

## Week 7

# Time-Independent Problems in 3D

## 7.1 Three-Dimensional Harmonic Oscillator

2/12: • Last time.

- We discussed some of the problems we face in 3D.
- The Hamiltonian is now

$$\hat{H} = -\frac{\hbar^2}{2m} \left[ \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right] + V(x, y, z)$$

- Derivatives in three coordinates.
- The potential is time-independent.
- If the potential does not depend on anything more specific (e.g., is not central, for instance), then only  $\hat{H}$  is conserved.
- We solve

$$\hat{H}\psi(x, y, z) = E\psi(x, y, z)$$

for  $\psi, E$ .

- There are three compatible operators:

$$\hat{H}, \hat{L}^2, \hat{L}_z$$

- The  $z$ -angular momentum operator, in particular, has the form

$$\hat{L}_z = -i\hbar \frac{\partial}{\partial \phi}$$

which is analogous to the form  $\hat{p}_z = -i\hbar(\partial/\partial z)$ .

- The potential is central, i.e.,

$$V(x, y, z) = V(r) = V(\sqrt{x^2 + y^2 + z^2})$$

- If the potential is depends on  $r$ , we solve the ODE in polar coordinates  $(r, \theta, \phi)$ .
- There are also many cases when we only have

$$V(x, y, z) = V(\sqrt{x^2 + y^2})$$

- In this case,  $\hat{H}, \hat{L}_z, \hat{p}_z$  will all be compatible.

- If the potential depends via

$$V(x, y, z) = V(\sqrt{x^2 + y^2}, z)$$

then we will conserve  $\hat{H}, \hat{L}_z$ .

- We will play with this in the problem set.

- Today, we begin with the **asymmetric harmonic oscillator**.
- **Asymmetric harmonic oscillator:** A particle subject to the following three-dimensional potential.  
*Constraint*

$$V(x, y, z) = \frac{M\omega_1^2 x^2}{2} + \frac{M\omega_2^2 y^2}{2} + \frac{M\omega_3^2 z^2}{2}$$

- This potential is special in the sense that it allows us to solve by separation of variables.
- In other words, since we can write the ODE in the form

$$\left[ -\frac{\hbar^2}{2M} \frac{\partial^2 \psi}{\partial x^2} + \frac{M\omega_1^2 x^2}{2} \psi \right] + \left[ -\frac{\hbar^2}{2M} \frac{\partial^2 \psi}{\partial y^2} + \frac{M\omega_2^2 y^2}{2} \psi \right] + \left[ -\frac{\hbar^2}{2M} \frac{\partial^2 \psi}{\partial z^2} + \frac{M\omega_3^2 z^2}{2} \psi \right] = E\psi$$

we may write

$$\psi(x, y, z) = X(x)Y(y)Z(z)$$

- This allows us to algebraically manipulate the ODE into the form

$$\frac{1}{X} \left[ -\frac{\hbar^2}{2M} \frac{d^2 X}{dx^2} + \frac{M\omega_1^2 x^2}{2} X \right] + \frac{1}{Y} \left[ -\frac{\hbar^2}{2M} \frac{d^2 Y}{dy^2} + \frac{M\omega_2^2 y^2}{2} Y \right] + \frac{1}{Z} \left[ -\frac{\hbar^2}{2M} \frac{d^2 Z}{dz^2} + \frac{M\omega_3^2 z^2}{2} Z \right] = E$$

- We switch from partial to total derivatives here because now each function is only a function of one variable (e.g.,  $X(x)$  depends only on  $x$ )!
- Since the sum of these three independent terms is equal to a constant, each term must equal a constant!
- Splitting the above equation into three, we obtain

$$\begin{aligned} -\frac{\hbar^2}{2M} \frac{d^2 X}{dx^2} + \frac{M\omega_1^2 x^2}{2} X &= E_{n_1} X \\ -\frac{\hbar^2}{2M} \frac{d^2 Y}{dy^2} + \frac{M\omega_2^2 y^2}{2} Y &= E_{n_2} Y \\ -\frac{\hbar^2}{2M} \frac{d^2 Z}{dz^2} + \frac{M\omega_3^2 z^2}{2} Z &= E_{n_3} Z \end{aligned}$$

