

Week 10

Finals Week

10.1 Final Exam Review

- 3/4: • Final.
- Thursday, 5:30pm-7:30pm.
 - May release before 5:30, but it should be returned at 7:30 *sharp* so they can grade in the evening.
 - Doable in 2 hours (he believes).
 - Open book.
 - Largely conceptual; minimal calculation required.
 - No tricky questions.
 - Heavily based on problem sets and the midterm.
 - Simply a test of what we've learned in the course.
 - For those who have poor PSet grades, the final will acquire more relevance.
 - He will try to have a rule along the lines of “If you do better in the final than the problem sets, then we'll boost your grade by some amount.”
 - For the multiple choice questions, be sure to read *all* of the answer choices before selecting one because they're looking for the *best* answer even when multiple may be partially correct.
 - Will the final focus more on the topics covered since the midterm?
- There are notes for a review class posted on Canvas.
 - We'll go for 40-45 minutes today.
 - We now begin the review.
 - In this course, we've looked at a particle that satisfies the equation
$$\left[-\frac{\hbar^2}{2M} \vec{\nabla}^2 + V(\vec{r}, t) \right] \psi = i\hbar \frac{\partial \psi}{\partial t}$$
 - Specifically, we've focused on time-independent potentials
$$V(\vec{r}, t) = V(\vec{r})$$
 - For time-independent potentials, we do not have the possibility that the expected value of the energy is zero.

- Additionally, energy is conserved.

$$\frac{d}{dt} \left(\langle \psi | \hat{H} | \psi \rangle \right) = \frac{i}{\hbar} \langle \psi | [\hat{H}, \hat{H}] | \psi \rangle + 0 = 0$$

- For generic Hermitian operators, if we have the following, then we know that the corresponding quantity is conserved.

$$\frac{d}{dt} \left(\langle \psi | \hat{O} | \psi \rangle \right) = \frac{i}{\hbar} \underbrace{\langle \psi | [\hat{H}, \hat{O}] | \psi \rangle}_0 + \underbrace{\langle \psi | \frac{\partial \hat{O}}{\partial t} | \psi \rangle}_0 = 0$$

- There are certain operators to be aware of.
 - The Hamiltonian operator gives energy.

$$\hat{H} |\psi_n\rangle = E_n |\psi_n\rangle$$

- Other operators: Momentum (\hat{p}), position (\hat{r}), and potential ($\hat{V}(\vec{r})$).
- Every operator can be expressed as a function $F(\hat{p}, \hat{r})$ of the momentum and position operators.
 - Thus, you can also determine if an operator is conserved using its decomposition in terms of position and momentum operators:

$$\frac{d}{dt} \left(\langle \psi_n | F(\hat{p}, \hat{r}) | \psi_n \rangle \right) = 0$$

- A generic wave function can be expressed as a linear combination of a basis of eigenfunctions.

$$|\psi\rangle(t) = \sum_n c_n |\psi_n\rangle e^{-iE_n t/\hbar}$$

- Using such a decomposition, we can calculate the expected energy as follows.

$$\langle \psi | \hat{H} | \psi \rangle = \sum_n |c_n|^2 E_n$$

- Note that each $|c_n|^2$ is the probability of finding E_n when you take a measurement of the particle.
- Another important decomposition is that of the expected position.

$$\langle \psi | \hat{r} | \psi \rangle = \sum_{m,n} c_m^* c_n \langle \psi_m | \hat{r} | \psi_n \rangle e^{i(E_m - E_n)t/\hbar}$$

- The quantity $|\psi(\vec{r}, t)|^2$ is the probability density of the particle.
 - The probability that the particle will be *somewhere* in space is certain.

$$\int d^3\vec{r} |\psi|^2 = 1$$

- If we have a potential, the energy of the particle should always be greater than the minimum of the potential.
 - Additionally, quantum particles can penetrate somewhat into regions where the potential is greater than their energy.
 - Quantum particles can also **tunnel** through finitely long regions of high potential.
- In this course, we spent a lot of time trying to find the bound states of energy.
- The only normalizable states are those for which the energy is quantized.

