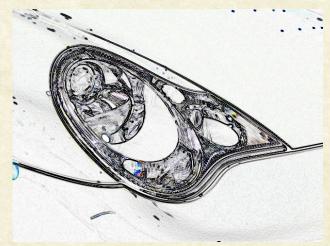




# CS7.505: Computer Vision

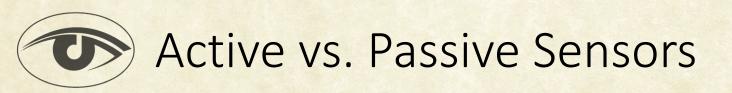
#### Spring 2024: Active Illumination Methods







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#### **Passive Sensor**

- Record the interaction between the scene and existing energy sources
- Does not change the world (other than sensor presence)
- Advantages:
  - Safety
  - Energy efficiency
  - Stealth
  - Multi-spectral imaging

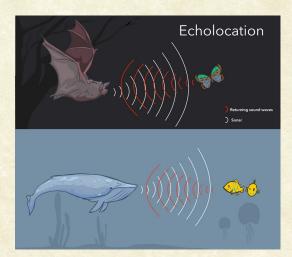
#### **Active Sensors**

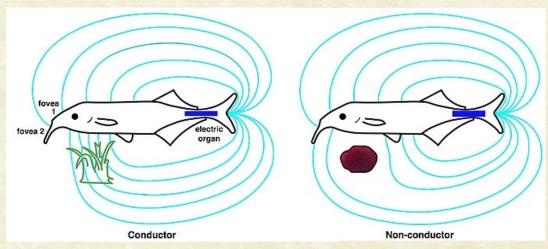
- Send energy/probe out into the world and record its interaction with the world
- Modifies the world through introduction of probe
- Advantages:
  - · Works in 'dark'
  - Accuracy
  - Reliability / Predictability
  - See through occlusions
  - Sense several scene characteristics



## Active Sensors in Nature

- Echolocation
  - Bats, Whales
- Electrolocation
  - Electric eel
- Mechanical
  - Antennae
  - Whiskers
- Chemical
  - Slime Mold
- Bioluminesence
  - Firefly









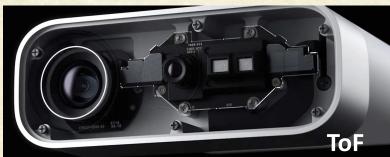




# Active Sensing Devices



















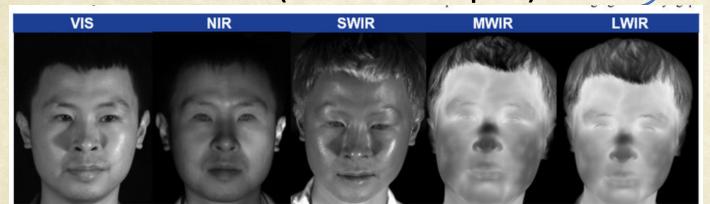




- Reflection
- Transmission
- Emission
- Induction
- Capacitance

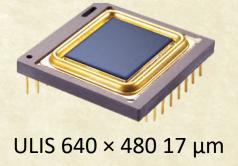


- Near Infrared (700 nm 1.4  $\mu$ m):
  - Most camera sensors are sensitive to this range (sans IR-cut filter)
- Short-Wave Infrared  $(1.4 3 \mu m)$ :
- Mid-Wave Infrared (3 8 μm):
- Long-Wave Infrared (8 15 μm):
- Far-Infrared (15 1000 μm):





Thermal Camera







#### Photometric Stereo

Based on the Lectures:
First Principles of Computer Vision
By Shree K Nayar

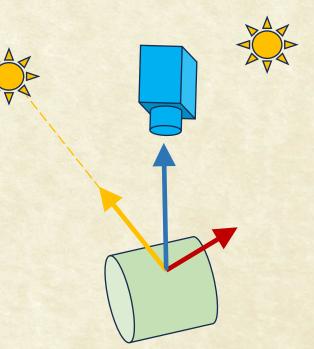


## Photometric Stereo

- A method for recovering 3D shape from images measured using multiple known light sources.
- The image intensity, I, at a point is:

I = f(Source, Normal, Reflectance)

- Image intensities are measured
- Source directions, distances and intensities are known
- Reflectance is characterized by a known (often simplified) BRDF function



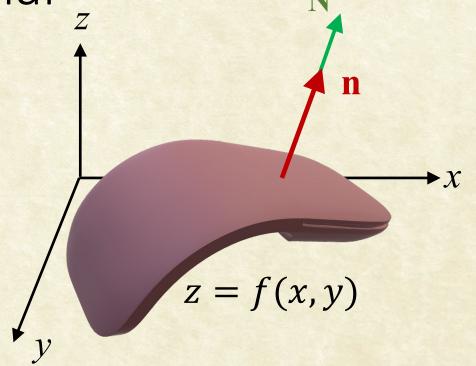


## Surface Gradient and Normal

• Gradient: 
$$\left(-\frac{dz}{dx}, -\frac{dz}{dy}\right) = (p, q)$$

• Normal: 
$$\left(-\frac{dz}{dx}, -\frac{dz}{dy}, 1\right) = (p, q, 1)$$

- The surface normal, N, can be normalized to unit length to get: n.
- The surface normal is represented by two parameters: (p,q)
- Source direction can also be expressed in (p, q) space.

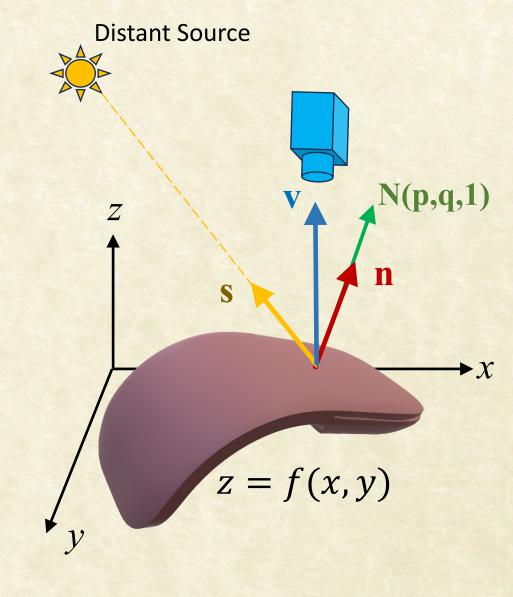




# Reflectance Map: R(p,q)

- The Reflectance Map, R(p,q), gives the image intensity at any point (x,y), computed as a function of the source direction, surface gradient (p,q), and surface reflectance (BRDF).
- For a Lambertian (matte) surface, the image intensity is independent of viewing direction.

$$I = c \frac{\rho}{\pi} \frac{J}{r^2} \cos \theta_i = c \frac{\rho}{\pi} k(\boldsymbol{n}.\boldsymbol{s}) = \boldsymbol{n}.\boldsymbol{s}$$

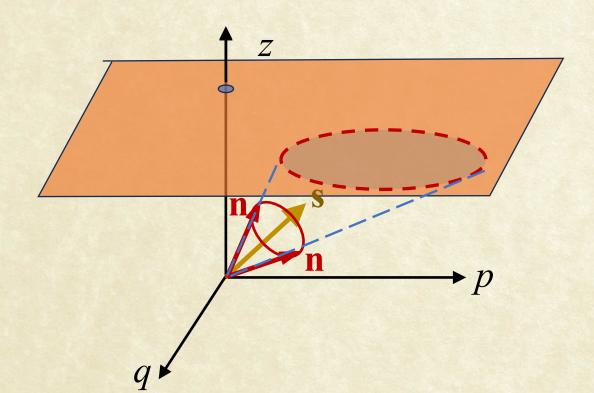


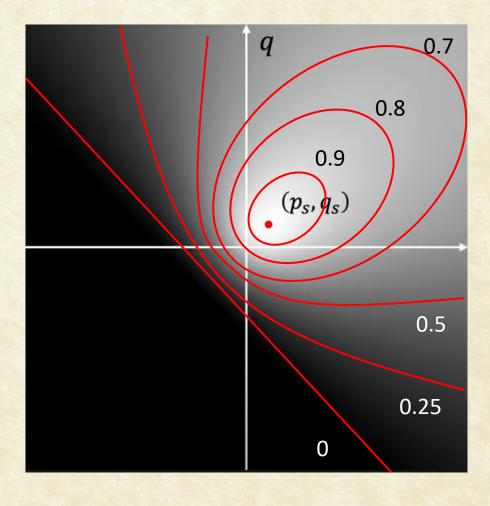
$$= \frac{pp_s + qq_s + 1}{\|N\| \|S\|} = R(p, q)$$



