

```
\newcommand{\lI}{\mathrm{l}}
\newcommand{\lIw}{\mathrm{i}\omega}
\newcommand{\lWmax}{\omega_{\mathrm{max}}}
\newcommand{\dd}{\mathrm{d}}
\newcommand{\tau_k}{\bar{\tau}_k}
\newcommand{\wk}{\bar{\omega}_k}
\newcommand{\vk}{\bar{\nu}_k}
\newcommand{\hatFmat}{\hat{\mathbf{F}}}
\newcommand{\Fmat}{\mathbf{F}}
```

虚時間グリーン関数に対するスパースモデリング入門 (1)

品岡寛 (埼玉大学)



自己紹介

- 専門: 量子多体理論、第一原理計算、幾何学的フラストレート磁性...
- 計算物理バックエンド (数値計算手法、ライブラリ)の開発に興味があります。
ALPS量子モンテカルロコード、スパースモデリング...
- 海外との連携: ウィーン工科大学、ミシガン大、フリブール大 (スイス)、ミュンヘン大、King's College London...

グループウェブサイト

経歴

1. 博士 (工学) 2009年3月@東大物工
2. ポスドク 2009年3月～2015年9月 (東大、産総研、チューリッヒ連邦工科大)
3. 埼玉大 2015年10月～

宣伝

- 学術変革領域B「量子古典融合アルゴリズムが拓く計算物質科学」(代表: 品岡)
 - 2023～2025年度
 - 計画研究班代表: 品岡、大久保、水上
 - スパースモデリング、テンソルネットワーク、動的平均場理論、密度汎関数理論、変分波動関数理論、量子情報・・・

近日、ポスドク・学生等の公募が出る予定！→詳しくはポスターで

- JST創発「2粒子レベルの量子埋め込み理論に基づく新規第一原理計算手法の開発と実証」2024年度から基本7年
 - RA (博士課程学生支援あり)

前提知識

- 虚時間形式グリーン関数の基礎
- Python or Juliaの基礎知識
 - 基本文法
 - 多次元配列
 - ...

何ができる？

虚時間・松原形式に基づく「数値」計算の高速・省メモリ化

超伝導転移温度の第一原理計算

T. Wang *et al.*, PRB 102, 134503
(2020), Nb

- メモリの使用量が40分の1に!
[松原周波数 4096点 → 103点]
- 計算速度が20倍に!

他の応用例



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背後にある技術

- 虚時間グリーン関数のコンパクトな中間表現基底
 - $G(\tau) = \sum_{l=0}^{L-1} U_l(\tau) g_l + \epsilon$
 - $L \propto \log \beta W$ (β : inverse temperature, W : band width)
 - $\epsilon \propto \exp(-aL)$ (ϵ : truncation error, $a > 0$)
- 虚時間・虚周波数におけるスパースメッシュ: # of points $\simeq L$.
- SparseIR.jl (Julia), sparse-ir (Python)

参考資料

- 固体物理 2021年6月 温度グリーン関数の情報圧縮に基づく高速量子多体計算法
- ↑の英語訳・加筆 + 新ライブラリsparse-irに更新
[H. Shinaoka et al., SciPost Phys. Lect. Notes 63 \(2022\)](#)
- sparse-ir tutorials (大量のサンプルコード)
<https://spm-lab.github.io/sparse-ir-tutorial/index.html>
IR基底の基礎、フーリエ変換、2次摂動、FLEX、DMFTなどなど
Python, Julia (Jupyter notebook) + Fortran

虚時間以外を圧縮したい！: Quantics tensor trains

- 一般の時空依存性 (実時間、波数依存性など)の圧縮
- 指数的に異なるエネルギー・長さスケール間の低エンタングルメント構造を仮定
- Julia実装 (ITensors.jlベース)

 QTT center

[arXiv:2210.12984](https://arxiv.org/abs/2210.12984) (to appear in PRX)

概要

- Part I
 - i. 虚時間グリーン関数の性質のまとめ
 - ii. 中間表現基底
 - iii. スパースサンプリング法

Imaginary-time Green's functions

Also known as Matsubara Green's functions:

$$G(\tau) = -\langle T_\tau A(\tau) B(0) \rangle,$$

where

- $A(\tau)$, $B(\tau)$ are operators in the Heisenberg picture ($A(\tau) = e^{\tau H} A e^{-\tau H}$).
- $\langle \dots \rangle = \text{Tr}(e^{-\beta H} \dots)$, where $\beta = 1/T$ ($k_B = 1$).

