

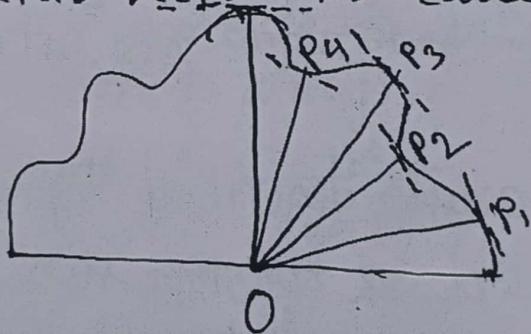
Central Forces

- i) simple harmonic motion
- ii) projectile motion
- iii) uniformly accelerated motion
- iv) others, like electrostatic, magnetostatic forces, etc.

(19)

Central Force: A force which is always directed towards a fixed point is called a central force and the path described by the particle under the action of a central force is called the central orbit. The fixed point is called the centre of force. Examples of central forces i) Uniform circular motion, ii) force due to gravitation

Apse: In a central orbit, the point at which the radius vector is maximum or minimum is called Apse. Otherwise, in a central orbit, the point where the tangent is perpendicular to the radius vector is called Apse.



Here, P_1, P_2, P_3, \dots - - - Apse.

Apsidal distance: The length of the radius vector at an apse is called the apsidal distance. Otherwise, the distance of an apse from the centre of force is called an apsidal distance.

Here $OP_1, OP_2, OP_3, OP_4, \dots$ - - - apsidal distance

Apsidal angle: The angle between two consecutive lines of apsidal distance is called apsidal angle. Otherwise, the angle between the radius vectors of two consecutive apse's.

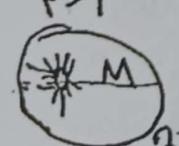
Here, $\angle P_0 P_2$, $\angle P_2 P_3$, $\angle P_3 P_4 \dots$ apsidal angle.

(v) Kepler's laws: The astronomer Kepler, after many years of patient labour, discovered three laws connecting the motions of the various planets about the sun. They are the following:



Focus
Equal areas in equal time
planet
orbital path
area S_1
area S_2

1. Each planet describes an ellipse having the sun in one of its foci. [সূর্য একটি অক্ষে
কার পথের মধ্যে পার্শ্ব দৃষ্টিকোণে
পরিসরিত হয়ে থাকে]



$P = \text{period}$
 $P \propto M^{\frac{3}{2}}$
M=mass
Major axis

2. The areas described by the radii drawn from the planet to the sun are in the same orbit, proportional to the times of describing them. [একই কালাম্বের প্রতি ব্যূহের অবস্থায়ে সময়ে
অন্তরিম ক্ষেত্র পরিসরিত হয়ে থাকে
ক্ষেত্রের একান্ত পরিসরের সময়ের সমান্তরালি]

3. The squares of the periodic times of the various planets are proportional to the cubes of the major axes of their orbits. [পরিসর
সময়ের ক্ষেত্র পরিসরের কাছে একইভাবে প্রতি অন্তরিম
ক্ষেত্রের সময়ের সমান্তরালি]

In a central orbit $v^r = h^r \left[u^r + \left(\frac{du}{d\theta} \right)^2 \right]$.

From differential calculus we know in any curve $\frac{1}{pr} = \frac{1}{r^2} + \frac{1}{r^4} \left(\frac{dr}{d\theta} \right)^2$

$$= u^r + \left(\frac{du}{d\theta} \right)^2 \text{ where } u = \frac{1}{r}$$

Putting $r = \frac{h}{p} \Rightarrow \frac{1}{r} = \frac{1}{\frac{h}{p}} = \frac{p}{h}$

$\Rightarrow \frac{1}{\frac{h}{p}r} = u^r + \left(\frac{du}{d\theta} \right)^2$

$\Rightarrow v^r = h^r \left[u^r + \left(\frac{du}{d\theta} \right)^2 \right]$

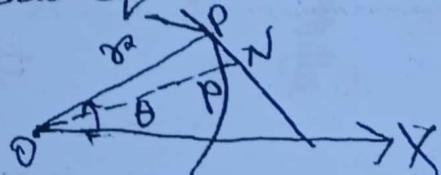
$p = \text{distance}$
 $v = \text{velocity}$
 $h = \text{constant}$

Theorem: A particle moves in a path so that the acceleration is always directed to a fixed point and is equal to $\frac{u}{(\text{distance})^2}$. Show that its path

is a conic. Discuss the different cases.

Solution:

Consider, pole O be the fixed point and at time t the particle stay at a distance r from O and F be the acceleration of the particle. If p be the perpendicular distance from the Origin to the tangent, then we have the pedal equation.