- It follows that

$$E = E_{n_1} + E_{n_2} + E_{n_3}$$

- We already know the solution to each of these three ODEs! They are just quantum harmonic oscillators. Thus,

$$E_{n_i} = \hbar\omega_i \left( n_i + \frac{1}{2} \right)$$

and

$$E = E_{n_1 n_2 n_3} = \hbar\omega_1 \left( n_1 + \frac{1}{2} \right) + \hbar\omega_2 \left( n_2 + \frac{1}{2} \right) + \hbar\omega_3 \left( n_3 + \frac{1}{2} \right)$$

- Additionally, it follows that the wave functions of each direction are of the form (for example)

$$X_{n_1}(x) = \left( \frac{M\omega_1}{\hbar\pi} \right)^{1/4} \frac{H_{n_1}(\xi_1)}{\sqrt{2^{n_1} n_1!}} \exp \left[ -\frac{\xi_1^2}{2} \right]$$

where  $\xi_1 = x\sqrt{M\omega_1/\hbar}$ .

- What happens to  $X_{n_1}(x), Y_{n_2}(y)$  in the limiting case that  $n_1 \rightarrow n_2$ ,  $x \rightarrow y$ , and  $\omega_1 \rightarrow \omega_2$ ?
  - We start approaching something interesting.
  - We need to go a bit further, though.

- Now consider the limiting case where

$$\omega_1 = \omega_2 = \omega_3 = \omega$$

- Herein, the Hamiltonian becomes

$$\begin{aligned}\hat{H} &= -\frac{\hbar^2}{2M}\vec{\nabla}^2 + \frac{M\omega^2}{2}(x^2 + y^2 + z^2) \\ &= -\frac{\hbar^2}{2M}\vec{\nabla}^2 + \frac{M\omega^2 r^2}{2}\end{aligned}$$

- In this *central potential*, recall that we have

$$\hat{L}^2 Y_{\ell m}(\theta, \phi) = \hbar^2 \ell(\ell + 1) Y_{\ell m}(\theta, \phi)$$

and

$$\hat{L}_z Y_{\ell m}(\theta, \phi) = \hbar m Y_{\ell m}(\theta, \phi)$$

and

$$-\frac{\hbar^2}{2M} \frac{d^2}{dr^2} [U_{n\ell}(r)] + \underbrace{\left[ V(r) + \frac{\hbar^2 \ell(\ell + 1)}{2Mr^2} \right]}_{V_{\text{eff}}(r)} U_{n\ell}(r) = E_{n\ell} U_{n\ell}(r)$$

- This leads directly into our discussion of the **spherically symmetric harmonic oscillator**.
- **Spherically symmetric harmonic oscillator:** A particle subject to the following one-dimensional effective potential. *Constraint*

$$V_{\text{eff}}(r) = \frac{M\omega^2 r^2}{2} + \frac{\hbar^2 \ell(\ell + 1)}{2Mr^2}$$

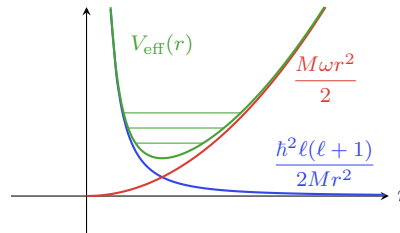


Figure 7.1: Spherically symmetric harmonic oscillator potential.

- The problem we have to solve here is

$$-\frac{\hbar^2}{2M} \frac{d^2}{dr^2} [U_{n\ell}(r)] + \left[ \frac{M\omega r^2}{2} + \frac{\hbar^2 \ell(\ell + 1)}{2Mr^2} \right] U_{n\ell}(r) = E_{n\ell} U_{n\ell}(r)$$

- Recall that

$$\psi(r, \theta, \phi) = R_{n\ell}(r) Y_{\ell m}(\theta, \phi) \quad R_{n\ell}(r) r = U_{n\ell}(r)$$

- In the effective potential, we have the interplay of two peaking potentials as in Figure 7.1.

