if any Cauchy sequence converges to a vector in the space.

2.15 Hilbert space

A complete linear vector space, finite or infinite dimensional, (and hence @@ @@@enclosed@ @ @ with a norm and metric induced by the scalar product), is called a Hilbert space.

A finite-dimensional vector space is always complete. So, a finite dimensional linear vector space in which a scalar product is defined is a Hilbert space.

An infinite dimensional vector space with a scalar product may or may not be complete. Whether or not an infinite dimensional vector space is complete depends upon how exactly the vector space is defined and on the metric.

2.16 Basis vectors in a Hilbert space

2.16.1 Finite dimensional space

In a finite dimensional vector space of dimension n , any set of linearly independent vectors $\psi_1, \psi_2, \dots, \psi_n$ spans the entire space. In other words any vector ψ in the space can be expressed as linear combination of $\psi_1, \psi_2, \dots, \psi_n$, i.e.,

$$\psi = \sum_{i=1}^n a_i \psi_i \quad (2.96)$$

The vectors $\psi_1, \psi_2, \dots, \psi_n$ form a complete basis for the vector space. The vectors $\psi_1, \psi_2, \dots, \psi_n$, even if linearly independent, may not be orthogonal to each other. It is more convenient to use a set of orthonormal vectors $\phi_1, \psi_2, \dots, \phi_n$ as the basis. Being orthogonal, the vectors $\phi_1, \psi_2, \dots, \phi_n$ are automatically linearly independent. The orthonormal set of basis vectors $\{\phi_i\}$ $i = 1, 2, \dots, n$ can be

constructed from the set $\{\psi_i\} i = 1, 2, \dots, n$ by using the schmidt orthonormalization procedure. Choosing the orthonormal set as the basis, any vector ψ in the vector space can be written as

$$\psi = \sum_{i=1}^n a_i \phi_i \quad (2.97)$$

where

$$(\phi_i, \phi_j) = \delta_{ij} \quad (2.98)$$

using equation 2.98 we have

$$a_i = (\phi_i, \psi) \quad (2.99)$$

2.16.2 Infinite dimensional vector space

In an infinite dimensional vector space the number of basis vectors is infinity. Let $\{\phi_1, \phi_2, \dots\}$ be an infinite set of orthonormal basis vectors spanning the infinite dimensional Hilbert space. This set of basis vectors is said to be complete if any vector ψ in the Hilbert space can be expanded as a linear combination of the basis vectors, i.e.,

$$\psi = \sum_{i=1}^{\infty} a_i \phi_i \quad (2.100)$$

In an infinite dimensional vector space, choosing an infinite number of basis vectors may not ensure that the basis set is complete. It may so happen that there are other linearly independent vectors, may be infinite in numbers, which have been missed in the first choice of the basis vectors.

Whenever we have an infinite sum, as in equation 2.100 is to be understood in the sense that the sequences consisting of the partial sums

$$f_n = \sum_{i=1}^n a_i \phi_i \quad (2.101)$$

$$n = 1, 2, 3, \dots$$

converges to ψ , i.e.,

$$\lim_{n \rightarrow \infty} \|\psi - f_n\| \rightarrow 0 \quad (2.102)$$

Since the vector ψ must have a finite norm, we must have

$$\|\psi\|^2 = (\psi, \psi) = \sum_{i=1}^{\infty} |a_i|^2 < \infty \text{ (finite)} \quad (2.103)$$

If the basis vectors $\{\phi_i\}$ are orthonormal, we have

$$a_i = (\phi_i, \psi) \quad (2.104)$$

So that equation 2.103 can be written as

$$\sum_{i=1}^{\infty} |(\phi_i, \psi)|^2 < \infty \quad (2.105)$$

The scalar a_i can be regarded as the components of ψ in the 'directions' of ϕ_i

2.16.3 Example

Question : Show that the set of all square integrable functions, i.e., set of all functions f such that

$$\int_{-\infty}^{\infty} f^*(x)f(x)dx < \infty \text{ (i.e., finite)} \quad (2.106)$$

belong to a Hilbert space. This Hilbert space is denoted as $L^2(-\infty, \infty)$.

To show this, consider the following

1. If f and g are square integrable functions, so is $f + g$, and hence $f + g$ also belongs to the Hilbert space

$$\|f + g\| \leq \|f\| + \|g\| \quad (2.107)$$

2. We can define the scalar product between f and g as follows:

$$(f, g) \stackrel{def}{=} \int_{-\infty}^{\infty} f^*(x)g(x)dx \quad (2.108)$$

That the scalar product exists follows from the Schwarz inequality

$$|(f, g)| \leq \|f\| \cdot \|g\|, < \infty \quad (2.109)$$

3. It can also be shown that any Cauchy sequence of square integrable functions converges to a limit which is also square integrable. In other words, the space of all square integrable functions is complete.

Hence the linear vector space consisting of all square integrable functions is indeed a Hilbert space.

2.17 Dirac notation

(Cohen-Tannoudji; page 109)

"ket" vectors and "bra" vectors.

Notation

Any element, or vector of a vector space V is called a ket vector, or more simply, a ket. It is represented by the symbol $|\rangle$, inside which is placed a distinctive sign which enables us to distinguish between different kets, for example $|\psi\rangle$.

2.17.1 Scalar product

With each pair of kets $|\phi\rangle$ and $|\psi\rangle$, taken in this order, we associate a complex number, which is their scalar product ($|\phi\rangle, |\psi\rangle$) and which satisfies various properties discussed earlier (section 2.3).

2.17.2 Dual vector space

Linear functional: A linear functional χ is a linear operation on the kets such that χ operating on a ket $|\psi\rangle$ gives a complex scalar:

$$\chi|\psi\rangle \rightarrow \text{scalar where } |\psi\rangle \in V \quad (2.110)$$

and

$$\chi(\lambda_1|\psi\rangle + \lambda_2|\psi_2\rangle) = \lambda_1\chi|\psi_1\rangle + \lambda_2\chi|\psi_2\rangle \quad (2.111)$$

The set of all linear functions defined in the kets of a vector space V themselves form a linear vector space called the dual space of V and symbolized by V^* .