We use the Hamiltonian formalism throughout this lecture.

Imaginary-time Green's functions

$$G(\tau) = -\langle T_\tau A(\tau) B(0) \rangle$$

- A and B are fermionic operators $\rightarrow G(\tau) = -G(\tau + \beta)$
- A and B are bosonic operators $\rightarrow G(\tau) = G(\tau + \beta)$

In general, $G(\tau)$ has a discontinuity at $\tau = n\beta$ ($n \in \mathbb{N}$).

Imaginary-frequency (Matsubara) Green's functions

Matsubara Green's function:

$$G(i\omega) = \int_0^\beta d\tau e^{i\omega\tau} G(\tau).$$

From $G(\tau + \beta) = \mp G(\tau)$,

- $\omega = (2n + 1)T\pi$ (fermion)
- $\omega = 2nT\pi$ (boson)

$(n \in \mathbb{N})$

These discrete imaginary frequencies are denoted as Matsubara frequencies.

Spectral/Lehmann representation

$$G(z) = \int_{-\infty}^{+\infty} d\omega' \frac{\rho(\omega')}{z - \omega'},$$

where $\rho(\omega)$ is a spectral function.

- $z = i\omega \rightarrow$ Matsubara Green's function
- $z = \omega + i0^+ \rightarrow$ Retarded Green's function (not used in this lecture)

How Green's function look like in τ ?

Example (single pole): $\rho(\omega) = \delta(\omega - \omega_0)$, $\omega_0 > 0$

$$G(i\omega) = \frac{1}{i\omega - \omega_0}$$

$$G(\tau) = -\frac{e^{-\tau\omega_0}}{1 + e^{-\beta\omega_0}} \quad (0 < \tau < \beta)$$

At $\tau \approx 0$, $G(\tau) \propto e^{-\tau\omega_0}$.

For $\beta\omega_0 \gg 1$, coexisting two time scales : $1/\omega_0 \ll \beta$

How Green's function look like in Matsubara frequency space

Example (single pole): $\rho(\omega) = \delta(\omega - \omega_0)$, $\omega_0 > 0$

$$G(i\omega) = \frac{1}{i\omega - \omega_0}$$

$$G(\tau) = -\frac{e^{-\tau\omega_0}}{1 + e^{-\beta\omega_0}} \quad (0 < \tau < \beta)$$

At high frequencies $|\omega| \gg |\omega_0|$, $G(i\omega) \approx 1/(i\omega)$.

For $\beta\omega_0 \gg 1$, coexisting two energy scales: $\omega_0 \ll T = 1/\beta$

Difficulties in numerical simulations

If band width W and temperature T differ by orders of magnitudes as $\beta W \gg 1$:

- Slow power-law decay at high frequencies \rightarrow Large truncation errors
- Uniform dense mesh in τ requires a huge number of points $\propto \beta W$.

Example:

- Band width 10 eV, superconducting temperature 1 K ≈ 0.1 meV $\rightarrow \beta W = 10^5$.

We need a compact basis with exponential convergence.

Compact representations

- **Intermediate representation** (sparse-ir)
 - *Ab initio* calculations (Eliashberg theory, *GW*, Lichtenstein formula)
 - Diagrammatic calculations (FLEX)
- Discrete Lehmann representation (implemented in sparse-ir as well)
- Minmax method (from Kresse's group)

Mathematical background: singular value decomposition (SVD)

Any complex-valued matrix A of size $M \times N$ can be decomposed as

$$A = U \Sigma V^\dagger,$$

where

$$U = (u_1, u_2, \dots, u_L) : M \times L,$$

$$V = (v_1, v_2, \dots, v_L) : N \times L,$$

where $u_i^\dagger u_j = \delta_{ij}$, $v_i^\dagger v_j = \delta_{ij}$, $L = \min(M, N)$. Σ is a diagonal matrix with non-negative diagonal elements $s_1 \geq s_2 \geq \dots \geq s_L \geq 0$.

- Unique up to a phase if the singular values s_i are non-degenerate.
- If A is a real matrix, U and V are also real orthogonal matrices.