■ The particle will have certain energy states within the well.

- In the limiting case that  $r$  is small ( $r \rightarrow 0$ ), we can approximate the potential as giving us

$$-\frac{\hbar^2}{2M} \frac{d^2 U_{n\ell}}{dr^2} + \frac{\hbar^2 \ell(\ell + 1)}{2Mr^2} U_{n\ell} + \dots = 0$$

- In this case, the solution is proportional to

$$U_{n\ell} \propto Cr^{\ell+1}$$

- This is because

$$\begin{aligned}\frac{d}{dr}(Cr^{\ell+1}) &= (\ell+1)Cr^{\ell} \\ \frac{d^2}{dr^2}(Cr^{\ell+1}) &= \ell(\ell+1)C\frac{r^{\ell+1}}{r^2}\end{aligned}$$

- In the limiting case that  $r$  is large ( $r \rightarrow \infty$ ), we can approximate the potential as giving us

$$-\frac{\hbar^2}{2M}\frac{d^2U_{n\ell}}{dr^2} + \frac{M\omega^2r^2}{2}U_{n\ell} + \dots = 0$$

- In this case, the solution is proportional to

$$U_{n\ell} = Ce^{-M\omega r^2/2\hbar}$$

- Thus, we combine the two partial solutions to propose the overall ansatz

$$U_{n\ell} = f_{n\ell}r^{\ell+1}e^{-M\omega r^2/2\hbar}$$

- Substituting back into the original ODE, we obtain the differential equation

$$f_{n\ell}'' + 2\left(\frac{\ell+1}{r} - \frac{M\omega r}{\hbar}\right)f_{n\ell}' + \left[\frac{2ME_{n\ell}}{\hbar^2} - \frac{(2\ell+3)M\omega}{\hbar}\right]f_{n\ell} = 0$$

- As we have previously, propose that

$$f_{n\ell}(r) = \sum_j a_j r^j$$

- But there's a problem:  $f_{n\ell}'(r=0) = a_1$ , and this would allow the  $(\ell+1)/r$  term to diverge and make the differential equation blow up.
- Thus, we choose  $a_1 = 0$  and proceed.

- Substituting this power series into the differential equation, we obtain

$$\sum_j j(j-1)a_j r^{j-2} + 2\left(\frac{\ell+1}{r} - \frac{M\omega r}{\hbar}\right)\sum_j ja_j r^{j-1} + \left[\frac{2ME_{n\ell}}{\hbar^2} - \frac{(2\ell+3)M\omega}{\hbar}\right]\sum_j a_j r^j = 0$$

- Make a change of variables  $j \rightarrow j+2$  so that we can start the sum from zero.

$$\begin{aligned}\sum_{j=0}^{\infty} (j+2)(j+1)a_{j+2}r^j + 2\left(\frac{\ell+1}{r} - \frac{M\omega r}{\hbar}\right)\sum_{j=0}^{\infty} (j+2)a_{j+2}r^{j+1} \\ + \left[\frac{2ME_{n\ell}}{\hbar^2} - \frac{(2\ell+3)M\omega}{\hbar}\right]\sum_{j=0}^{\infty} a_j r^j = 0\end{aligned}$$

- We will finish this derivation on Wednesday.

## 7.2 Office Hours (Yunjia)

2/13: • PSet 2, Q2c.

- If we can get up to Equation 12 in the answer key, that's full credit.
- The thing with  $\kappa_{II}^{-1}$  is the idea that if we have a value that's very large (like  $\kappa_{II}$  will be as  $V_0 \rightarrow \infty$  since  $\kappa_{II} \propto V_0^{1/2}$ ), then we can Taylor expand in its reciprocal.
  - We cannot Taylor expand in the large values; we can only Taylor expand in small values.
  - This technique is called **perturbation theory** and will be a major topic of QMechII; Yunjia's use of it here was admittedly a bit extra.

• A brief introduction to perturbation theory.

- Suppose we seek to solve an equation

$$f(x, \epsilon) = 0$$

where  $\epsilon$  is small.