2.17.3 Bra notation for the vectors of V^*

Any element, or vector, of the space V^* is called a "bra vector", or more simply, a bra. It is symbolized by $\langle \cdot |$. For example, the bra $\langle \chi |$ designates the linear functional χ we shall henceforth use the notation $\langle \chi | \psi \rangle$ to denote the number obtained by causing the linear functional $\langle \chi | \in V^*$ to act on the ket $|\psi\rangle \in V$. Thus

$$\chi(|\psi\rangle) = \langle \chi | \psi \rangle \quad (2.112)$$

2.17.4 Correspondence between kets and bras

The existence of the scalar product in V will now enable us to show that we can associate with every ket $|\phi\rangle \in V$ and element of V^* , that is a bra, which will be denoted by $\langle \phi |$.

The ket $|\phi\rangle$ does indeed enable us to define a linear functional, the one which associates with each $|\psi\rangle \in V$ a complex number which is equal to the scalar product $(|\phi\rangle, |\psi\rangle)$. Let $\langle \phi |$ be this linear functional. It is thus defined by the relation

$$\langle \phi | \psi \rangle = (|\phi\rangle, |\psi\rangle) \quad (2.113)$$

2.17.5 The correspondence in anti-linear

Let $\lambda_1 |\phi_1\rangle + \lambda_2 |\phi_2\rangle$ be a ket. Then

$$(\lambda_1 |\phi_1\rangle + \lambda_2 |\phi_2\rangle, \psi) = \lambda_1^* (\langle \phi_1 |, \psi) + \lambda_2^* (\langle \phi_2 |, \psi) \quad (2.114)$$

$$= \lambda_1^* \langle \phi_1 | \psi \rangle + \lambda_2^* \langle \phi_2 | \psi \rangle \quad (2.115)$$

$$= (\lambda_1^* \langle \phi_1 | + \lambda_2^* \langle \phi_2 |) |\psi\rangle \quad (2.116)$$

Thus

$$\lambda_1 |\phi_1\rangle + \lambda_2 |\phi_2\rangle \xrightarrow{dc} \lambda_1^* \langle \phi_1 | + \lambda_2^* \langle \phi_2 | \quad (2.117)$$

Where "dc" is short for dual correspondence.

comment

If λ is a complex number and $|\psi\rangle$ is a ket, then $\lambda |\psi\rangle$ is also a ket. We are sometimes led to write $\lambda |\psi\rangle$ as $|\lambda \psi\rangle$:

$$|\lambda \psi\rangle = \lambda |\psi\rangle \quad (2.118)$$

One must be careful to remember that $\langle \lambda \psi |$ represents the bra associated with the ket $|\lambda \psi\rangle$. since the correspondence between a bra and a ket is anti-linear we have

$$\langle \lambda \psi | = \lambda^* \langle \psi | \quad (2.119)$$

2.17.6 Dirac notation for the scalar product

We now have at our disposal two distinct notations for designating the scalar product of $|\psi\rangle$ by $|\phi\rangle$, namely, $(|\phi\rangle, |\psi\rangle)$ and $\langle \phi | \psi \rangle$, $\langle \phi |$ being the bra associated with the ket $|\phi\rangle$. We shall mostly use the Dirac notation $\langle \phi | \psi \rangle$. In the table below we summarize, in Dirac notation, the properties of the scalar product.

$$\langle \phi | \psi \rangle = \langle \psi | \phi \rangle^* \quad (2.120)$$

$$\langle \phi | \lambda_1 \psi_1 + \lambda_2 \psi_2 \rangle = \lambda_1 \langle \phi | \psi_1 \rangle + \lambda_2 \langle \phi | \psi_2 \rangle \quad (2.121)$$

$$\langle \lambda_1 \phi_1 + \lambda_2 \phi_2 | \psi \rangle = \lambda_1^* \langle \phi_1 | \psi \rangle + \lambda_2^* \langle \phi_2 | \psi \rangle \quad (2.122)$$

$$(2.123)$$

$\langle \psi | \psi \rangle$ is real, positive; zero if and only if $|\psi\rangle = \phi$ (null)

3. sheet-3 : Linear Vector Space (continued)

3.1 Operators in a Hilbert Space

An operator is a prescription by which every vector ψ_a in a Hilbert space H is associated with another vector ψ_b in the space:

$$\hat{A} : \psi_1 \rightarrow \psi_b \quad (3.1)$$

for $\psi_a, \psi_b \in H$. We usually employ the notation

$$\psi_b = \hat{A}\psi_a \quad (3.2)$$

In Dirac notation, we write

$$|b\rangle = \hat{A}|a\rangle \quad (3.3)$$

where both $|a\rangle$ and $|b\rangle$ belong to the ket-space. An operator can also act on a bra vector (bra-space is also a Hilbert space; it is dual to the ket space) changing it to another bra-vector. The notation we employ is

$$\langle\psi| = \langle\phi|\hat{A} \quad (3.4)$$

Here the operator \hat{A} acts on the bra-vector $\langle\phi|$ to produce the bra vector $\langle\psi|$. We place the bra-vector on which the operator acts on the left of the operator.

