

The construction of the consistent self-gravitating galaxy model with specific rotation curve

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Background

- How to set the initial conditions for the galaxy simulation
 - > Simulate for some time to achieve a dynamic equilibrium
 - > Virialize initially

$$\boxed{\sigma_R^2} \quad \sigma_\theta^2 \quad \overline{v}_R^2 \quad \overline{v}_\theta^2$$

Binney & Tremaine (1987)

Applied by Hernquist(1993) , Springel (2005), etc.

$$\sigma_R \leftarrow Q \equiv \frac{\sigma_R \kappa}{3.36 G \Sigma} \quad \sigma_\theta = \sigma_R \frac{\kappa}{2\Omega} \quad \overline{v}_\theta^2 - \overline{v}_c^2 = \sigma_R^2 \left(1 - \frac{\kappa^2}{4\Omega^2} - 2 \frac{R}{R_d} \right)$$

Epicyclic approximation

Main idea

- Obtain a consistent distribution function

$$f(\vec{v}, \vec{x}) \longrightarrow \sigma_R^{-2} \quad \sigma_\theta^{-2} \quad \overline{v}_R \quad \overline{v}_\theta$$

$$f(E, L_z) \longleftarrow \text{Jeans Theorem}$$

$$\frac{df}{dt} = \frac{\partial f}{\partial E} \frac{dE}{dt} + \frac{\partial f}{\partial L_z} \frac{dL_z}{dt} = 0$$

Main idea

- A typical profile of distribution function (Entropy maximized)

$$f(E, L_z) = P(L_z) \exp\left[\frac{-(E - E_c)}{\sigma^2}\right]$$

$$\sigma \longleftarrow Q \equiv \frac{\sigma\kappa}{3.36G\Sigma} = 1$$

$$\begin{aligned} & \int \int P(L_z) \exp\left[\frac{-(E - E_c)}{\sigma^2}\right] dP_R dL_z = R \cdot \Sigma(R) \\ & \downarrow \\ & P(L_z) \longleftarrow \int_0^{R_{h,0}(R)} K_0(R, R_h) R_h S(R_h) dR_h = R \cdot \Sigma(R) \\ & R_h^2 \Omega(R_h) = L_z \end{aligned}$$

Main idea

- Volterra-type integral equation

$$\int_0^{R_{h,0}(R)} K_0(R, R_h) \cdot R_h \cdot S(R_h) dR_h = R \cdot \Sigma(R)$$

Weight Function Same Dimension

Main methods

- Discretize the integration into several short intervals, within which apply some kind of integral formula

$$\int \dots \rightarrow \sum \dots$$

- Expand the unknown function into a series of orthogonal polynomials

$$F(x) \rightarrow \{A_j\}_{j=0,1,2,\dots}$$

Method I

- The Gauss-Legendre integral formula

$$\int_{-1}^1 f(x) dx \approx I_*(f) = f(1/\sqrt{3}) + f(-1/\sqrt{3})$$

- Discretize the integration into several short intervals

$$\int_0^{R_{h,0}(R)} K_0(R, R_h) \cdot R_h \cdot S(R_h) dR_h = \sum_{j=0}^{N(R)} \int_{R_h^{(j)}}^{R_h^{(j+1)}} K_0(R, R_h) \cdot R_h \cdot S(R_h) dR_h \quad R_h^{(j)} = jh, j \in [0, N]$$

- In each interval, apply the G-L integral formula

$$\begin{aligned} & \int_{R_h^{(j)}}^{R_h^{(j+1)}} K_0(R, R_h) \cdot R_h \cdot S(R_h) dR_h \\ &= \left[K_0\left(R, R_{h,-}^{(j)}\right) \cdot f\left(R_{h,-}^{(j)}\right) + K_0\left(R, R_{h,+}^{(j)}\right) \cdot f\left(R_{h,+}^{(j)}\right) \right] \cdot \left(R_h^{(j+1)} - R_h^{(j)} \right) / 2 \end{aligned}$$

$$R_{h,-}^{(j)} = \left((-1/\sqrt{3} + 1)/2 + j \right) h, \quad R_{h,+}^{(j)} = \left((1/\sqrt{3} + 1)/2 + j \right) h$$