- We can approximate the solution in the form

$$f^{(0)}(x) + f^{(1)}(x)\epsilon + f^{(2)}(x)\epsilon^2 = 0$$

where the digit superscripts in parentheses just denote different functions, not derivatives or anything like that. For example, we could equally well have used the notation  $f, g, h$ ; it's just that this is less general.

- To solve the original equation, we first solve

$$f^{(0)}(x_0) = 0$$

for  $x_0$ .

- Then we solve

$$f^{(0)}(x_0 + \epsilon x_1) + \epsilon f^{(1)}(x_0) = 0$$

for  $x_1$ .

- Continuing in this fashion, our solution takes on the following form and is progressively refined as more terms are calculated.

$$x = x_0 + \epsilon x_1 + \epsilon^2 x_2 + \dots$$

## 7.3 Spherically Symmetric Harmonic Oscillator

2/14: • Review.

- Recall that the 3D case we're considering corresponds to the Hamiltonian

$$\hat{H} = -\frac{\hbar^2}{2M} \left( \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right) + \frac{M\omega^2}{2}(x^2 + y^2 + z^2)$$

- For this Hamiltonian, we are trying to solve the Eigenvalue equation

$$\hat{H}\psi(x, y, z) = E\psi(x, y, z)$$

- The solution may be obtained in Cartesian coordinates as a limiting case of the asymmetric harmonic oscillator, i.e., via the separation of variables

$$\psi(x, y, z) = X(x)Y(y)Z(z)$$

- This results in the solutions

$$\psi(x, y, z) = \prod_{i=1}^3 H_{n_i}(\xi_i) e^{-\xi_i^2/2} c_{n_i} \quad E_{n_1 n_2 n_3} = \hbar\omega \left( n_1 + n_2 + n_3 + \frac{3}{2} \right)$$

where  $\xi_i = x_i \sqrt{m\omega/\hbar}$  and  $x_1, x_2, x_3 = x, y, z$ , respectively.

- Recall also the polar coordinates  $r, \theta, \phi$ . The solution may be obtained here as well.

- In polar coordinates, we can see that the potential described above is central.
- Thus, we have that

$$\hat{L}^2 Y_{\ell m}(\theta, \phi) = \hbar^2 \ell(\ell+1) Y_{\ell m}(\theta, \phi)$$

and

$$\hat{L}_z Y_{\ell m}(\theta, \phi) = \hbar m Y_{\ell m}(\theta, \phi)$$

and

$$-\frac{\hbar^2}{2M} \frac{d^2}{dr^2} [U_{n\ell}(r)] + \underbrace{\left[ V(r) + \frac{\hbar^2 \ell(\ell+1)}{2Mr^2} \right]}_{V_{\text{eff}}(r)} U_{n\ell}(r) = E_{n\ell} U_{n\ell}(r)$$

where  $-\ell \leq m \leq \ell$  and thus there is a  $2\ell+1$  degeneracy of  $E_{n\ell}$  associated with different  $m$ .

➤ Recall that

$$R_{n\ell}(r) = \frac{U_{n\ell}(r)}{r}$$

- Substituting in

$$V(r) = \frac{M\omega^2 r^2}{2}$$

we obtain the effective potential described in Figure 7.1.

- Limiting cases then lead us to construct the ansatz

$$U_{n\ell} = f_{n\ell} r^{\ell+1} e^{-M\omega r^2/2\hbar}$$

- Now propose that

$$f_{n\ell}(r) = \sum_j a_j r^j$$

- Recall that we may obtain the differential equation

$$f_{n\ell}'' + 2 \left( \frac{\ell+1}{r} - \frac{M\omega r}{\hbar} \right) f_{n\ell}' + \left[ \frac{2ME_{n\ell}}{\hbar^2} - \frac{(2\ell+3)M\omega}{\hbar} \right] f_{n\ell} = 0$$