- A very important case into which we looked is the harmonic oscillator.

$$V(x) = \frac{m\omega^2 x^2}{2}$$

- The energy eigenvalues are

$$E_n = \hbar\omega \left(n + \frac{1}{2} \right)$$

- The uncertainty principle is what requires that $1/2$ term, since it implies we must have

$$\sigma_x \sigma_p \geq \frac{\hbar}{2}$$

because

$$\sigma_A \cdot \sigma_B \geq \frac{1}{2} |\langle \psi | [\hat{A}, \hat{B}] | \psi \rangle|$$

and

$$[[\hat{x}, \hat{p}]] = \hbar$$

- Diving deeper into the harmonic oscillator, we defined the raising and lowering operators

$$a_{\pm} = \frac{1}{\sqrt{2\hbar M\omega}} [\mp i\hat{p} + M\omega\hat{x}]$$

- These operators have the properties that

$$a_+ |n\rangle \propto |n+1\rangle \qquad a_- |n\rangle \propto |n-1\rangle$$

- We can also use these to prove that

$$\langle n | \hat{x} | n \rangle = 0 \qquad \langle n | \hat{p} | n \rangle = 0$$

- The eigenfunctions of the harmonic oscillator form an orthonormal basis of the function space.

$$\langle n | m \rangle = \int dx \psi_n^* \psi_m = \delta_{nm}$$

- Energy is shared between the kinetic and potential.

$$\left\langle n \left| \frac{M\omega\hat{x}^2}{2} \right| n \right\rangle = \frac{\hbar\omega}{2} \left(n + \frac{1}{2} \right) \qquad \left\langle n \left| \frac{\hat{p}^2}{2M} \right| n \right\rangle = \frac{\hbar\omega}{2} \left(n + \frac{1}{2} \right)$$

- This is what we needed to know for the midterm.
- After that, we went into three dimensions.
- In particular, we looked at central potentials, which have the form

$$V(\vec{r}) = V(r)$$

- We introduced the angular momentum operators \hat{L}_z, \hat{L}^2 .
- We look at these operators in spherical coordinates (r, θ, ϕ) .
- For the component operators, we have the commutativity relation

$$[\hat{L}_i, \hat{L}_j] = i\hbar\epsilon_{ijk}\hat{L}_k$$

- The z -direction angular momentum operator has the special property that

$$\hat{L}_z = -i\hbar \frac{\partial}{\partial \phi}$$

- The two main operators introduced above satisfy the eigenvalue equations

$$\hat{L}_z Y_{\ell m} = \hbar m Y_{\ell m} \qquad \hat{L}^2 Y_{\ell m} = \hbar^2 \ell(\ell+1) Y_{\ell m}$$

- The solutions are of the form

$$Y_{\ell m}(\theta, \phi) = P_{\ell, m}(\theta) e^{im\phi}$$

where m takes on the $2\ell + 1$ values from $-\ell \leq m \leq \ell$.

- The overall wave function has the angular components from above and also a radial component so that

$$\psi_{n\ell m}(\vec{r}) = R_{n\ell}(r) Y_{\ell m}(\theta, \phi)$$

- If we define

$$U_{n\ell}(r) = r R_{n\ell}(r)$$

then we obtain the effective 1D Schrödinger equation

$$\left[-\frac{\hbar^2}{2M} \frac{d^2}{dr^2} + \underbrace{\frac{\hbar^2 \ell(\ell+1)}{2Mr^2}}_{V_{\text{eff}}(r)} + V(r) \right] U_{n\ell} = E_{n\ell} U_{n\ell}$$

- We then used these techniques to address the harmonic oscillator in three dimensions.

- We found the energies to be

$$E_{n\ell} = \hbar\omega \left(\underbrace{N + \ell}_n + \frac{3}{2} \right) = \hbar\omega \left(\underbrace{n_1 + n_2 + n_3}_n + \frac{3}{2} \right)$$

■ N is the degree of the polynomial.

■ N is even.

- The bound states are a battle between the two terms in the effective potential energy, as summarized in Figure 7.1.

- We also used 3D central potential techniques to tackle the hydrogen atom.

- We found the energies to be

$$E_{n\ell} = -\frac{\text{Ry}}{\underbrace{(N + \ell + 1)}_n^2}$$

- There are n^2 solutions for each each n , which we determine by summing the $2\ell + 1$ degeneracy over $\ell = 0, \dots, n-1$.