MR-P-5778: 7/10/02:

$$\text{elliptical } u = \frac{1}{r} + \frac{e \cos \theta}{r^2} \left(\frac{du}{d\theta} + u \right)$$
$$f = h^r u^r \left(-\frac{1}{r^2} \frac{du}{d\theta} + u \right)$$

$$f = h^r u^r \left(-\frac{e}{r^2} \cos \theta + \frac{1}{r} + \frac{e}{r^2} \cos \theta \right)$$
$$f = \frac{h^r}{r} \cdot u^r = mu^r =$$

$$\frac{1}{r^2} = \frac{1}{r^2} + \frac{1}{r^4} \left(\frac{dr}{d\theta} \right)^2$$

$$= u^r + u^r \left(-\frac{1}{u^r} \frac{du}{d\theta} \right)^2, u = \frac{1}{r}$$

$$\Rightarrow \left(\frac{du}{d\theta} \right)^2 + u^r = \frac{1}{r^2}$$

Differentiating w.r.t. θ

$$2 \left(\frac{du}{d\theta} \right) \frac{d^2 u}{d\theta^2} + 2u \frac{du}{d\theta} = -\frac{2}{r^3} \frac{dp}{d\theta}$$

$$\Rightarrow 2 \left(\frac{du}{d\theta} \right) \left(\frac{d^2 u}{d\theta^2} + u \right) = -\frac{2}{r^3} \frac{dp}{d\theta}$$

$$\Rightarrow 2 \frac{du}{d\theta} \cdot \frac{F}{mu^2} = -\frac{2}{r^3} \frac{dp}{d\theta} \quad \left[\because \frac{d^2 u}{d\theta^2} + u = \frac{F}{mu^2} \right]$$

$$\Rightarrow \frac{h^r}{r^3} \frac{dp}{d\theta} + \frac{F}{u^2} \frac{du}{d\theta} = 0$$

$$\Rightarrow \frac{h^r}{r^3} \frac{dp}{d\theta} + Fr^r \left(-\frac{1}{r^2} \frac{dr}{d\theta} \right) = 0$$

$$\Rightarrow \frac{h^r}{r^3} \frac{dp}{d\theta} = F \quad \dots \quad (i)$$

$$\text{Here, } F = \frac{u}{r^2}$$

$$\therefore \frac{h^r}{r^3} \frac{dp}{d\theta} = \frac{u}{r^2}$$

$$\Rightarrow h^r \frac{dp}{r^3} = u \cdot \frac{1}{r^2} dr \quad p.R.O$$

Integrating

$$h^r \cdot \left(-\frac{1}{2pr} \right) = u \left(-\frac{1}{r} \right) + C_1; C_1 = \text{constant}$$

$$\Rightarrow \frac{h^r}{pr} = \frac{2u}{r} + C$$

$$\Rightarrow vr = \frac{h^r}{pr} = \frac{2u}{r} + C \quad \left[\because r = \frac{h}{p} \right] \quad (2)$$

But the pedal equations of parabola, ellipse and hyperbola referred to focus \odot are

$$\frac{b^r}{pr} = \frac{2a}{r}, \quad \frac{b^r}{pr} = \frac{2a}{r} - 1 \quad \text{and} \quad \frac{b^r}{pr} = \frac{2a}{r} + 1$$

If $C = 0$, $c < 0$ and $e > 0$, then the equation of (2) represents parabola, ellipse and hyperbola.

Again,

(3)

$$\frac{b^r}{pr} = \frac{2a}{r} - 1 \quad \text{and} \quad \frac{h^r}{pr} = \frac{2u}{r} + C \quad (4)$$

$\underline{(4)}$

Comparing the equations (3) and (4)
we get,

$$\frac{\frac{h^r}{pr}}{\frac{b^r}{pr}} = \frac{\frac{2u}{r}}{\frac{2a}{r}} = \frac{c}{-1}$$

p.t.o.

$$\Rightarrow \frac{h^r}{b^r} = \frac{M}{a} = \frac{c}{-1}$$

From 1st and 2nd part we get,

$$h^r = \frac{b^r M}{a}$$

$$\Rightarrow h = \sqrt{M \times \frac{b^r}{a}}$$

$$\Rightarrow h = \sqrt{M \times \text{semi latus rectum}}$$

and from 2nd and 3rd part we get,

$$c = -\frac{M}{a}$$

Similarly, comparing the equations

$$\frac{b^r}{p_2} = \frac{2a}{r} + 1 \text{ and } \frac{h^2}{p_2} = \frac{2M}{r} + c$$

$$\text{we get, } c = \frac{M}{a}$$

Putting the value of c in (2) we get,

$$V^2 = \frac{2M}{r}, \text{ the equation of Parabola.}$$

$$V^2 = \frac{2M}{r} - \frac{M}{a}, \frac{a}{r} = \frac{2M}{r} + \frac{M}{a} \text{ " ellipse - hyperbola.}$$