#### From Intensities to Surface Normals

- Given a Reflectance map for a point source, the iso-intensity contours are conics
- Given an intensity there are several possible normals

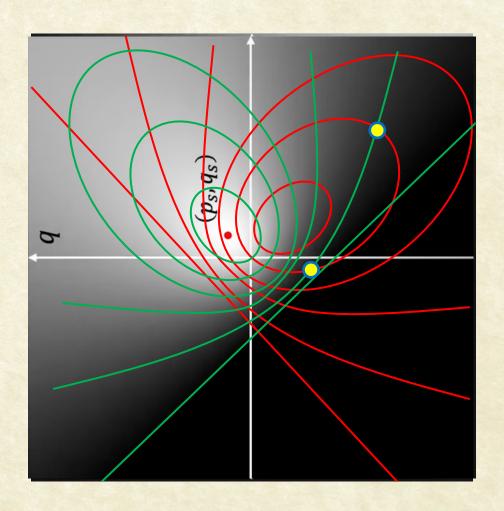






# Computing Unique Normals

- Use multiple light sources
  - Note: We assume the viewing direction and source directions are same for all points
- The correct normal should satisfy all reflectance maps
- Can be uniquely solved by three light sources (or more)
  - In-fact can compute albedo also with 3 sources





#### Photometric Stereo: A Naïve Algorithm

- 1. Capture k images with k known light sources
- 2. Using the known source details and BRDF, compute the reflectance map for each source.:  $R_1(p,q), \ldots, R_k(p,q)$ .
- 3. Quantize the R-Map and for each value of (p,q), enter the k intensities and corresponding normal into a lookup table.
- 4. For each pixel (x,y), lookup table row that best matches the intensities:  $I_1$ ,  $I_2$ ,  $I_3$ , ,  $I_k$  and choose the corresponding (p,q).

#### Notes:

- All light sources may not illuminate all points
- We assume the surface is Lambertian and Isotropic.
- What happens if the above are not satisfied?

### Computing Normals: Lambertian Case

$$\bullet I_1 = \frac{\rho}{\pi} n. s_1, \quad I_2 = \frac{\rho}{\pi} n. s_2, \quad I_3 = \frac{\rho}{\pi} n. s_3,$$
 where  $n = \begin{bmatrix} n_x \\ n_y \\ n_z \end{bmatrix}, \quad \text{and} \quad s_i = \begin{bmatrix} s_{xi} \\ s_{yi} \\ s_{zi} \end{bmatrix}$ 

We can write the three equations together in matrix form:

$$\begin{bmatrix} I_1 \\ I_2 \\ I_3 \end{bmatrix} = \frac{\rho}{\pi} \begin{bmatrix} S_{x1} & S_{y1} & S_{z1} \\ S_{x2} & S_{y2} & S_{z2} \\ S_{x3} & S_{y3} & S_{z3} \end{bmatrix} n$$

$$I = \frac{\rho}{\pi} S. n; \text{ or } I = S. N \text{ and } N = S^{-1}I;$$

$$\text{Normal: } n = \sqrt[N]{\|N\|}; \text{ Albedo: } \rho = \pi \|N\|$$



## Extending to k Light Sources

$$\begin{bmatrix} I_1 \\ I_2 \\ \vdots \\ I_k \end{bmatrix} = \frac{\rho}{\pi} \begin{bmatrix} S_{\chi 1} & S_{y1} & S_{z1} \\ S_{\chi 2} & S_{y2} & S_{z2} \\ \vdots & \vdots & \vdots \\ S_{\chi k} & S_{yk} & S_{zk} \end{bmatrix} n$$

