THE EFFECTS OF TELLURIC LINES IN RADIAL VELOCITY SEARCHES FOR PLANETS WITH IODINE CELL AS CALIBRATORS 1

Sharon Xuesong Wang (王雪凇)^{2,3}, Jason T. Wright^{2,3}, Chad Bender^{2,3}, Andrew W. Howard⁴, Geoffrey W. Marcy⁵, Howard Isaacson⁵, and Suvrath Mahadevan^{2,3}

ABSTRACT

Tellurics are bad and you really don't want them. Here's how to get rid of them. Subject headings: instrumentation

1. INTRODUCTION

We are going to cite Artigau et al. (2014); Cunha et al. (2014) and Rothman et al. (2013).

2. METHODOLOGY

Introduction to Keck/HIRES CPS data,

2.1. Simulated Keck/HIRES Spectra

ZZZ Plots:

- comparison between simulated and observed Keck/HIRES spectra: blaze, and zoom in.
- comparison between SME stellar template and empirical, deconvolved template.

2.2. Injection of Telluric Lines

Description of TERRASPEC.

ZZZ Plots:

- telluric line atlas in the optical? just to demonstrate where they are.
- simulated observation and template with injected tellurics, overplotted with clean simulated spectra, which would also demonstrate that due to BC, lines will pull on each other.

2.3. Extraction of Radial Velocities

Introduction to Doppler code and relevant adoption part for the simulation.

How the end product RVs are computed (vanking, or other weighting schemes).

3. RESULTS

3.1. Effects of Telluric Lines on RV Precision

- telluric free simulations to demonstrate RV precision, vs. year, vs. BC

¹ Based on observations observations obtained at the Keck Observatory, which is operated by the University of California. The Keck Observatory was made possible by the generous financial support of the W. M. Keck Foundation.

² Department of Astronomy and Astrophysics, 525 Davey Laboratory, The Pennsylvania State University, University Park, PA 16802, USA; Send correspondence to xxw131@psu.edu and itwright@astro.psu.edu

jtwright@astro.psu.edu \$^3\$ Center for Exoplanets and Habitable Worlds, 525 Davey Laboratory, The Pennsylvania State University, University Park, PA 16802, USA

⁴ Hawaii, USA

 $^5\,\mathrm{Department}$ of Astronomy, University of California, Berkeley, CA 94720, USA

- telluric injected simulations to demonstrate effects, vs. vear, vs. BC
- RV vs. BC for chunks with and without telluric lines before and after injection

3.2. Remedies and Their Effectiveness

ZZZ Plots:

- demonstration of effectiveness and ineffectiveness of double masking.
- demonstration of effectiveness of a clean template and forward modeling tellurics.
- demonstration of precision required for modeling tellurics, and maximum tolerance for template 'cleaness'.

4. SUMMARY AND FUTURE WORK

Work on actual observations!

The authors thank John A. Johnson for providing a copy of his Doppler code and his help with our incorporation of the code into the HET pipeline. The authors also thank Debra Fischer for her assistance in this regard.

This work was partially supported by funding from the Center for Exoplanets and Habitable Worlds, which is supported by the Pennsylvania State University, the Eberly College of Science, and the Pennsylvania Space Grant Consortium.

The authors appreciate the significant Keck observing time and associated funding support from NASA for the study of long period planets and mulitplanet systems. J.T.W. and S.X.W. acknowledge support from NASA Origins of Solar Systems grant NNX10AI52G.

The work herein is based on observations obtained at the W. M. Keck Observatory, which is operated jointly by the University of California and the California Institute of Technology. The Keck Observatory was made possible by the generous financial support of the W.M. Keck Foundation. We wish to recognize and acknowledge the very significant cultural role and reverence that the summit of Mauna Kea has always had within the indigenous Hawaiian community. We are most fortunate to have the opportunity to conduct observations from this mountain.

The Hobby-Eberly Telescope is a joint project of the University of Texas at Austin, the Pennsylvania State University, Stanford University, Ludwig Maximillians Universität München, and Georg August Universität Göttingen. The HET is named in honor of its principal benefactors, William P. Hobby and Robert E. Eberly.

This work has made use NASA' s Astrophysics Data

System Bibliographic Services.

REFERENCES

Artigau, É., Astudillo-Defru, N., Delfosse, X., et al. 2014, in Society of Photo-Optical Instrumentation Engineers (SPIE) Conference Series, Vol. 9149, Society of Photo-Optical Instrumentation Engineers (SPIE) Conference Series, 5 Cunha, D., Santos, N. C., Figueira, P., et al. 2014, A&A, 568, A35

Rothman, L. S., Gordon, I. E., Babikov, Y., et al. 2013, J. Quant. Spec. Radiat. Transf., 130, 4