Intermediate representation

Shinaoka *et al.* Phys. Rev. B 96, 035147 (2017)

Analytic continuation kernel

Fermion & boson:

$$G(\backslash \mathbf{i} \mathbf{v}) = \int_{-\infty}^{\infty} \backslash \mathbf{d} \mathbf{d} \omega \underbrace{\frac{1}{\backslash \mathbf{i} \mathbf{v} - \omega}}_{\equiv K(\backslash \mathbf{i} \mathbf{v}, \omega)} A(\omega)$$

$K(\backslash \mathbf{i} \mathbf{v}, \omega)$ is system independent and $A(\omega) = -\backslash \mathbf{i} \mathbf{i} (G^R(\omega) - G^A(\omega))$.

Analytic continuation kernel

$$G(\tau) = - \int_{-\infty}^{\infty} d\omega K(\tau, \omega) A(\omega)$$

$$K(\tau, \omega) \equiv -\frac{1}{\beta} \sum_{\mathbf{iv}} e^{-\mathbf{iv}\tau} K(\mathbf{iv}, \omega) = \begin{cases} \frac{e^{-\tau\omega}}{1+e^{-\beta\omega}} & (\text{fermion}) \\ \frac{e^{-\tau\omega}}{1-e^{-\beta\omega}} & (\text{boson}) \end{cases}$$

where $0 < \tau < \beta$.

For bosons, $|K(\tau, \omega)| \rightarrow +\infty$ at $\omega \rightarrow 0$. We want to use the same kernel for fermion & boson. How?

Logistic kernel

$$G(\tau) = - \int_{-\infty}^{\infty} d\omega K^{\text{L}}(\tau, \omega) \rho(\omega)$$

where $K^{\text{L}}(\tau, \omega)$ is the "logistic kernel" defined as

$$K^{\text{L}}(\tau, \omega) = \frac{e^{-\tau\omega}}{1 + e^{-\beta\omega}},$$

and $\rho(\omega)$ is the modified spectral function

$$\rho(\omega) \equiv \begin{cases} A(\omega) & (\text{fermion}), \\ \frac{A(\omega)}{\tanh(\beta\omega/2)} & (\text{boson}). \end{cases}$$

This trick has been widely used in the lattice QCD community for a long time. This was introduced into condensed matter physics in J. Kaye *et al.* (2022).

Singular value expansion

We introduce an ultraviolet $0 < \omega_{\text{max}} < \infty$ and a dimensionless parameter $\Lambda \equiv \omega_{\text{max}}\beta$

Because $K^{\text{L}} \in C^\infty$ and $\in L^2$:

$$K^{\text{L}}(\tau, \omega) = \sum_{l=0}^{\infty} U_l(\tau) S_l V_l(\omega),$$

for $-\omega_{\text{max}} \leq \omega \leq \omega_{\text{max}}$ and $0 \leq \tau \leq \beta$.

Singular functions: $\int_{-\omega_{\text{max}}}^{\omega_{\text{max}}} d\omega V_l(\omega) V_{l'}(\omega) = \delta_{ll'}$ and $\int_0^\beta d\tau U_l(\tau) U_{l'}(\tau) = \delta_{ll'}$.

→ Intermediate-representation basis functions

Singular values: $\omega_{\max} = 1$

- Exponential decay
- Number of relevant S_l grows as $O(\log \Lambda)$ (only numerical evidence)

Basis functions: $\omega_{\max} = 1$ and $\beta = 100$

- Even/odd functions for even/odd l
- l roots
- Converge to Legendre polynomials at $\Lambda \rightarrow 0$

Basis functions in Matsubara frequency

$$U_l(\mathbf{iv}) \equiv \int_0^\beta d\tau e^{i\mathbf{v}\tau} U_l(\tau).$$

Fourier transform can be done numerically.

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Expansion in IR

$$G(\tau) = \sum_{l=0}^{L-1} G_l U_l(\tau) + \epsilon_L,$$

$$\hat{G}(i\nu) = \sum_{l=0}^{L-1} G_l \hat{U}_l(i\nu) + \hat{\epsilon}_L,$$

where ϵ_L , $\hat{\epsilon}_L \approx S_L$. The expansion coefficients G_l can be determined from the spectral function as

$$G_l = -S_l \rho_l,$$

where

$$\rho_l = \int_{-\omega_{\max}}^{\omega_{\max}} d\omega \rho(\omega) V_l(\omega).$$

Convergence

$|G_l|$ converges as fast as S_l .

