

Satellite Orbit Tracker

Orbital calculations and 3D visualization
entirely in the browser with JavaScript.

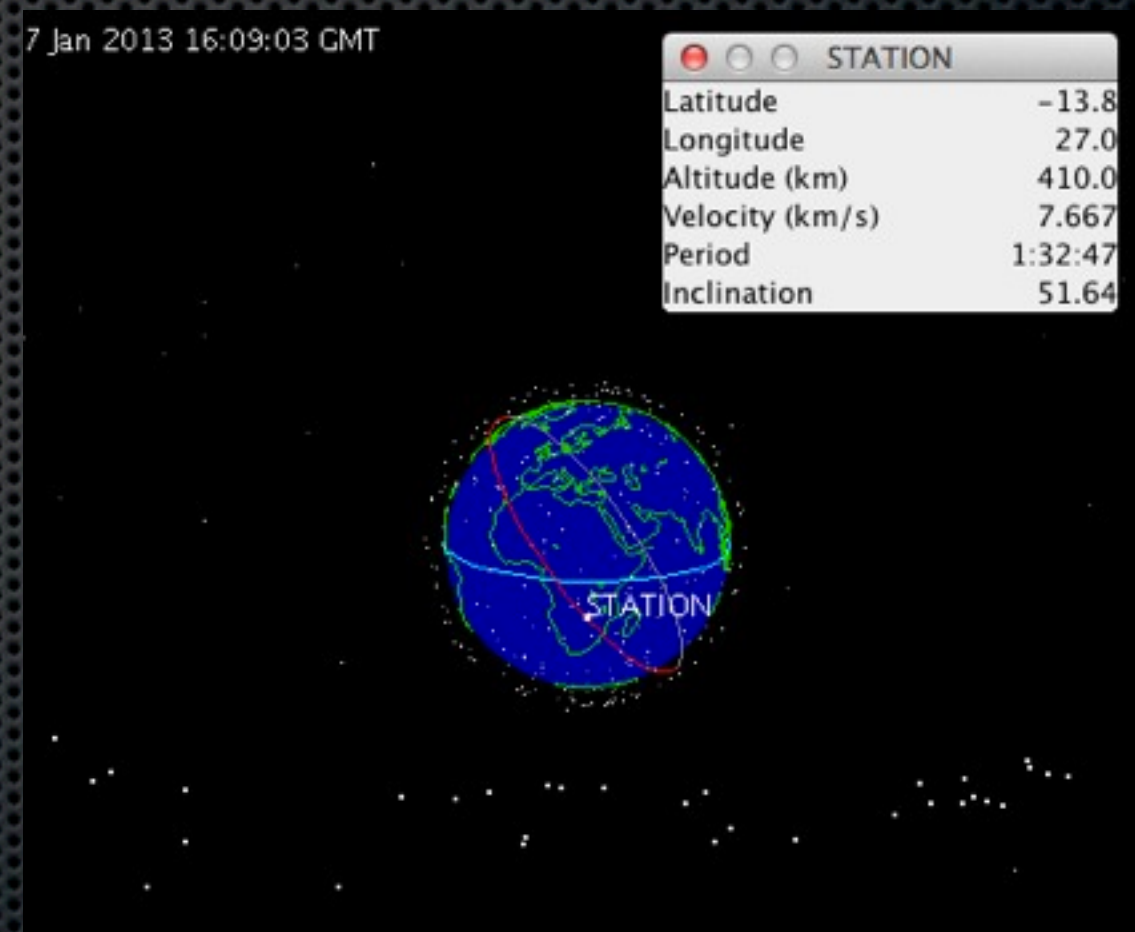
A modern approach to interactive web applications.