3.2 Linear operators

An operator \hat{A} is said to be a linear operator if it has the following property :

1. For any vector $|a\rangle$ and $|b\rangle$ and any complex number λ_1 and λ_2 , we have

$$\hat{A}(\lambda_1|a\rangle + \lambda_2|b\rangle) = \lambda_1\hat{A}|a\rangle + \lambda_2\hat{A}|b\rangle \quad (3.5)$$

A linear vector operator can act on a bra vector also

$$(\lambda_1\langle a| + \lambda_2\langle b|)\hat{A} = \lambda_1\langle a|\hat{A} + \lambda_2\langle b|\hat{A} \quad (3.6)$$

2. The operator \hat{A} is anti linear if

$$\hat{A}(\lambda_1 |a\rangle + \lambda_2 |b\rangle) = \lambda_1^* \hat{A}|a\rangle + \lambda_2^* \hat{A}|b\rangle \quad (3.7)$$

3. Two operators \hat{A} and \hat{B} are equal if

$$\hat{A}|\psi\rangle = \hat{B}|\psi\rangle \quad (3.8)$$

for all $|\psi\rangle$ in the vector space.

4. Sum of two operators \hat{A} and \hat{B} is defined as

$$(\hat{A} + \hat{B})|\psi\rangle = \hat{A}|\psi\rangle + \hat{B}|\psi\rangle \quad (3.9)$$

5. Product of two operators \hat{A} and \hat{B} is defined as

$$(\hat{A}\hat{B})|\psi\rangle = \hat{A}(\hat{B}|\psi\rangle) \quad (3.10)$$

This equation says that the operator $\hat{A}\hat{B}$ acting on $|\psi\rangle$ produces the same vector which would be obtained if we first let \hat{B} act on $|\psi\rangle$ and then \hat{A} acts on the resultant of the previous operation. In general $\hat{A}\hat{B} \neq \hat{B}\hat{A}$, although in exceptional cases we may have $\hat{A}\hat{B} = \hat{B}\hat{A}$.

3.3 Commutator of two operators

The commutator of two operators \hat{A} and \hat{B} is defined as

$$[\hat{A}, \hat{B}] \stackrel{\text{def}}{=} \hat{A}\hat{B} - \hat{B}\hat{A} \quad (3.11)$$

In general $[\hat{A}, \hat{B}] \neq 0$ (null operator). If $[\hat{A}, \hat{B}] = 0$, we say that \hat{A} and \hat{B} commute with each other.

3.3.1 Some properties of commutators

$$[\hat{A}, \lambda\hat{B}] = \lambda[\hat{A}, \hat{B}] \quad (3.12)$$

$$[\lambda\hat{A}, \hat{B}] = \lambda[\hat{A}, \hat{B}] \quad (3.13)$$

$$[\hat{A}, \hat{B} + \hat{C}] = [\hat{A}, \hat{B}] + [\hat{A}, \hat{C}] \quad (3.14)$$

$$[\hat{A}, \hat{B}] = -[\hat{B}, \hat{A}] \quad (3.15)$$

$$[\hat{A}, \hat{B}\hat{C}] = \hat{B}[\hat{A}, \hat{C}] + [\hat{A}, \hat{B}]\hat{C} \quad (3.16)$$

$$[\hat{A}\hat{B}, \hat{C}] = \hat{A}[\hat{B}, \hat{C}] + [\hat{A}, \hat{C}]\hat{B} \quad (3.17)$$

3.4 Projection operator

(An important example of a linear operator).

Consider the operator \hat{P}_a defined as

$$\hat{P}_a = |a\rangle\langle a| \quad (3.18)$$

where

$$\langle a|b\rangle = 1 \quad (3.19)$$

Operating by \hat{P}_a on an arbitrary ket $|\psi\rangle$, we have

$$\hat{P}_a |\psi\rangle = |a\rangle \langle a|\psi\rangle \quad (3.20)$$

i.e., \hat{P}_a projects the ket $|\psi\rangle$ along $|a\rangle$. The complex number $\langle a|\psi\rangle$ is the component of $|\psi\rangle$ along $|a\rangle$. Now, $|P_a\rangle$ is a linear operator. To show this consider

$$\hat{P}_a (\lambda_1 |\psi_1\rangle + \lambda_2 |\psi_2\rangle) = |a\rangle \langle a| (\lambda_1 |\psi_1\rangle + \lambda_2 |\psi_2\rangle) \quad (3.21)$$

$$= \lambda_1 |a\rangle \langle a| |\psi_1\rangle + \lambda_2 |a\rangle \langle a| |\psi_2\rangle \quad (3.22)$$

$$= \lambda_1 \hat{P}_a |\psi_1\rangle + \lambda_2 \hat{P}_a |\psi_2\rangle \quad (3.23)$$

Another important property of the projection operator is

$$\hat{P}_a^2 = \hat{P}_a \quad (3.24)$$

To prove this allow \hat{P}_a^2 to act on a ket.

$$\hat{P}_a^2 |\psi\rangle = \hat{P}_a \hat{P}_a |\psi\rangle \quad (3.25)$$

$$= \hat{P}_a |a\rangle \langle a|\psi\rangle \quad (3.26)$$

$$= |a\rangle \langle a|a\rangle \langle a|\psi\rangle \quad (3.27)$$

$$= |a\rangle \langle a|\psi\rangle \quad (3.28)$$

$$= \hat{P}_a |\psi\rangle \quad (3.29)$$

Exercise Six operator are defined as follows:

$$\hat{A}_1 \psi(x) = [\psi(x)]^2 \quad (3.30)$$

$$\hat{A}_2 \psi(x) = \frac{d}{dx} \psi(x) \quad (3.31)$$

$$\hat{A}_3 \psi(x) = \int_a^x \psi(x') dx' \quad (3.32)$$

$$\hat{A}_4 \psi(x) = x^2 \psi(x) \quad (3.33)$$

$$\hat{A}_5 \psi(x) = \sin(\psi(x)) \quad (3.34)$$

$$\hat{A}_6 \psi(x) = \frac{d^2}{dx^2} \psi(x) \quad (3.35)$$

which of these operators \hat{A}_i are linear operator.

