

PHYS 23410 (Quantum Mechanics I) Problem Sets

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1 Formalism of Quantum Mechanics

1/12: 1. The Schrödinger equation is given by

$$\left(-\frac{\hbar^2}{2m}\vec{\nabla}^2 + V(\vec{r}, t)\right)\psi(\vec{r}, t) = i\hbar\frac{\partial\psi(\vec{r}, t)}{\partial t}$$

a) Use this equation, and its complex conjugate, to demonstrate the **continuity equation**

$$\frac{\partial|\psi(\vec{r}, t)|^2}{\partial t} + \vec{\nabla} \cdot \left(\frac{i\hbar}{2m}\right) (\psi\vec{\nabla}\psi^* - \psi^*\vec{\nabla}\psi) = 0$$

where the first term, with $|\psi|^2 = \psi^*\psi$, is the partial derivative of the probability density, while the second term is the divergence of the probability current density.

Answer. Multiply both sides of the Schrödinger equation by $-i/\hbar$:

$$\frac{\partial\psi}{\partial t} = \left(\frac{i\hbar}{2m}\vec{\nabla}^2 - \frac{i}{\hbar}V\right)\psi$$

We may then obtain the complex conjugate of the above equation by replacing all instances of i with its complex conjugate $-i$ and ψ with its complex conjugate ψ^* . In a nutshell, this works because we have a case of multiplying some complex number $a + bi = \psi(\vec{r})$ on the left by i , which gives $i(a + bi) = -b + ai$, and the complex conjugate of this number is $-b - ai = -i(a - bi) = -i(a + bi)^*$.

$$\frac{\partial\psi^*}{\partial t} = \left(-\frac{i\hbar}{2m}\vec{\nabla}^2 + \frac{i}{\hbar}V\right)\psi^*$$

We will use the above two equations to substitute into the following algebraic derivation, which yields the desired result.

$$\begin{aligned}\frac{\partial|\psi|^2}{\partial t} &= \frac{\partial}{\partial t}(\psi^*\psi) \\ &= \psi^*\frac{\partial\psi}{\partial t} + \psi\frac{\partial\psi^*}{\partial t} \\ &= \psi^*\left(\frac{i\hbar}{2m}\vec{\nabla}^2 - \frac{i}{\hbar}V\right)\psi + \psi\left(-\frac{i\hbar}{2m}\vec{\nabla}^2 + \frac{i}{\hbar}V\right)\psi^* \\ &= \frac{i\hbar}{2m}(\psi^*\vec{\nabla}^2\psi - \psi\vec{\nabla}^2\psi^*) \\ &= \frac{i\hbar}{2m}\left[(\vec{\nabla}\psi^*\vec{\nabla}\psi + \psi^*\vec{\nabla}^2\psi) - (\vec{\nabla}\psi\vec{\nabla}\psi^* + \psi\vec{\nabla}^2\psi^*)\right] \\ &= \frac{i\hbar}{2m}\vec{\nabla} \cdot (\psi^*\vec{\nabla}\psi - \psi\vec{\nabla}\psi^*) \\ &= -\vec{\nabla} \cdot \left(\frac{i\hbar}{2m}\right) (\psi\vec{\nabla}\psi^* - \psi^*\vec{\nabla}\psi)\end{aligned}$$

□

b) Discuss the physical interpretation of this equation. What happens if you integrate the first and second terms over a region of space, defined by a finite volume V and separated by the rest of space by a boundary area S ?

Hint: Use the analogy with the case of electromagnetic charge. If you integrate over the whole volume of space, the continuity equation leads to charge conservation. The probability current density, just as the charge current density, is assumed to vanish sufficiently fast at infinity, so that there is no flow of probability (charge) at infinity.

