

(step 0: retrieve data)

<http://archive.stsci.edu/hst/search.php>

Proposal IDs:

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12679,12879,13101,13334,13335,13344,13678,13686,13928,14062,14206

Filters/Gratings:

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F160W\*

Got 248 entries returned.

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*query\_result.csv*

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**(1) Select the Cepheids in our sample and convert the calendar dates to Julian dates.**

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**(2) If the same object was observed several times within two hours, take them as only one epoch. Derive the phase of each epoch.**

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**(3) Calculate three types of sigma:**

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- (3.1.1) Uncertainty from the model:

The model reads

$$m_t = M + L \cdot [A_0 + \sum_{i=1}^7 A_i \cos(2\pi i(\phi_t + \psi) + \Phi_i)] + \sigma_t \epsilon$$

.....

which yields

$$\begin{aligned}\sigma_1 &= \frac{\partial m}{\partial M} \cdot \sigma_M + \frac{\partial m}{\partial L} \cdot \sigma_L + \frac{\partial m}{\partial \psi} \cdot \sigma_\psi \\ &= \sigma_M + \sigma_L \cdot |A_0 + \sum_{i=1}^7 A_i \cos(2\pi i(\phi_t + \psi) + \Phi_i)| + L \cdot \sigma_\psi \cdot |\sum_{i=1}^7 A_i \sin(2\pi i(\phi_t + \psi) + \Phi_i) \cdot 2\pi i|\end{aligned}$$

- (3.1.2) Uncertainty from model fit residuals (since template is not the true light curve)

$\sigma_2 = \sigma$  where  $\sigma$  is the model-measurement scatter from [Table 4 of Inno+ \(2015\)](#), or maybe better, take  $\sigma$  as the standard deviation of our model fit residuals, since we fit the model for individual Cepheids and have some freedom in the template to reduce  $\sigma$ .

- (3.1.3) Uncertainty from the uncertainty of the period

$$\begin{aligned}\sigma_3 &= \frac{\partial m}{\partial \phi} \cdot \frac{\partial \phi}{\partial P} \cdot \sigma_P \\ &= L \cdot |\sum_{i=1}^7 A_i \sin(2\pi i(\phi_t + \psi) + \Phi_i) \cdot 2\pi i \cdot \frac{N_{cyc}}{P}| \cdot \sigma_P\end{aligned}$$

Where  $N_{cyc}$  is the number of cycles between HST and ground-based observations, and it can be negative if the HST observation is prior to the ground-based observation.

**(3.2.1) If there is only one HST observation, then  $\sigma = \sqrt{\sigma_1^2 + \sigma_2^2 + \sigma_3^2}$ .**

**(3.2.2) If there are multiple HST observations:**

$$\sigma_{1,total} = \sigma_M + \frac{\sigma_L}{N} \cdot |\sum_{j=1}^N \{A_0 + \sum_{i=1}^7 A_i \cos(2\pi i(\phi_{t,j} + \psi) + \Phi_i)\}| + \frac{L \cdot \sigma_\psi}{N} \cdot |\sum_{j=1}^N \{\sum_{i=1}^7 A_i \sin(2\pi i(\phi_{t,j} + \psi) + \Phi_i) \cdot 2\pi i\}|$$

The  $\sigma_{1,total}$  denotes the total model uncertainty. It does not go down with the square root of number of observations  $N$ , but the differences in phase might beat the last two terms down.

$$\sigma_{2,total} = \frac{\sigma_2}{\sqrt{N}}$$

This assumes that the model residuals of ground-based measurements are the same as that of HST-based measurements. This is true if the residuals of the model fit of true light curves are relatively large. If not,  $\sigma_{2,total}$  would over estimate the residual uncertainty.

$$\sigma_{3,total} = \frac{L}{N} \cdot |\sum_{j=1}^N \sum_{i=1}^7 \{A_i \sin(2\pi i(\phi_{t,j} + \psi) + \Phi_i) \cdot 2\pi i \cdot \frac{N_{cyc}}{P}\}| \cdot \sigma_P$$

Similar to  $\sigma_{1,total}$ ,  $\sigma_{3,total}$  does not go down with  $N$  substantially. It decrease only when the HST observations distribute

both before and after the ground-based observations.

Finally, the total phase correction uncertainty of multiple HST observations should be