3.5 Representation of vectors and operators

Let ϕ_i be a complete orthonormal basis set in a Hilbert space. Since the basis is orthonormal, we must have

$$(\phi_i, \phi_j) = \delta_{ij} \quad (3.36)$$

An arbitrary vector ψ_a can be written as a linear combination of the basis vectors.

We write

$$\psi_a = \sum_i a_i \phi_i \quad (3.37)$$

Where the scalars a_i are the components of the vector ψ_a along the basis vectors ϕ_i . Using the orthonormality of the basis vectors we immediately have

$$a_i = (\phi_i, \psi_a) \quad (3.38)$$

We can arrange these numbers as a column matrix

$$\begin{pmatrix} a_1 \\ a_2 \\ \vdots \end{pmatrix} \equiv \begin{pmatrix} (\phi_1, \psi_a) \\ (\phi_2, \psi_a) \\ \vdots \end{pmatrix} \quad (3.39)$$

This column matrix is called the representation of the vector ψ_a with respect to the given basis $\{\phi_i\}$. In Dirac notation we represent the vector ψ_a as $|a\rangle$ and the basis vectors ϕ_i are written as $|i\rangle$. We can expand a general ket $|a\rangle$ as a linear combination of the basis kets:

$$|a\rangle = \sum_{i=1}^{\infty} a_i |i\rangle \quad (3.40)$$

Orthonormality of the basis kets can be written as

$$\langle i|j \rangle = \delta_{ij} \quad (3.41)$$

The complex scalar a_i are called the components of the ket $|a\rangle$ along $|i\rangle$. Using the orthonormality condition of the basis vectors (3.40), we have

$$a_i = \langle i|\psi \rangle \quad (3.42)$$

These scalars a_1, a_2, \dots arranged as a column matrix is called the representation of $|a\rangle$ in the basis $\{|i\rangle\}, i = 1, 2, \dots$

Thus

$$|a\rangle \rightarrow \begin{pmatrix} a_1 \\ a_2 \\ \vdots \end{pmatrix} \equiv \begin{pmatrix} \langle 1|a \rangle \\ \langle 2|a \rangle \\ \vdots \end{pmatrix} \quad (3.43)$$

We can write down the representation of any one of the basis vectors in the same basis as

$$|1\rangle \rightarrow \begin{pmatrix} 1 \\ 0 \\ 0 \\ \vdots \end{pmatrix} \quad |2\rangle \rightarrow \begin{pmatrix} 0 \\ 1 \\ 0 \\ \vdots \end{pmatrix} \quad |3\rangle \rightarrow \begin{pmatrix} 0 \\ 0 \\ 1 \\ \vdots \end{pmatrix} \quad (3.44)$$

and so on.

Now, using equation ?? in equation ?? we have

$$|a\rangle = \sum_i a_i |i\rangle \quad (3.45)$$

$$= \sum_i \langle i|a \rangle |i\rangle \quad (3.46)$$

$$= \sum_i |i\rangle \langle i|a \rangle \quad (3.47)$$

$$= \left(\sum_i \hat{P}_i \right) |a\rangle \quad (3.48)$$

where

$$\hat{P}_i = |i\rangle \langle i| \quad (3.49)$$

is the projection operator along $|i\rangle$. Since equation 3.45 is true for all $|a\rangle$ in the vector space (this is because $\{|i\rangle\}$ form a complete set), we must have

$$\sum_i \hat{P}_i = \sum_i |i\rangle \langle i| = \hat{\mathbb{I}} \quad (3.50)$$

where \mathbb{I} is the identity operator. Equation 3.50 is called the **completeness condition** for the basis vectors.

3.6 Matrix representation of ket and bra vectors

The ket vector $|a\rangle$ is represented by a column matrix

$$\begin{pmatrix} a_1 \\ a_2 \\ \vdots \end{pmatrix} \equiv \begin{pmatrix} \langle 1|a\rangle \\ \langle 2|a\rangle \\ \vdots \end{pmatrix} \quad (3.51)$$

in a basis $\{|i\rangle\}$. The dual of ket $|a\rangle$ is the $\langle a|$. But what is the matrix representation of the bra $\langle a|$ in the same basis? To see this we can expand $\langle a|$ as

$$\langle a| = \sum_i \langle a|i\rangle \langle i| \quad (3.52)$$

The $\langle a|$ is represented by a row vector:

$$\langle a| \rightarrow (\langle a|1\rangle \quad \langle a|2\rangle \quad \dots) = (a_1^* \quad a_2^* \quad \dots) \quad (3.53)$$

Then the scalar product becomes a number. Thus

$$\langle a|a\rangle = (a_1^* \quad a_2^* \quad \dots) \begin{pmatrix} a_1 \\ a_2 \\ \vdots \end{pmatrix} \quad (3.54)$$

$$= a_1^* a_1 + a_2^* a_2 + \dots \quad (3.55)$$

$$= \sum_i a_i^* a_i \quad (3.56)$$

$$= \sum_i |a_i|^2 \quad (3.57)$$

Here the quantity $|a_i|^2$ is just a number.

In general

$$\langle b|a\rangle = \sum_i \langle b|i\rangle \langle i|a\rangle \quad (3.58)$$

$$= (\langle b|1\rangle \quad \langle b|2\rangle \quad \dots) \begin{pmatrix} \langle 1|a\rangle \\ \langle 2|a\rangle \\ \vdots \end{pmatrix} \quad (3.59)$$

$$= (b_1^* \quad b_2^* \quad \dots) \begin{pmatrix} a_1 \\ a_2 \\ \vdots \end{pmatrix} \quad (3.60)$$

$$= b_1^* a_1 + b_2^* a_2 + \dots \quad (3.61)$$

$$= \sum_i b_i^* a_i \quad (3.62)$$

Here the quantity $b_i^* a_i$ is complex number.

