\newcommand{\Fmat}{{\mathbf{F}}}

虚時間グリーン関数に対するスパースモデリング入門 (1)

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center

## 前提知識

- 虚時間形式グリーン関数の基礎
- Python or Juliaの基礎知識
  - 。 基本文法
  - 。多次元配列

0 ...

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### 何ができる?

虚時間・松原形式に基づく「数値」計算の高速・省メモリ化

# 超伝導転移温度の第一 原理計算

- T. Wang *et al.*, PRB 102, 134503 (2020), Nb
  - メモリの使用量が40分の1に! [松原周波数 4096点→ 103点]
  - 計算速度が20倍に!

# 虚時間グリーン関数に対するスパースモデリング入門(1) **他の応用例**



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### 背後にある技術

- 虚時間グリーン関数のコンパクトな中間表現基底
  - $\circ~G( au) = \sum_{l=0}^{L-1} U_l( au) g_l + \epsilon$
  - $\circ L \propto \log \beta W$  ( $\beta$ : inverse temperature, W: band width)
  - $\circ \; \epsilon \propto \exp(-aL)$  ( $\epsilon$ : truncation error, a>0)
- 虚時間・虚周波数におけるスパースメッシュ: # of points  $\simeq L$ .
- SparseIR.jl (Julia), sparse-ir (Python)

## 参考資料

- 固体物理 2021年6月 温度グリーン関数の情報圧縮に基づく高速量子多体計算法
- ↑の英語訳・加筆 + 新ライブラリsparse-irに更新 H. Shinaoka *et al.*, SciPost Phys. Lect. Notes 63 (2022)
- sparse-ir tutorials (大量のサンプルコード)
   https://spm-lab.github.io/sparse-ir-tutorial/index.html

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## 概要

- Part I
  - i. 虚時間グリーン関数の性質のまとめ
  - ii. 中間表現基底
  - iii. スパースサンプリング法

### **Imaginary-time Green's functions**

Also known as Matsubara Green's functions:

$$G( au) = -\langle T_{ au} A( au) B(0) \rangle,$$

where

- A( au), B( au) are operators in the Heisenberg picture ( $A( au) = e^{ au H} A e^{- au H}$ ).
- ullet  $\langle \cdots 
  angle = {
  m Tr}(e^{-eta H} \cdots)$  , where eta = 1/T ( $k_{
  m B} = 1$ ).

We use the Hamiltonian formalism throughout this lecture.

### **Imaginary-time Green's functions**

$$G( au) = -\langle T_ au A( au) B(0) 
angle$$

- ullet A and B are fermionic operators o G( au) = -G( au+eta)
- ullet A and B are bosonic operators o G( au) = G( au + eta)

In general, G( au) has a discontinuity at au=neta ( $n\in\mathbb{N}$ ).

# Imaginary-frequency (Matsubara) Green's functions

Matsubara Green's function:

$$G(\mathrm{i}\omega) = \int_0^eta \mathrm{d} au e^{\mathrm{i}\omega au} G( au).$$

From 
$$G(\tau+eta)=\mp G( au)$$
,

- $\omega = (2n+1)T\pi$  (fermion)
- $\omega = 2nT\pi$  (boson)

$$(n \in \mathbb{N})$$

These discrete imaginary frequencies are denoted as Matsubara frequencies.

### Spectral/Lehmann representation

$$G(z) = \int_{-\infty}^{+\infty} \mathrm{d}\omega' rac{
ho(\omega')}{z-\omega'},$$

where  $\rho(\omega)$  is a spectral function.

- $z=\mathrm{i}\omega o$  Matsubara Green's function
- $z = \omega + \mathrm{i}0^+ \to \mathrm{Retarded}$  Green's function (not used in this lecture)

### How Greeen's function look like in $\tau$ ?

Example (single pole):  $ho(\omega)=\delta(\omega-\omega_0)$ ,  $\omega_0>0$ 

$$G(\mathrm{i}\omega)=rac{1}{\mathrm{i}\omega-\omega_0}$$

$$G( au) = -rac{e^{- au\omega_0}}{1+e^{-eta\omega_0}} \; (0< au$$

At aupprox 0 ,  $G( au)\propto e^{- au\omega_0}$  .

For  $eta\omega_0\gg 1$  , coexisting two time scales :  $1/\omega_0\ll eta$ 

# How Greeen's function look like in Matsubara frequency space

Example (single pole):  $ho(\omega)=\delta(\omega-\omega_0)$ ,  $\omega_0>0$ 

$$G(\mathrm{i}\omega)=rac{1}{\mathrm{i}\omega-\omega_0}$$

$$G( au) = -rac{e^{- au\omega_0}}{1+e^{-eta\omega_0}} \; (0< au$$

At high frequencies  $|\omega|\gg |\omega_0|$  ,  $G(\mathrm{i}\omega)pprox 1/(\mathrm{i}\omega)$  .