2013-01-08 Presentation to HITSS AppDev

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In the Beginning: JTrack3D

- ✧ > 10 years old applet at MSFC
- ✧ Migrated from MSFC to HQ
- ✧ Resurrected by decompiling
- ✧ Java: not working well now
- ✧ Educators disappointed

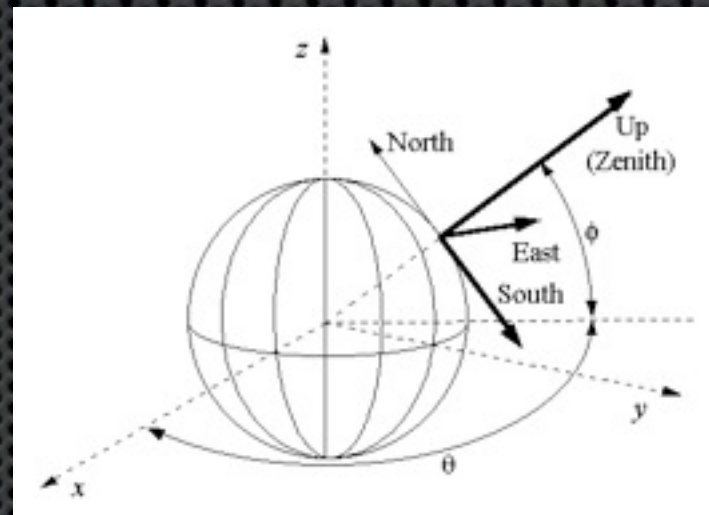


Tuesday, January 8, 13

Java no longer installed by Apple. 64-bit Java from Oracle won't run in Chrome as it's 32-bit; have to use FireFox, Safari. When we migrated Science.nasa.gov from MSFC to HQ's NASAScience.nasa.gov we were initially told not to bother with JTrack3D due to the delay it would have caused.

SGP: Orbit Calculations

- ✦ Simplified General Perturbation model
- ✦ Developed in 1959
- ✦ FORTRAN in 1988
- ✦ MATLAB, others
- ✦ Miura's PhD thesis in 2009
- ✦ TLEs used by NORAD, NASA



Appendix C – TEME Coordinate System

This section describes the equations necessary to implement the nutation equations for the TEME approach. There are two approaches – using the GMST, and using the equation of the equinoxes. For sidereal time, GMST is needed. GMST is found using UT1. From McCarthy (1992:30)

$$\theta_{\text{GMST1982}} = 67,310,548.41^\circ + (876,600^h + 8,640,184.812 \text{ } 866^s) T_{\text{UT1}} + 0.093 \text{ } 104 T_{\text{UT1}}^2 - 6.2 \times 10^{-6} T_{\text{UT1}}^3 \quad (\text{C-1})$$

The transformation to ITRF is found using the polar motion (x_p, y_p) values and the GMST. Note that “PEF” implies the *pseudo-Earth-fixed* frame, where polar motion has not yet been applied (Vallado, 2004: 217).

$$\begin{aligned} [\mathbf{W}]_{\text{ITRF-PEF}} &= \text{ROT1}(y_p) \text{ROT2}(x_p) \\ \mathbf{r}_{\text{ITRF}} &= [\mathbf{W}]^T [\text{ROT3}(\theta_{\text{GMST1982}})]^T \mathbf{r}_{\text{TEME}} \\ \mathbf{v}_{\text{ITRF}} &= [\mathbf{W}]^T \left\{ [\text{ROT3}(\theta_{\text{GMST1982}})]^T \mathbf{v}_{\text{TEME}} + \dot{\theta}_p \times \mathbf{r}_{\text{PEF}} \right\} \end{aligned} \quad (\text{C-2})$$

