

# Exam August 2018

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## Problem 1: Compute PSD from data using Pisarenko (sinusoid with additive white noise)

Here's a sequence of data from a measurement.

$$\{-11, 21, 18, 62, 34, -5, -14, -64, -49, -35, -1, 29, 37, 49, 32, 12\}$$

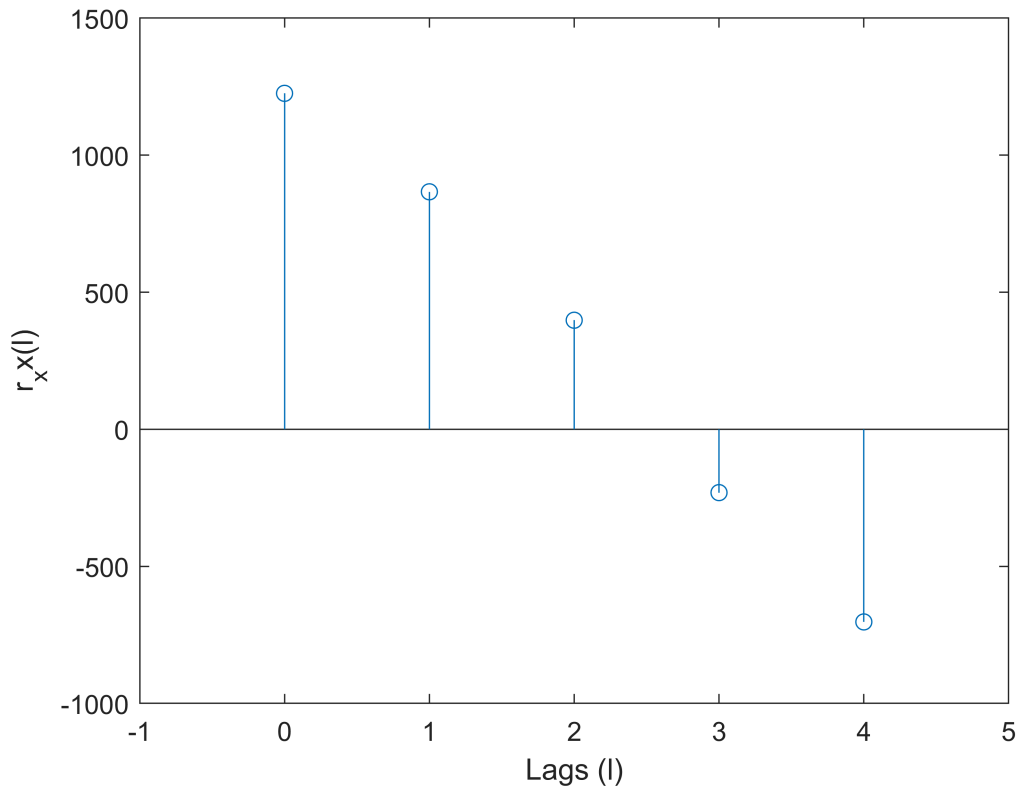
### [✓] 1) Compute and plot the autocorrelation for lags 0 to 4

```
x = [-11, 21, 18, 62, 34, -5, -14, -64, -49, -35, -1, 29, 37, 49, 32, 12];  
  
% Estimate autocorrelation using data  
[r_xx, e11] = xcorr(x, 'biased');  
  
% Print out the lags 0 to 4  
mid = floor(numel(e11)/2)+1;
```

```
r_xx(mid:mid+4)
```

```
ans = 1x5  
103 ×  
1.2256 0.8664 0.3979 -0.2310 -0.7027
```

```
% Plot the results  
stem(0:4, r_xx(mid:mid+4))  
xlim([-1, 5])  
xlabel('Lags (l)')  
ylabel('r_x(l)')
```



## [✓] 2) Compute the power spectral density assuming a Pisarenko model.

Assume that the signal is a sinusoid in additive white Gaussian noise.