3.7 Representation of an operator in a basis

Consider the equation

$$|b\rangle = \hat{A} |a\rangle \quad (3.63)$$

Let $\{|i\rangle\}$ where $i = 1, 2, \dots$ be a complete set of orthonormal basis states. Taking the component of equation 3.63 along $|i\rangle$, we have

$$\langle i|b\rangle = \langle i|\hat{A}|a\rangle \quad (3.64)$$

$$= \sum_j \langle i|\hat{A}|j\rangle \langle j|a\rangle \quad (3.65)$$

In matrix notation

$$b_i = \sum_j A_{ij} a_j \quad (3.66)$$

where

$$b_i \equiv \langle i|b\rangle \quad (3.67)$$

$$a_j \equiv \langle j|a\rangle \quad (3.68)$$

$$A_{ij} \equiv \langle i|\hat{A}|j\rangle \quad (3.69)$$

Writing in full, equation 3.66 becomes

$$\begin{pmatrix} b_1 \\ b_2 \\ \vdots \\ b_n \end{pmatrix} = \begin{pmatrix} A_{11} & A_{12} & \dots \\ A_{21} & A_{22} & \dots \\ \vdots & \ddots & \dots \\ \vdots & \vdots & \dots \\ \vdots & \ddots & \dots \end{pmatrix} \begin{pmatrix} a_1 \\ a_2 \\ \vdots \\ a_n \end{pmatrix} \quad (3.70)$$

The matrix $[A]$ with elements $A_{ij} = \langle i | \hat{A} | j \rangle$ is called the matrix representation of the operator \hat{A} with respect to the given basis $\{|i\rangle\}$. Using a basis set, the operator \hat{A} can also be written as

$$\hat{A} = \hat{\mathbb{I}} \hat{A} \hat{\mathbb{I}} \quad (3.71)$$

$$= \left(\sum_i |i\rangle \langle i| \right) \hat{A} \left(\sum_j |j\rangle \langle j| \right) \quad (3.72)$$

$$= \sum_{i,j} |i\rangle \langle i| \hat{A} |j\rangle \langle j| \quad (3.73)$$

$$= \sum_{i,j} |i\rangle A_{ij} \langle j| \quad (3.74)$$

3.8 Matrix representation of the sum and product of two operators

Let

$$\hat{C} = \hat{A} + \hat{B} \quad (3.75)$$

Then

$$C_{ij} = \langle i | \hat{C} | j \rangle \quad (3.76)$$

$$= \langle i | \hat{A} + \hat{B} | j \rangle \quad (3.77)$$

$$= \langle i | \hat{A} | j \rangle + \langle i | \hat{B} | j \rangle \quad (3.78)$$

$$= A_{ij} + B_{ij} \quad (3.79)$$

Next, let \therefore

$$C_{ij} = \langle i | \hat{C} | j \rangle \quad (3.80)$$

$$= \langle i | \hat{A} \hat{B} | j \rangle \quad (3.81)$$

$$= \langle i | \hat{A} \hat{\mathbb{I}} \hat{B} | j \rangle \quad (3.82)$$

$$= \sum_k \langle i | \hat{A} | k \rangle \langle k | \hat{B} | j \rangle \quad (3.83)$$

$$= \sum_k A_{ik} B_{kj} \quad (3.84)$$

In full matrix form

$$[C] = [A][B] \quad (3.85)$$

where

$$[A] = \begin{pmatrix} \langle 1 | \hat{A} | 1 \rangle & \langle 1 | \hat{A} | 2 \rangle & \dots \\ \langle 2 | \hat{A} | 1 \rangle & \langle 2 | \hat{A} | 2 \rangle & \dots \\ \vdots & \vdots & \vdots \end{pmatrix} \quad (3.86)$$

and similar expression applies for $[B]$ and $[C]$. This result shows that the matrix of an operator product is equal to the product of the matrices representing the operators, taken in the same order.

Example 1 Using a basis set $\{|i\rangle\}$ write down $\langle b|\hat{A}|a\rangle$ as a matrix product.

Ans

$$\langle b|\hat{A}|a\rangle = \sum_{i,j} \langle b|i\rangle \langle i|\hat{A}|j\rangle \langle j|a\rangle \quad (3.87)$$

$$= \sum_{i,j} b_i^* A_{ij} a_j \quad (3.88)$$

$$= [b]^\dagger [A][a] \quad (3.89)$$

where $[b]^\dagger$ is the matrix representation of $\langle b|$.

$$[b]^\dagger = (b_1^* \ b_2^* \ \dots) \quad (3.90)$$

$[A]$ is the matrix representation of the operator \hat{A}

$$[A] = \begin{bmatrix} A_{11} & A_{12} & \dots \\ A_{21} & A_{22} & \dots \\ \vdots & \vdots & \vdots \end{bmatrix} \quad (3.91)$$

and $[a]$ is the matrix representation of the ket $|a\rangle$.

$$[a] = \begin{pmatrix} a_1 \\ a_2 \\ \vdots \end{pmatrix} \quad (3.92)$$

writing in full

$$\langle b|\hat{A}|a\rangle = (b_1^* \ b_2^* \ \dots) \begin{pmatrix} A_{11} & A_{12} & \dots \\ A_{21} & A_{22} & \dots \\ \vdots & \vdots & \vdots \end{pmatrix} \begin{pmatrix} a_1 \\ a_2 \\ \vdots \end{pmatrix} \quad (3.93)$$

4. sheet-4 : Linear Vector Space (continued)

4.1 Adjoint operator

Consider the equation

$$|b\rangle = \hat{K}|a\rangle \quad (4.1)$$

The operator \hat{K} carries the ket $|a\rangle$ to the ket $|b\rangle$. The dual of $|a\rangle$ and $|b\rangle$ are the bras $\langle a|$ and $\langle b|$ respectively. Then the operator which carries $\langle a|$ to $\langle b|$ is called the adjoint of \hat{K} and is denoted by \hat{K}^\dagger . Thus in dual space 4.1 is

$$\langle b| = \langle a|\hat{K}^\dagger \quad (4.2)$$

$$\text{ket space } |a\rangle \xrightarrow{\hat{K}} |b\rangle \quad (4.3)$$

$$\text{bra space } \langle a| \xrightarrow{\hat{K}^\dagger} \langle b| \quad (4.4)$$

Here bra space is the dual of ket space and

$$|a\rangle \xrightarrow{dc} \langle a| \quad (4.5)$$

$$|b\rangle \xrightarrow{dc} \langle b| \quad (4.6)$$

dc \equiv dual correspondence.