If the equation of the equinox approach is taken, you must find the nutation parameters. The IAU-80 nutation uses so-called Delaunay variables and coefficients to calculate nutation in longitude ($\Delta\psi_{1980}$) and nutation in the obliquity of the ecliptic ($\Delta\epsilon_{1980}$). (McCarthy, 1992:32)

$$\begin{aligned} M_1 &= 134.962 \text{ } 981 \text{ } 39^\circ + 1,717,915,922.6330^\circ T_{\text{JT}} + 31.31 T_{\text{JT}}^2 + 0.064 T_{\text{JT}}^3 \\ M_0 &= 357.527 \text{ } 723 \text{ } 33^\circ + 129,596,581.2240^\circ T_{\text{JT}} - 0.5777 T_{\text{JT}}^2 + 0.012 T_{\text{JT}}^3 \\ \mu_1 &= 93.271 \text{ } 910 \text{ } 28^\circ + 1,739,527,263.1370^\circ T_{\text{JT}} - 13.257 T_{\text{JT}}^2 - 0.011 T_{\text{JT}}^3 \\ D_0 &= 297.850 \text{ } 363 \text{ } 06^\circ + 1,602,961,601.3280^\circ T_{\text{JT}} - 6.891 T_{\text{JT}}^2 + 0.019 T_{\text{JT}}^3 \\ \Omega_1 &= 125.044 \text{ } 522 \text{ } 22^\circ - 6,962,890.5390^\circ T_{\text{JT}} + 7.455 T_{\text{JT}}^2 + 0.008 T_{\text{JT}}^3 \end{aligned} \quad (\text{C-3})$$

The nutation parameters are then found using (McCarthy, 1992:33)

$$\begin{aligned} a_{p1} &= a_{s1} M_1 + a_{s2} M_0 + a_{s3} \mu_1 + a_{s4} D_0 + a_{s5} \Omega_1 \\ \Delta\psi &= \sum_{i=1}^{100} (A_{p1} + A_{p2} T_{\text{JT}}) \sin(a_{p1}) \quad \Delta\epsilon = \sum_{i=1}^{100} (A_{e1} + A_{e2} T_{\text{JT}}) \cos(a_{e1}) \end{aligned} \quad (\text{C-4})$$

Corrections to the nutation parameters ($\delta\Delta\psi_{1980}$ and $\delta\Delta\epsilon_{1980}$) supplied as Earth Orientation Parameters (EOP) from the IERS are simply added to the resulting values in Eq. 4 to provide compatibility with the newer IAU 2000 Resolutions (Kaplan, 2005). These corrections also include effects from Free Core Nutation (FCN) that correct errors in the IAU-76 precession and IAU-80 nutation. However for TEME, these corrections do not appear to be used. The nutation parameters let us find the true obliquity of the ecliptic, ϵ . (McCarthy, 1992:29–31)

$$\begin{aligned} \Delta\psi_{1980} &= \Delta\psi + \delta\Delta\psi_{1980} \quad \Delta\epsilon_{1980} = \Delta\epsilon + \delta\Delta\epsilon_{1980} \\ \bar{\epsilon} &= 84,381.448^\circ - 46.8150 T_{\text{JT}} - 0.000 \text{ } 59 T_{\text{JT}}^2 + 0.001 \text{ } 813 T_{\text{JT}}^3 \\ \epsilon &= \bar{\epsilon} + \Delta\epsilon_{1980} \end{aligned} \quad (\text{C-5})$$

The equation of the equinoxes (EQ_{eq1980}) can then be found. The last two terms in the EQ_{eq1980} are probably not included in AFSPC formulations. From McCarthy (1992:30)

$$\begin{aligned} EQ_{\text{eq1980}} &= \Delta\psi_{1980} \cos(\bar{\epsilon}) + 0.002 \text{ } 64^\circ \sin(\Omega_1) + 0.000 \text{ } 063 \sin(2\Omega_1) \\ \theta_{\text{GMST1982}} &= 67,310.54841^\circ + (876,600^h + 8,640,184.812 \text{ } 866^s) T_{\text{UT1}} + 0.093 \text{ } 104 T_{\text{UT1}}^2 - 6.2 \times 10^{-6} T_{\text{UT1}}^3 \\ \theta_{\text{GMST1982}} &= \theta_{\text{GMST1982}} + EQ_{\text{eq1980}} \end{aligned} \quad (\text{C-6})$$