#

periodic times: Since h is equal to twice the area described in a unit time, it follows that if T be the ^{whole arc of the} time the particle takes to describe the ellipse, then

$$\frac{1}{2} h \times T = \text{area of the ellipse} = \pi ab$$

$$\therefore T = \frac{2\pi ab}{h}$$

[यद्यपि h एक मात्राये रूप से विद्युत, अर्थात् कोई उपर्युक्त वर्णन नहीं हो सकता, तो]

[उपर्युक्त एवं इसके लिये जब तक तकनीकी तरफ़ आवश्यक नहीं हो, तो]

$$T = \frac{\text{उपर्युक्त विषय}}{\frac{h}{2}}$$

$$\Rightarrow T = \frac{2\pi ab}{h}$$

Q. A particle describe an ellipse under a force $\frac{\mu}{(\text{distance})^2}$ towards the focus. If it was projected with velocity v from the point distance r from the centre of force, show that its periodic time is

$$\frac{2\pi}{\sqrt{\mu}} \left[\frac{2}{r} - \frac{v^2 r^2}{\mu} \right]^{3/2}$$

solution:

$$V^2 = \frac{2\mu}{r} - \frac{\mu}{a}$$

$$\Rightarrow \frac{v^2}{\mu} = \frac{2}{r} - \frac{1}{a}$$

$$\Rightarrow \frac{1}{a} = \frac{2}{r} - \frac{v^2}{\mu}$$

$$\Rightarrow a = \left[\frac{2}{r} - \frac{v^2}{\mu} \right]^{-1} \quad \dots \text{(1)}$$

If T be the time the particle takes to describe the whole arc of the ellipse, then

$$T = \frac{\text{area of the ellipse}}{\frac{h}{2}}$$

$$\Rightarrow T = \frac{2\pi ab}{h}$$

$$\Rightarrow \text{Also } h = \sqrt{\mu r} = \sqrt{\mu \cdot \frac{b^2}{a}} = \sqrt{\frac{\mu}{a}} \cdot b$$

$$\Rightarrow T = \frac{2\pi ab}{\sqrt{\frac{\mu}{a}} \cdot b}$$

$$\Rightarrow T = \frac{2\pi a^{\frac{3}{2}}}{\sqrt{\mu}}$$

P.T.O.

$$\therefore T = \frac{2\pi}{\sqrt{\mu}} \left[\frac{2}{r} - \frac{v^2}{\mu} \right]^{-\frac{3}{2}}$$

(Proved) :-

(2019) If a particle moves with a central acceleration $\left[= \frac{\mu}{(\text{distance})^2} \right]$; it is projected with velocity v at a distance R . Show that its path is a rectangular hyperbola if the angle of projection is $\sin^{-1} \left[\frac{\mu}{VR(v^2 - \frac{2\mu}{R})^{\frac{1}{2}}} \right]$.

Solution: Let the angle of projection be α . Then $p = r \sin \phi = r \sin \alpha$ and $v = \frac{1}{p} \Rightarrow h = vp = VR \sin \alpha \quad \text{--- (i)}$

Now in the case of rectangular hyperbola, we have $b = a$,

$$\therefore h = \sqrt{vr} = \sqrt{\mu \cdot \frac{b^2}{a^2}} = \sqrt{\mu \cdot \frac{a^2}{a}} = \sqrt{\mu a} \quad \text{--- (ii)}$$

Also we know that when the path is a rectangular hyperbola, then

$$v^2 = \mu \left(\frac{2}{R} + \frac{1}{a} \right)$$

P.T.O.

$$\Rightarrow V^2 - \frac{2\mu}{R} = \frac{\mu}{a}$$

$$\therefore a = \frac{\mu}{V^2 - \frac{2\mu}{R}} \quad \text{--- (iii)}$$

Combining (i), (ii), and (iii), we get

$$VR \sin \alpha = h = \sqrt{\mu a} = \sqrt{\mu} \cdot \sqrt{\frac{\mu}{(V^2 - \frac{2\mu}{R})}} \\ = \frac{\mu}{\sqrt{V^2 - \frac{2\mu}{R}}}$$

$$\Rightarrow \sin \alpha = \frac{VR \sqrt{(V^2 - \frac{2\mu}{R})}}{\mu}$$

$$\therefore \alpha = \sin^{-1} \left[\frac{1}{VR \left(V^2 - \frac{2\mu}{R} \right)^{\frac{1}{2}}} \right]$$

(Showed) ✓ b

(P)

2020

Q. A particle moves with central acceleration $\mu [3au^4 - 2(a^2 - b^2)u^5]$, $a > b$, and is projected from an apse at a distance $(a+b)$ with velocity $\sqrt{\mu}/(a+b)$. Show that its orbit is $r = a + b \cos\theta$.