The Pisarenko method can be used to recover the sinusoidal frequencies of a corrupted signal  $x(n)$  given two assumptions:

- The signal  $x(n)$  consists of  $p$  sinusoids that has been corrupted by white noise.
- The autocorrelation matrix of size  $(p + 1) \times (p + 1)$  is known or can be estimated from data

The Pisarenko method consists of the following steps:

**Step 1:** Compute the autocorrelation matrix  $\mathbf{R}_{xx}$

**Step 2:** Find the eigenvector corresponding to the smallest minimum eigenvalue. The elements of this eigenvector is the parameters of the ARMA(2p, 2p) model

**Step 3:** Find the frequencies  $\{f_i\}$  of the sinusoids. This can be done by computing the roots of the polynomial  $A(z)$  in Eq. (14.5.4) in the book. This polynomial has  $2p$  poles on the unit circle which correspond to the frequencies of the system.

$$A(z) = 1 + \sum_{m=1}^{2p} a_m z^{-m} \quad (14.5.4)$$

**Step 4:** Solve Eq. (14.5.11) for the signal powers  $\{P_i\}$

$$\begin{bmatrix} \cos 2\pi f_1 & \cos 2\pi f_2 & \cdots & \cos 2\pi f_p \\ \cos 4\pi f_1 & \cos 4\pi f_2 & \cdots & \cos 4\pi f_p \\ \vdots & \vdots & & \vdots \\ \cos 2\pi p f_1 & \cos 2\pi p f_2 & \cdots & \cos 2\pi p f_p \end{bmatrix} \begin{bmatrix} P_1 \\ P_2 \\ \vdots \\ P_p \end{bmatrix} = \begin{bmatrix} \gamma_{yy}(1) \\ \gamma_{yy}(2) \\ \vdots \\ \gamma_{yy}(p) \end{bmatrix} \quad (14.5.11)$$

where

- $\gamma_{yy}(1), \gamma_{yy}(2), \dots, \gamma_{yy}(p)$  are the estimated autocorrelation values
- $P_i = \frac{A_i^2}{2}$  is the average power of the  $i$ th sinusoid and  $A_i$  is the corresponding amplitude

**Step 5:** Estimate the amplitude  $A_i = \sqrt{2P_i}$

These steps are coded up in the `pisarenko()` function (see at end of this document):

```
[F, A, P, lambda_min] = pisarenko(r_xx(mid:mid+4), 1)
```

```
F = 0.0938
A = 45.6585
P = 1.0423e+03
lambda_min = 183.2154
```

We can describe the signal as:

$$x(n) = 45.6585 \cos(2\pi \cdot 0.0938n) + w(n)$$

where  $w(n)$  is white noise with variance  $\sigma_w^2 = 183.2154$

To compute the power spectral density, we need to find the autocorrelation function of the sinusoid.

In ADSI Problem 4.4, we found that the autocorrelation of a real sinusoid given

by  $y(n) = A \cos(\omega n + \phi)$  where  $A$  and  $\omega$  are real constants and  $\phi$  is a random variable with  $\phi \sim U(0, 2\pi)$  is:

$$r_{yy}(\ell) = \frac{A^2}{2} \cos(\omega \ell)$$

The autocorrelation function of white noise with variance  $\sigma_w^2$  is given by:

$$r_{ww}(\ell) = \sigma_w^2 \delta(\ell)$$

Combining these results, the general autocorrelation function of our signal is:

$$r_{xx}(\ell) = \frac{A^2}{2} \cos(\omega \ell) + \sigma_w^2 \delta(\ell)$$

The power spectral density is the Fourier transform the autocorrelation function which is given by:

$$S(\omega) = \frac{A^2}{2} \pi [\delta(\omega - \omega_0) + \delta(\omega + \omega_0)] + \sigma_w^2$$

A, F, lambda\_min

```
A = 45.6585  
F = 0.0938  
lambda_min = 183.2154
```

Using the values found via the Pisarenko, the PSD is:

$$S(\omega) = \frac{(45.6585)^2}{2} \pi [\delta(\omega - 2\pi \cdot 0.0938) + \delta(\omega + 2\pi \cdot 0.0938)] + 183.2154$$