ISS (ZARYA)

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1 25544U 98067A 12312.59429472 .00006293 00000-0 11296-3 0 6057
2 25544 51.6473 129.0708 0015313 255.0735 149.5553 15.51208285800293
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SGP: from a known point, velocity, etc, calculate forward in time to next position, velocity, etc; repeat. Math from: <http://celestrak.com/columns/v02n02/>, <http://celestrak.com/publications/AIAA/2006-6753/>. The SATRAK code used by JSC is SECRET due to NORAD restrictions.

Miura's thesis indicated the MATLAB implementation was the most accurate, so this is what I started with. TLE: Two line element set giving last verified position, velocity, time, other math attributes.

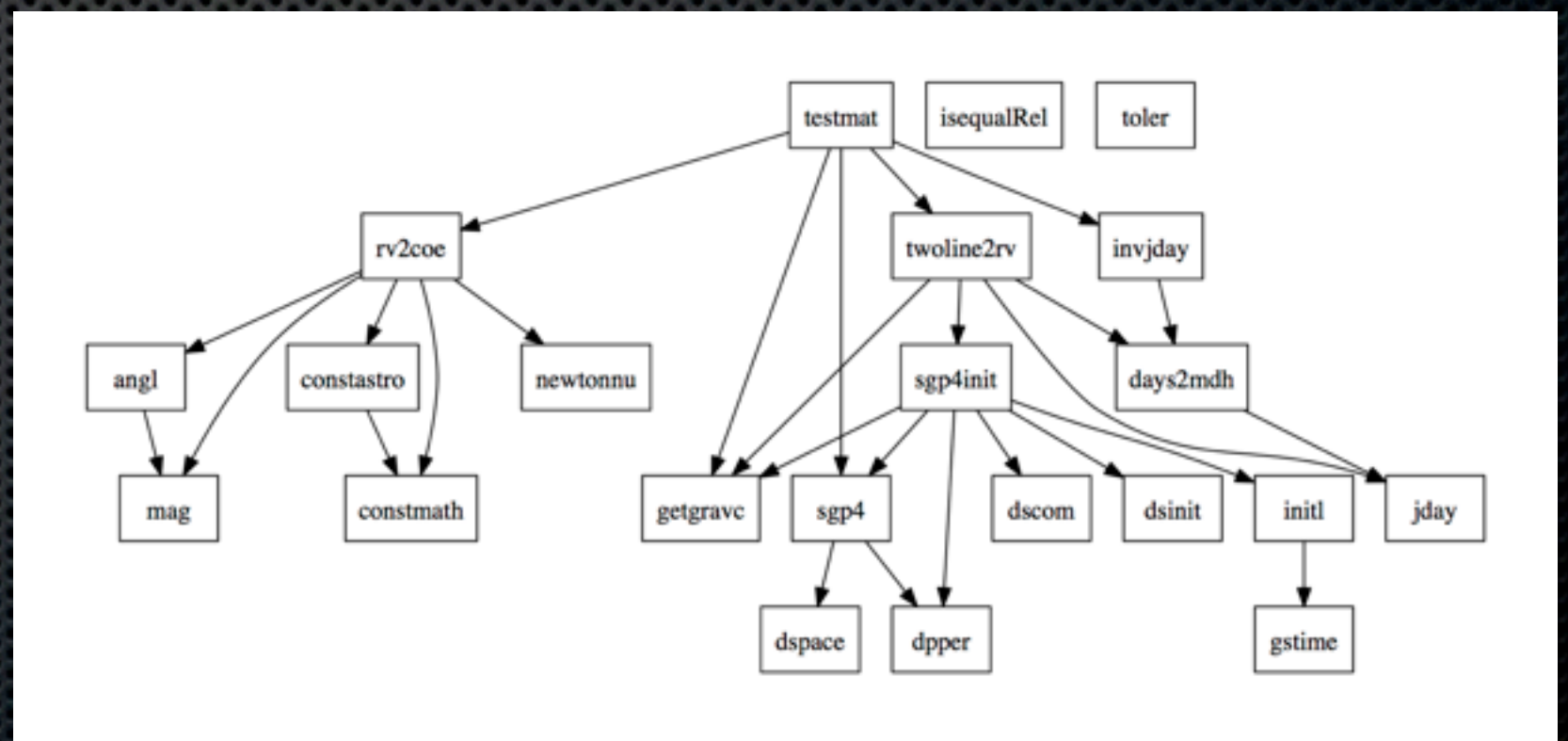
JavaScript is Powerful, Fast

- ✦ JavaScript is the language of the web
- ✦ Run entirely in browser, no servers needed
- ✦ Performance isn't everything, but matters here
- ✦ JavaScript is expressive, a real (powerful) language

Language	CPU time			Slower than	
	User	System	Total	C++	previous
C++ (<i>optimized with -O2</i>)	1,520	0,188	1,708	-	-
Java (<i>non-std lib</i>)	2,446	0,150	2,596	52%	52%
C++ (<i>not optimized</i>)	3,208	0,184	3,392	99%	31%
Javascript (nodejs)	4,068	0,544	4,612	170%	36%
Java	8,521	0,192	8,713	410%	150%
Python	27,886	0,168	28,054	1543%	108%
Perl	41,671	0,100	41,771	2346%	49%
PHP 5.3	94,622	0,364	94,986	5461%	127%

