

- 1 (a) State what is meant by a *gravitational force*.

.....
 [1]

- (b) A binary star system consists of two stars S_1 and S_2 , each in a circular orbit.

The orbit of each star in the system has a period of rotation T .

Observations of the binary star from Earth are represented in Fig. 1.1.

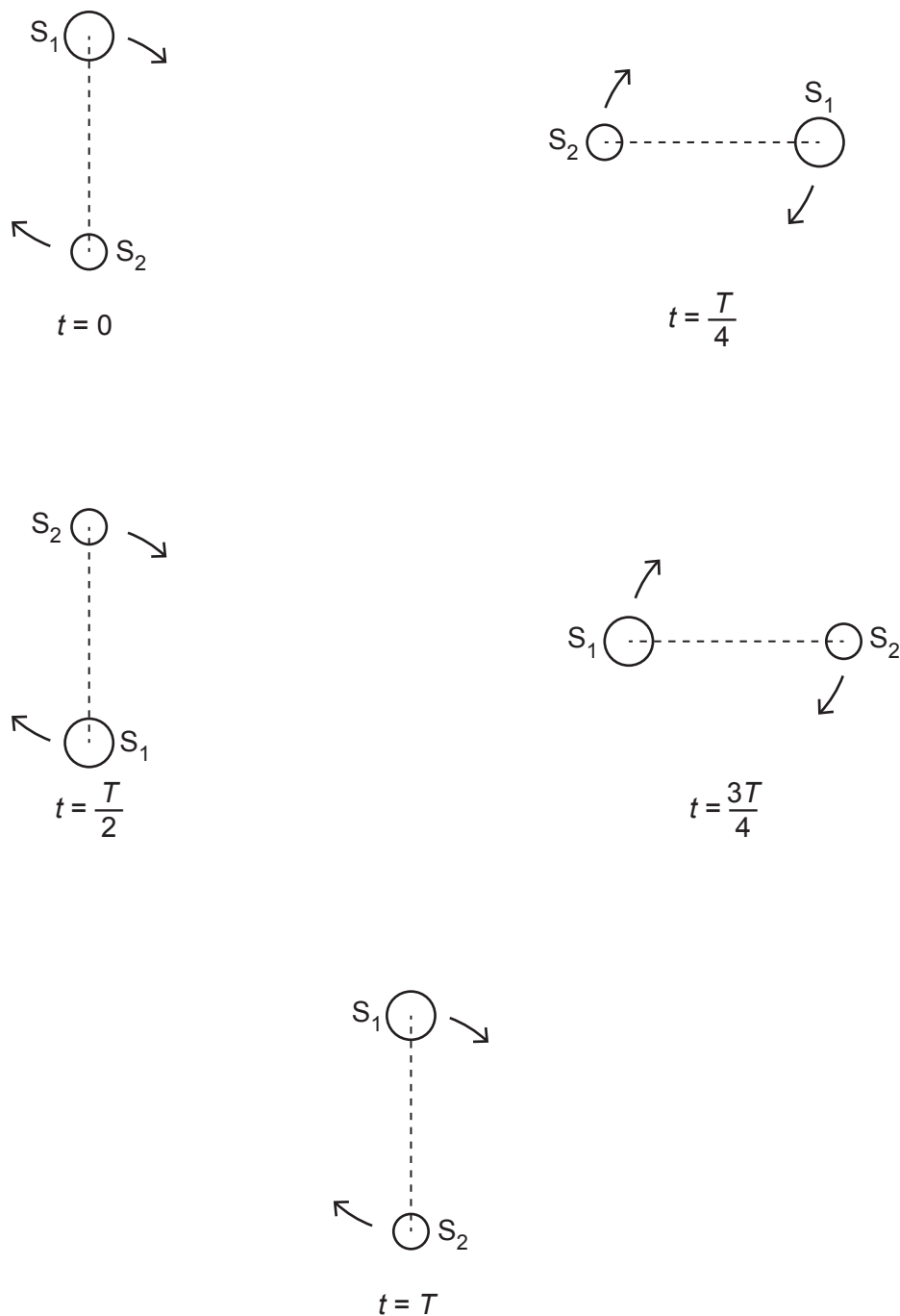


Fig. 1.1 (not to scale)

Observed from Earth, the angular separation of the centres of S_1 and S_2 is 1.2×10^{-5} rad. The distance of the binary star system from Earth is 1.5×10^{17} m.

Show that the separation d of the centres of S_1 and S_2 is 1.8×10^{12} m.

[1]

- (c) The stars S_1 and S_2 rotate with the same angular velocity ω about a point P, as illustrated in Fig. 1.2.

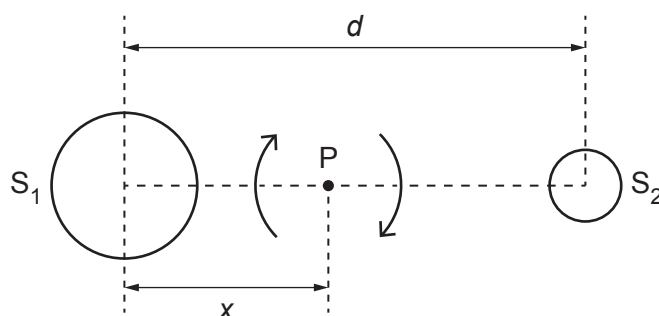


Fig. 1.2 (not to scale)

Point P is at a distance x from the centre of star S_1 . The period of rotation of the stars is 44.2 years.

- (i) Calculate the angular velocity ω .

$$\omega = \dots \text{ rad s}^{-1} \quad [2]$$

- (ii) By considering the forces acting on the two stars, show that the ratio of the masses of the stars is given by

$$\frac{\text{mass of } S_1}{\text{mass of } S_2} = \frac{d-x}{x}.$$

[2]

- (iii) The mass M_1 of star S_1 is given by the expression

$$GM_1 = d^2(d-x)\omega^2$$

where G is the gravitational constant.

The ratio in (ii) is found to be 1.5.

Use data from (b) and your answer in (c)(i) to determine the mass M_1 .

$$M_1 = \dots\dots\dots \text{ kg [3]}$$