- We must set  $a_1 = 0$ .
- Moving on, we obtain

$$\sum_j j(j-1) a_j r^{j-2} + 2 \left( \frac{\ell+1}{r} - \frac{M\omega r}{\hbar} \right) \sum_j j a_j r^{j-1} + \left[ \frac{2ME_{n\ell}}{\hbar^2} - \frac{(2\ell+3)M\omega}{\hbar} \right] \sum_j a_j r^j = 0$$

- We now begin on new content, continuing the same derivation from above.
- We can further simplify the above equation by solving for  $a_{j+2}$  in terms of  $a_j$ .
- Begin by bringing all  $r$ 's into the summations and running all sums from 0 to  $\infty$  with no terms that go to zero so that every term is in  $r^j$ .

$$\begin{aligned} \sum_{j=0}^{\infty} (j+2)(j+1) a_{j+2} r^j + 2(\ell+1) \sum_{j=0}^{\infty} (j+2) a_{j+2} r^j - \frac{2M\omega}{\hbar} \sum_{j=0}^{\infty} j a_j r^j \\ + \left[ \frac{2ME_{n\ell}}{\hbar^2} - \frac{(2\ell+3)M\omega}{\hbar} \right] \sum_{j=0}^{\infty} a_j r^j = 0 \end{aligned}$$

- Combine the summations.

$$\sum_{j=0}^{\infty} \left[ (j+1)(j+2)a_{j+2} + 2(j+2)(\ell+1)a_{j+2} - \frac{2jM\omega}{\hbar}a_j + \frac{2ME_{n\ell}}{\hbar^2}a_j - \frac{(2\ell+3)M\omega}{\hbar}a_j \right] r^j = 0$$

- Simplify and combine terms.

$$\sum_{j=0}^{\infty} \left[ (j+2)(j+2\ell+3)a_{j+2} + \left( \frac{2ME_{n\ell}}{\hbar^2} - \frac{M\omega}{\hbar}(2j+2\ell+3) \right) a_j \right] = 0$$

- Because each term in the above summation is affixed to a different power of  $r$ , meaning that no two terms can cancel, not only is the entire sum above equal to zero, but each individual term in it is equal to zero, too.

- Thus, for all  $j \in \mathbb{Z}_{\geq 0}$ ,

$$0 = (j+2)(j+2\ell+3)a_{j+2} + \left( \frac{M\omega}{\hbar}(2j+2\ell+3) - \frac{2ME_{n\ell}}{\hbar^2} \right) a_j$$

$$a_{j+2} = \frac{\frac{2ME_{n\ell}}{\hbar^2} - \frac{M\omega}{\hbar}(2j+2\ell+3)}{(j+2)(j+2\ell+3)} a_j$$

- This combined with the fact that  $a_1 = 0$  means that *all odd  $a_j$  equal zero*.
  - It follows that  $f_{n\ell}$  can be viewed as a function of  $r^2$ , not just  $r$ , since this fact means that the power series will be of the form

$$f_{n\ell}(r) = a_0 + a_2r^2 + a_4r^4 + a_6r^6 + \cdots + a_{2n}r^{2n} + \cdots$$

- Now observe that in the limit of large  $j$  (i.e., as  $j \rightarrow \infty$ ),

$$a_{j+2} \approx \frac{\frac{M\omega}{\hbar}(2j)}{j^2 + 2j}$$

and thus<sup>[1]</sup>

$$f_{n\ell}(r) \approx e^{M\omega r^2/\hbar}$$

- This, in turn, would lead to an exponential growth of  $U_{n\ell}$  as  $r \rightarrow \infty$  and hence a non-renormalizable solution.
- Consequently, there must be some maximum value of  $j$  which we will denote by  $n := j_{\max}$ .
- In particular,  $n$  will be the value of  $j$  such that the numerator of the expression above giving  $a_{j+2}(a_j)$  equals zero. This will guarantee that  $a_{n+2} = 0$  and hence all  $a_j = 0$  for  $j > n$ .
- Solving for this  $n$ , we have that

$$\frac{2ME_{n\ell}}{\hbar^2} = \frac{M\omega}{\hbar}(2n+2\ell+3)$$

$$E_{n\ell} = \hbar\omega \left( n + \ell + \frac{3}{2} \right)$$

- Recall that  $n$  is even;  $n \geq 0$ ;  $\ell \geq 0$ ; and for each  $\ell$ , we have  $2\ell+1$  solutions with  $-\ell \leq m \leq \ell$  where  $\hbar m$  are the eigenvalues of  $\hat{L}_z$ .