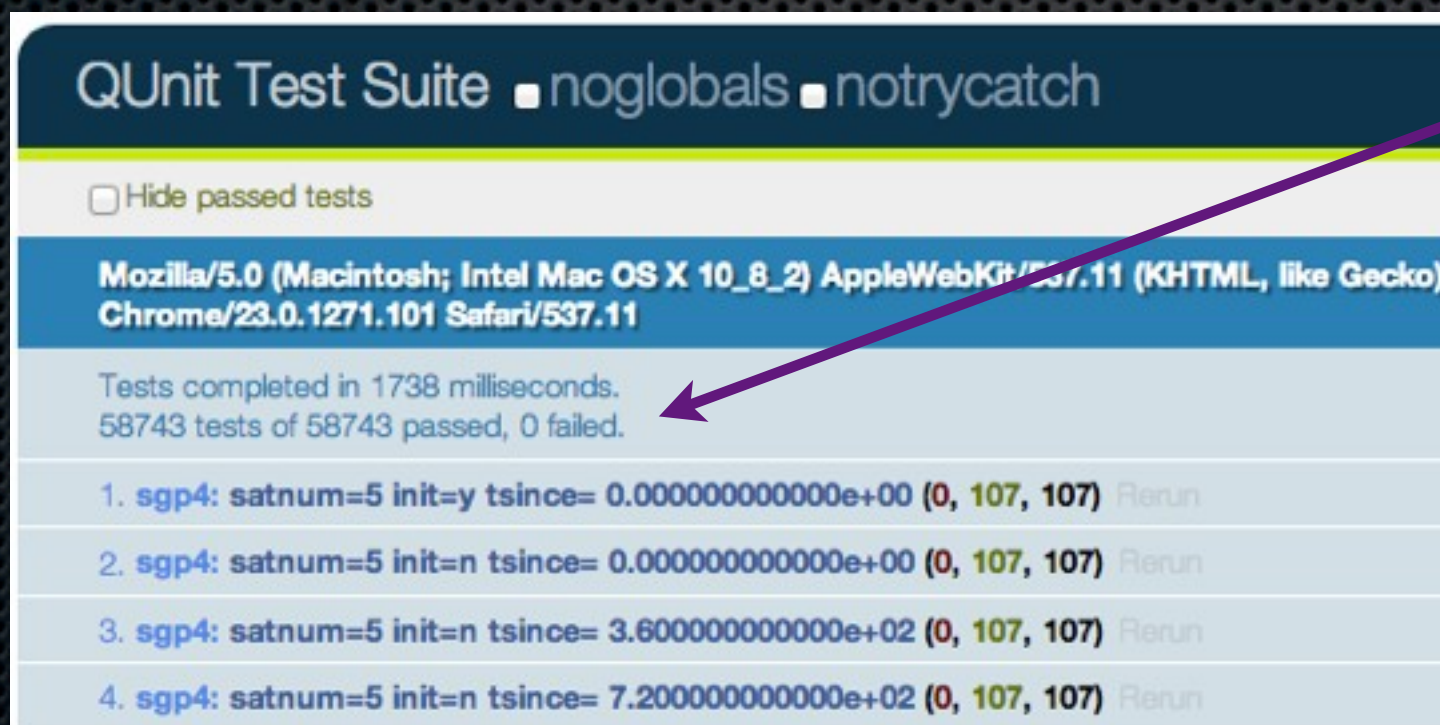
Translating MATLAB to JS

- ✧ Problem: I didn't know MATLAB
- ✧ Problem: I didn't know JavaScript
- ✧ Graph dependencies
- ✧ Translate literally



Tests: Only Way to Verify

- ✦ Create tests for MATLAB, capture results
- ✦ Recreate for JavaScript to verify results
- ✦ Single module had almost 60,000 tests
- ✦ Exercised all other modules
- ✦ I told you JavaScript was fast: 1.738 seconds



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That's 38,000 tests per second exercising all the math modules!