Method I

- Input a series of data points

$$R \in \left\{ R_i \right\}_{i=0,1,\dots N}$$

- The integration turns into a matrix multiplication

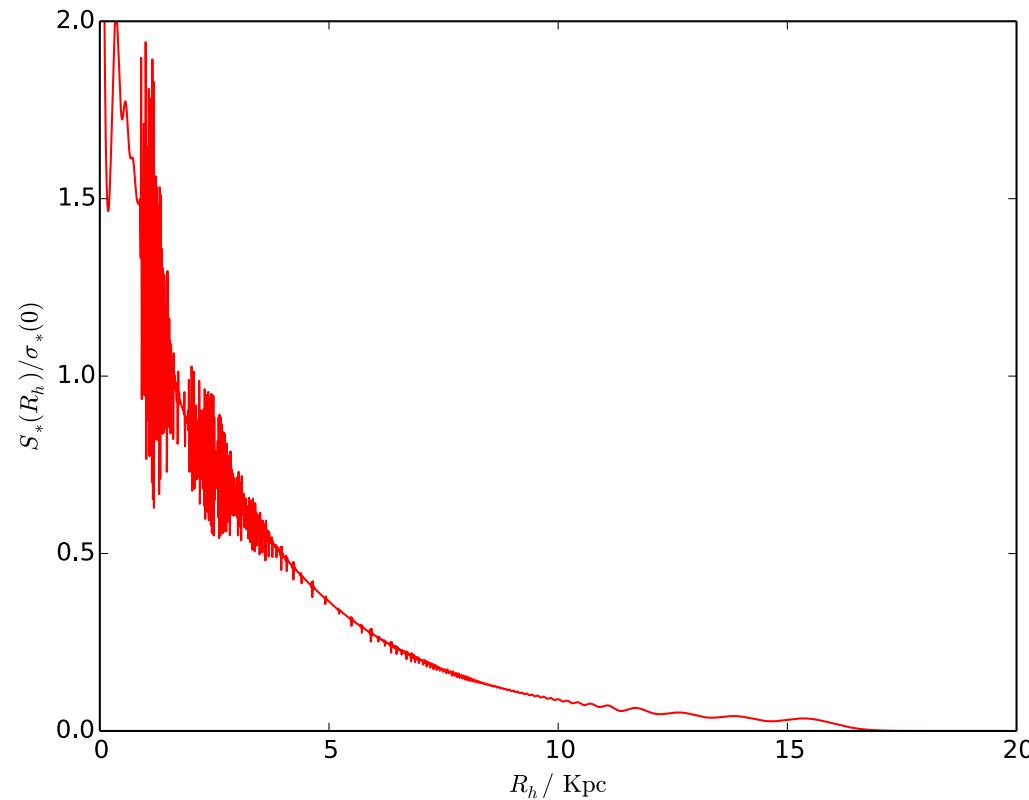