- Notice the remarkable similarity between the energy equations for the spherically symmetric harmonic oscillator in Cartesian coordinates (left below) and polar coordinates (right below).

$$E_{n_1 n_2 n_3} = \hbar\omega \left( \bar{n} + \frac{3}{2} \right) \qquad E_{\bar{n}} = \hbar\omega \left( \bar{n} + \frac{3}{2} \right)$$

<sup>1</sup>How did we get this transformation to exponential growth??

- On the left above,  $\bar{n} = n_1 + n_2 + n_3$ . On the right above,  $\bar{n} = n + \ell$ .
- Now let's investigate some particular solutions in both cases.
- $\bar{n} = 0$ .
  - Cartesian: The only possible values are  $n_1 = n_2 = n_3 = 0$ , corresponding to
 
$$e^{-M\omega(x^2+y^2+z^2)/2\hbar}$$
  - Polar: The only possible values are  $n = \ell = 0$ , corresponding to
 
$$e^{-M\omega r^2/2\hbar}$$
  - In both cases, there is only one solution, and the solutions are mathematically equivalent.
- $\bar{n} = 1$ .
  - Cartesian: We could have  $n_1 = 1, n_2 = n_3 = 0$ ;  $n_2 = 1, n_1 = n_3 = 0$ ; or  $n_3 = 1, n_2 = n_1 = 0$ ; corresponding to
 
$$xe^{-M\omega r^2/2\hbar} \qquad ye^{-M\omega r^2/2\hbar} \qquad ze^{-M\omega r^2/2\hbar}$$
  - Polar: We have  $n = 0$ ;  $\ell = 1$ ; and  $m = 1, m = 0$ , or  $m = -1$ ; corresponding to
 
$$re^{-M\omega r^2/2\hbar} \underbrace{\sin \theta e^{i\phi}}_{(x+iy)/r} \qquad re^{-M\omega r^2/2\hbar} \cos \theta \qquad re^{-M\omega r^2/2\hbar} \underbrace{\sin \theta e^{-i\phi}}_{(x-iy)/r}$$
  - In both cases, there are three solutions, and the solutions are mathematically equivalent (up to linear combinations).
- A pattern is emerging: Naturally, it makes sense that the coordinate system chosen should not affect the solutions.
- $\bar{n} = 2$ .

$n_1$	$n_2$	$n_3$
0	0	2
0	2	0
2	0	0
0	1	1
1	0	1
1	1	0

(a) Cartesian coordinates.

$n$	$\ell$	$m$
0	2	2
0	2	1
0	2	0
0	2	-1
0	2	-2
2	0	0

(b) Spherical coordinates.

Table 7.1: Spherically symmetric harmonic oscillator solutions ( $\bar{n} = 2$ ).

- In both cases, there are six solutions.
- Note that we do not consider the case where  $n = \ell = 1$  in Table 7.1b because this would mean that  $j_{\max} = n = 1$  is an odd number, which is not allowed.



## 7.4 Hydrogen Atom: Energy Eigenvalues and Eigenstates

2/16: • Today: The hydrogen atom.

- The central potential is

$$V(r) = -\frac{e^2}{4\pi\epsilon_0 r}$$

- The problem is an electron revolving around a proton.
- The proton and electron have very different masses.

$$M_p c^2 \approx 1 \text{ GeV}$$

$$m_e c^2 \approx 511 \text{ keV}$$

- The ratio is

$$\frac{M_p}{m_e} \approx 2000$$

- This justifies assuming that the proton is fixed (the Born-Oppenheimer approximation).