Answer. This equation implies that probability density is *locally* conserved. In its current differential form, it states that the change in probability density (first term) is exactly offset by the divergence or the probability current density (second term). This interpretation is clarified upon integrating over a volume V contained within a boundary surface S . In the integral form, the first term becomes the change in probability density within V , i.e., the rate at which the particle becomes more or less likely to exist within V . Moreover, this is equal to the opposite of the second term, which when integrated gives the flux of the probability density through S . Essentially, if the particle is going to become less likely to exist within V , then a corresponding amount of probability density must flow out of V through S (and vice versa if the particle is going to become more likely to exist within V).

Additionally, note that if you take the limit as $V \rightarrow \mathbb{C}^3$, both terms go to zero (because probability density is conserved within the entire space). Since the sum of two terms equal to zero is zero, this is another way of verifying the equality in the continuity equation. \square

2. Consider the expectation value of the operator $\hat{p}^2 = -\hbar^2 \vec{\nabla}^2$, namely

$$\int d^3\vec{r} \psi(\vec{r}, t)^* \left(-\hbar^2 \vec{\nabla}^2 \psi(\vec{r}, t) \right)$$

- a) Using integration by parts, demonstrate that this can be rewritten as the integral of the modulus square of the gradient of $\hbar\psi(\vec{r}, t)$ and that it is therefore a positive quantity. As before, assume that the wave functions go sufficiently fast to zero at infinity.

Answer. To begin, we need to derive a three-dimensional analogy to the standard one-dimensional integration by parts formula. Let $u, v : \mathbb{C}^3 \rightarrow \mathbb{C}^3$ be functions with smoothness constraints analogous to ψ . As in the one-dimensional case, we'll start with the product rule and integrate.

$$\begin{aligned} \vec{\nabla} \cdot (u \vec{\nabla} v) &= (\vec{\nabla} u)(\vec{\nabla} v) + u \vec{\nabla}^2 v \\ \int_{\Omega} \vec{\nabla} \cdot (u \vec{\nabla} v) &= \int_{\Omega} (\vec{\nabla} u)(\vec{\nabla} v) + \int_{\Omega} u \vec{\nabla}^2 v \\ \int_{\Omega} \vec{\nabla} \cdot (u \vec{\nabla} v) - \int_{\Omega} (\vec{\nabla} u)(\vec{\nabla} v) &= \int_{\Omega} u \vec{\nabla}^2 v \end{aligned}$$

We can then apply this result and simplify to yield the final result.

$$\begin{aligned} \int d^3\vec{r} \psi^* \left(-\hbar^2 \vec{\nabla}^2 \psi \right) &= - \int d^3\vec{r} \underbrace{\hbar\psi^*}_u \underbrace{\vec{\nabla}^2(\hbar\psi)}_{\vec{\nabla}^2 v} \\ &= - \underbrace{\int_{\Omega} \vec{\nabla} \cdot [(\hbar\psi^*) \vec{\nabla}(\hbar\psi)]}_0 + \int d^3\vec{r} \vec{\nabla}(\hbar\psi^*) \vec{\nabla}(\hbar\psi) \\ &= \int d^3\vec{r} |\vec{\nabla}(\hbar\psi)|^2 \end{aligned}$$

To clarify, we know that the first term given by integration by parts goes to zero because of the divergence theorem. In particular, let Ω be a finite region of space encapsulated by $d\Omega$; we will take the limit as $\Omega \rightarrow \mathbb{C}^3$ and $d\Omega$ approaches the boundary of \mathbb{C}^3 . Then

$$\int_{\Omega} \vec{\nabla} \cdot [\hbar\psi^* \vec{\nabla}(\hbar\psi)] = \int_{d\Omega} [\hbar\psi^* \vec{\nabla}(\hbar\psi)] \cdot \hat{n} dS$$

Essentially, this means that the original integral is equal to an integral of an integrand containing ψ^* (which goes to zero at the “boundary” of \mathbb{C}^3) at a surface approaching the boundary through the aforementioned limit. This means that the second integral — under the limit that $d\Omega$ approaches the “boundary” of \mathbb{C}^3 — is zero, justifying the original substitution. \square

- b) Now consider the expectation value of the Hamiltonian $\hat{p}^2/2m + V(\vec{r}, t)$ and assume that the function $\psi(\vec{r}, t)$ is an eigenfunction of the Hamiltonian. In such a case,

$$\hat{H}\psi(\vec{r}, t) = E\psi$$

and the particle therefore has a well-defined energy, equal to E . Demonstrate, based on the result of part (a), that E must be larger than the minimum value of $V(\vec{r}, t)$.