### [✓] 3. Discuss whether the Pisarenko model can be considered appropriate for the given data.

We can describe the signal as:

$$x(n) = 45.6585 \cos(2\pi \cdot 0.0938n) + w(n)$$

where  $w(n)$  is white noise with variance  $\sigma_w^2 = 183.2154$

The general autocorrelation function of that signal is:

$$r_{xx}(\ell) = \frac{A^2}{2} \cos(\omega \ell) + \sigma_w^2 \delta(\ell)$$

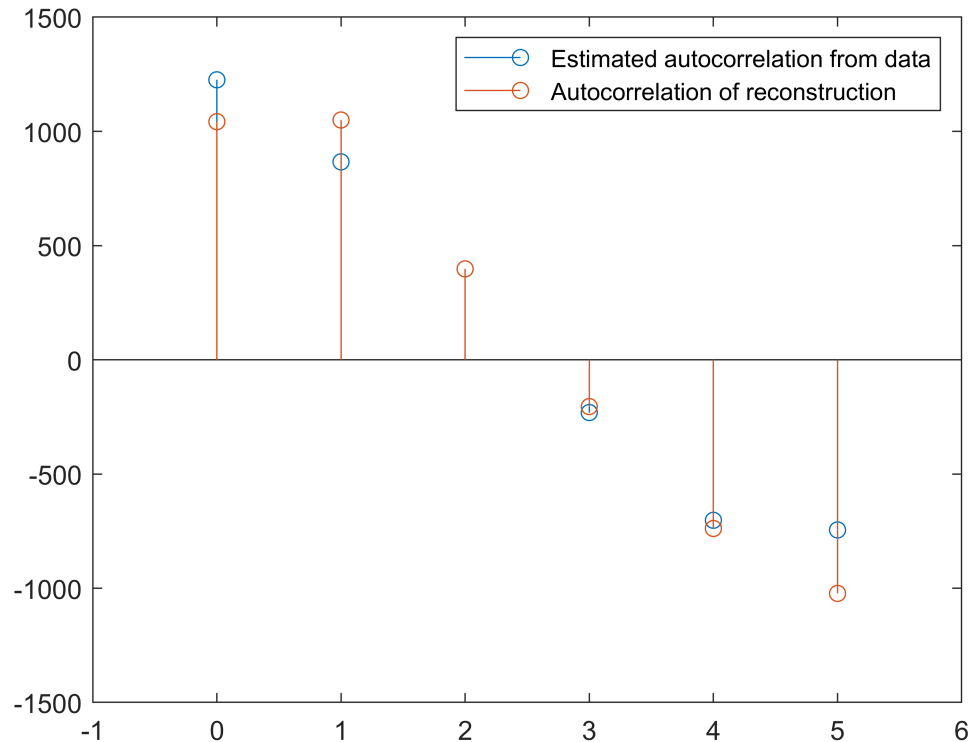
Considering the estimated autocorrelation of the data and the autocorrelation of the reconstruction, we can say that the Pisarenko model is appropriate.

```
% Plot the autocorrelation function of the sinusoid  
ell = 0:5;  
r_ww = lambda_min * (ell == 1);  
r_xx_p = (A^2/2) * cos(2*pi*F*ell) + r_ww;  
stem(ell, r_xx(mid:mid+max(ell)))
```

```

hold on;
stem(e11, r_xx_p)
legend('Estimated autocorrelation from data', 'Autocorrelation of reconstruction')
xlim([min(e11)-1, max(e11)+1])
hold off;

```



In addition, plotting the limited amount of samples, we see a sinusoidal shape. Since Pisarenko model assumes sinusoidal signal that has been corrupted by white noise then it can be appropriate model.

## Problem 2: True/False Questions

For the statements given below, state whether they are true or false and justify your answer for each statement.

[✓] 1) Is a filter invertible if the poles of the system function lies within the unit circle?

A filter with a rational system function  $H(z) = B(z)/A(z)$  is invertible if the poles of the system function lies within the unit circle.

Answer: FALSE.

A filter  $H(z) = \frac{B(z)}{A(z)}$  is said to be stable and causal if all the poles of  $H(z)$  are inside the unit circle.