$$U F_* = G_*$$

- Operate the matrix inversion

$$F_* = U^{-1} G_*$$

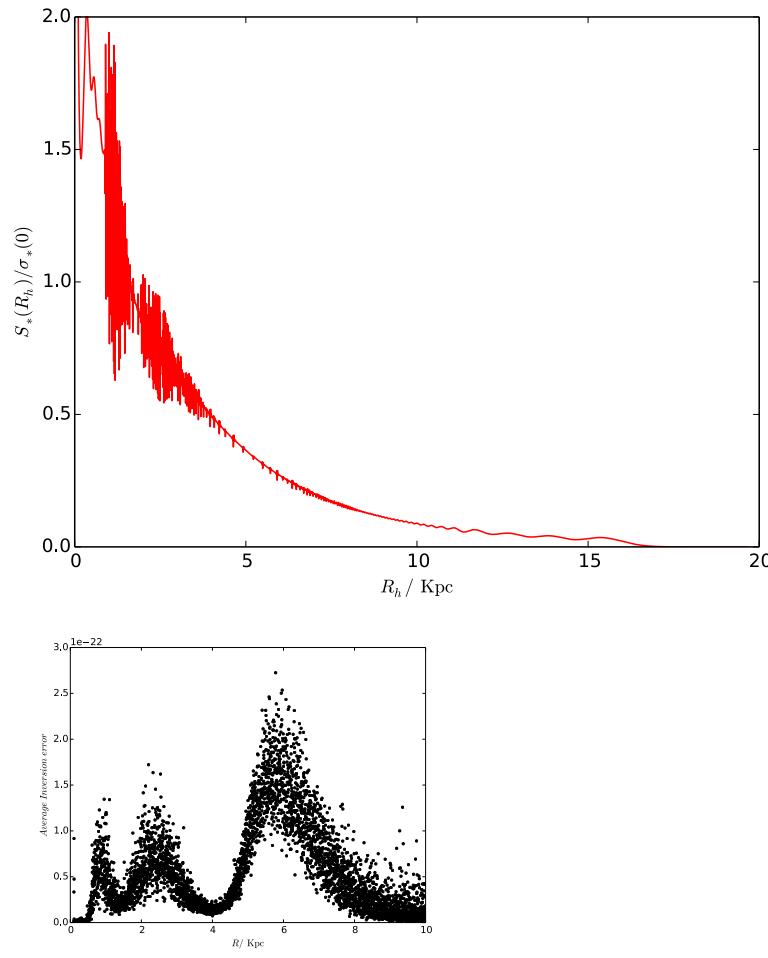
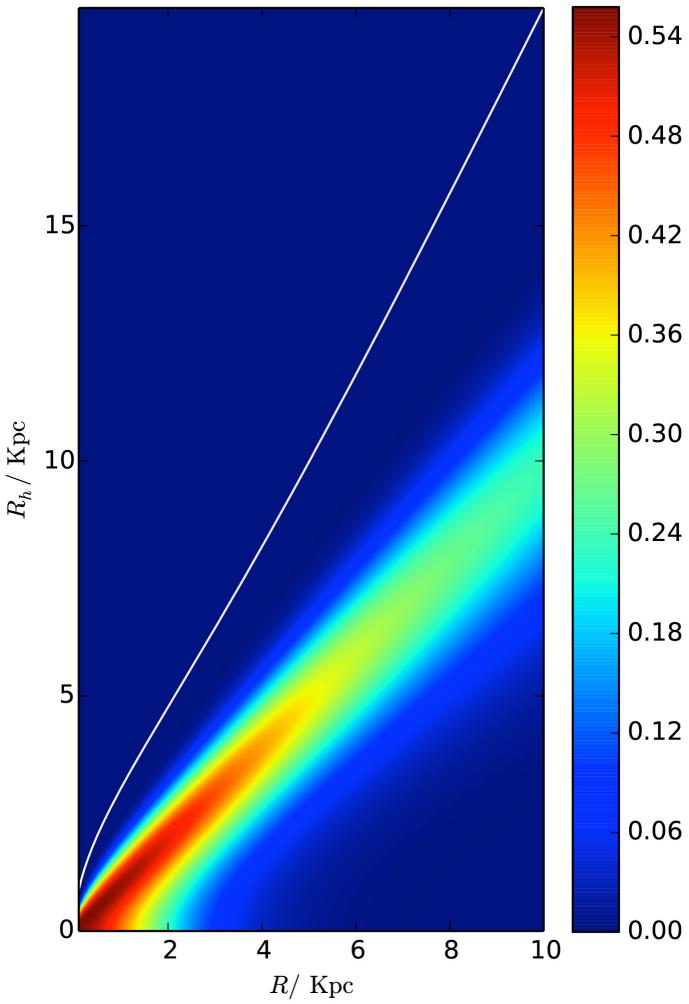
Result I