Hint: Use $(\hat{p}^2/2m)\psi = (E - V)\psi$, and the fact that the mean value of V should be larger than its minimum value.

The lesson is that the particle can enter regions of space where its energy is lower than the potential, but this is not possible everywhere in space. The fact that the particle can go through regions of space where its energy is lower than the potential (the wave function does not vanish in those regions of space) leads to the famous phenomenon of tunneling, namely a particle can go through a *finite region of space where the potential is higher than its energy* and has a probability of being transmitted to the other side.

Answer. Taking the hint, we have that

$$\begin{aligned} E - \langle V \rangle &= E \cdot 1 - \langle V \rangle \\ &= E \cdot \int d^3\vec{r} \psi^* \psi - \int d^3\vec{r} \psi^* V \psi \\ &= \int d^3\vec{r} \psi^* (E - V) \psi \\ &= \int d^3\vec{r} \psi^* \left(\frac{\hat{p}^2}{2m} \right) \psi \\ &= \frac{1}{2m} \int d^3\vec{r} \psi^* (-\hbar^2 \vec{\nabla}^2) \psi \\ &= \frac{1}{2m} \int d^3\vec{r} |\vec{\nabla}(\hbar\psi)|^2 && \text{Part (a)} \\ &> 0 \end{aligned}$$

Taking the hint again, we have that

$$\langle V \rangle > V_{\min}$$

Therefore, by transitivity, we have that

$$\begin{aligned} E - \langle V \rangle &> 0 \\ E &> \langle V \rangle > V_{\min} \end{aligned}$$

as desired. □

3. We shall define **Hermitian operators** as those ones \hat{O} satisfying the property

$$\int d^3\vec{r} \psi_m^*(\vec{r}, t) \hat{O} \psi_n(\vec{r}, t) = \left(\int d^3\vec{r} \psi_n^*(\vec{r}, t) \hat{O} \psi_m(\vec{r}, t) \right)^* \quad (1.1)$$

where ψ_m is a particular solution of the Schrödinger equation. Observe that when you identify $\psi_n = \psi_m$, you obtain that the mean value of a Hermitian operator is real and thus can be associated with an observable.

Observe also that, in general, this could be written as

$$\int d^3\vec{r} \psi_m^*(\vec{r}, t) \hat{O} \psi_n(\vec{r}, t) = \int d^3\vec{r} [\hat{O} \psi_m(\vec{r}, t)]^* \psi_n(\vec{r}, t)$$

Therefore, in the case of a Hermitian operator, I can “transfer” the application of the operator from the right to the left.

- a) Use this property to demonstrate that if you take two different Hermitian operators and you transfer them in the proper order, then

$$\langle \psi_m | \hat{O}_1 \hat{O}_2 | \psi_n \rangle = \left(\langle \psi_n | \hat{O}_2 \hat{O}_1 | \psi_m \rangle \right)^*$$

where I used the Dirac notation. Observe that if I take \hat{O}_2 to be a real constant, this equation reduces to Equation 1.1.