Example:

$$\rho(\omega) = \frac{1}{2}(\delta(\omega - 1) + \delta(\omega + 1))$$

$$\begin{aligned}\rho_l &= \int_{-\omega_{\max}}^{\omega_{\max}} d\omega \rho(\omega) V_l(\omega) \\ &= \frac{1}{2}(V_l(1) + V_l(-1))\end{aligned}$$

$$\beta = 100, \text{ } \backslash \text{wmax} = 1.$$

Sparse sampling

Li, Wallerberger, Chikano, Yeh, Gull, and Shinaoka, Phys. Rev. B 101, 035144 (2000)

Sparse time and frequency meshes

Solving Dyson equation for given $\Sigma(\backslash \mathbf{i} \mathbf{w})$:

$$G(\backslash \mathbf{i} \mathbf{w}) = (G_0^{-1}(\backslash \mathbf{i} \mathbf{w}) + \Sigma(\backslash \mathbf{i} \mathbf{w}))^{-1}$$
$$G_l = \sum_{n=-\infty}^{+\infty} U_l^*(\backslash \mathbf{i} \mathbf{w}_n) G(\backslash \mathbf{i} \mathbf{w}_n)$$

Q. Need to compute $G(\backslash \mathbf{i} \mathbf{w})$ on ALL Matsubara frequencies to determine L IR coefficients G_l ?

A. No, we need to know $G(\backslash \mathbf{i} \mathbf{w})$ on *appropriately chosen* ($\approx L$) sampling frequencies.

Dense mesh in τ ?

Second-order self-energy (Hubbard U):

$$\Sigma(\tau) \propto U G^2(\tau) G(\beta - \tau)$$

$$G_l = \int_0^\beta d\tau U_l(\tau) G(\tau)$$

Q: Need to compute $G(\tau)$ on a dense mesh of τ ?

A: No, we need to know $G(\tau)$ on *appropriately chosen* ($\approx L$) sampling points?

Sampling points

Simple rule: extrema (or somewhere in between two adjacent roots) of U_L

$\beta = 10$, $\backslash \text{wmax} = 10$, $L = 30$:

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Sampling points

Simple rule: extrema (or somewhere in between two adjacent roots) of U_L

$\beta = 10$, $\backslash \text{wmax} = 10$, $L = 30$:

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Transform from time/frequency to IR

- Well conditioned fitting problem
- Implemented in sparse-ir as stable linear transform

$$\begin{aligned} G_l &= \operatorname{argmin}_{G_l} \sum_k \left| G(\backslash \mathbf{tauk}) - \sum_{l=0}^{N_{\text{smp}}-1} U_l(\backslash \mathbf{tauk}) G_l \right|^2 \\ &= (\backslash \mathbf{Fmat}^+ \mathbf{G})_l, \end{aligned} \tag{1}$$

where we define $(\backslash \mathbf{Fmat})_{kl} = U_l(\backslash \mathbf{tauk})$ and $\backslash \mathbf{Fmat}^+$ is its pseudo inverse.

Condition number

- Small condition number of $\backslash \mathbf{Fmat}$
- If condition number is 10^p , you may lose p digits in transformation (three out of 16 digits)

Numerical demonstration

Two-pole model: $\beta = 100$, $\backslash \mathbf{w}_{\max} = 1$: Almost 16 significant digits!

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Stable and efficient numerical transform

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QA sessions

How to implement diagrammatic equations

Second-order perturbation theory

- Solving Dyson equation in frequency space
- Evaluating the self-energy in frequency space

Implementation of second-order perturbation theory

[Online tutorial](#)

Hubbard model on a square lattice:

$$\mathcal{H} = -t \sum_{\langle i,j \rangle} c_{i\sigma}^\dagger c_{j\sigma} + U \sum_i n_{i\uparrow} n_{i\downarrow} - \mu \sum_i (n_{i\uparrow} + n_{i\downarrow}),$$

where $t = 1$ and $\mu = U/2$ (half filling). $c_{i\sigma}^\dagger$ ($c_{i\sigma}$) a creation (annihilation) operator for an electron with spin σ at site i .

Non-interacting band dispersion:

$$\epsilon(\mathbf{k}) = -2(\cos k_x + \cos k_y),$$

where $\mathbf{k} = (k_x, k_y)$.