- $S(R_h)$



Problem I

- Matrix inversion error



Solution I

- Modify the weight function $K_0(R, R_h)$

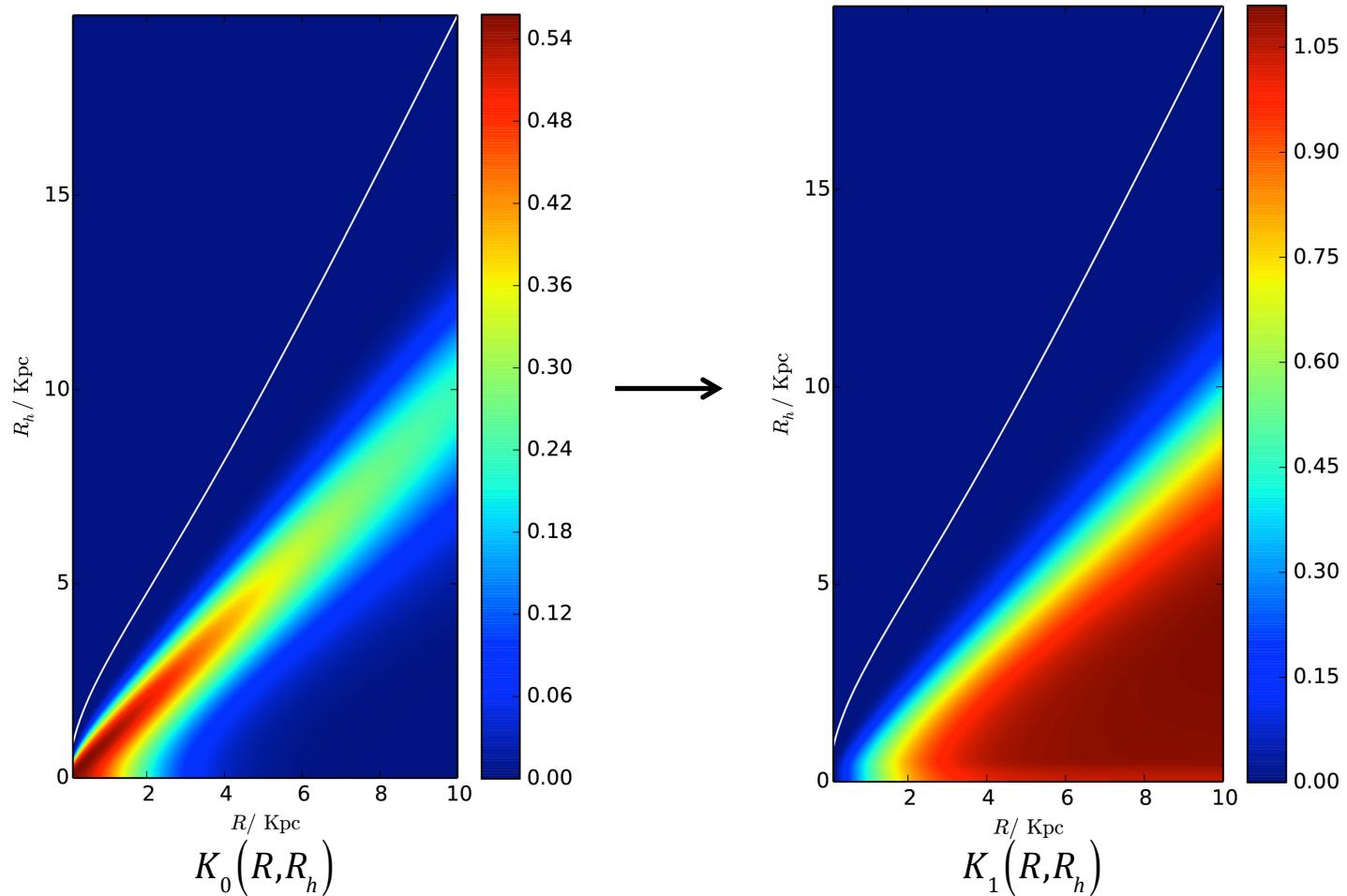
$$\int_0^R dR \int_0^{R_{h,0}(R)} dR_h K_0(R, R_h) \cdot R_h \cdot S(R_h) = \int_0^R dR R \Sigma(R)$$

$$\int_0^{R_{h,0}(R)} dR_h K_1(R, R_h) \cdot R_h \cdot S(R_h) = \int_0^R dR R \Sigma(R)$$