Solution:

Consider the particle starts from $A(a+b, 0)$

with the velocity $\sqrt{\mu}/(a+b)$ and after time t it reached at $P(r, \theta)$.

The attractive force along $p_0 = \mu u \{3au^4 - 2(a^2 - b^2)u^5\}$
" acceleration along $p_0, F = \mu \{3au^4 - 2(a^2 - b^2)u^5\}$.

The equation of motion is

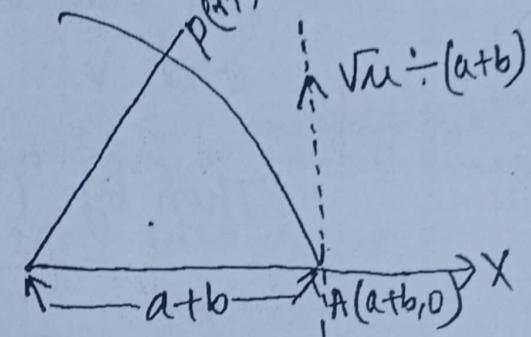
$$\lambda r u^2 \left(u + \frac{du}{dr}\right) = F = \mu u \{3au^4 - 2(a^2 - b^2)u^5\} \quad \boxed{F = ma} \\ \Rightarrow \lambda r u^2 \left(u + \frac{du}{dr}\right) = \mu \{3au^4 - 2(a^2 - b^2)u^5\} \quad \boxed{a = \frac{F}{m}}$$

$$\Rightarrow \lambda r u^2 \left(u + \frac{du}{dr}\right) = \mu \{3au^4 - 2(a^2 - b^2)u^5\}$$

Multiplying by $2 \frac{du}{d\theta}$

$$\lambda r \left[2u \frac{du}{d\theta} + 2 \frac{du}{dr} \frac{du}{d\theta} \right] = 2\mu \left[3au^4 - 2(a^2 - b^2)u^5 \right] \frac{du}{d\theta}$$

$$\Rightarrow \lambda r \left[\frac{d}{d\theta} (u^2) + \frac{d}{d\theta} \left(\frac{du}{d\theta} \right)^2 \right] = 2\mu \left[3au^4 - 2(a^2 - b^2)u^5 \right] \frac{du}{d\theta}$$



Integrating w.r.t. to θ

$$v^r = \frac{1}{2} \left[u^2 + \left(\frac{du}{d\theta} \right)^2 \right] = 2\mu \left[au^3 - \frac{(a^2 - b^2)u^4}{2} \right] + A \quad \text{--- (i)}$$

Initially $r = a+b = \frac{1}{u}$, $\frac{du}{d\theta} = 0$

$$\text{and } v = \frac{\sqrt{\mu}}{a+b}$$

Then by (i)

$$\frac{\mu}{(a+b)^2} = \frac{1}{(a+b)^2} = 2\mu \left[\frac{a}{(a+b)^3} - \frac{(a^2 - b^2)}{2(a+b)^4} \right] + A$$

and then

$$\frac{\mu}{(a+b)^2} = \frac{1}{(a+b)^2} \Rightarrow 1 = \mu$$

$$\text{and } \frac{\mu}{(a+b)^2} = 2\mu \left[\frac{a}{(a+b)^3} - \frac{(a^2 - b^2)}{2(a+b)^4} \right] + A$$

$$\Rightarrow A = \frac{\mu}{(a+b)^2} - \mu \left[\frac{2a}{(a+b)^3} - \frac{(a^2 - b^2)}{(a+b)^4} \right]$$

$$= \frac{\mu}{(a+b)^2} \left[1 - \frac{2a}{a+b} + \frac{(a+b)(a-b)}{(a+b)^2} \right]$$

$$= \frac{\mu}{(a+b)^2} \left[-\frac{a-b}{a+b} + \frac{a-b}{a+b} \right]$$

$$= 0$$

P.T.O.