- The relevant Schrödinger equation is

$$-\frac{\hbar^2}{2m_e} \vec{\nabla}^2 \psi_{n\ell m}(r, \theta, \phi) - \frac{e^2}{4\pi\epsilon_0 r} \psi_{n\ell m}(r, \theta, \phi) = E_{n\ell} \psi_{n\ell m}(\theta, \phi)$$

- Note that  $E$  does not depend on  $m$  because  $m$  corresponds to the  $2\ell + 1$  degeneracy in energy.
  - Moreover,  $m$  only specifies orientation in space, which should intuitively not affect energy because space is isotropic and affine.
  - This is something we should absolutely know!!
- Recall that

$$\psi_{n\ell m}(r, \theta, \phi) = R_{n\ell}(r) Y_{\ell m}(\theta, \phi)$$

$$\hat{L}^2 Y_{\ell m}(\theta, \phi) = \hbar^2 \ell(\ell + 1) Y_{\ell m}(\theta, \phi)$$

$$\hat{L}_z Y_{\ell m}(\theta, \phi) = \hbar m Y_{\ell m}(\theta, \phi)$$

- Recall also polar coordinates

$$z = r \cos \theta$$

$$x = r \sin \theta \cos \phi$$

$$y = r \sin \theta \sin \phi$$

- Making the substitution

$$U_{n\ell}(r) = r R_{n\ell}(r)$$

can simplify the Schrödinger equation to the following equivalent effective potential and 1D problem.

$$-\frac{\hbar^2}{2m_e} \frac{d^2}{dr^2} [U_{n\ell}(r)] + \left[ \frac{\hbar^2 \ell(\ell + 1)}{2m_e r^2} - \frac{e^2}{4\pi\epsilon_0 r} \right] U_{n\ell}(r) = E_{n\ell} U_{n\ell}(r)$$

- This is the problem that started the whole game of quantum mechanics; it has enormous consequences in particle physics.
- As with the discussion associated with Figure 7.1, we have two competing potentials here (see Figure 7.2).
  - We are interested in finding the **bound states**.

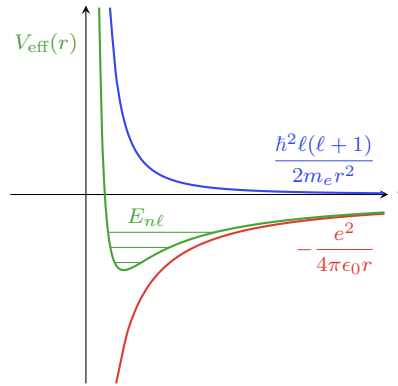


Figure 7.2: Hydrogen atom potential.

- In the limiting case that  $r$  is small ( $r \rightarrow 0$ ), we can approximate the potential as with Figure 7.1 and take

$$U_{n\ell} \propto C r^{\ell+1}$$

- In the limiting case that  $r$  is large ( $r \rightarrow \infty$ ), we can approximate the potential as going to zero and take

$$U_{n\ell} \propto A e^{\pm k_{n\ell} r}$$

where

$$|E_{n\ell}| = \frac{\hbar^2 k_{n\ell}^2}{2m_e}$$

- Thus, we combine the two partial solutions to propose the overall ansatz

$$U_{n\ell}(r) = f_{n\ell}(r) r^{\ell+1} e^{-k_{n\ell} r}$$

- Note that we choose the negative exponent so the solution does not blow up at large  $r$ .

- Following the algebra in the notes, we obtain the following ODE determining  $f_{n\ell}$ .

$$f_{n\ell}''(r) + f_{n\ell}'(r) \left[ \frac{2(\ell+1)}{r} - 2k_{n\ell} \right] + f_{n\ell}(r) \left[ -\frac{2k_{n\ell}(\ell+1)}{r} + \frac{2m_e}{\hbar^2} \frac{e^2}{4\pi\epsilon_0 r} \right] = 0$$

- Aside: The prefactor to the rightmost  $1/r$  term above (excepting the 2 coefficient) is typically written as follows.

$$\frac{m_e c}{\hbar} \frac{e^2}{4\pi\epsilon_0 \hbar c}$$

- The right fraction is the **electromagnetic fine structure constant**.