$$K_0(R, R_h) \longrightarrow K_1(R, R_h)$$

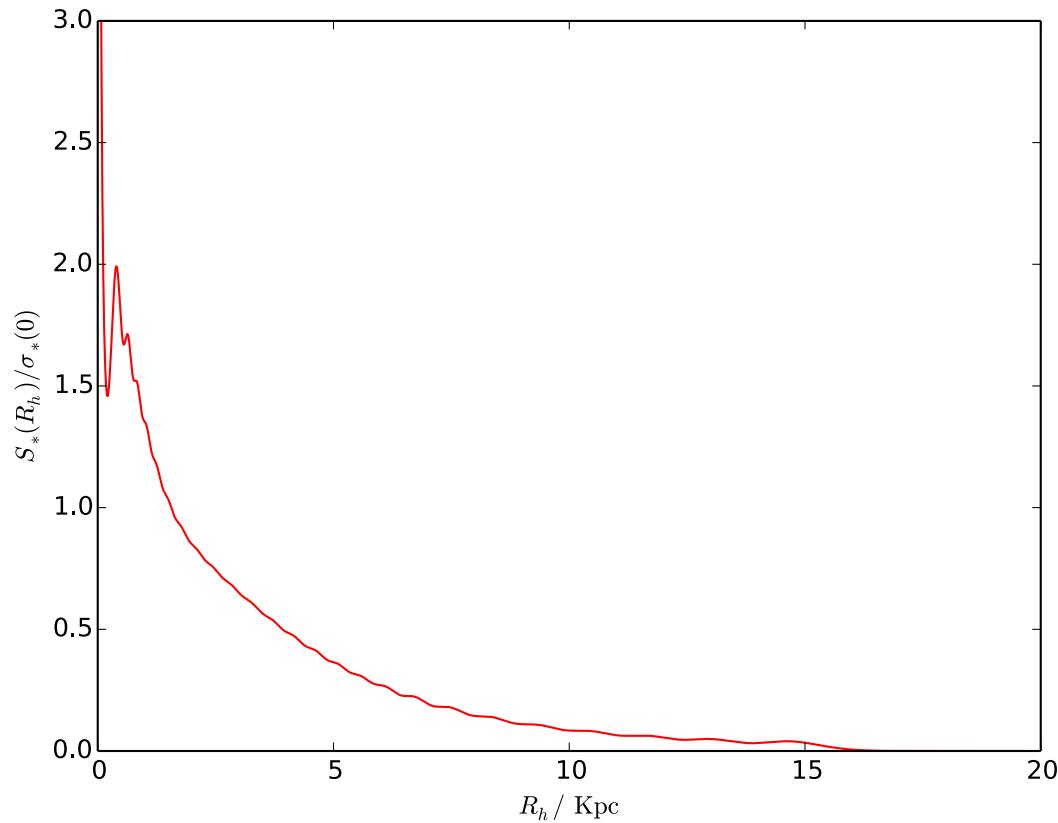
Solution I

- Modified weight function $K_1(R, R_h)$



Result I.I

- $S(R_h)$



Method II

- The Legendre polynomials

$$P_j(x) = \frac{1}{2^j j!} \frac{d^j}{dx^j} \left[(x^2 - 1)^j \right]$$

- Expand the unknown function

$$S(R_h) = \sum_{j=0}^N A_j P_j(x)$$

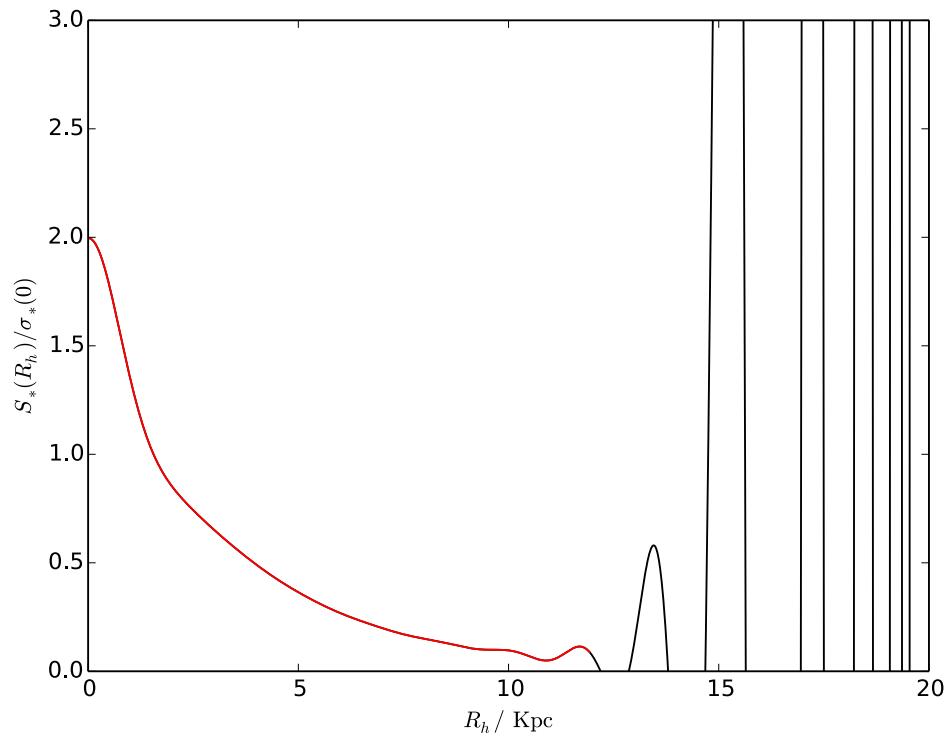
- Calculate the integration for each basis and turn the integral equation into a series of linear equations

$$\sum_{j=0}^N A_j \int_0^{R_{h,0}(R)} K_0(R, R_h) \cdot R_h \cdot P_j[x(R_h)] dR_h = R \cdot \Sigma(R)$$