So (i) becomes

$$u \left[u^r + \left(\frac{du}{d\theta} \right)^2 \right] = 2u \left[au^3 - \frac{(a^r - b^r) u^4}{2} \right]$$

$$\Rightarrow u^r + \left(\frac{du}{d\theta} \right)^2 = 2au^3 - (a^r - b^r)u^4.$$

But $u = \frac{1}{r} \Rightarrow \frac{du}{dr} = -\frac{1}{r^2} \Rightarrow du = -\frac{1}{r^2} dr$

$$\frac{1}{r^2} + \frac{1}{r^4} \left(\frac{dr}{d\theta} \right)^2 = \frac{2a}{r^3} - \frac{(a^r - b^r)}{r^4}$$

$$\Rightarrow r^r + \left(\frac{dr}{d\theta} \right)^2 = 2ar - (a^r - b^r)$$

$$\begin{aligned} \Rightarrow \left(\frac{dr}{d\theta} \right)^2 &= 2ar - (a^r - b^r) - r^r \\ &= -(r^r - 2ar + a^r - a^r) - (a^r - b^r) \\ &= -(r-a)^r + a^r - (a^r - b^r) \end{aligned}$$

$$\Rightarrow \frac{dr}{d\theta} = \sqrt{b^r - (r-a)^r}$$

$$\Rightarrow \frac{dr}{\sqrt{b^r - (r-a)^r}} = d\theta$$

Integrating

$$\sin^{-1} \left(\frac{r-a}{b} \right) = \theta + B \quad \text{--- (ii)}$$

P.T.Q

Initially $r = a+b$, $\theta = 0$ and
we have

$$\sin^{-1}\left(\frac{b}{r}\right) \phi = B$$

$$\Rightarrow B = \sin^{-1} \sin\left(\frac{\pi}{2}\right)$$

$$\Rightarrow B = \frac{\pi}{2}$$

and (ii) becomes

$$\sin^{-1}\left(\frac{r-a}{r}\right) = \theta + \frac{\pi}{2}$$

$$\Rightarrow \frac{r-a}{r} = \sin\left(\frac{\pi}{2} + \theta\right)$$

$$\Rightarrow r-a = r \cos \theta$$

$$\therefore r = a + r \cos \theta$$

(Showed) /.

Tangents and Normals (in POLAR)

Angle between a tangent and radius vector on a curve : If ~~dr = f(θ)~~ be the curve at ~~any~~ point If ϕ be the angle between a tangent and a radius vector at any point on a curve then

$$\tan \phi = r \frac{d\theta}{dr}, \cot \phi = \frac{1}{r} \frac{dr}{d\theta},$$

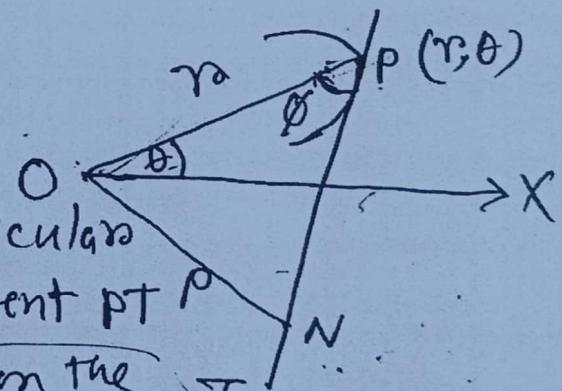
$$\sin \phi = r \frac{d\theta}{ds}, \cos \phi = \frac{dr}{ds}$$

perpendicular from a pole on a Tangent :

Let O be the pole.

$ON = p$ be the perpendicular length on the tangent pt P

at the point P on the curve.



$$\sin \phi = \frac{ON}{OP}$$

$$\Rightarrow ON = OP \sin \phi$$

$$\therefore p = r \sin \phi \Rightarrow \frac{1}{p} = \frac{1}{r} \csc \phi$$

$$\Rightarrow \frac{1}{pr} = \frac{1}{r^2} \csc \phi$$

$$\Rightarrow \frac{1}{pr} = \frac{1}{r^2} (1 + \cot^2 \phi)$$

$$= \frac{1}{r^2} \left\{ 1 + \frac{1}{r^2} \left(\frac{dr}{d\theta} \right)^2 \right\} \quad \left[\because \cot \phi = \frac{1}{r} \frac{dr}{d\theta} \right]$$

$$= \frac{1}{r^2} + \frac{1}{r^4} \left(\frac{dr}{d\theta} \right)^2$$

$$u = \frac{1}{r} \Rightarrow \frac{du}{d\theta} = -\frac{1}{r^2} \frac{dr}{d\theta}$$

$$\therefore \frac{1}{pr} = u^2 + \left(\frac{du}{d\theta} \right)^2$$

1-P-23v
H.W.

Q. A particle, subject to a central force per unit of mass equal to $\mu^2(a^r + b^r)u^2[3a^2u^2]$, is projected at a distance a with a velocity $(\sqrt{\mu}/a)$ in a direction at right angles to the initial distance. Show that the path is the curve: $r^2 = a^2 \cos^2 \theta + b^2 \sin^2 \theta$.

Q. A particle moves with a central acceleration $\mu(r^5 - c^4 r)$ being projected from an apse at distance c with a velocity $\sqrt{\frac{2\mu}{3}} c^3$. Show that it describes the path $x^4 + y^4 = c^4$.

