## **Spectrum of Blackbody**

 ${\sf Sudheendra}\ {\sf B.R.}$ 

Suhas P.K.

Shahsank K.K.

Guided by V.H. Belvadi Yuvaraja's College, Mysuru

#### Overview

#### Part 1

Introduction to the Black Body problem

Charged oscillators

Rayleigh's intensity function

The Ultra-Violet catastropy

Plank's emphirical formula

Part 2

Intensity of a Black Body spectrum and the problem

Einstein's work on Black Bodies

Part 3

## Part 1

#### Introduction

An idealised object that can **absorb** and subsequently **emit**, all the radiation incident on it is called a **Blackbody**.

- Consider gaseous medium, in which we have a charged oscilator.
- Due to random motion of the gaseous atoms, it often collides with the oscillator.

- Consider gaseous medium, in which we have a charged oscilator.
- Due to random motion of the gaseous atoms, it often collides with the oscillator.
- As a result the gaseous atoms imparts some of it's kinetic energy to the oscillator.

- Mean kinetic energy acquired by the oscillator is,  $\frac{1}{2}kT$ .
- And the total kinetic energy will be kT.

• Since the charge is radiating energy the system cannot attain equilibrium.

• How do we tackle this situation?

- How do we tackle this situation?
- We confine the system.

- Inner walls of the box is totally reflective.
- $\bullet\,$  After some time the system will attains thermal equilibrium

So, the energy radiated by the oscillator per second can be given by,

$$\frac{1}{Q} = \frac{\left(\frac{dW}{dt}\right)}{\omega_0 W}$$

Q is the radiation reaction.

 $\omega_0$  is the natural frequency of vibration.

W is the total energy content of the oscillator.

Also, we have

$$\frac{1}{Q} = \frac{\gamma}{\omega_0}$$

Where  $\gamma$  is the dampping constant.

Remember the mean energy of the oscillator is kT? for three degrees of freedom, it is 3kT.

Average energy loss due to radiation is given by,

$$\left\langle \frac{dW}{dt} \right\rangle = \gamma kT$$

- Let,  $I(\omega)d\omega$  be the incident light energy with in a frequeny range  $d\omega$ .
- $\bullet$  Assuming all the radiation which falls on a cross section ' $\sigma_s$  ' is absorbed completely,
- Scattered light must be the product of  $I(\omega)d\omega$  and  $\sigma_s$ .

Expression for the cross section is given by,

$$\sigma_s = \frac{8\pi r_0^2}{3} \left[ \frac{\omega^4}{(\omega^2 - \omega_0^2)^2 + \gamma^2 \omega^2} \right]$$

When  $\omega = \omega_0$ ,

$$\sigma_s = \frac{2\pi r_0^2 3\omega_0^2}{3(\omega - \omega_0)^2 + \frac{\gamma^2}{4}}$$

$$\frac{dW_s}{dt} = \int_0^\infty \mathbf{I}(\omega) \, \sigma_s \, d\omega = \int_0^\infty \frac{2\pi r_0^2 \omega_0^2 \, \mathbf{I}(\omega) \, d\omega}{3 \left[ (\omega - \omega_0)^2 + \frac{\gamma^2}{4} \right]}$$

since we're considering the equilibrium condition ( $\omega = \omega_0$ ), equation reduces to,

$$rac{dW_s}{dt} = rac{2}{3}\pi r_0^2 \omega_0^2 \, \mathbf{I}(\omega_0) \int_{-\infty}^{\infty} rac{d\omega}{(\omega - \omega_0)^2 + \left(rac{\gamma}{2}
ight)^2}$$

Since the integral reduces to arctan function we get,  $\frac{2\pi}{\gamma}$ . and as

$$\frac{dW_s}{dt} = \frac{dW}{dt}$$

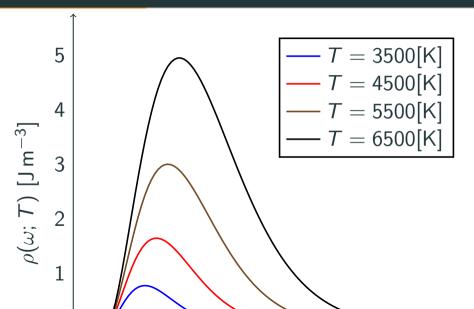
,

$$\mathbf{I}(\omega) = \frac{9\gamma^2 \, kT}{4\pi^2 r_0^2 \omega_0^2}$$

The value of  $\gamma$  wa found to be equal to,

$$frac23 \frac{r_0 \omega_0^2}{c}$$

$$I(\omega) = \frac{\omega^2 \, kT}{\pi^2 c^2}$$



• Theoretical proof was needed for to explain the obtained experimental results.

- The harmonic oscillator can take up energies only in the multiples of  $\hbar\omega$ .
- ullet The probability of occupying an energy level **E** is  ${f P}({f E})=lpha e^{rac{-\hbar\omega}{kT}}$  .