- Apply the least square method to determine each coefficient

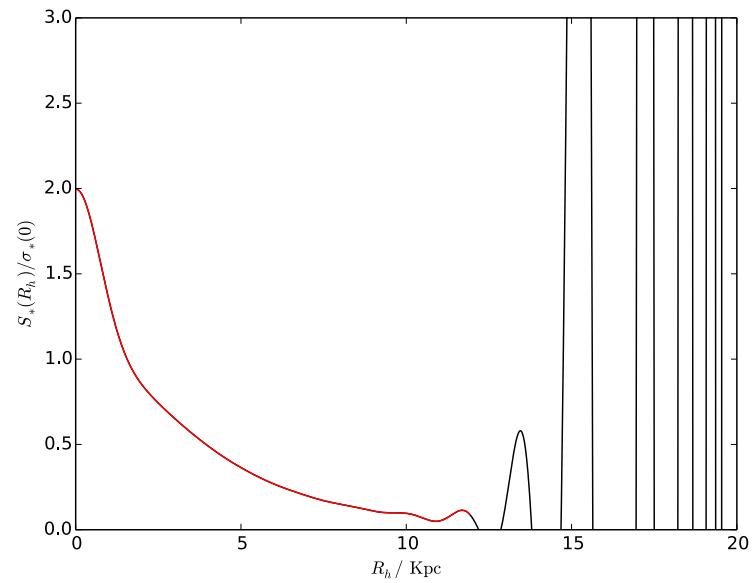
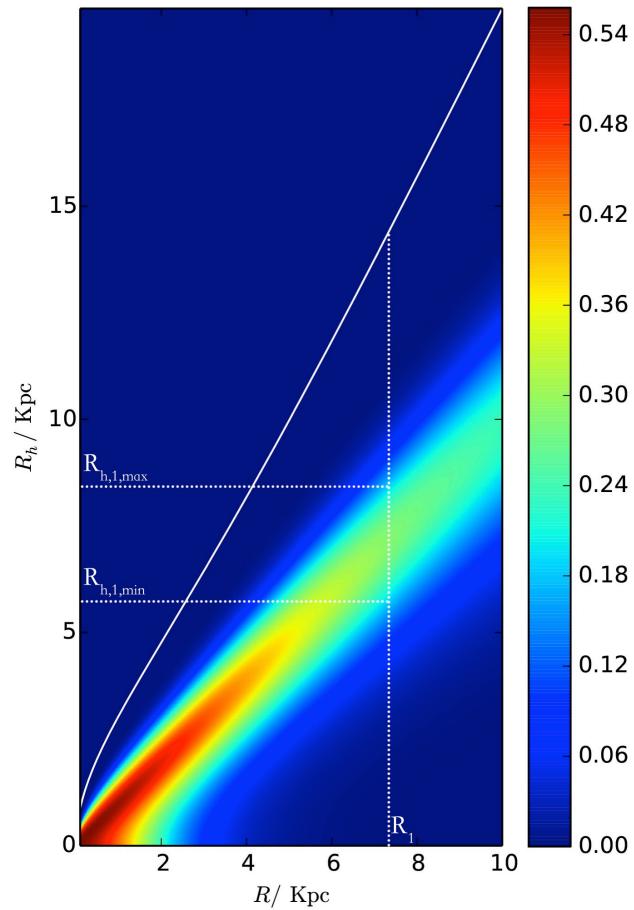
Result II