From equation 4.1 and 4.2 it follows that

$$\langle c|b\rangle = \langle c|\hat{K}|a\rangle \quad (4.7)$$

and

$$\langle b|c\rangle = \langle a|\hat{K}^\dagger|c\rangle \quad (4.8)$$

since

$$\langle b|c\rangle = \langle c|b\rangle^* \quad (4.9)$$

we have

$$\langle c|\hat{K}|a\rangle = \langle a|\hat{K}^\dagger|c\rangle \quad (4.10)$$

equation 4.10 is the defining equation for the adjoint \hat{K}^\dagger of the operator \hat{K} . In scalar product notation

$$\langle c|b\rangle \equiv (\psi_c, \psi_b) \quad (4.11)$$

equation 4.10 can be written as

$$(\psi_c, \hat{K}\psi_a) = (\psi_a, \hat{K}^\dagger\psi_c)^* \quad (4.12)$$

$$= (\hat{K}^\dagger\psi_c, \psi_a) \quad (4.13)$$

In particular, if we take $|c\rangle$ and $|a\rangle$ as the basis states $|i\rangle$ and $|j\rangle$, equation 4.10 becomes

$$\langle i|\hat{K}|j\rangle = \langle j|\hat{K}^\dagger|i\rangle^* \quad (4.14)$$

$$K_{ij} = (\hat{K}^\dagger)_{ji}^* \quad (4.15)$$

$$K_{ji}^\dagger = K_{ij}^* \quad (4.16)$$

$$K_{ij}^\dagger = (K_{ji})^* \quad (4.17)$$

$$[\hat{K}^\dagger] = [K]^\dagger \quad (4.18)$$

i.e., The matrix representation of the adjoint operator is the hermitian conjugate of the matrix representation of \hat{K} .

4.2 Hermitian or Self-adjoint operator

If $\hat{K}^\dagger = \hat{K}$, then \hat{K} is said to be a self-adjoint or a hermitian operator. For a hermitian operator

$$[K] = [K^\dagger] = [K]^\dagger \quad (4.19)$$

or

$$K_{ij} = K_{ji}^* \quad (4.20)$$

i.e., $[K]$ is a hermitian matrix.

Example 1 show that

$$(AB)^\dagger = B^\dagger A^\dagger \quad (4.21)$$

Answer:

$$(\psi_a, \hat{A}\hat{B}\psi_b) = ((\hat{A}\hat{B})^\dagger \psi_a, \psi_b) \quad (4.22)$$

Also

$$(\psi_a, \hat{A}\hat{B}\psi_b) = (\hat{A}^\dagger \psi_a, \hat{B}\psi_b) = (\hat{B}^\dagger \hat{A}^\dagger \psi_a, \psi_b) \quad (4.23)$$

Comparing these equation (equation 4.22 and 4.23) we get,

$$(\hat{A}\hat{B})^\dagger = \hat{B}^\dagger \hat{A}^\dagger \quad (4.24)$$

4.3 Inverse operator

An operator \hat{B} is said to be the inverse of \hat{A} if

$$\hat{A}\hat{B} = \hat{B}\hat{A} = \hat{\mathbb{I}} \quad (4.25)$$

Obviously, if \hat{B} is the inverse of \hat{A} , then \hat{A} is the inverse of \hat{B} . We write

$$\hat{B} = \hat{A}^{-1} \quad (4.26)$$

or

$$\hat{A} = \hat{B}^{-1} \quad (4.27)$$

If equation 4.25 is satisfied.

4.4 Unitary operator

An operator \hat{U} is said to be unitary if

$$\hat{U}\hat{U}^\dagger = \hat{U}^\dagger\hat{U} = \hat{\mathbb{I}} \quad (4.28)$$