- There are k equations in 3 unknowns
  - Find the least-squared-error solution
- Mathematically,
  - S is a kx3 matrix, which cannot be inverted.
- Solution:
  - Pre-multiply both sides by S<sup>-1</sup> and solve.
  - The method of pseudo-inverse



### Other Considerations

- Dealing with non-point light sources
- Dealing with non-Lambertian surfaces
- Dealing with color objects
- Dealing with inter-reflections
- Other errors in assumption

By Calibration

# Computing Depth from Normals

- Find a set of z values that minimizes the error between measured normals and normals computed from z values.
  - Naïve approach: Incrementally compute z from its derivatives (n)
  - Does not work well in presence of noise
- Minimize the squared error between the two. i.e, find z that minimizes:

$$D = \iint_{Image} \left(\frac{dz}{dx} + p\right)^2 + \left(\frac{dz}{dy} + q\right)^2 dx dy$$

Note, Gradient: 
$$\left(-\frac{dz}{dx}, -\frac{dz}{dx}\right) = (p, q)$$

#### From Normals to Depth: Frankot & Chellappa

- Key Idea: Minimize D in Fourier domain
- Let Z(u,v), P(u,v) and Q(u,v) be the Fourier transforms of z(x,y), p(x,y) and q(x,y) respectively.
- Find Z(u, v) that minimizes D by setting:  $\frac{\partial D}{\partial z} = 0$ .
- Solution:

$$\tilde{Z}(u,v) = \frac{iuP(u,v) + ivQ(u,v)}{u^2 + v^2}$$

• Compute  $\mathcal{F}^{-1}\left(\tilde{Z}(u,v)\right)$  to get  $\tilde{z}(x,y)$ .



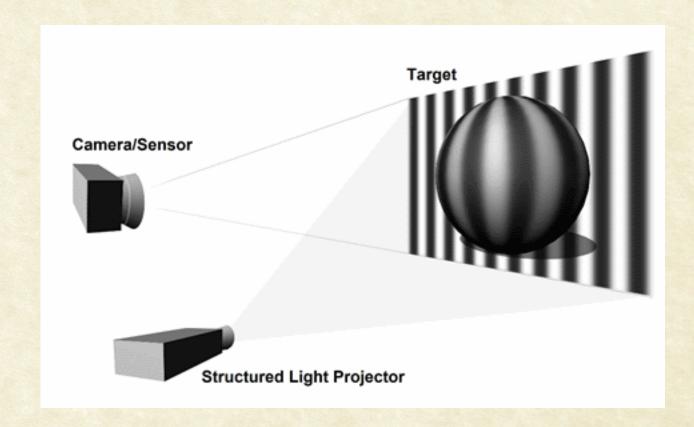


# Structured Light Sensing



#### Projector-Camera Systems

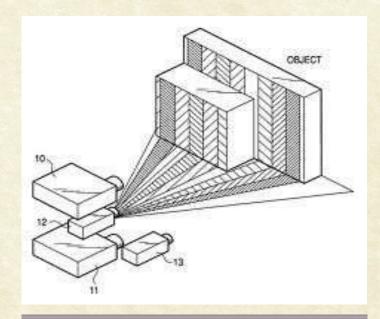
- Projector is a camera
- Projector and camera forms a stereo pair.
- We can control what we project
  - Single point at a time
  - Binary Coding
  - Vertical/Horizontal stripes

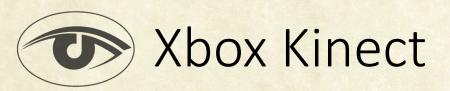




### Structured Lighting with Stereo Cameras

- Finding correspondences is hard by itself
- Can we help it by projecting patterns onto the world?
- Light-strip range finders, etc.
- Combination of sinusoids sometimes to get dense matches
- Active vision, as it changes the appearance
- The light projected need not be in the visible spectrum





#### IR-based range sensor for Xbox

- Aligned depth and RGB images at 640 × 480
- Original goal: Interact with games in full 3D
- Computer vision happy with real-time depth and image
  - Games, HCI, etc
  - Action recognition
  - Image based modelling of dynamic scenes
- Fastest selling electronic appliance ever!!
- Other products that use PrimeSense sensor