- $S(R_h)$



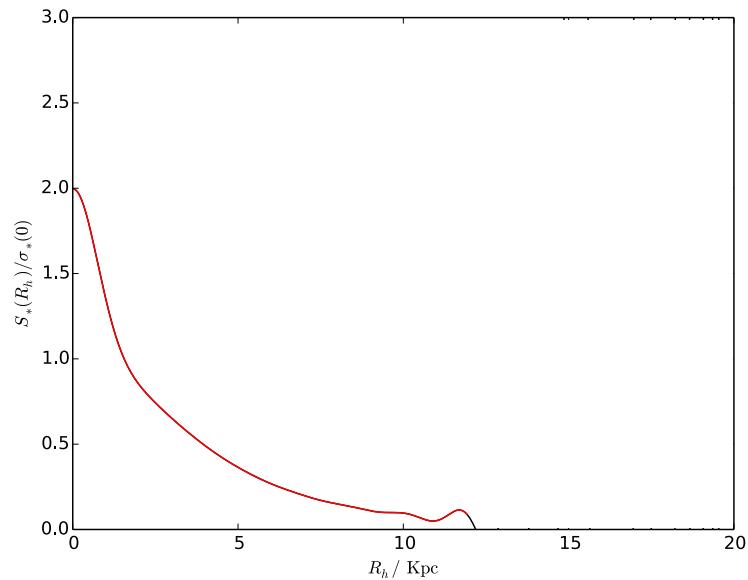
Problem II

- The singularity of weight function $K_0(R, R_h)$



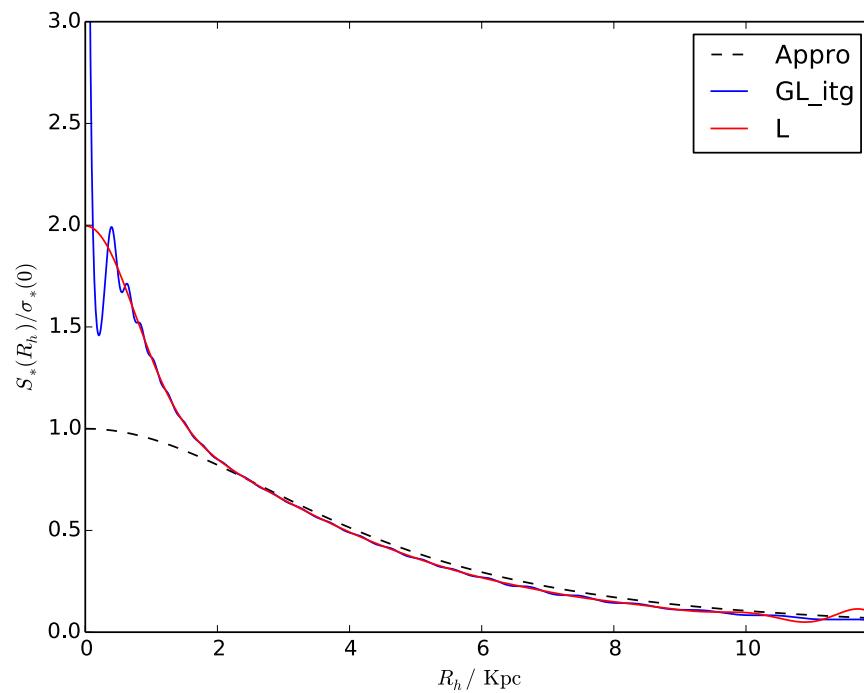
Solution II

- Cut it off!



Comparison

- Shu's appromixation $S(R_h) \approx \Sigma(R_h)$
- My derivation: $S(R_h) \approx 0.92 \Sigma(R_h)$



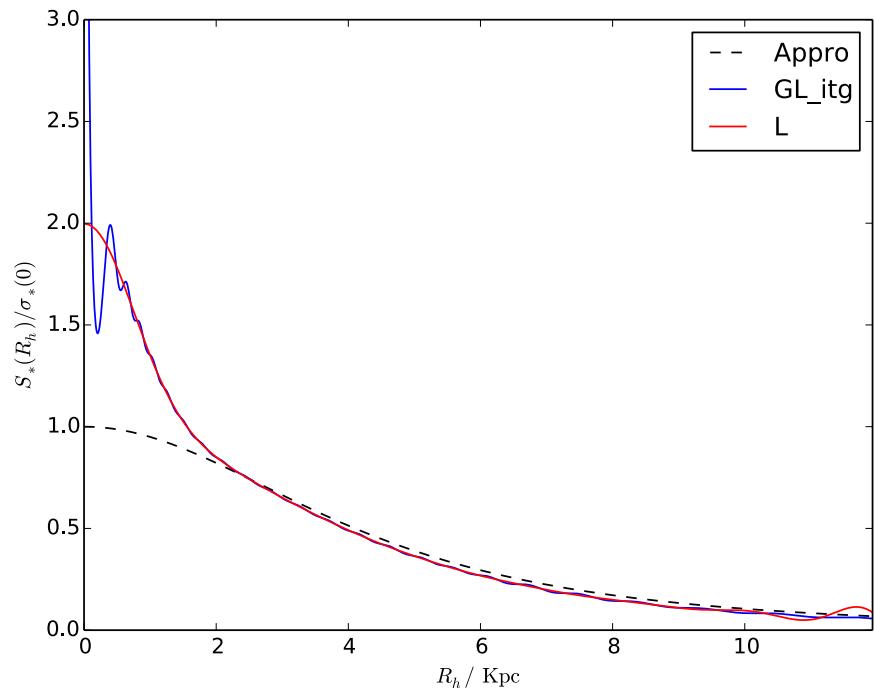
Problem

- $S(0) = ?$
- My derivation: $S(0) \approx \frac{\pi^{3/2} x_T}{2 \chi_s} \Sigma(0) = 2.0001 \Sigma(0)$

$$\chi_s = \int_0^1 s \exp\left[-\frac{s^4}{(\pi x_T)^2}\right] \operatorname{erf}\left[\frac{(1-s^4)^{1/2}}{\pi x_T}\right] ds$$