If the inverse filter  $H_{\text{inv}}(z) = \frac{A(z)}{B(z)}$  has to be stable and causal, then all the poles of  $H_{\text{inv}}(z)$  must be inside the unit circle or equivalently all zeros of  $H(z)$  must be inside the unit circle.

So if we want the inverse filter to be stable, then the answer is false. Typically, we want an inversable filter to be minimum-phase which means that the filter *and* its inverse must be causal and stable.

### [✓] 2) If a signal is scaled by 2, does its autocorrelation also scale by 2?

If a signal  $x(n)$  is scaled to become  $y(n) = 2x(n)$ , the autocorrelation is scaled similarly and  $r_y(l) = 2r_x(l)$ .

Answer: FALSE.

The autocorrelation of a signal  $x(n)$  is defined as:

$$r_x(\ell) = E[x(n)x(n - \ell)]$$

We can compute the autocorrelation of signal  $y(n) = 2x(n)$  as follows:

$$r_y(\ell) = E[y(n)y(n - \ell)]$$

$$r_y(\ell) = E[2x(n)2x(n - \ell)]$$

$$r_y(\ell) = 4 \cdot E[x(n)x(n - \ell)]$$

$$r_y(\ell) = 4r_x(\ell)$$

[?] 3) Under stationary conditions, a least mean square (LMS) adaptive filter with proper choice of the step-size will converge to the Wiener-Hopf solution.

## Problem 3: Matched Filters

A deterministic signal is given by

$n$	$s(n)$
0	1
1	-1
2	1
3	-1

The signal is distorted by additive low frequency noise with autocorrelation  $r_v(l) = 0.8^{|l|}$ .

```
clear variables;
```

**[✓] 1) Design a matched filter to improve the signal to noise ratio and comment on the improvement.**

The impulse response of the matched filter is given by:

$$\mathbf{h} = \kappa \mathbf{R}_v^{-1} \mathbf{s}. \quad (14.97)$$

where  $\mathbf{R}_v$  is autocorrelation matrix of noise and  $\kappa$  is the normalisation factor.

Although the maximum SNR can be obtained by any choice of constant  $\kappa$ , we choose the constant by requiring that:

- (a)  $\mathbf{h}^T \mathbf{s} = 1$ , which yields  $\kappa = 1/\mathbf{s}^T \mathbf{R}_v^{-1} \mathbf{s}$
- (b)  $E(v_0^2[n]) = \mathbf{h}^T \mathbf{R}_v \mathbf{h} = 1$ , which yields  $\kappa = 1/\sqrt{\mathbf{s}^T \mathbf{R}_v^{-1} \mathbf{s}}$ .

```
s = [1, -1, 1, -1]';
p = numel(s); % Signal length

% The autocorrelation matrix must be p x p since
% its inverse is multiplied by a p-tap signal s(n)
ell = 0:p-1;
r_vv = 0.8.^ell;
R_vv = toeplitz(r_vv);

% Compute normalisation factor (b)
k = 1/sqrt(s'*(R_vv\s)); % Same as 1/sqrt(s'*inv(R_vv)*s)

% Compute the filter
h = k*(R_vv\s); % Same as k*inv(R_vv)*s

% Print the matched filter coefficients
h
```

```
h = 4x1
    0.9449
```

```
-1.7008
1.7008
-0.9449
```

The optimum SNR is given by:

$$\text{SNR}_o = a^2 \tilde{s}^T \tilde{s} = a^2 s^T R_v^{-1} s. \quad (14.98)$$

Assuming the attenuation factor  $a = 1$ :

```
a = 1;
SNR = a^2 * s' * (R_vv\s)
```

```
SNR = 28.0000
```

## [✓] 2) Can the signal to noise ratio be improved by using more than two blocks?

2. The  $s(n)$  signal consists of two blocks each containing 1 and -1. Can the signal to noise ratio be improved by using more than two blocks?