- Note that

$$\text{Ry} = 13.6 \text{ eV}$$

is a very important number.

- Brief review of emitting electromagnetic radiation.

- Last thing: Spin.

- Every particle has it; we don't know why.

- The quantity

$$\hat{\vec{S}} = \hat{S}_x \hat{x} + \hat{S}_y \hat{y} + \hat{S}_z \hat{z}$$

is an angular momentum that has nothing to do with angular direction. It has nothing to do with any direction in spacetime; rather, it is an *intrinsic* property of the particle.

- We have the commutation relation

$$[\hat{S}_x, \hat{S}_y] = i\hbar \hat{S}_z$$

- We can introduce ladder operators

$$\hat{S}_{\pm} = \hat{S}_x \pm i\hat{S}_y$$

- As in real angular momentum, we have analogous eigenvalue expressions

$$\begin{aligned}\hat{\vec{S}}^2 |s, m_s\rangle &= \hbar^2 s(s+1) |s, m_s\rangle \\ \hat{S}_z |s, m_s\rangle &= \hbar m_s |s, m_s\rangle\end{aligned}$$

- $2s + 1$ states implies that $s = 0, 1/2, 1, 3/2, 2, \dots$ can take on half-integer values.
 - Almost all elementary particles have spin $1/2$ (protons, neutrons, the quarks that make them up, electrons, leptons).
 - Force carriers (photon, gluon) have spin 1.
 - The graviton (if it exists; “it exists”) has spin 2.
 - Higgs has spin 0.

- In the hydrogen atom, there are $2n^2$ states for each n given by

$$|n, \ell, m, \frac{1}{2}, \pm \frac{1}{2}\rangle$$

- This is the 2 s -orbital positions, 8 $s + p$ -orbital positions, and on and on.

- We can modify the total degeneracy of the spin with a magnetic field. Use the Hamiltonian

$$-\vec{\mu} \cdot \vec{B} = -\gamma \cdot \hat{\vec{S}} \cdot \vec{B}$$

- If $\vec{B} = B\hat{z}$, then the above also equals $-\gamma B \hat{S}_z$.

- In a generic state, we can introduce spinors

$$\chi = \begin{pmatrix} \chi_+ \\ \chi_- \end{pmatrix}$$

- The probability of finding the particle with spin up or spin down in the z -direction must be certain.

$$|\chi_+|^2 + |\chi_-|^2 = 1$$

- $|\chi_+|^2$ is the probability of spin up.
 - $|\chi_-|^2$ is the probability of spin down.

- In the magnetic field case, we have to solve the Schrödinger equation

$$-\gamma B \hat{S}_z \begin{pmatrix} \chi_+ \\ \chi_- \end{pmatrix} = i\hbar \frac{\partial}{\partial t} \begin{pmatrix} \chi_+ \\ \chi_- \end{pmatrix}$$

- The three spin components can be represented by the matrices

$$\hat{S}_x = \frac{\hbar}{2} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \quad \hat{S}_y = \frac{\hbar}{2} \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \quad \hat{S}_z = \frac{\hbar}{2} \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

- The matrices are called the Pauli matrices σ_i .
- These matrices have the properties that

$$\sigma_i^2 = I \qquad \hat{S}_i = \frac{\hbar}{2}\sigma_i \qquad \hat{S}_i^2 = \frac{\hbar^2}{4}I$$

- Now suppose we let

$$|\chi_+\rangle(0) = \cos\left(\frac{\theta_s}{2}\right) \qquad |\chi_-\rangle(0) = \sin\left(\frac{\theta_s}{2}\right)$$

- It then follows that

$$\begin{aligned} \langle \chi | \hat{S}_z | \chi \rangle &= \frac{\hbar}{2} \cos(\theta_s) \\ \langle \chi | \hat{S}_x | \chi \rangle &= \frac{\hbar}{2} \sin(\theta_s) \cos(\gamma B t + \phi_+ - \phi_-) \end{aligned}$$

- Then the probability of spin up or spin down is

$$P_{\pm} = \frac{1}{2} [1 \pm \sin(\theta_s) \cos(\gamma B t + \phi_+ - \phi_-)]$$