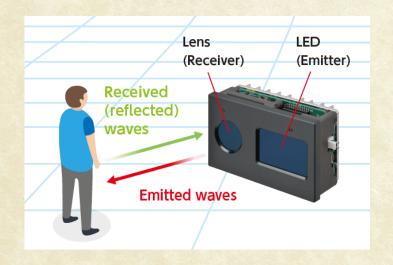


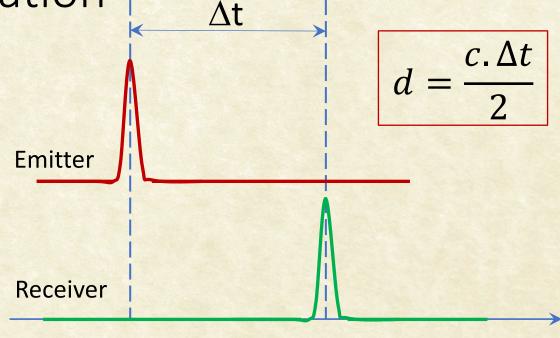
# Optical Time-of-Flight (ToF) Sensors



#### Direct ToF or Pulse Modulation

- Send a short light pulse
- Measure the time taken for reflection
- Compute distance based on time elapsed



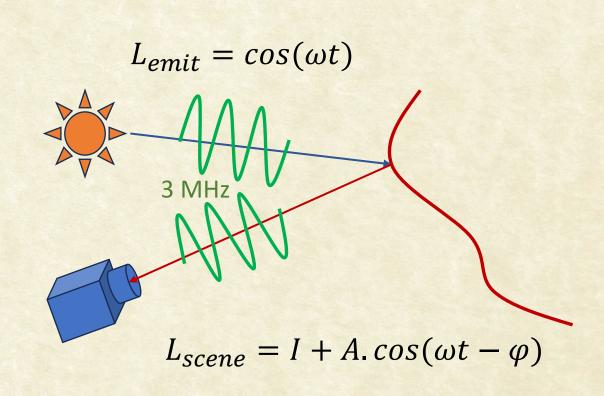


- Need very short pulse for accuracy
- Shorter pulses should be very bright for effective sensing (referred to as the flash method)



#### Indirect ToF or Continuous Modulation

- Use a temporally modulated light source
- Find the phase difference of the reflected light
- Distance is proportional the phase difference
- Terminology:
  - ω: Modulation Frequency
  - I: Ambient Illumination
  - φ: Phase difference





# Computing Depth

• 
$$L_{scene} = I + A.cos(\omega t - \varphi)$$

• 
$$S_{ref} = cos(\omega t - \delta_i)$$

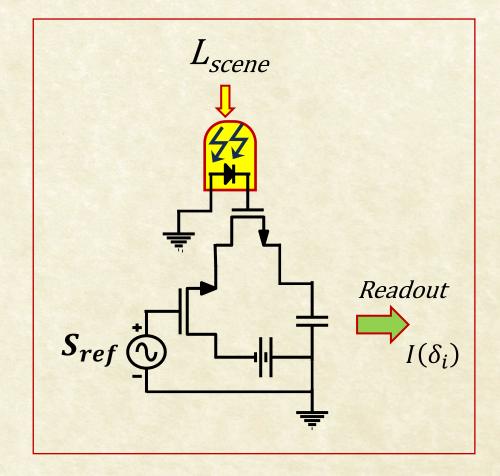
• 
$$I(\delta_i) = \int_0^t L_{scene}(t) \times S_{ref}(t, \delta_i) dt$$

Substituting and simplifying, we get:

$$I(\delta_i) = \mathbf{P} + \mathbf{Q} \cos(\delta_i - \boldsymbol{\varphi})$$

• Measure  $I(\delta_i)$  for three different  $\delta_i$  to solve for the unknowns and get  $\varphi$ .

• E.g., 
$$d(f = 3MHz, \varphi = \pi) = 25m$$



$$d(f,\varphi) = \frac{c.\,\varphi}{4\pi f}$$

$$f = \omega/2\pi$$



Questions?