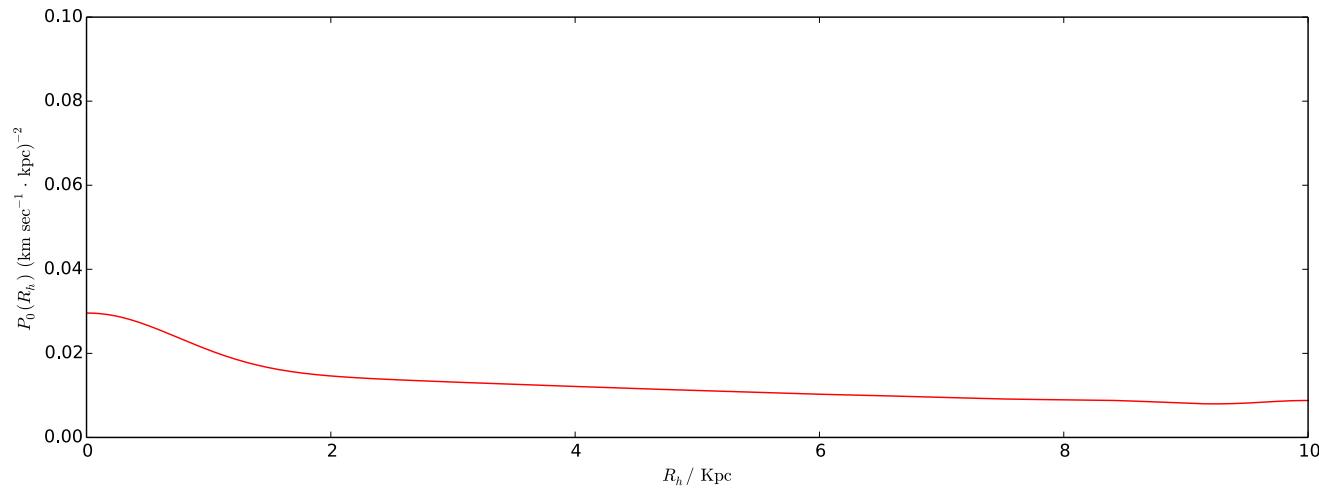
$$x_T = 0.08507$$

Constant from Toomre's criterion



Other result

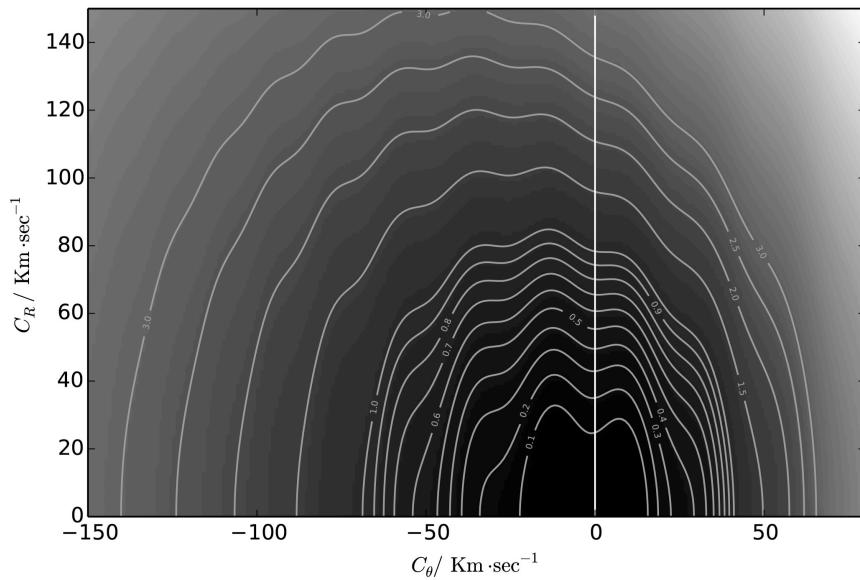
- $P_0(R_h)$



Other result

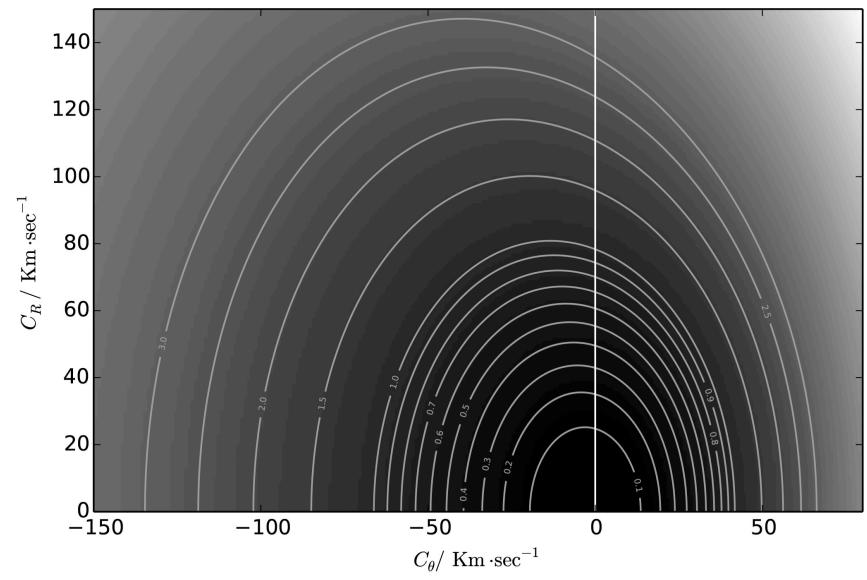
- Distribution function

$$f(E, L_z) \longrightarrow f(R, C_\theta, C_R)$$



G-L

$R = 10 \text{ kpc}$



Appro

Thanks for your time