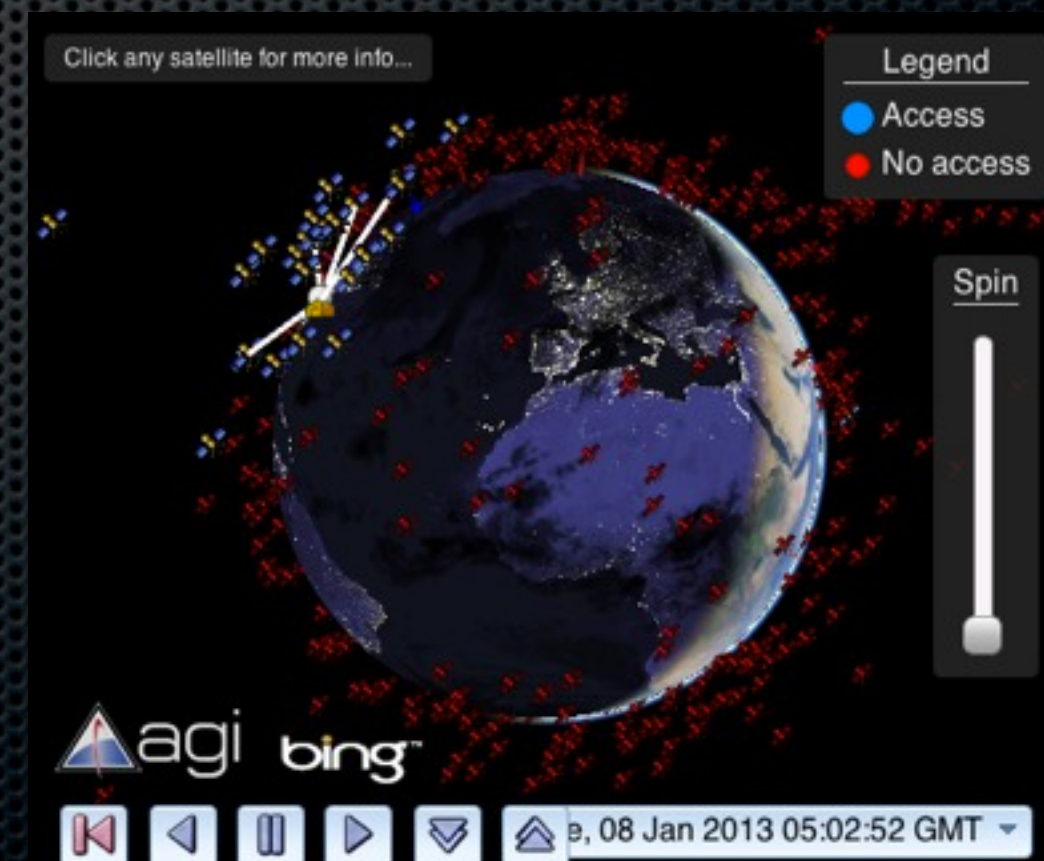
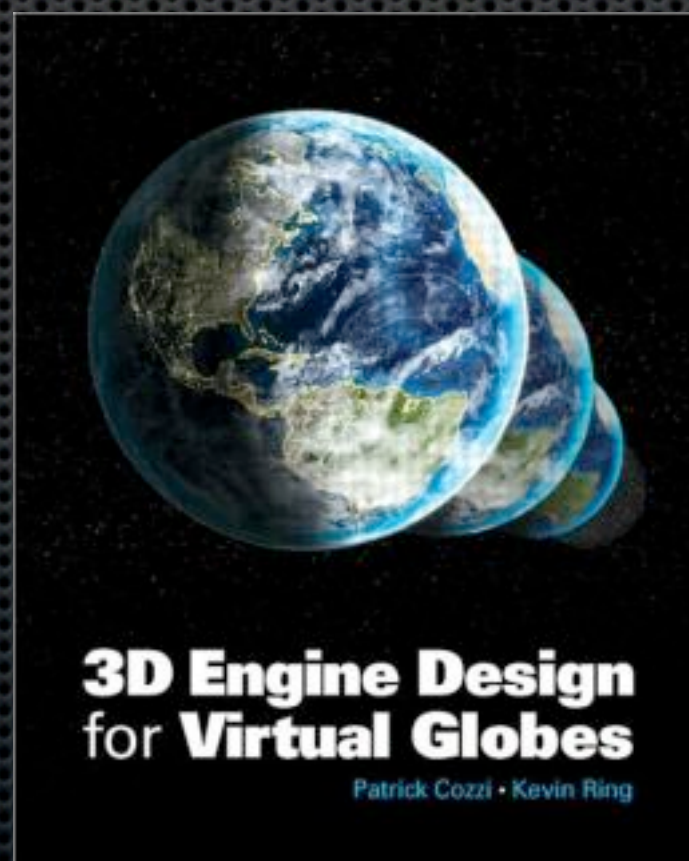
Visualization?

- ✦ How do we visualize the computed positions? Earth?
- ✦ Lots of libraries in various states of readiness
- ✦ I don't want to have to invent 3D visualization :-(
- ✦ I don't know how to invent 3D visualization!



Cesium

- ✦ A chance meeting at NASA Open Source Summit 2012
- ✦ AGI has lots of smart PhDs
- ✦ Open Source
- ✦ Pure JavaScript
- ✦ WebGL-based



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Check out AGI's mind-blowing "Lots of Satellites" live demo; the position data is calculated by servers in the cloud, streamed to the browser-based Cesium rendering engine. Code: <https://github.com/AnalyticalGraphicsInc/cesium>

WebGL-Capable Platforms

- ✦ Desktops supported now on most browsers
- ✦ Mobile coming along quickly

Desktop	Android	iOS
IE: with ChromeFrame FireFox: yes Chrome:yes Safari: if enabled	Default Browser: no FireFox: yes Chrome:?	No: disabled by Apple except for advertising partners :-(

SOT: Satellite Orbit Tracker

- ✧ Calculate with SGP
- ✧ Visualize with Cesium
- ✧ All JavaScript
- ✧ Entirely in Browser
- ✧ Danger: live demo