i.e., if

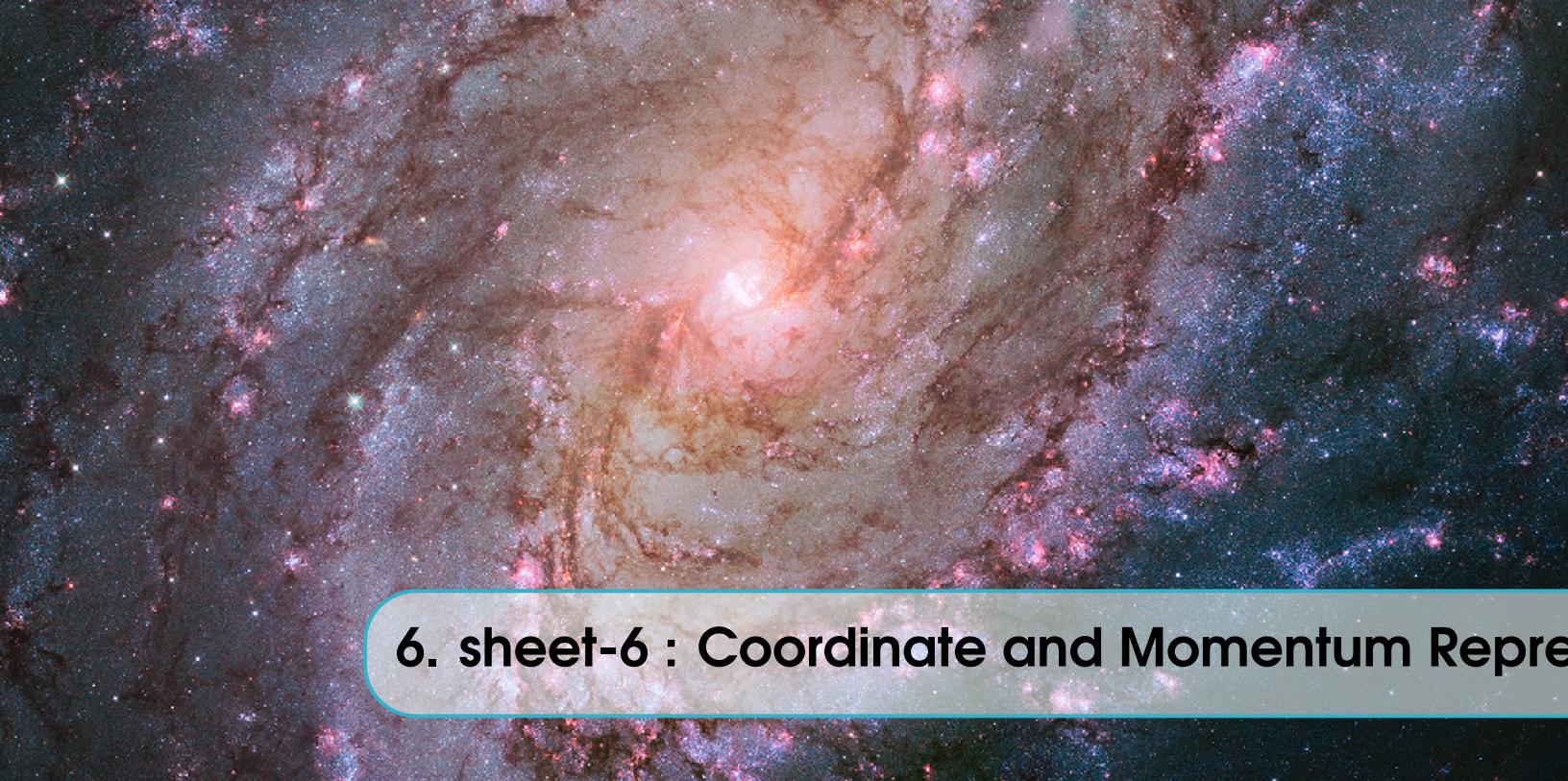
$$\hat{U}^\dagger = \hat{U}^{-1} \quad (4.29)$$

Thus, for a unitary operator, its adjoint is @@@@ @@@@ @@@@ @@@@ @@@@ @@@@ @@@@ @