Answer. We have that

$$\begin{aligned} \langle \psi_m | \hat{O}_1 \hat{O}_2 | \psi_n \rangle &= \int d^3\vec{r} \psi_m^* \hat{O}_1 \hat{O}_2 \psi_n \\ &= \int d^3\vec{r} (\hat{O}_1 \psi_m)^* \hat{O}_2 \psi_n \\ &= \int d^3\vec{r} [\hat{O}_2 (\hat{O}_1 \psi_m)]^* \psi_n \\ &= \left(\int d^3\vec{r} \psi_n^* [\hat{O}_2 (\hat{O}_1 \psi_m)] \right)^* \\ &= \left(\int d^3\vec{r} \psi_n^* \hat{O}_2 \hat{O}_1 \psi_m \right)^* \\ &= \left(\langle \psi_n | \hat{O}_2 \hat{O}_1 | \psi_m \rangle \right)^* \end{aligned}$$

as desired. □

- b) Use this relation to demonstrate that the mean value of the **commutator** of two Hermitian operators, which is given by

$$[\hat{O}_1, \hat{O}_2] = \hat{O}_1 \hat{O}_2 - \hat{O}_2 \hat{O}_1$$

is a pure imaginary number, and hence (unless it is multiplied by an imaginary factor), cannot be associated with a physical observable. In the particular example of momentum and position, for instance, $[\hat{p}_i, \hat{r}_j] = -i\hbar\delta_{ij}$, where δ_{ij} is the **Kronecker delta**. Do it for the mean value of the commutator in a particular state with wave function ψ_n .

Answer. We have that

$$\begin{aligned} \langle \psi | [\hat{O}_1, \hat{O}_2] | \psi \rangle &= \langle \psi | \hat{O}_1 \hat{O}_2 - \hat{O}_2 \hat{O}_1 | \psi \rangle \\ &= \langle \psi | \hat{O}_1 \hat{O}_2 | \psi \rangle - \langle \psi | \hat{O}_2 \hat{O}_1 | \psi \rangle \\ &= \underbrace{\langle \psi | \hat{O}_1 \hat{O}_2 | \psi \rangle}_{a+bi} - \underbrace{(\langle \psi | \hat{O}_1 \hat{O}_2 | \psi \rangle)^*}_{a-bi} \end{aligned}$$

Since taking the difference of a complex number $a + bi$ and its complex conjugate $a - bi$ yields the purely imaginary number $2bi$, we have the desired result.

Now consider the explicit case where $\hat{O}_1 = \hat{p}_i$, $\hat{O}_2 = \hat{r}_j$, and $\psi = \psi_n$. Then we have that

$$\begin{aligned} \langle \psi_n | [\hat{p}_i, \hat{r}_j] | \psi_n \rangle &= \langle \psi_n | -i\hbar\delta_{ij} | \psi_n \rangle \\ &= -i\hbar\delta_{ij} \int d^3\vec{r} \psi_n^* \psi_n \\ &= -i\hbar\delta_{ij} \int d^3\vec{r} |\psi_n|^2 \end{aligned}$$

Therefore, the mean value is equal to an imaginary number times an integral that will be real, so the final answer is, indeed, purely imaginary. □

- c) Demonstrate that the above relation remains true if you compute the mean value of the commutator for an arbitrary wave function

$$\Psi = \sum_n c_n \psi_n$$

Hint: Organize the total sum that you obtain in pairs that share the same values of m and n , and demonstrate that when you add up the two terms with the same values of m and n , you obtain a purely imaginary number. The terms in the total sum for which $m = n$ don't have pairs, but you can use the result of part (b) to show that they are indeed imaginary.

Answer. We are interested in computing

$$\langle \Psi | [\hat{O}_1, \hat{O}_2] | \Psi \rangle$$

Taking the hint, we break it into two sums as follows.