For  $eta\omega_0\gg 1$ , coexisting two energy scales:  $\omega_0\ll T=1/eta$ 

### Difficulties in numerical simulations

If band width W and temperature T differ by orders of magnitudes as  $eta W\gg 1$ :

- ullet Slow power-law decay at high frequencies o Large truncation errors
- ullet Uniform dense mesh in au requires a huge number of points  $\propto eta W$ .

#### Example:

ullet Band width 10 eV, superconducting temperature 1 K pprox 0.1 meV  $ightarrow eta W = 10^5$  .

We need a compact basis with exponetial convergence.

### **Compact representations**

- Intermediate represenation (sparse-ir)
  - Ab initio calculations (Eliashberg theory, GW, Lichtenstein formula)
  - Diagrammatic calculations (FLEX)
- Discrete Lehmann representation (implemented in sparse-ir as well)
- Minmax method (from Kresse's group)

# 虚時間グリーン関数に対するスパースモデリング入門(1) Mathematical background: singular value decomposition (SVD)

Any complex-valued matrix A of size M imes N can be decomposed as

$$A=U\Sigma V^{\dagger},$$

where

$$U=(u_1,u_2,\cdots,u_L):M imes L,$$

$$V=(v_1,v_2,\cdots,v_L):N imes L,$$

where  $u_i^\dagger u_i = \delta_{ij}$ ,  $v_i^\dagger v_j = \delta_{ij}$ ,  $L = \min(M,N)$ .  $\Sigma$  is a diagonal matrix with nonnegative diagonal elements  $s_1 > s_2 > \cdots > s_L > 0$ .

- Unique up to a phase if the singular values  $s_i$  are non-degenerate.
- ullet If A is a real matrix, U and V are also real orthogonal matrices.

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### Intermediate representation

Shinaoka et al. Phys. Rev. B 96, 035147 (2017)

### **Analytic continuation kernel**

Fermion & boson:

$$G(ackslash{\mathbf{iv}}) = \int_{-\infty}^{\infty} ackslash{\mathrm{dd}} \omega rac{1}{\underbrace{ackslash{\mathbf{iv}} - \omega}} A(ackslash{\omega})$$

 $K(\mathbf{iv},\omega)$  is system independent and  $A(\omega)=-\mathbf{ii}(G^R(\omega)-G^A(\omega))$ .

### **Analytic continuation kernel**

$$G( au) = -\int_{-\infty}^{\infty} ackslash \mathrm{d} \mathrm{d} \omega K( au,\omega) A(\omega) \, ,$$

$$K( au,\omega) \equiv -rac{1}{eta} \sum_{egin{subarray}{c} {
m i}{
m v} \end{array}} e^{-igl({
m i}{
m v} au} K(igl({
m i}{
m v},\omega)) = egin{cases} rac{e^{- au\omega}}{1+e^{-eta\omega}} & ext{(fermion)} \ rac{e^{- au\omega}}{1-e^{-eta\omega}} & ext{(boson)} \end{cases}$$

where  $0 < \tau < \beta$ .

For bosons,  $|K(\tau,\omega)| \to +\infty$  at  $\omega \to 0$ . We want to use the same kernel for fermion & boson. How?

### **Logistic kernel**

$$G( au) = -\int_{-\infty}^{\infty} ackslash ext{dd} \omega K^{ ext{L}}( au,\omega) 
ho(\omega),$$

where  $K^{\mathrm{L}}( au,\omega)$  is the "logistic kernel" defined as

$$K^{
m L}( au,\omega) = rac{e^{- au\omega}}{1+e^{-eta\omega}},$$

and  $ho(\omega)$  is the modified spectral function

$$ho(\omega) \equiv egin{cases} A(\omega) & ext{(fermion),} \ rac{A(\omega)}{ anh(eta\omega/2)} & ext{(boson).} \end{cases}$$

This trick has been widely used in the lattice QCD community for a long time. This was introduced into condensed matter physics in J. Kaye *et al.* (2022).