Solution: The equation of motion is

$$F = \mu r^2(u + \frac{du}{d\theta}) = \mu(r^5 - c^4 r)$$

$$\Rightarrow \mu r(u + \frac{du}{d\theta}) = \frac{\mu}{u^2} \left(\frac{1}{u^5} - \frac{c^4}{u} \right), [r = \frac{1}{u}]$$

Multiplying by $2 \frac{du}{d\theta}$

$$\Rightarrow \mu^2 \left(2u \frac{du}{d\theta} + 2 \frac{du}{d\theta} \frac{d^2 u}{d\theta^2} \right) = 2\mu \left(\frac{1}{u^7} - \frac{c^4}{u^3} \right) \frac{du}{d\theta}$$

$$\Rightarrow \mu^2 \left[\frac{d}{d\theta} (u^2) + \frac{d}{d\theta} \left(\frac{du}{d\theta} \right)^2 \right] = 2\mu \left(\frac{1}{u^7} - \frac{c^4}{u^3} \right) \frac{du}{d\theta}$$

Integrating w.r.t. θ

$$\mu^2 \left[u^2 + \left(\frac{du}{d\theta} \right)^2 \right] = 2\mu \left(-\frac{1}{6u^6} + \frac{c^4}{2u^3} \right) + A$$

$$\Rightarrow v^2 = \mu^2 \left[u^2 + \left(\frac{du}{d\theta} \right)^2 \right] = \mu \left(\frac{c^4}{u^2} - \frac{1}{3u^6} \right) + A$$

At an apse $r_0 = c$ --- (i)

$$\Rightarrow \frac{1}{u} = c,$$

$$\frac{du}{d\theta} = 0 \text{ and } v = \sqrt{\frac{2\mu}{3}} c^3$$

Then by (i)

$$\frac{2mc^6}{3} = h^r \frac{1}{c^r} = m \left(c^6 - \frac{c^6}{3} \right) + A$$

$$\Rightarrow \frac{2mc^6}{3} = \frac{h^r}{c^r} = \frac{2mc^6}{3} + A$$

which implies that

$$\frac{2mc^6}{3} = \frac{h^r}{c^r} \Rightarrow h^r = \frac{2mc^8}{3}$$

$$\text{and } \frac{2mc^6}{3} = \frac{2mc^6}{3} + A \Rightarrow A = 0$$

So by (i) we have

$$\frac{2mc^8}{3} \left[u^r + \left(\frac{du}{d\theta} \right)^2 \right] = m \left(\frac{c^4}{u^r} - \frac{1}{3u^6} \right)$$

$$\Rightarrow u^r + \left(\frac{du}{d\theta} \right)^2 = \frac{3}{2c^8} \left(\frac{c^4}{u^r} - \frac{1}{3u^6} \right)$$

$$\Rightarrow u^r + \left(\frac{du}{d\theta} \right)^2 = \frac{3}{2c^8} \left(\frac{3c^4 u^4 - 1}{3u^6} \right)$$

$$\Rightarrow u^r + \left(\frac{du}{d\theta} \right)^2 = \frac{3c^4 u^4 - 1}{2c^8 u^6}$$

$$\Rightarrow \left(\frac{du}{d\theta} \right)^2 = \frac{3c^4 u^4 - 1}{2c^8 u^6} - u^r$$

$$\begin{aligned}
 \Rightarrow \left(\frac{du}{d\theta} \right)^r &= \frac{3c^4 u^4 - 1 - 2c^8 u^8}{2c^8 u^6} \\
 &= \frac{-2}{2c^8 u^6} \left(c^8 u^8 - \frac{3}{2} c^4 u^4 + \frac{1}{2} \right) \\
 &= -\frac{1}{c^8 u^6} \left[(c^4 u^4)^2 - 2 \cdot c^4 u^4 \cdot \frac{3}{4} + \left(\frac{3}{2}\right)^2 \right]^{\frac{1}{2}} \\
 &= \frac{1}{c^8 u^6} \left[\left(\frac{1}{4}\right)^2 - \left(c^4 u^4 - \frac{3}{2}\right)^2 \right]^{\frac{1}{2}}
 \end{aligned}$$

$$\Rightarrow \frac{du}{d\theta} = \frac{\sqrt{\left(\frac{1}{4}\right)^2 - \left(c^4 u^4 - \frac{3}{2}\right)^2}}{c^4 u^3}$$

$$\Rightarrow \frac{c^4 u^3}{\sqrt{\left(\frac{1}{4}\right)^2 - \left(c^4 u^4 - \frac{3}{2}\right)^2}} du = d\theta$$