$$\begin{aligned} \langle \Psi | [\hat{O}_1, \hat{O}_2] | \Psi \rangle &= \left\langle \sum_n c_n \psi_n \left| [\hat{O}_1, \hat{O}_2] \right| \sum_n c_n \psi_n \right\rangle \\ &= \sum_{n,m} \langle c_n \psi_n | [\hat{O}_1, \hat{O}_2] | c_m \psi_m \rangle \\ &= \sum_m \sum_{n=m}^n c_n^* c_n \langle \psi_n | [\hat{O}_1, \hat{O}_2] | \psi_n \rangle + \sum_m \sum_{n \neq m}^n c_m^* c_n \langle \psi_m | [\hat{O}_1, \hat{O}_2] | \psi_n \rangle \\ &= \sum_n c_n^* c_n \langle \psi_n | [\hat{O}_1, \hat{O}_2] | \psi_n \rangle \\ &\quad + \sum_{m < n} (c_m^* c_n \langle \psi_m | [\hat{O}_1, \hat{O}_2] | \psi_n \rangle + c_n^* c_m \langle \psi_n | [\hat{O}_1, \hat{O}_2] | \psi_m \rangle) \\ &= \sum_n c_n^* c_n \langle \psi_n | [\hat{O}_1, \hat{O}_2] | \psi_n \rangle \\ &\quad + \sum_{m < n} [c_m^* c_n \langle \psi_m | [\hat{O}_1, \hat{O}_2] | \psi_n \rangle - c_n^* c_m (\langle \psi_m | [\hat{O}_1, \hat{O}_2] | \psi_n \rangle)^*] \\ &= \sum_n c_n^* c_n \langle \psi_n | [\hat{O}_1, \hat{O}_2] | \psi_n \rangle \\ &\quad + \sum_{m < n} [c_m^* c_n \langle \psi_m | [\hat{O}_1, \hat{O}_2] | \psi_n \rangle - (c_m^* c_n \langle \psi_m | [\hat{O}_1, \hat{O}_2] | \psi_n \rangle)^*] \end{aligned}$$

From part (b), the first sum is a sum of purely imaginary numbers. Additionally, by the same logic as in part (b), the second sum is a sum of purely imaginary numbers. Thus, the overall term is a sum of purely imaginary numbers, and is thus a purely imaginary number, as desired^[1]. \square

¹Note that in the above derivation, I derive that terms across the diagonal from each other are complex conjugates of each other. However, I don't even need to do this because the commutator is Hermitian, so we have this by definition!

2 Infinite Well Motion and Quantum Tunneling

- 1/19: 1. In class, we demonstrated that given a certain time-independent potential, one can find solutions to the Schrödinger equation such that

$$-\frac{\hbar^2}{2m}\vec{\nabla}^2\psi_n(\vec{r}) + V(\vec{r})\psi_n(\vec{r}) = E_n\psi_n(\vec{r}) \quad (2.1)$$

Assume now that we are in one dimension, with the potential being a square well:

$$\begin{aligned} V(x) &\rightarrow \infty & \text{for } x \leq 0 \text{ and } x \geq a \\ V(x) &\rightarrow 0 & \text{for } 0 < x < a \end{aligned} \quad (2.2)$$

Show that in such a case, the solutions are given by

$$\psi_n(x) = \sqrt{\frac{2}{a}} \sin\left(\frac{n\pi x}{a}\right) \quad (2.3)$$

due to the fact that in order to get a finite mean energy value, the wave function must vanish at $x = 0, a$. The energy eigenstates are given by

$$E_n = \frac{n^2\pi^2\hbar^2}{2ma^2} \quad (2.4)$$

The factor $\sqrt{2/a}$ comes from the requirement of a good normalized solution, i.e., one with $\langle\psi_n|\psi_n\rangle = 1$. Now, imagine that at $t = 0$, the particle is in the state

$$\psi(x, 0) = \frac{A}{\sqrt{a}} \sin\left(\frac{\pi x}{a}\right) + \sqrt{\frac{3}{5a}} \sin\left(\frac{3\pi x}{a}\right) + \frac{1}{\sqrt{5a}} \sin\left(\frac{5\pi x}{a}\right) \quad (2.5)$$

where A is a real constant.

- a) Find the value of A such that $\psi(x, 0)$ is normalized. (Hint: Use $\langle\psi_n|\psi_m\rangle = \delta_{nm}$.)