Yes, the SNR can be improved. With three blocks, the SNR is increased to 46:

```
s = [1, -1, 1, -1, 1, -1]';
p = numel(s);
ell = 0:p-1;
r_vv = 0.8.^ell;
R_vv = toeplitz(r_vv);
k = 1/sqrt(s'*(R_vv\s));
h = k*(R_vv\s);
```

```
a = 1;
SNR = a^2 * s' * (R_vv\s)
```

```
SNR = 46.0000
```

## Problem 4: Wiener Filtering

A system given by  $H(z) = 4 + z^{-1}$  is excited by unit variance white Gaussian noise  $w(n)$  to give the signal  $s(n)$ .

### [✓] 1) Determine the autocorrelation function, $r_s(l)$



The output signal of the system is given as:

$$s(n) = 4w(n) + w(n-1)$$

The autocorrelation of this signal is:

$$\begin{aligned} r_s(\ell) &= E[s(n)s(n-\ell)] \\ &= E[(4x(n) + x(n-1))(4x(n-\ell) + x(n-\ell-1))] \\ &= E[4x(n)4x(n-\ell)] + E[4x(n)x(n-\ell-1)] + E[x(n-1)4x(n-\ell)] + E[x(n-1)x(n-\ell-1)] \\ &= 16E[x(n)x(n-\ell)] + 4E[x(n)x(n-\ell-1)] + 4E[x(n-1)x(n-\ell)] + E[x(n-1)x(n-\ell-1)] \end{aligned}$$

```
syms n l w(n)
expand((4*w(n) + w(n-1)) * (4*w(n-l) + w(n-l-1)))
```

$$\text{ans} = 16 w(n-l) w(n) + 4 w(n-l-1) w(n) + 4 w(n-l) w(n-1) + w(n-l-1) w(n-1)$$

$$= 16r_w(\ell) + 4r_w(\ell-1) + 4r_w(\ell+1) + r_w(\ell)$$

Since  $r_w(\ell-1) = r_w(\ell+1)$

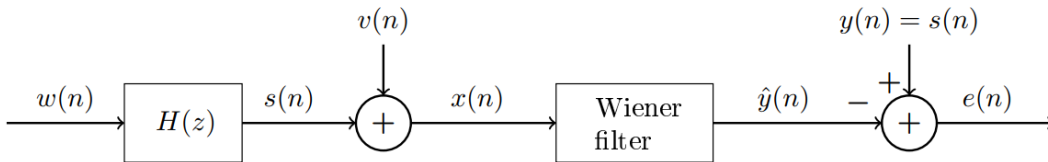
$$r_s(\ell) = 17r_w(\ell) + 8r_w(\ell-1)$$

Since the autocorrelation of unit variance white noise is  $r_{xx}(\ell) = \sigma_x^2 \delta(\ell) = \delta(\ell)$ :

$$r_s(\ell) = 17\delta(\ell) + 8\delta(\ell-1)$$

## [✓] 2) Solve the Wiener-Hopf Equation

As shown in the block diagram below, the signal is corrupted by additive white Gaussian noise with  $\sigma_v^2 = 3$  giving  $x(n) = s(n) + v(n)$ . It is desired to recover the signal with a 3-tap Wiener filter.



- Determine the  $3 \times 3$  autocorrelation matrix  $R_x$  and the  $3 \times 1$  crosscorrelation vector  $\mathbf{g} = E[\mathbf{x}(n)y(n)]$  and use these to solve the Wiener-Hopf equation.

We know that  $x(n) = s(n) + v(n)$  and  $y(n) = s(n) = 4w(n) + w(n-1)$

Assuming  $w(n)$  and  $v(n)$  are uncorrelated, the autocorrelation for  $x(n)$ :

$$\begin{aligned}
r_x(\ell) &= E[x(n)x(n-\ell)] \\
&= E[(s(n) + v(n))(s(n-\ell) + v(n-\ell))] \\
&= E[s(n)s(n-\ell) + s(n)v(n-\ell) + v(n)s(n-\ell) + v(n)v(n-\ell)] \\
&= E[s(n)s(n-\ell)] + E[s(n)v(n-\ell)] + E[v(n)s(n-\ell)] + E[v(n)v(n-\ell)] \\
&= r_s(\ell) + 2r_{sv}(\ell) + r_v(\ell) \\
&= r_s(\ell) + r_v(\ell) \text{ (because } w(n) \text{ and } v(n) \text{ are uncorrelated so } r_{sv}(\ell) = 0) \\
&= 17\delta(\ell) + 8\delta(\ell-1) + 3\delta(\ell) \text{ -- } r_v(\ell) = 3\delta(\ell) \text{ since } \sigma_v^2 = 3 \\
&= 20\delta(\ell) + 8\delta(\ell-1)
\end{aligned}$$