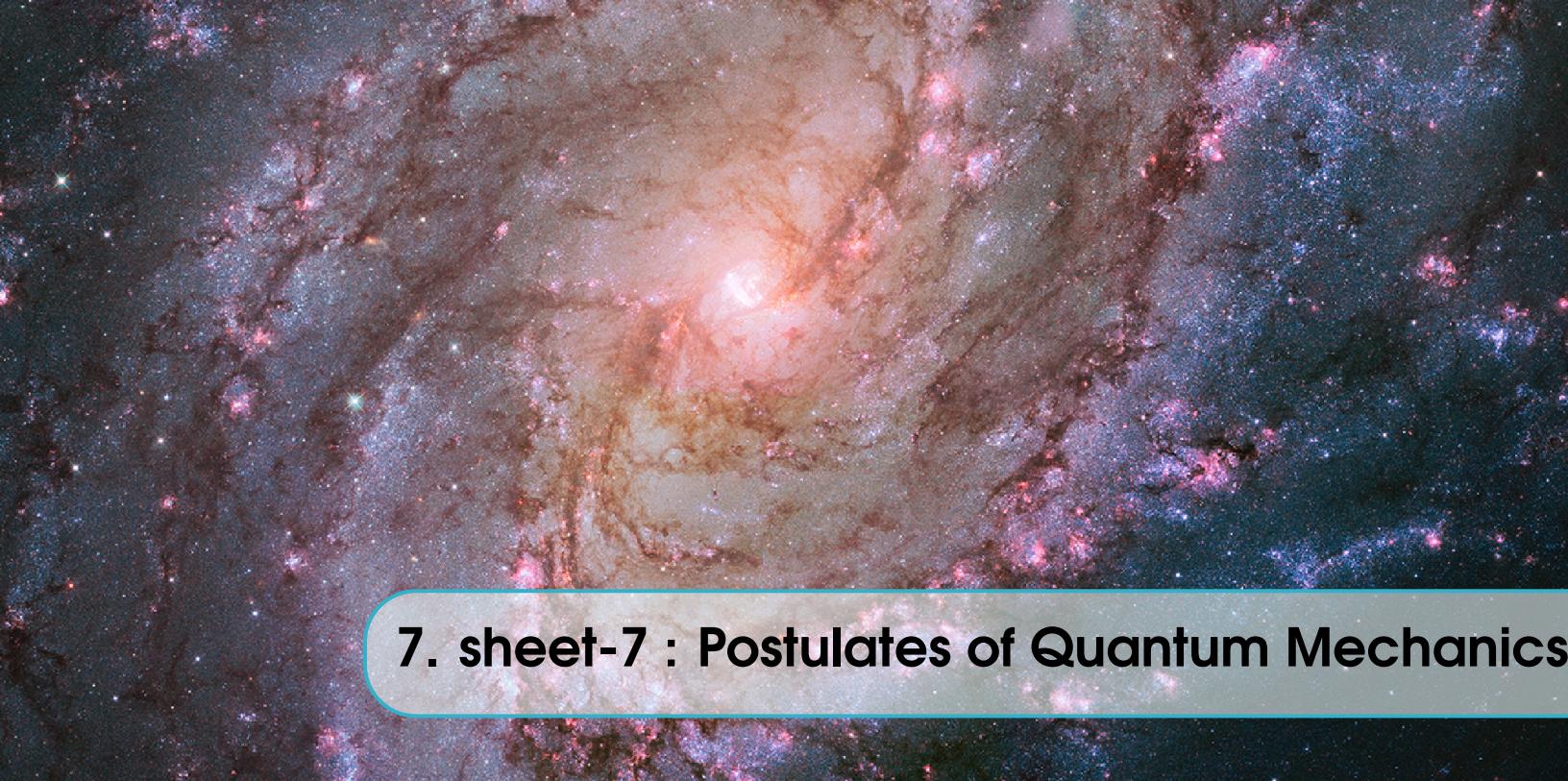
4.5 Function of Operators



5. sheet-5 : Eigenvalue and Eigenvectors of Operators



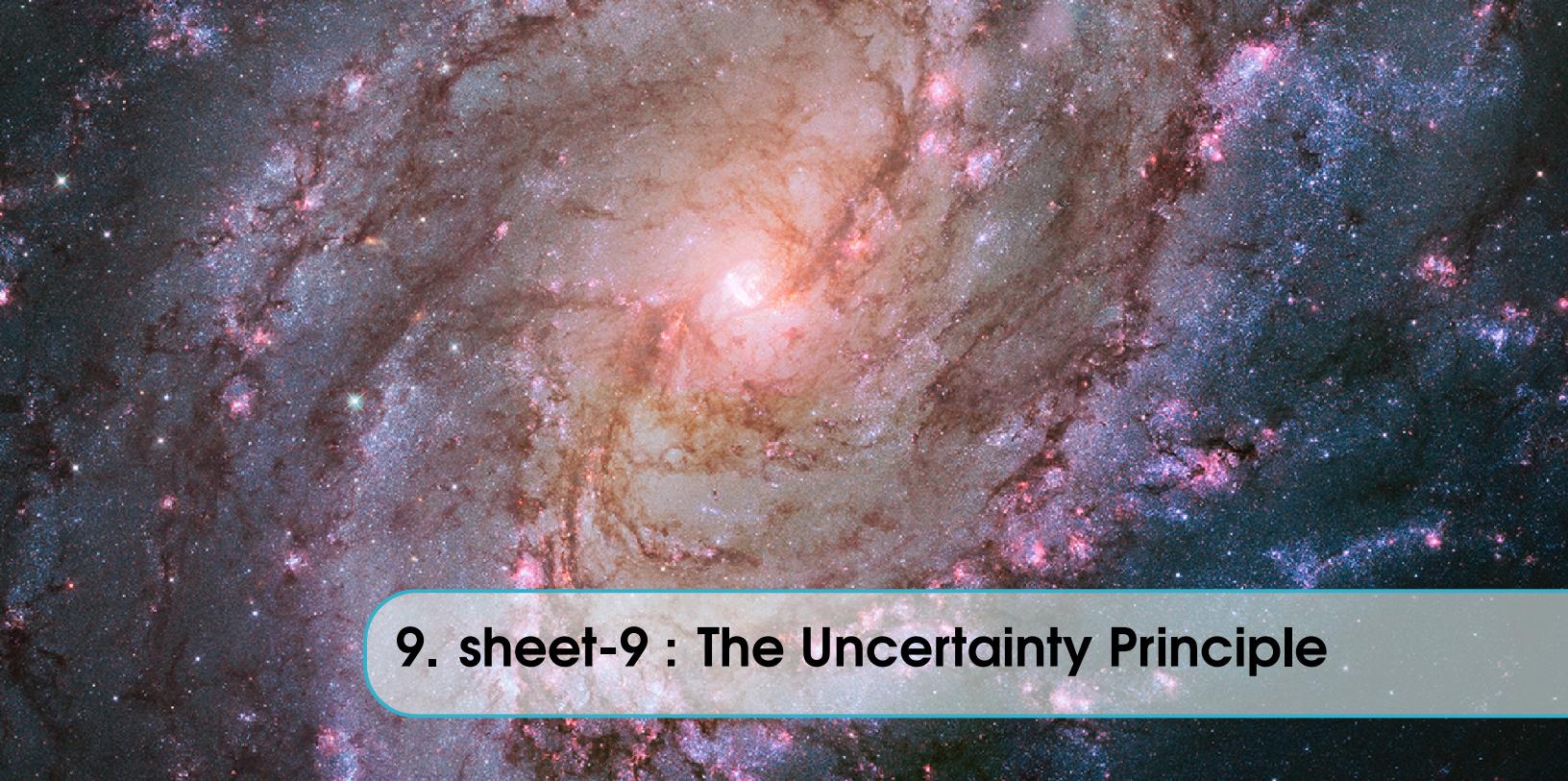
6. sheet-6 : Coordinate and Momentum Repre



7. sheet-7 : Postulates of Quantum Mechanics



8. sheet-8 : Compatibility of Observable



9. sheet-9 : The Uncertainty Principle



10. sheet-10 : Harmonic Oscillator



11. sheet-11 : Quantum Dynamics



12. sheet-12 : The Path Integral Formulation of



13. sheet-13 : Potential Wells and Barriers



14. sheet-14 : Particle Motion in 3-Dimension



15. sheet-15 : The Hydrogen Atom



16. sheet-16 : Angular Momentum 1



17. sheet-17 : Angular Momentum 2: spin 1/2



18. sheet-18 : Rotations and Angular Momentum



19. sheet-19 : Addition of Angular Momentum



20. sheet-20 : Time Independent Perturbation



21. sheet-21 : Variational Method



22. sheet-22 : WKB approximation



23. sheet-23 : Time Dependent Perturbation



24. sheet-24 : Scattering Theory



25. sheet-25 : Integral Equation for Scattering



26. sheet-26 : Identical Particles