Answer. For $\psi(x, 0)$ to be normalized, it must satisfy

$$\langle\psi(x, 0)|\psi(x, 0)\rangle = 1$$

Now recognize that $\psi(x, 0)$ is of the form

$$\psi = c_1\psi_1 + c_3\psi_3 + c_5\psi_5$$

Thus, we have that

$$\begin{aligned} 1 &= \langle\psi|\psi\rangle \\ &= \langle c_1\psi_1 + c_3\psi_3 + c_5\psi_5 | c_1\psi_1 + c_3\psi_3 + c_5\psi_5 \rangle \\ &= \langle c_1\psi_1 | c_1\psi_1 \rangle + \langle c_3\psi_3 | c_3\psi_3 \rangle + \langle c_5\psi_5 | c_5\psi_5 \rangle + 2 \underbrace{\langle c_1\psi_1 | c_3\psi_3 \rangle}_0 + 2 \underbrace{\langle c_1\psi_1 | c_5\psi_5 \rangle}_0 + 2 \underbrace{\langle c_3\psi_3 | c_5\psi_5 \rangle}_0 \\ &= \langle c_1\psi_1 | c_1\psi_1 \rangle + \langle c_3\psi_3 | c_3\psi_3 \rangle + \langle c_5\psi_5 | c_5\psi_5 \rangle \\ &= \int_0^a \frac{A^2}{a} \sin^2\left(\frac{\pi x}{a}\right) dx + \int_0^a \frac{3}{5a} \sin^2\left(\frac{3\pi x}{a}\right) dx + \int_0^a \frac{1}{5a} \sin^2\left(\frac{5\pi x}{a}\right) dx \\ &= \frac{A^2}{2} + \frac{3}{10} + \frac{1}{10} \\ &= \frac{5A^2 + 4}{10} \end{aligned}$$

$A = \sqrt{\frac{6}{5}}$

□

- b) If measurements of the energy are carried out, what are the values that will be found and what are the probabilities of measuring such energies? Calculate the average energy.

Answer. As mentioned above, $\psi(x, 0)$ is of the form

$$\psi(x, 0) = c_1\psi_1(x) + c_3\psi_3(x) + c_5\psi_5(x)$$

Thus, the energies that will be found are

$$\boxed{E_1 = \frac{\hbar^2\pi^2}{2ma^2} \quad E_3 = \frac{3^2\hbar^2\pi^2}{2ma^2} \quad E_5 = \frac{5^2\hbar^2\pi^2}{2ma^2}}$$

Moreover, the probabilities of measuring such energies are given by the integrals calculated in part (a). In other words, the probabilities P_i of measuring energy E_i are

$$\boxed{P_1 = \frac{6}{10} \quad P_3 = \frac{3}{10} \quad P_5 = \frac{1}{10}}$$

The average energy could be calculated by evaluating $\langle\psi|\hat{H}|\psi\rangle$, or by calculating

$$\langle E \rangle = E_1P_1 + E_3P_3 + E_5P_5$$

$$\boxed{\langle E \rangle = \frac{29\hbar^2\pi^2}{10ma^2}}$$

□

- c) Find the expression of the wave function at a later time t . (Hint: What is $\psi_n(x, t)$?)

Answer. Taking the hint, we know that

$$\psi_n(x, t) = \psi_n(x)e^{-iE_nt/\hbar} = \sqrt{\frac{2}{a}}\sin\left(\frac{n\pi x}{a}\right)e^{-iE_nt/\hbar}$$

We can rewrite $\psi(x, 0)$ in a form relatable to the above as follows.

$$\begin{aligned} \psi(x, 0) &= \sqrt{\frac{6}{5a}}\sin\left(\frac{\pi x}{a}\right) + \sqrt{\frac{3}{5a}}\sin\left(\frac{3\pi x}{a}\right) + \frac{1}{\sqrt{5a}}\sin\left(\frac{5\pi x}{a}\right) \\ &= \sqrt{\frac{3}{5}}\sqrt{\frac{2}{a}}\sin\left(\frac{\pi x}{a}\right) + \sqrt{\frac{3}{10}}\sqrt{\frac{2}{a}}\sin\left(\frac{3\pi x}{a}\right) + \frac{1}{\sqrt{10}}\sqrt{\frac{2}{a}}\sin\left(\frac{5\pi x}{a}\right) \\ &= \sqrt{\frac{3}{5}}\psi_1(x, 0) + \sqrt{\frac{3}{10}}\psi_3(x, 0) + \frac{1}{\sqrt{10}}\psi_5(x, 0) \end{aligned}$$