- Additionally, the other factor  $\hbar/m_e c$  decomposes into  $(\hbar/m_e c) \cdot (1/2\pi)$  where we may recall from the first lecture that  $\hbar/m_e c$  is the **Compton wavelength**  $\lambda_c$ .

- The overall quantity is equal to the inverse of the **Bohr radius**.

- Thus, we can simplify the above equation to

$$f_{n\ell}''(r) + f_{n\ell}'(r) \left[ \frac{2(\ell+1)}{r} - 2k_{n\ell} \right] + f_{n\ell}(r) \left[ -\frac{2k_{n\ell}(\ell+1)}{r} + \frac{2}{a_{\text{Br}}} \right] = 0$$

- As per usual, we propose that

$$f_{n\ell}(r) = \sum_j a_j r^j$$

and collapse functions that diverge.

- Substituting this power series into the differential equation, we obtain

$$\begin{aligned}
 0 &= \sum_j a_j r^{j-2} j(j-1) + \sum_j a_j j r^{j-1} \left[ \frac{2(\ell+1)}{r} - 2k_{n\ell} \right] + \sum_j a_j r^j \left[ -\frac{2k_{n\ell}(\ell+1)}{r} + \frac{2}{a_B r} \right] \\
 &= \sum_j a_j r^{j-1} j(j-1) + \sum_j a_j j r^{j-1} [2(\ell+1) - 2k_{n\ell} r] + \sum_j a_j r^j \left[ -2k_{n\ell}(\ell+1) + \frac{2}{a_B} \right] \\
 &= \sum_j a_{j+1} r^j j(j+1) + \sum_j a_{j+1} (j+1) r^j 2(\ell+1) - \sum_j a_j j r^j k_{n\ell} 2 + \sum_j a_j r^j \left[ -2k_{n\ell}(\ell+1) + \frac{2}{a_B} \right]
 \end{aligned}$$

- From line 1 to line 2, we multiplied through this function equal to zero by  $r$ .
- From line 2 to line 3, we reindex some terms on the left from  $j \rightarrow j+1$ .
- It follows just like last class that

$$a_{j+1}(j+1)[j+2(\ell+1)] = a_j \left[ 2k_{n\ell}j + 2k_{n\ell}(\ell+1) - \frac{2}{a_B} \right]$$

- Thus, we get that

$$a_{j+1} = \frac{k_{n\ell}(2j+2\ell+2) - \frac{2}{a_B}}{(j+1)[j+2(\ell+1)]} a_j$$

- Once again, for similar reasons, there will also be some  $j_{\max} = n$ .
- Then

$$(n+\ell+1)k_{n\ell} = \frac{1}{a_B}$$

which means that

$$k_{n\ell} = \frac{1}{a_B(n+\ell+1)}$$

- Then we get that

$$E_{n\ell} = -\frac{\hbar^2}{2m_e a_B^2 (n+\ell+1)^2}$$

where everything except the quantum numbers is the **Rydberg constant**.

- Consequently, in this case, we may define  $\bar{n} = n + \ell + 1$ .

- **Bound state:** A state in which the electron can escape to  $\infty$ .

- We tend to suppress bound states so that the wave function does not have probability at  $\infty$ .

- **Electromagnetic fine structure constant:** The constant defined as follows. Denoted by  $\alpha$ . Given by

$$\alpha = \frac{e^2}{4\pi\epsilon_0 \hbar c} \approx \frac{1}{137}$$

- **Bohr radius:** The most probable distance from the nucleus of a hydrogen atom for its electron to exist. Denoted by  $a_B$ . Given by

$$a_B = \frac{4\pi\epsilon_0 \hbar^2}{m_e e^2} \approx \frac{137}{2\pi} \lambda_c = 5.3 \times 10^{-11} \text{ m}$$

- Note that we get the approximation from the aside's note that

$$a_B^{-1} = \lambda_c^{-1} 2\pi\alpha$$

- **Rydberg constant:** The constant defined as follows. Denoted by **Ry**. Given by

$$\text{Ry} = \frac{\hbar^2}{2m_e a_B^2} = 13.6 \text{ eV}$$

- Wagner is Argentinian.
- We'll continue on Monday.