$$\sigma = \sqrt{\sigma_{1,total}^2 + \sigma_{2,total}^2 + \sigma_{3,total}^2}$$

Results

#	id	sigma.1	sigma.2	sigma.3	sigma.total	mag.corr
	adpup	0.03144	0.02413	0.00035	0.03963	-0.14656
	aqcar	0.00745	0.00782	0.00000	0.01080	-0.03681
	aqpup	0.01110	0.01294	0.00022	0.01705	0.00957
	betad	0.03286	0.03251	0.00000	0.04622	-0.13521
	bnpup	0.06742	0.04923	0.00007	0.08348	-0.05277
	crcar	0.00987	0.01672	0.00000	0.01942	0.03029
	drvel	0.02510	0.02501	0.00017	0.03543	0.03932
	hwcar	0.00576	0.00580	0.00000	0.00818	-0.00794
	kkcen	0.00838	0.01874	0.00006	0.02053	-0.13490
	kncen	0.01871	0.01356	0.00012	0.02311	-0.04146
	lcarl	0.03923	0.03513	0.00035	0.05266	-0.10137
	rysco	0.00959	0.00966	0.00000	0.01361	-0.03411
	ryvel	0.01398	0.01742	0.00009	0.02234	-0.01858
	s-nor	0.02993	0.03157	0.00000	0.04350	0.03132
	sscma	0.01817	0.01313	0.00056	0.02243	-0.00256
	svvel	0.01919	0.01125	0.00000	0.02225	0.11542
	synor	0.01760	0.01638	0.00000	0.02404	0.02580
	t-mon	0.03821	0.02592	0.00000	0.04618	0.14277
	u-car	0.01656	0.01493	0.00035	0.02230	0.18328
	uumus	0.00838	0.00575	0.00041	0.01017	-0.07461
	vjara	0.00864	0.01437	0.00000	0.01676	0.09615
	vjcen	0.02584	0.02908	0.00004	0.03890	0.08094
	vwcen	0.00991	0.01248	0.00002	0.01594	-0.18063
	vycar	0.01417	0.00936	0.00082	0.01700	0.04012
	vzpup	0.03728	0.02650	0.00110	0.04575	-0.26534
	w-sgr	0.02044	0.03140	0.00000	0.03747	-0.00706
	wxpup	0.01431	0.01205	0.00025	0.01871	0.08612
	wzsg	0.00513	0.00362	0.00024	0.00629	-0.01021
	x-pup	0.02486	0.01564	0.00119	0.02940	-0.03171
	xxcar	0.02473	0.02645	0.00019	0.03621	0.00050
	xycar	0.00868	0.00667	0.00000	0.01095	-0.03409
	xzcar	0.01006	0.00880	0.00012	0.01336	0.04828
	yzcar	0.02715	0.01639	0.00005	0.03171	-0.07605
	yzsgr	0.00824	0.02115	0.00001	0.02270	0.04179

(4) Update the table in the draft paper

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\begin{deluxetable*}{lrrrrrrrrrc}
\tabletypesize{\scriptsize}
\tablecaption{Light Curve Parameters \label{tbl:par}}
\tablewidth{0pt}
\tablehead{
\multicolumn{1}{c}{Object} & \multicolumn{1}{c}{ $P$ } & \multicolumn{1}{c}{ $\sigma_P$ } & \multicolumn{1}{c}{ $t_0$ } & \multicolumn{1}{c}{ $M$ } & \multicolumn{1}{c}{ $\sigma_M$ } & \multicolumn{1}{c}{ $L$ } & \multicolumn{1}{c}{ $\sigma_L$ } & \multicolumn{1}{c}{ $\psi$ } & \multicolumn{1}{c}{ $\sigma_{\psi}$ } & \multicolumn{1}{c}{ $\sigma_{\mathrm{corr}}$ } \\
\multicolumn{1}{c}{} & \multicolumn{1}{c}{(d)} & \multicolumn{1}{c}{([1])} & \multicolumn{1}{c}{(d)} & \multicolumn{1}{c}{(mag)} & \multicolumn{3}{c}{( $10^{-4}$  mag)} & \multicolumn{2}{c}{([2])} & \multicolumn{1}{c}{( $10^{-4}$  mag)} \\
\startdata
\input{tables/pars.tex}
\enddata
\tablecomments{[1]: units of  $10^{-6}$  s; [2]: units of  $10^{-4}$  rad/2 $\pi$ .}
\end{deluxetable*}

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W~Sgr	&	7.585536	&	1768	&	7187	&	2.886	&	97	&	2902	&	365	&	677
5	&	153	&	375	&		&		&		&		&		&	
WX~Pup	&	8.935991	&	422	&	6967	&	6.690	&	48	&	2571	&	224	&	330
8	&	83	&	187	&		&		&		&		&		&	
HW~Car	&	9.199488	&	39	&	6759	&	6.753	&	35	&	1501	&	234	&	659
7	&	113	&	82	&		&		&		&		&		&	
V339~Cen	&	9.466540	&	76	&	6961	&	5.809	&	108	&	2649	&	329	&	682
0	&	194	&	389	&		&		&		&		&		&	
YZ~Sgr	&	9.553551	&	290	&	6900	&	4.941	&	44	&	2558	&	116	&	597
7	&	71	&	227	&		&		&		&		&		&	
S~Nor	&	9.754615	&	122	&	7416	&	4.384	&	170	&	2292	&	446	&	482
5	&	243	&	435	&		&		&		&		&		&	
CR~Car	&	9.758552	&	119	&	7179	&	8.211	&	43	&	2079	&	121	&	905
1	&	95	&	194	&		&		&		&		&		&	
AQ~Car	&	9.769427	&	119	&	6762	&	6.704	&	44	&	2023	&	138	&	23
4	&	169	&	108	&		&		&		&		&		&	
$\beta$ ~Dor	&	9.842865	&	3365	&	6910	&	1.974	&	116	&	2661	&	337	&	102
9	&	194	&	462	&		&		&		&		&		&	
DR~Vel	&	11.199240	&	86	&	6996	&	5.983	&	94	&	2759	&	258	&	290
4	&	147	&	354	&		&		&		&		&		&	
UU~Mus	&	11.636093	&	156	&	7038	&	6.959	&	27	&	3109	&	91	&	112
4	&	31	&	102	&		&		&		&		&		&	
KK~Cen	&	12.182794	&	135	&	7042	&	8.106	&	30	&	2714	&	80	&	451
0	&	58	&	205	&		&		&		&		&		&	
SS~Cma	&	12.353912	&	571	&	6958	&	6.836	&	106	&	2463	&	374	&	812
7	&	150	&	224	&		&		&		&		&		&	
XY~Car	&	12.436119	&	23	&	6735	&	6.425	&	60	&	3104	&	174	&	122
1	&	74	&	110	&		&		&		&		&		&	
SY~Nor	&	12.645111	&	94	&	7181	&	6.009	&	76	&	3163	&	265	&	378
9	&	79	&	240	&		&		&		&		&		&	

[illegible]