Therefore,

$$\begin{aligned} \psi(x, t) &= \sqrt{\frac{3}{5}}\psi_1(x, t) + \sqrt{\frac{3}{10}}\psi_3(x, t) + \frac{1}{\sqrt{10}}\psi_5(x, t) \\ &= \sqrt{\frac{3}{5}}\sqrt{\frac{2}{a}}\sin\left(\frac{\pi x}{a}\right)e^{-iE_1t/\hbar} + \sqrt{\frac{3}{10}}\sqrt{\frac{2}{a}}\sin\left(\frac{3\pi x}{a}\right)e^{-iE_3t/\hbar} \\ &\quad + \frac{1}{\sqrt{10}}\sqrt{\frac{2}{a}}\sin\left(\frac{5\pi x}{a}\right)e^{-iE_5t/\hbar} \end{aligned}$$

$$\boxed{\psi(x, t) = \sqrt{\frac{6}{5a}}\sin\left(\frac{\pi x}{a}\right)e^{-iE_1t/\hbar} + \sqrt{\frac{3}{5a}}\sin\left(\frac{3\pi x}{a}\right)e^{-iE_3t/\hbar} + \frac{1}{\sqrt{5a}}\sin\left(\frac{5\pi x}{a}\right)e^{-iE_5t/\hbar}}$$

□

- d) Is the mean value of the position operator independent of time? What about the mean value of the momentum? (Hint: Use symmetry properties with respect to the central point of the well.)

Answer. To determine whether or not the position operator is independent of time, it will suffice to evaluate

$$\frac{d}{dt} \left(\langle \psi(x, t) | \hat{x} | \psi(x, t) \rangle \right)$$

If the above expression is equal to zero, then the position operator is independent of time, and if it is not equal to zero, then the position operator is not independent of time. Let's begin.

We have that

$$\begin{aligned} & \langle \psi(x, t) | \hat{x} | \psi(x, t) \rangle \\ &= \left\langle \frac{3}{5} \psi_1 e^{-iE_1 t/\hbar} + \frac{3}{10} \psi_3 e^{-iE_3 t/\hbar} + \frac{1}{10} \psi_5 e^{-iE_5 t/\hbar} \left| \hat{x} \right| \frac{3}{5} \psi_1 e^{-iE_1 t/\hbar} + \frac{3}{10} \psi_3 e^{-iE_3 t/\hbar} + \frac{1}{10} \psi_5 e^{-iE_5 t/\hbar} \right\rangle \\ &= \frac{9}{25} e^{-2iE_1 t/\hbar} \langle \psi_1 | \hat{x} | \psi_1 \rangle + \frac{9}{100} e^{-2iE_3 t/\hbar} \langle \psi_3 | \hat{x} | \psi_3 \rangle + \frac{1}{100} e^{-2iE_5 t/\hbar} \langle \psi_5 | \hat{x} | \psi_5 \rangle \\ &\quad + \frac{9}{50} e^{-i(E_1 + E_3)t/\hbar} \langle \psi_1 | \hat{x} | \psi_3 \rangle + \frac{3}{50} e^{-i(E_1 + E_5)t/\hbar} \langle \psi_1 | \hat{x} | \psi_5 \rangle + \frac{3}{100} e^{-i(E_3 + E_5)t/\hbar} \langle \psi_3 | \hat{x} | \psi_5 \rangle \end{aligned}$$

Now, computing integrals, we have the following.

$$\langle \psi_1 | \hat{x} | \psi_3 \rangle = \langle \psi_1 | \hat{x} | \psi_5 \rangle = \langle \psi_3 | \hat{x} | \psi_5 \rangle = 0$$

