

- 2 The Sun may be considered as a uniform sphere with a mass of  $1.99 \times 10^{30}$  kg and a surface temperature of 5780 K.

A probe with a mass of 2.63 kg moves in a straight line towards the Sun.

When it is at a distance  $x$  from the centre of the Sun, the probe measures the gravitational field strength  $g$  due to the Sun and the radiant flux intensity  $F$  of radiation from the Sun.

- (a) Define gravitational field.

.....  
..... [1]

- (b) For the position of the probe where  $x = 1.47 \times 10^{11}$  m:

- (i) calculate  $g$

$$g = \dots \text{ N kg}^{-1} \quad [2]$$

- (ii) determine the gravitational potential energy  $E_P$  of the probe.

$$E_P = \dots \text{ J} \quad [2]$$

- (c) (i) Show that, for any particular value of  $x$ , the numerical values of  $g$  and  $F$  are related by

$$g = \frac{4\pi GM}{L} F$$

where  $M$  is the mass of the Sun,  $L$  is the luminosity of the Sun and  $G$  is the gravitational constant.

[3]





(ii) Fig. 2.1 shows the variation of  $g$  with  $F$ .

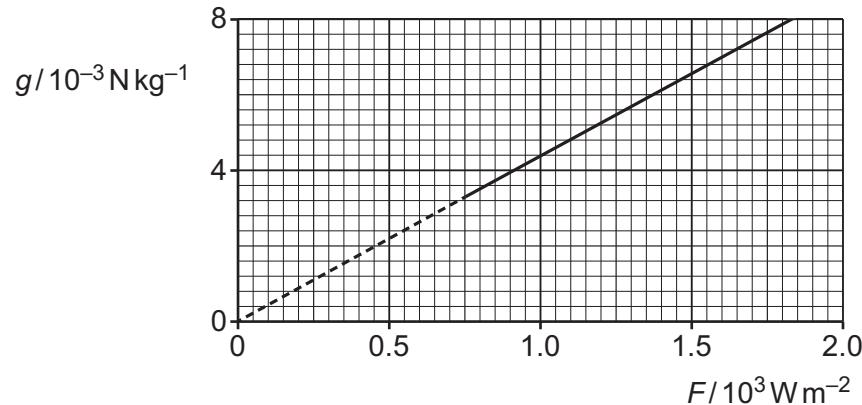


Fig. 2.1

Determine a value for the luminosity  $L$  of the Sun. Give a unit with your answer.

$L = \dots \text{ unit } \dots [2]$

(iii) Use your answer in (c)(ii) to determine the radius  $r$  of the Sun.

$r = \dots \text{ m } [2]$