### Singular value expansion

We introduce an ultraviolet  $0<\bigvee\max<\infty$  and a dimensionless parameter  $\Lambda\equiv\bigvee\max\beta$ 

Because  $K^{\mathrm{L}} \in C^{\infty}$  and  $\in L^2$ :

$$K^{
m L}( au,\omega) = \sum_{l=0}^{\infty} U_l( au) S_l V_l(\omega),$$

for  $-\backslash \text{wmax} \leq \omega \leq \backslash \text{wmax}$  and  $0 \leq \tau \leq \beta$ .

Singular functions:  $\int_{-\backslash \text{wmax}}^{\backslash \text{wmax}} \backslash \text{dd}\omega V_l(\omega) V_{l'}(\omega) = \delta_{ll'}$  and  $\int_0^\beta \backslash \text{dd}\tau U_l(\tau) U_{l'}(\tau) = \delta_{ll'}$ .

→ Indermediate-represetation basis functions

# Singular values: $\omega_{\mathrm{max}}=1$

- Exponential decay
- Number of relevant  $S_l$  grows as  $O(\log \Lambda)$  (only numerical evidence)

# Basis functions: $\omega_{ m max}=1$ and eta=100

- ullet Even/odd functions for even/odd l
- *l* roots
- ullet Converge to Legendre polynomials at  $\Lambda o 0$

### **Basis functions in Matsubara frequency**

$$U_l(ackslash{ ext{iv}}) \equiv \int_0^eta ackslash{ ext{dd}} au e^{ackslash{ ext{iv}} au} U_l( au).$$

Fourier transform can be done numerically.

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# 虚時間グリーン関数に対するスパースモデリング入門(1) **EXPANSION IN IR**

$$G( au) = \sum_{l=0}^{L-1} G_l U_l( au) + \epsilon_L,$$

$$\hat{G}(\mathrm{i}
u) = \sum_{l=0}^{L-1} G_l \hat{U}_l(\mathrm{i}
u) + \hat{\epsilon}_L,$$

where  $\epsilon_L$ ,  $\hat{\epsilon}_L \approx S_L$ . The expansion coefficients  $G_L$  can be determined from the spectral function as

$$G_l = -S_l \rho_l$$

where

$$ho_l = \int_{-\omega_{
m max}}^{\omega_{
m max}} {
m d}\omega 
ho(\omega) V_l(\omega).$$

### Convergence

 $|G_l|$  converges as fast as  $S_l$ .

#### Example:

$$egin{align} 
ho(\omega) &= rac{1}{2}(\delta(\omega-1)+\delta(\omega+1)) \ 
ho_l &= \int_{-\omega_{ ext{max}}}^{\omega_{ ext{max}}} \mathrm{d}\omega 
ho(\omega) V_l(\omega) \ &= rac{1}{2}(V_l(1)+V_l(-1)). \end{align}$$

$$\beta = 100$$
,  $\backslash \text{wmax} = 1$ .

### **Sparse sampling**

Li, Wallerberger, Chikano, Yeh, Gull, and Shinaoka, Phys. Rev. B 101, 035144 (2000)

### Sparse time and frequency meshes

Solving Dyson equation for given  $\Sigma(\mathbf{iw})$ :

$$egin{align} G(ackslash\mathbf{iw}) &= (G_0^{-1}(ackslash\mathbf{iw}) + \Sigma(ackslash\mathbf{iw})) \ G_l &= \sum_{n=-\infty}^{+\infty} U_l^*(ackslash\mathbf{iw}_n) G(ackslash\mathbf{iw}_n) \ \end{pmatrix}$$

Q. Need to compute  $G(\mathbf{iw})$  on ALL Mastubara frequencies to determine L IR coefficients  $G_l$ ?

A. No, we need to know  $G(\mathbf{w})$  on appropriately chosen  $(\approx L)$  sampling frequencies.

### Dense mesh in $\tau$ ?

Second-order self-energy (Hubbard U):

$$\Sigma( au) \propto UG^2( au)G(eta- au)$$

$$G_l = \int_0^eta d au U_l( au) G(\pi) \, .$$

Q: Need to compute  $G(\tau)$  on a dense mesh of  $\tau$ ?

A: No, we need to know G( au) on appropriately chosen (pprox L) sampling points?

### **Sampling points**

Simple rule: extrema (or somewhere in between two adjacent roots) of  $U_L$ 

$$\beta = 10$$
,  $\backslash \text{wmax} = 10$ ,  $L = 30$ :

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## **Sampling points**

Simple rule: extrema (or somewhere in between two adjacent roots) of  $U_{L}$ 

$$eta=10$$
,  $ackslash$ max  $=10$ ,  $L=30$ :

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### Transform from time/frequency to IR

- Well conditioned fitting problem
- Implemented in sparse-ir as stable linear transform

$$egin{aligned} G_l &= rgmin_{G_l} \sum_{k} \left| G(ackslash ext{tauk}) - \sum_{l=0}^{N_{ ext{smpl}}-1} U_l(ackslash ext{tauk}) G_l 
ight|^2 \ &= (ackslash ext{Fmat}^+ oldsymbol{G})_l, \end{aligned}$$

where we define  $(\mathbf{Fmat})_{kl} = U_l(\mathbf{tauk})$  and  $\mathbf{Fmat}^+$  is its pseudo inverse.