Integrating we get

$$\frac{1}{4} \sin^{-1} \left[\frac{c^4 u^4 - \frac{3}{2}}{\frac{1}{4}} \right] = \theta + B$$

P.T.O.

$$\Rightarrow \frac{1}{4} \sin^{-1} (4c^4 u^4 - 3) = \theta + B \quad \text{---(ii)}$$

At an apse $r = \infty$

$$\Rightarrow \frac{1}{u} = 0 \text{ and } \theta = 0.$$

Then (ii) gives

$$\frac{1}{4} \sin^{-1} 1 = B$$

$$\Rightarrow \frac{1}{4} \sin^{-1} \sin \frac{\pi}{2} = B$$

$$\Rightarrow B = \frac{\pi}{8}$$

Hence (ii) becomes

$$\frac{1}{4} \sin^{-1} (4c^4 u^4 - 3) = \theta + \frac{\pi}{8}$$

$$\Rightarrow \sin^{-1} (4c^4 u^4 - 3) = 4\theta + \frac{\pi}{2}$$

$$\Rightarrow 4c^4 u^4 - 3 = \sin(\frac{\pi}{2} + 4\theta)$$

$$\Rightarrow 4c^4 u^4 = 3 + \cos 4\theta$$

$$= 3 + 2 \cos^2 2\theta - 1$$

$$= 2 + 2 \cos^2 2\theta$$

$$= 2(1 + \cos^2 2\theta)$$

$\rho \cdot T \cdot \sigma$

$$\begin{aligned}
 \Rightarrow 2c^4 u^4 &= 1 + \cos^2 2\theta \\
 &= (\cos^2 \theta + \sin^2 \theta)^2 + (\cos^2 \theta - \sin^2 \theta)^2 \\
 &= 2[(\cos^2 \theta)^2 + (\sin^2 \theta)^2] \\
 \Rightarrow \frac{c^4}{r^4} &= \cos^4 \theta + \sin^4 \theta, \quad [r = \frac{1}{u}] \\
 \Rightarrow (r \cos \theta)^4 + (r \sin \theta)^4 &= c^4 \\
 \Rightarrow x^4 + y^4 &= c^4. \\
 &\text{(Showed) /}
 \end{aligned}$$

Q. A particle moves under a central repulsive force $m\mu/(\text{distance})^3$, and is projected from an apse at a distance a with velocity v . Show that the equation to the path is $r \cos \theta = a$, and that the angle described in time t is $\frac{1}{p} \tan^{-1} \left(\frac{pv}{a} t \right)$, where

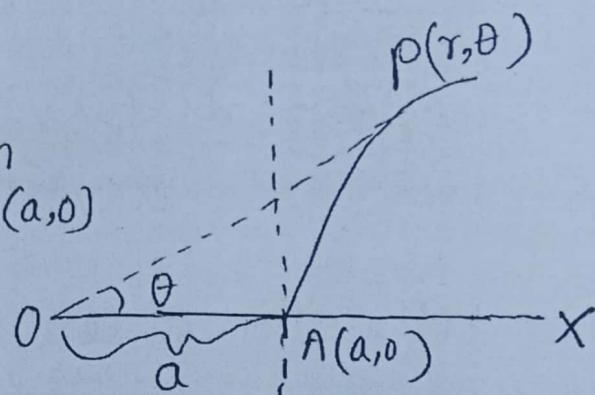
P.T.O

$$P^r = \frac{\mu + a^r v^2}{a^r v^2}$$

Q. In a central orbit the force is $\mu u^3(3 + 2a^r u^2)$; if the particle be projected at a distance a with velocity $\sqrt{5\mu/a^r}$ in a direction making an angle $\tan^{-1}(1/2)$ with the radius, show that the equation to the path is $au = \tan(\theta + \xi)$, or the path is $r = a \tan \theta$.

Solution:

Let at initial position
of the particle is at $A(a, 0)$
on OX at a distance
(a) from O .



And at time (t) its position be $p(r, \theta)$ for the
central repulsive force. \therefore The repulsive force
along OP at $P = \frac{mu}{r^3}$.