We know that  $x(n) = s(n) + v(n)$  and  $y(n) = s(n) = 4w(n) + w(n-1)$ .

Compute the cross-correlation vector:

$$\begin{aligned}
r_{xy}(\ell) &= E[x(n)y(n-\ell)] \\
&= E[(s(n) + v(n))(4w(n-\ell) + w(n-\ell-1))] \\
&= 4E[s(n)w(n-\ell)] + E[s(n)w(n-\ell-1)] + 4E[v(n)w(n-\ell)] + E[v(n)w(n-\ell-1)] \\
&= 4r_{sw}(\ell) + r_{sw}(\ell-1) + 4r_{vw}(\ell) + r_{vw}(\ell-1)
\end{aligned}$$

We assume that  $w(n)$  and  $v(n)$  are uncorrelated so  $r_{vw}(\ell) = 0$ :

$$= 4r_{sw}(\ell) + r_{sw}(\ell-1) + 0 + 0$$

According to Eq. 13.100, the cross-correlation  $r_{sw}(\ell) = \sigma_w^2 h(\ell) = h(\ell)$  since  $\sigma_w^2 = 1$ . We know that the impulse response  $h(\ell) = [4, 1]$ :

$$= 4h(\ell) + h(\ell-1)$$

The third order Wiener filter for estimating the signal  $y(n)$  is given by:

$$\hat{y}(n) = h_1 x(n) + h_2 x(n-1) + h_3 x(n-2)$$

The optimum Wiener filter to estimate a random process is given by Eq. 14.109:

$$\mathbf{h}_0 = \mathbf{R}_x^{-1} \mathbf{g}, \quad (14.109)$$

where  $\mathbf{R}_x$  is the correlation matrix of a random vector  $\mathbf{x}$  and  $\mathbf{g}$  is the cross-correlation vector between  $\mathbf{x}$  and  $\mathbf{y}$

```
r_xx = [20, 8, 0];
R_xx = toeplitz(r_xx)
```

```
R_xx = 3x3
    20     8     0
     8    20     8
     0     8    20
```

```
g = [4, 1, 0]'
```

```
g = 3x1
     4
     1
     0
```

```
h_opt = R_xx\g
```

```
h_opt = 3x1
    0.2176
   -0.0441
    0.0176
```

The optimal Wiener filter is given by:

$$H(z) = 0.2471 - 0.1176z^{-1} + 0.0471z^{-2}$$

**[?] 3) Discuss whether 3 taps is an optimum choice for this problem?**

```
r_xx = [20, 8];
R_xx = toeplitz(r_xx);
g = [4, 1]';
h_opt = R_xx\g
```

```
h_opt = 2x1
    0.2143
   -0.0357
```

## Problem 5: Probability Density Function

Consider the following probability density function

$$f_X(x) = \begin{cases} \alpha x^2 + \frac{1}{4} & \text{for } -1 < x < 1 \\ 0 & \text{elsewhere} \end{cases}$$

where  $\alpha$  is a real number.

