

6

(a)

State what is meant by a *field of force* around a mass and hence define *gravitational field strength*.

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[2]

**(b)**

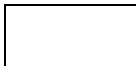
Using Newton's laws, show that, for a circular orbit of an object about the centre of a planet,

$$T^2 = \frac{4\pi^2 r^3}{GM}$$

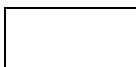
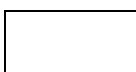
where  $T$  is the orbital period of the object,

$M$  is the mass of the planet, and

$r$  is the distance between the centre of mass of the object and the planet.



[3]



(c)

The planet Jupiter has several moons. Deriving from the expression in (b), data for some of these moons are plotted on a graph of  $\log_{10}(T/\text{days})$  against  $\log_{10}(r/m)$  on Fig. 6.1.

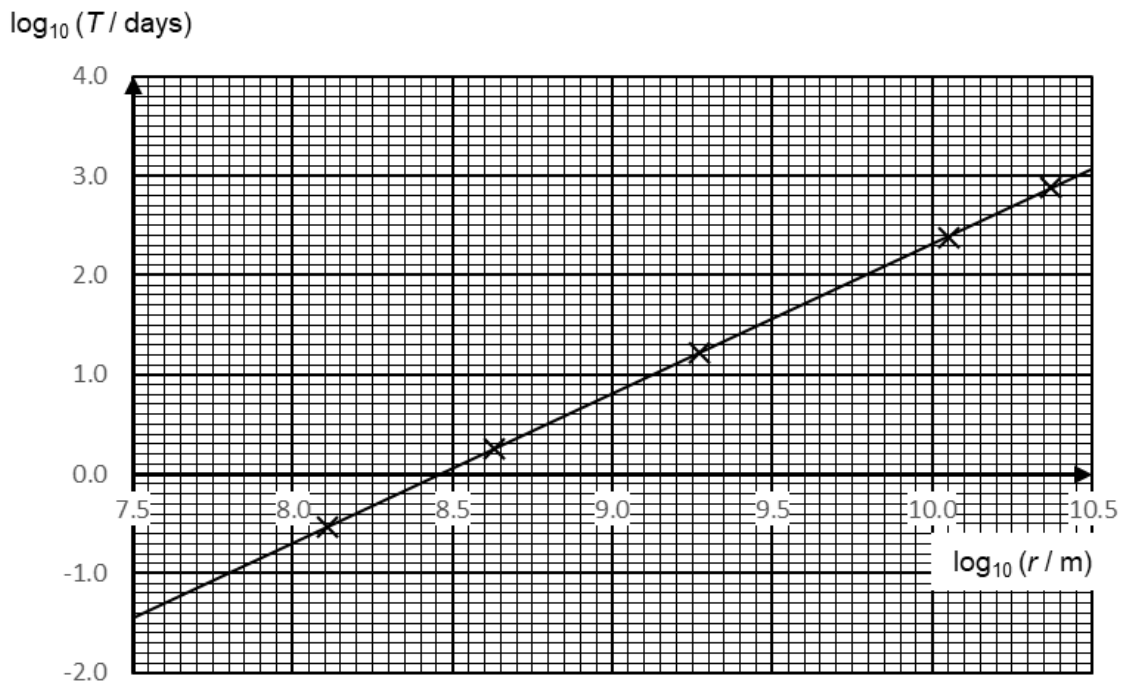
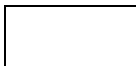
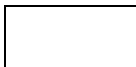
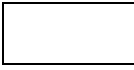


Fig. 6.1



Observations show that the moon Ganymede orbits Jupiter with a period of 7.16 Earth days. Use Fig. 6.1 to estimate the orbital radius of Ganymede.



orbital radius = ..... m

[2]

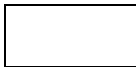
(d)

It is reported in the media that the moon Thebe is discovered to orbit Jupiter once every 16.2 hours at a height of  $2.22 \times 10^5$  km above the surface of Jupiter.

Referring to Fig. 6.1, comment on the accuracy of this media report.

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(e)

Fig. 6.2 shows how the gravitational potential,  $\phi$ , vary due to the Earth at three orbits, A, B and C. These orbits could be occupied by a 53.2 kg satellite in circular orbit around the Earth, E, of radius  $6.38 \times 10^3$  km.

orbit	$\phi$ / MJ kg <sup>-1</sup>	distance from surface / km
A	- 42.5	3000
B	- 47.6	2000
C	- 54.1	1000



Fig. 6.2



(i)

Determine the change in gravitational potential energy (GPE) from orbit B to orbit A. Leave your answer in standard form.

change in GPE = ..... J

[2]

**(ii)**

Using data from Fig. 6.2, show that the gravitational potential is inversely proportional to the radial distance.

[1]