- ullet The number of oscillators in ground state is  $N_0$ .
- The number of oscillator in first state is  $N_1 = N_0 e^{-\hbar\omega/kT}$ .
- ullet Then  $N_n=N_0x^n$  , where  $x=e^{-\hbar\omega/kT}$

- The energy in the ground state is zero.
- The energy in the first state is  $N_1\hbar\omega$  or  $N_0\hbar\omega x$
- Total energy is given by

$$E_{(total)=N_0\hbar\omega(0+x+2x^2+\cdots)}$$

# Part 2

• The intensity of radiation for frequency  $\omega$  is given by

$$I(\omega)d\omega = rac{\hbar\omega^3d\omega}{\pi^2c^2\left(e^{rac{\hbar\omega}{kT}}-1
ight)}$$

- Planck considered that the matter was quantized and not light.
- Einstein came up with an alternate idea to solve Black body problem.

 $\bullet$  Einstein considered that even the light was quantized, by taking that light is actually photons with energy  $\hbar\omega$ 

• a pic is needed.

- Einstein assumed that Planck's formula was right.
- Consider two energy levels , say , the  $n^{th}$  level and  $m^{th}$  level.
- A light of certain frequency is incident on atom and by absorbing a photon, transition from state n to state m occurs.

- The probability of the transition is proportional to intensity of the light.
- ullet The proportionality constant is  $B_{nm}$  and mathematically can be written as,

$$R_{n\to m} = N_n B_{nm} I(\omega) \tag{1}$$

- Spontaneous emission.
- Stimulated emission.

• The combined mathematical expression is,

$$R_{m\to n} = N_m [A_{mn} + B_{mn} I(\omega)] \tag{2}$$

- At thermal equilibrium, the number of atoms going to higher energy level ,ust be equal to the number of atoms coming to lower energy state.
- The ratio of  $N_m$  to  $N_n$  is given by  $e^{\frac{-\hbar\omega}{kT}}$

• At thermal equilibrium both  $R_{n\to m}$  and  $R_{m\to n}$  are equal.

• 
$$N_n B_{nm} I(\omega) = N_m [A_{mn} + B_{mn} I(\omega)]$$

 $\bullet$  On dividing the previous equation by  $N_m$  we get,

• 
$$B_{nm}I(\omega)e^{\frac{\hbar\omega}{kT}} = A_{mn} + B_{mn}I(\omega)$$

• So the equation that Einsteim got for intensity of radiation for frequency is

$$I(\omega) = \frac{A_{mn}}{B_{nm}e^{\frac{\hbar\omega}{kT}} - B_{mn}} \tag{3}$$

 Since Einstein considered that the formula given by Planck was correct, we should get ,

• 
$$B_{nm} = B_{mn}$$
 and also

$$\frac{A_{mn}}{B_{nm}} = \frac{\hbar \omega^3 d\omega}{\pi^2 c^2}$$

This is how Einstein rederived Planck's formula from a quantum mechanical viewpoint.

# Part 3

- ullet  $\phi = Probability amplitude$
- $\mathbf{P} = |\phi|^2$

$$\bullet \ \phi = \phi_1 + \phi_2$$

$$\bullet \ \mathbf{P} = \left| \phi_1 + \phi_2 \right|^2$$

#### Case A

- Detector 1 is set to detect only  $\alpha$  particles and Dectector 2 is set to detect only oxygen atoms.
- Probability amplitude of the scattering is given as  $f(\theta)$  when they are at an angle  $\theta$ .
- The probability of this event  $= |f(\theta)|^2$

#### Case B

- ullet Set up the dectectors such that the detectors would dectect either lpha particle or oxygen atom.
- We will not distinguish which particle is which entering the detector.
- This means that if oxygen atom in position  $\theta$  , then  $\alpha$  particle on the opposite side is at an angle  $\pi-\theta$ .

- ullet Probability amplitude of oxygen atom  $= f(\pi heta)$
- Probability amplitude of  $\alpha$  particle =  $f(\theta)$
- The probability of a particle being detected at detector  $1=\left|f(\theta)\right|^2+\left|f(\pi-\theta)\right|^2$

- ullet Consider if both are lpha particles,
- Then we would not know which particle entered the detector, so the total probability changes to,
- ullet The probability of a lpha particle being detected at detector  $1=\left|f( heta)+f(\pi- heta)
  ight|^2$

• If 
$$\theta = \frac{\pi}{2}$$
, then applying this to the expression  $\left| f(\theta) + f(\pi - \theta) \right|^2$  we get,

• Probability =  $4\left|f\left(\frac{\pi}{2}\right)\right|^2$  , if the particles are indistinguishible.

- Suppose the particles were distinguishible, then the probability for  $\theta = \frac{\pi}{2}$  when applied for  $\left| f(\theta) \right|^2 + \left| f(\pi \theta) \right|^2$  is given by,
- Probability =  $2 \left| f\left(\frac{\pi}{2}\right) \right|^2$
- This shows that the probability gets doubled for indistinguishible particles.