## [✓] 1) Find a valid probability density function

1. Determine  $\alpha$  so that  $f_X(x)$  is a valid probability density function.

A valid probability density function is given by:

$$\int_{-\infty}^{\infty} f_X(x) dx = 1$$

We need to compute:

$$\int_{-1}^1 \alpha x^2 + \frac{1}{4} dx$$

For convenience and to avoid silly mistakes, use MATLAB:

```
syms x a
int(a*x^2 + 1/4, x, -1, 1)
```

ans =

$$\frac{2a}{3} + \frac{1}{2}$$

Solve the equation for  $a$  in MATLAB:

```
solve(int(a*x^2 + 1/4, x, -1, 1) - 1)
```

ans =

$$\frac{3}{4}$$

Let us check the results:

```
int(3/4 * x^2 + 1/4, x, -1, 1)
```

ans = 1

For  $f_X(x)$  to be a valid probability density function,  $a$  must be:

$$a = \frac{3}{4}$$

## [✓] 2) Compute the probability given a probability density function?

2. What is the probability that  $1/4 < X \leq 3/4$ ?

```
p = int(3/4 * x^2 + 1/4, x, 1/4, 3/4)
```

```
p =  

$$\frac{29}{128}$$

```

```
vpa(p)
```

```
ans = 0.2265625
```

The answer is:

$$\Pr\left(\frac{1}{4} < X \leq \frac{3}{4}\right) = \frac{29}{128} = 0.2265625$$

### [✓] 3) Compute the expected value a function:

Let a function be given by  $g(x) = e^{-|x|}$ .

3. Compute  $E[g(x)]$ .

We need to compute:

$$E[g(x)] = \int_{-\infty}^{\infty} g(x)f_X(x) dx = \int_{-1}^1 e^{-|x|} \cdot \frac{3}{4}x^2 + \frac{1}{4} dx$$

```
syms x  
f = exp(-abs(x)) * 3/4 * x^2 + 1/4
```

```
f =  

$$\frac{3x^2e^{-|x|}}{4} + \frac{1}{4}$$

```

```
int(f , x, -1, 1 )
```

```
ans =  

$$\frac{7}{2} - \frac{15e^{-1}}{2}$$

```

```
vpa(int(f , x, -1, 1 ))
```

```
ans = 0.74090419121418258803357172378904
```

The answer is:

$$E[g(x)] \approx 0.74$$

## Functions

```
function [F, A, P, lambda_min]=pisarenko(r_xx, p)
% Estimates the frequencies and amplitudes using the Pisarenko method
% r_xx: autocorrelation sequence starting from zero.
% The length of ACRS must be at least 2p+1.
% p: assumed number of sinusoids in the signal
% F: normalised frequencies of the sinuoids
% A: amplitudes of the sinuoids
if numel(r_xx) < 2*p+1
    error(strcat('The length of ACRS must be at least ', int2str(2*p+1)));
end

% Compute the autocorrelation matrix
R_xx = toeplitz(r_xx(1:2*p+1));

% Perform the eigendecomposition
[eigvecs, eigvals] = eigs(R_xx, size(R_xx, 1), 'smallestreal');

eigvals = diag(eigvals);
lambda_min = eigvals(1);

% Find the eigenvector 'a' corresponding to the smallest eigenvalue.
% The function eigs() sorts eigenvectors, so just pick the first column.
a = eigvecs(:, 1);

% Ensure that a_0 = 1 (this is by definition)
a = a / a(1);

% The elements of this eigenvector corresponds to the parameters
% of an ARMA(2p, 2p) model: a_0, a_1, ..., a_2p where p: number of sinusoids
% The polynomial A(z) in (14.5.4) has 2p poles on the unit circle.
% Obtain the poles by finding the roots of the system.
z = roots(a);

% Estimate frequencies
F = zeros(p, 1);
for i = 1:p
    % The poles come in pairs. Each pair is complex conjugate
    % of one another. Only use one of them and find the absolute value.
    z_i = z(2*i);
    F(i) = abs(angle(z_i)) / (2*pi);
end

% Build the matrix of cosines
```

```

C = zeros(p);
for i = 1:p
    for j = 1:p
        C(i, j) = cos(i * 2*pi * F(j));
    end
end

% Start the ACRS from the second element according to equation (14.5.11)
gamma = r_xx(2:p+1);

% Solve equation (14.5.11) for P
P = C\gamma; % Same as `inv(C)*gamma` but faster and more accurate

% Since  $P=A^2/2$ , we can compute the amplitude  $A=\sqrt{2*P}$ 
A = sqrt(2 * P);
end

```