Acceleration along OP , $f = \frac{r^3}{r^3}$

We know,

$$\frac{du}{d\theta^2} + u = \frac{f}{hrur} = \frac{1}{hrur} \left(-\frac{mu}{r^3} \right) = -\frac{mu^3}{hrur^2}$$

$$\Rightarrow \frac{du}{d\theta^2} + u = -\frac{mu}{hr^2} \quad \dots (i)$$

[the negative sign is due to the fact
that force is repulsive]

Multiplying both sides of (i) by $2 \frac{du}{d\theta}$,
we get

$$2 \frac{du}{d\theta} \cdot \frac{du}{d\theta^2} + 2 \frac{du}{d\theta} \cdot u = -\frac{2mu}{hr^2} u \frac{du}{d\theta}$$
$$\Rightarrow \frac{d}{d\theta} \left\{ \left(\frac{du}{d\theta} \right)^2 + u^2 \right\} = \frac{d}{d\theta} \left(-\frac{mu^3}{hr^2} \right)$$

P.T.O

$$\Rightarrow h^r \left[\frac{d}{d\theta} \left(\frac{du}{dr} \right)^r + u^r \right] = \frac{d}{d\theta} (-mu^r)$$

Integrating both sides, we get

$$h^r \left[\left(\frac{du}{dr} \right)^r + u^r \right] - \cancel{\frac{mu^r}{r}} + c_1 \dots (2)$$

$$\Rightarrow v^r = -mu^r + c_1 \dots (3)$$

Now at an apse $r=a$

$$\text{i.e } u = \frac{1}{a}, \frac{du}{dr} = 0, v = V$$

$$\therefore v^r = -\frac{m}{a^r} + c_1$$

$$\Rightarrow c_1 = vr + \frac{m}{a^r}$$

$$\text{Given condition is } pr = \frac{u + a^r v^r}{a^r v^r}$$

$$\Rightarrow p^r v^r = \frac{u}{a^r} + v^r$$

$$\therefore c_1 = p^r v^r \dots (4)$$

Again since for an apse $p=r=a$, we get

$$h = pr = aV \Rightarrow h^r = a^r V^r \dots (5)$$

Putting the values of c_1 and h^r in (2), we get

$$a^r V^r \left[\left(\frac{du}{dr} \right)^r + u^r \right] = -mu^r + p^r V^r$$

$$\Rightarrow \left(\frac{du}{dr} \right)^r + u^r = \frac{p^r V^r - mu^r}{a^r V^r}$$

P.T.O.

$$\begin{aligned}
 \Rightarrow \left(\frac{du}{d\theta} \right)^r &= \frac{prV^r - u^r(u + arV^r)}{arV^r} \\
 &= \frac{prV^r}{arV^r} - \frac{u^r(u + arV^r)}{arV^r} \\
 &= \frac{pr}{ar} - u^r \left(\frac{u + arV^r}{arV^r} \right) \\
 &= \frac{pr}{ar} - u^r p^r \\
 &= \frac{pr}{ar} - \frac{pr}{ar} \cdot arur \\
 &= \frac{pr}{ar} (1 - arur)
 \end{aligned}$$

$$\Rightarrow \frac{du}{d\theta} = -\frac{p}{a} \sqrt{1 - arur}$$

$$\Rightarrow \frac{adu}{\sqrt{1 - (aru)^2}} = p d\theta$$

$$\Rightarrow \sin^{-1}(aru) = p\theta + C_2 \quad \text{[Integrating]} \\ \text{--- (6)}$$

P.T.O.

Initially, at an apse, $u = \frac{1}{a}$, $\theta =$

$$\therefore \sin^{-1} 1 = 0 + c_2 \Rightarrow c_2 = \sin^{-1} \sin \frac{\pi}{2}$$

$$\Rightarrow c_2 = \frac{\pi}{2}$$

Putting $c_2 = \frac{\pi}{2}$ in (6) we get,

$$\sin^{-1}(au) = p\theta + \frac{\pi}{2}$$

$$\Rightarrow au = \sin\left(\frac{\pi}{2} + p\theta\right) = \cos p\theta$$

$$\Rightarrow \frac{a}{r^p} = \cos p\theta \quad \left[\because u = \frac{1}{r^p} \right]$$

~~$r \cos p\theta = a$~~ which is the required path equation.

2nd part:

Again for t, we have

$$r^p \frac{d\theta}{dt} = l = av$$

$$\Rightarrow \frac{d\theta}{dt} = \frac{av}{r^p} = \frac{av}{ar} = \frac{v}{a} \cos p\theta$$

$$\Rightarrow \frac{d\theta}{\cos p\theta} = \frac{v}{a} dt$$

$$\Rightarrow \int \sec p\theta d\theta = \frac{v}{a} dt$$

Integrating both sides, we get

$$-\tan p\theta = \frac{v}{a} t + c_3 \dots (7)$$

Initially $t=0, \theta=0 \therefore c_3=0$

Thus from (7), we have

$$\frac{1}{p} \tan p\theta = \frac{V}{a} t$$

$$\Rightarrow p\theta = \tan^{-1}\left(\frac{PVt}{a}\right)$$

$$\therefore \theta = \frac{1}{p} \tan^{-1}\left(\frac{PV}{a} t\right) \text{ which is the}$$

required angle. (Showed)