Self-consistent equations

$$G(i\nu, \mathbf{k}) = \frac{1}{i\nu - \epsilon(\mathbf{k}) + \mu - \Sigma(\backslash \mathbf{i}\mathbf{v}, \mathbf{k})} \quad (1)$$

$$\downarrow \quad (1)$$

$$\downarrow (\backslash \mathbf{i}\mathbf{v} \rightarrow \tau, \mathbf{k} \rightarrow \mathbf{r}) \quad (1)$$

$$\downarrow \quad (1)$$

$$\Sigma(\tau, \mathbf{r}) = U^2 G^2(\tau, \mathbf{r}) G(\beta - \tau, \mathbf{r}) \quad (1)$$

$$\downarrow \quad (1)$$

$$\downarrow (\tau \rightarrow \backslash \mathbf{i}\mathbf{v}, \mathbf{r} \rightarrow \mathbf{k}) \quad (1)$$

$$\downarrow \quad (1)$$

$$\text{Go back to (1)} \quad (1)$$

Self-consistent equations (sparse sampling)

$$G(\backslash \textcolor{red}{i} \mathbf{v} \mathbf{k}, \mathbf{k}) = \frac{1}{\backslash \textcolor{red}{i} \mathbf{v} \mathbf{k} - \epsilon(\mathbf{k}) + \mu - \Sigma(\backslash \textcolor{red}{i} \mathbf{v} \mathbf{k}, \mathbf{k})} \quad (1)$$

$$\downarrow \quad (1)$$

$$\downarrow (\backslash \textcolor{red}{i} \mathbf{v} \mathbf{k} \rightarrow \textcolor{red}{I} \mathbf{R} \rightarrow \backslash \textcolor{red}{t} \mathbf{a} \mathbf{u} \mathbf{k}, \mathbf{k} \rightarrow \mathbf{r}) \quad (1)$$

$$\downarrow \quad (1)$$

$$\Sigma(\backslash \textcolor{red}{t} \mathbf{a} \mathbf{u} \mathbf{k}, \mathbf{r}) = U^2 G^2(\backslash \textcolor{red}{t} \mathbf{a} \mathbf{u} \mathbf{k}, \mathbf{r}) G(\beta - \backslash \textcolor{red}{t} \mathbf{a} \mathbf{u} \mathbf{k}, \mathbf{r}) \quad (1)$$

$$\downarrow \quad (1)$$

$$\downarrow (\backslash \textcolor{red}{t} \mathbf{a} \mathbf{u} \mathbf{k} \rightarrow \textcolor{red}{I} \mathbf{R} \rightarrow \backslash \textcolor{red}{i} \mathbf{v} \mathbf{k}, \mathbf{r} \rightarrow \mathbf{k}) \quad (1)$$

$$\downarrow \quad (1)$$

$$\text{Go back to (1)} \quad (1)$$

The whole calculaiton can be performed on sparse meshes.

Reconstruction of spectral function

Please read [our article in the sparse-ir tutorial!](#)

Q: Can you reconstruct a spectral function from numerical data of $G(\tau)$?

A: Very difficult

$$G(\tau) = G_{\text{exact}}(\tau) + \delta(\tau),$$

where $\delta(\tau)$ is noise.

$$\rho_l = -(S_l)^{-1} ((G_l)_{\text{exact}} + \delta_l),$$

where $(G_l)_{\text{exact}} = \int_0^\beta \text{d}\tau U_l(\tau) G_{\text{exact}}(\tau)$ and $\delta_l = \int_0^\beta \text{d}\tau U_l(\tau) \delta(\tau)$.

Reconstruction of spectral function

Q: Can you reconstruct a spectral function from numerical data of $G(\tau)$?

A: Very numerical unstable!

$$G(\tau) = G_{\text{exact}}(\tau) + \delta(\tau),$$

where $\delta(\tau)$ is noise.

$$\rho_l = -(S_l)^{-1} ((G_l)_{\text{exact}} + \delta_l),$$

where $(G_l)_{\text{exact}} = \int_0^\beta \text{d}\tau U_l(\tau) G_{\text{exact}}(\tau)$ and $\delta_l = \int_0^\beta \text{d}\tau U_l(\tau) \delta(\tau)$.

Noise is amplified by small singular values. \rightarrow ill-posed inverse problem. Needed a regularized solver: MaxEnt, SpM, Nevanlinna etc.

"Nevanlinna.jl: A Julia implementation of Nevanlinna analytic continuation", K. Nogaki, J. Fei, E. Gull, HS, [arXiv:2302.10476v1](https://arxiv.org/abs/2302.10476v1)