## **Evoltrans code details**

## Basic flowchart:

- Set defaults for various parameters and read in inputs from ./input/diskevol.inp
- Read the radial grid, alpha, sigma gas and set 2prsigma Sigma vapor also defined here
- Read Rosseland mean opacity table for dust [NOTE: not being used]
- Set dust grain size, gastodust and pahscalefactor
- Calculate L acc and G0, mu
- Call radialrtinit routine
- Calculate Tmid from flaring angle, set scaleheight amd disk temperature
- Main Run step over time
  - Calls photoevaporation routine to get sigmadotpe; Save data
  - Loop over time: for each step do
    - Set a non-adaptive timestep, but check if sigma is varying fast
    - Optional call to infall here [typically skip]
    - Call photoevaporation routine and slowly ramp up with time first
    - Optionally add disk gravity to star mass
    - Call to compute\_vr\_v0\_difcoef [ computes radial velocity and diffusion coefficient]
    - Call to update sigma diff [ does the diffusion with above coefficient]
    - Evolve dust in a similar way
      - Get the sigma dust fraction as function of mass for each size after a call to dustcollision [needs alpha, snowline radius and grid, a]
      - Scale the pahs with the smallest grains
      - Calculate Stokes for each size and radius
      - Check if the grain photoevaporates
      - Compute vr, D and diffuse this dust component as for gas above
      - Save this sigmadust and update the gastodust ratio
      - Do ice formation
        - Estimate midplane number density
        - Check for partial pressure for water ice (and co)
        - Calculate fractional mass in ice (all vapor at t=0) amd distribute this equally in all grain size bins as ice
        - Rest of the sigma is in vapor
      - Update gas/dust to account for ice
    - Check if planetesimal formation has occurred
      - Calculate mass weighted stokes [what is this]
      - See if gas/dust falls below critical value
      - For each grain with some minimum stokes, put all of its sigma in planetesimals
      - Adjust for the ice and solid dust fractions appropriately
    - Update gas to dust after planetesimal formation
  - Save data if needed