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### **Condition number**

- Small condition number of \Fmat
- If condition number is  $10^p$ , you may loos p digits in transformation (three out of 16 digits)

### **Numerical demonstration**

Two-pole model:  $\beta = 100$ ,  $\sqrt{\text{wmax}} = 1$ : Almost 16 significant digits!

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### Stable and efficient numerical transform

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### **QA** sessions

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### How to implement diagrammatic equations

### Second-order perturbation theory

- Solving Dyson equation in frequency space
- Evaluating the self-energy in frequency space

# Implementation of second-order perturbation theory

#### Online tutorial

Hubbard model on a square lattice:

$${\cal H} = -t \sum_{\langle i,j
angle} c^\dagger_{i\sigma} c_{j\sigma} + U \sum_i n_{i\uparrow} n_{i\downarrow} - \mu \sum_i (n_{i\uparrow} + n_{i\downarrow}),$$

where t=1 and  $\mu=U/2$  (half filling).  $c_{i\sigma}^{\dagger}$   $(c_{i\sigma})$  a creation (annihilation) operator for an electron with spin  $\sigma$  at site i.

Non-interacting band dispersion:

$$\epsilon(m{k}) = -2(\cos k_x + \cos k_y),$$

where  $oldsymbol{k}=(k_x,k_y)$  .

### **Self-consistent equations**

# 虚時間グリーン関数に対するスパースモデリング入門(1) **Self-consistent equations (sparse sampling)**

$$G(\ \ \mathbf{k}) = rac{1}{\langle \mathbf{ivk}, m{k} \rangle} = rac{1}{\langle \mathbf{ivk} - \epsilon(m{k}) + \mu - \Sigma(\ \ \mathbf{k}) \rangle} \qquad (1)$$
 $\downarrow \qquad \qquad (1)$ 
 $\downarrow \qquad \qquad (1)$ 
 $\downarrow \qquad \qquad (1)$ 
 $\Sigma(\ \ \mathbf{v}) = U^2 G^2(\ \ \mathbf{v}) G(\beta - \ \mathbf{v}) \qquad (1)$ 
 $\downarrow \qquad \qquad (1)$ 

The whole calculaiton can be performed on sparse meshes.

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### Reconstruction of spectral function

Please read our article in the sparse-ir tutorial!

Q: Can you reconstruct a spectral function from numerical data of  $G(\tau)$ ?

A: Very difficult

$$G( au) = G_{
m exact}( au) + \delta( au),$$

where  $\delta(\tau)$  is noise.

$$ho_l = -(S_l)^{-1}((G_l)_{\mathrm{exact}} + \delta_l),$$

where  $(G_l)_{\mathrm{exact}} = \int_0^\beta \backslash \mathrm{d}\mathrm{d} \tau U_l(\tau) G_{\mathrm{exact}}(\tau)$  and  $\delta_l = \int_0^\beta \backslash \mathrm{d}\mathrm{d} \tau U_l(\tau) \delta(\tau)$ .

### Reconstruction of spectral function

Q: Can you reconstruct a spectral function from numerical data of  $G(\tau)$ ?

A: Very numerical unstable!

$$G( au) = G_{
m exact}( au) + \delta( au),$$

where  $\delta(\tau)$  is noise.

$$ho_l = -(S_l)^{-1}((G_l)_{\mathrm{exact}} + \delta_l),$$

where 
$$(G_l)_{\mathrm{exact}} = \int_0^\beta \backslash \mathrm{d}\mathrm{d} \tau U_l(\tau) G_{\mathrm{exact}}(\tau)$$
 and  $\delta_l = \int_0^\beta \backslash \mathrm{d}\mathrm{d} \tau U_l(\tau) \delta(\tau)$ .

Noise is amplified by small singular values.  $\rightarrow$  ill-posed inverse problem. Needed a regularized solver: MaxEnt, SpM, Nevanlinna etc.

"Nevanlinna.jl: A Julia implementation of Nevanlinna analytic continuation", K. Nogaki,

J. Fei, E. Gull, HS, arXiv:2302.10476v1