- Can we apply the same logicto the electron-electron scattering?
- OBSERVATION: " When we have situation in which the identity of the electron which is arriving at a point is echanged with anothe one, the new amplitude interfere with old one with an opposite phase."
- In electrons case , the interfering amplitude for exchange interfere with a negative sign.

Probability for electron = 
$$|f(\theta) - f(\pi - \theta)|^2$$

# SPIN PROBABILITY TABLE

#### Identical Particles

**Identical particles**, also called **indistinguishable particle**, are particles that cannot be distinguished from one another.

## **Identical - Indistinguishable Particles**

- Consider particle 'a' and particle 'b'.
- Let the two particle collide and get sactttered in two different directions say '1' and '2' over a surface element  $ds_1$  and  $ds_2$  of the detector respectively.
- If the particles are indistinguishable then the amplitudes of these process will add up.
- Probability that the two particles arrive at  $ds_1$  and  $ds_2$  is  $|<1|a><2|b>+<2|a><1|b>|^2 ds_1 ds_2$

• Integrating over the area of the detector , if we let  $ds_1$  and  $ds_2$  range over the whole area ( $\triangle$  S) , we could count each part of the area twice since the expression  $|<1|a><2|b>+<2|a><1|b>|^2ds_1ds_2$  contains everything that can happen with any pair of surface elements  $ds_1$  and  $ds_2$ .

• Probability<sub>BOSE</sub> = 
$$\frac{\left(4\left|a\right|^{2}\left|b\right|^{2}\right)}{2}\left(\triangle S\right)$$

This is just twice what we got the probability for distinguished particles.

### State with n Bose particle

- Consider n particles say a, b, c... scattered in n direction say 1, 2, 3 ...
- Probability that each particle acting alone would go into an element of the surface ds of the detector is  $\left| <> \right|^2 ds$ .

- **Assumption** : All particles are distinguishable.
- Probability that n particles will be counted together in n different surface elements  $= \left| a_1 b_2 c_3 \dots \right|^2 ds_1 ds_2 \dots$
- If the amplitude does not depend on where ds is located in the detector , then the  $Probability = \left( \left| a \right|^2 \left| b \right|^2 ... \right) (ds_1 ds_2 ...)$

• Integrating each dS over the surface  $\triangle$  S of the dectector

$$(P_n)_{different} = \left( \left| a \right|^2 \left| b \right|^2 \dots \right) (\triangle S)^n$$

- Now suppose that all the particle are Bose particles.
- For n particles, there are n! different , but indistinguishable possibilities for which we must add the amplitudes.

- Probability that n particles will be counted on the n surface elements is given by  $Probability = \left( \left| a_1b_2c_3... + a_1b_3c_2... \right|^2 \right) (ds_1ds_2...)$
- Probability =  $\left(\left|n!abc...\right|^2\right) (ds_1ds_2...)$
- Integrate each ds over the area  $\triangle S$  of the detector  $(P_n)_{BOSE} = n! \left( \left| abc... \right|^2 \right) (\triangle S)^n$

 Compareing the probability when the particles are distinguishable and indistinguishable

$$(P_n)_{BOSE} = n! \left( \left| abc... \right|^2 \right) (\triangle S)^n$$
  
 $(P_n)_{different} = \left( \left| a \right|^2 \left| b \right|^2... \right) (\triangle S)^n$ 

• 
$$(P_n)_{BOSE} = n!(P_n)_{different}$$

What is the probability that a Bose particle will go into particular state when there are already n particles present ?

### **Emission and Absorption of photons**

- When the light is emitted, a photon is "created".
- Consider that there are some atom emitting n photons.
- OBSERVATION: The probability that an atom will wmit a photon into a
  particular final state is increased by the factor (n+1) if there are already n
  photons in that state.

# The Blackbody Spectrum

### The Blackbody Spectrum

- For each light frequency  $\omega$ , there are certain N number of atoms which have two energy states separated, given by the equation  $E = \omega \hbar$ .
- Let  $N_e$  and  $N_g$  be the average numbers of atoms that are in excited state and ground state.
- In thermal equilibrium at temperature T, from statistical mechanics
- NOTE: Each atom in the ground stae can absorb a photon and go into the
  excited state and each atom in the excited state can emit a photon and go to the
  ground state.

- At equilibrium, the rate of these two process must be equal.
- Rate is proportional to the probability of the event and the number of atoms present.
- $\bullet$   $\,\overline{n}$  is the average number of photons present in a given state with the frequency  $\omega$

- The absorption rate from the state is  $N_g \overline{n} \mid a \mid^2$ , and the emission rate into that state is  $N_e (\overline{n} + 1) \mid a \mid^2$ .
- ullet At equilibrium  $N_g\overline{n}\, \Big|\, a\, \Big|^{\,2}=N_e(\overline{n}+1)\, \Big|\, a\, \Big|^{\,2}$