One easy way to see this without direct computation is to observe, per the hint, that $(x - 0.5a)\psi_1\psi_3$ and $0.5a\psi_1\psi_3$ are both odd about the central point of the well and hence evaluate to zero. Thus,

$$\langle \psi_1 | \hat{x} | \psi_3 \rangle = \int_0^a x \psi_1 \psi_3 dx = \int_0^a (x - 0.5a) \psi_1 \psi_3 dx + \int_0^a 0.5a \psi_1 \psi_3 dx = 0 + 0 = 0$$

On the other hand, we may observe that $x\psi_i^2$ ($i = 1, 3, 5$) are all strictly positive on $(0, a)$ and hence evaluate to positive constants c_i . Consequently,

$$\langle \psi(x, t) | \hat{x} | \psi(x, t) \rangle = \frac{9}{25} c_1 e^{-2iE_1 t/\hbar} + \frac{9}{100} c_3 e^{-2iE_3 t/\hbar} + \frac{1}{100} c_5 e^{-2iE_5 t/\hbar}$$

This function clearly has a nonzero derivative with respect to time, meaning that the mean value of the position operator is not independent of time.

As to the second part of the question, we have that

$$\frac{d\langle \hat{p} \rangle}{dt} = \frac{d}{dt} \left(m \frac{d\langle \hat{x} \rangle}{dt} \right) \neq 0$$

Therefore, the mean value of the momentum operator is not independent of time, either. □

- e) Would the result of part (d) be different if we replaced ψ_3 by ψ_2 in Eq. 2.5?

Answer. If we replaced ψ_3 by ψ_2 , then we would additionally have

$$\langle \psi_1 | \hat{x} | \psi_2 \rangle \neq 0 \qquad \langle \psi_2 | \hat{x} | \psi_5 \rangle \neq 0$$

However, this would change neither result overall. □

2. a) Consider now the wave function $\Psi(x, t)$ of a particle moving in one dimension in a potential $V(x)$ such that

$$\begin{aligned} V(x) &\rightarrow \infty & \text{for } |x| \geq a/2 \\ V(x) &= 0 & \text{for } -a/2 < x < 0 \\ V(x) &= V_0 & \text{for } 0 \leq x < a/2 \end{aligned} \quad (2.6)$$

Considering that the wave function and its derivative are continuous at $x = 0$, and that the wave function vanishes at $x = \pm a/2$, try to find the equation that gives the possible energy states assuming $E_n > V_0$.

Hint: There are different combinations of sine and cosine functions for positive and negative values of x .

Answer. Taking the hint, split the total wave function $\psi(x)$ into the sum of two parts, $\psi_1(x)$ and $\psi_2(x)$, where $\psi_1(x) = 0$ for $x \geq 0$ and $\psi_2(x) = 0$ for $x \leq 0$. In general, we have

$$\psi_1(x) = A \sin(kx) + B \cos(kx) \quad \psi_2(x) = C \sin(k_2x) + D \cos(k_2x)$$

If $\psi = \psi_1 + \psi_2$ is to be continuous at $x = 0$, then we must have

$$\begin{aligned} \psi_1(0) &= \psi_2(0) \\ B &= D \end{aligned}$$

If $\psi = \psi_1 + \psi_2$ is to have a continuous first derivative at $x = 0$, then we must have

$$\begin{aligned} \psi'_1(0) &= \psi'_2(0) \\ kA &= k_2C \end{aligned}$$

Thus, altogether, we have that

$$\psi(x) = \begin{cases} A \sin(kx) + B \cos(kx) & x \leq 0 \\ \frac{kA}{k_2} \sin(k_2x) + B \cos(k_2x) & x > 0 \end{cases} \quad \text{for } |x| \leq a/2$$

The boundary condition is met when either...

- i. $B = 0$, $k = n\pi$, and $k_2 = m\pi/a$;
- ii. $A = 0$, $k = n\pi/2$, and $k_2 = m\pi/2a$.

□