6 - Time, Time Systems, and Canonical Units

Dmitry Savransky

Cornell University

MAE 4060

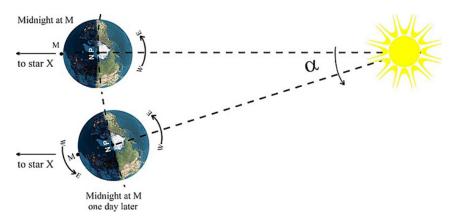
©Dmitry Savransky 2019

SI Seconds

- ▶ The second is the duration of 9,192,631,770 periods of the radiation corresponding to the transition between the two hyperfine levels of the ground state of the Cesium-133 atom.
 - -NIST (http://physics.nist.gov/cuu/Units/second.html)
- ► This definition refers to a cesium atom at rest at a temperature of 0 K. –BIPM

(http://www.bipm.org/en/publications/si-brochure/second.html)

Solar vs Sidereal Time



 $From: \ {\tt https://dept.astro.lsa.umich.edu/resources/ugactivities/Labs/Detroit/index.html}$

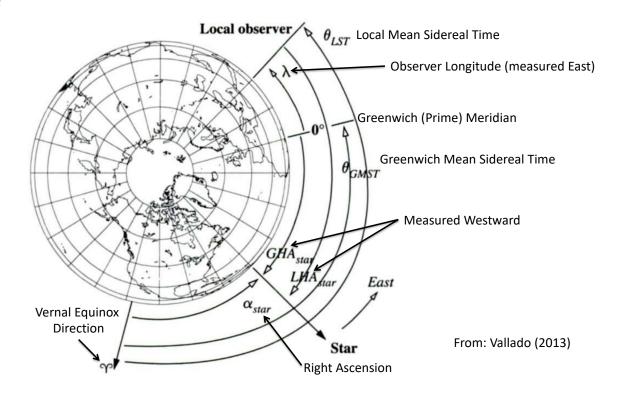
Mean Solar Day (d) 24 SI hours = 86400 SI seconds

 Solar (Tropical) Year
 365.242190402 d

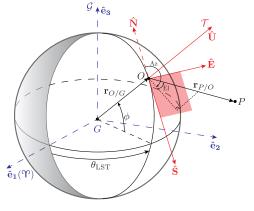
 Mean Sidereal Day
 23h56m4.09054s

 Sidereal Year
 365.256363004 d

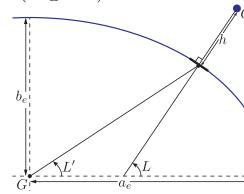
Hour Angle



Finding Where You Are (Again)



$$\begin{bmatrix} \mathbf{r}_{O/G} \end{bmatrix}_{\mathcal{G}} = \begin{bmatrix} x \cos \theta_{\text{LST}} \\ x \sin \theta_{\text{LST}} \\ z \end{bmatrix}_{\mathcal{G}}$$



A surface point O at Lon/Lat (λ, L) is at some height h above the reference geoid

Geoid described by a_e and e_e where: $e_e^2 = 2f - f^2$

$$x = \left(\frac{a_e}{\sqrt{1 - e_e^2 \sin^2 \phi}} + h\right) \cos \phi$$

$$y = \left(\frac{a_e(1 - e_e^2)}{\sqrt{1 - e_e^2 \sin^2 \phi}} + h\right) \sin \phi$$

Finding When You Are

$\theta_{\text{LST}} = \theta_{g0} + \omega_e(t - t_0) + \lambda_E$ OR Reference Value at Epoch

MEAN SIDEREAL TIME, 2019

Greenwich mean sidereal time at 0h UT

Greenwich mean sidereal time (GMST) on day d of month at hour t UT

= GMST at $0^{\rm h}$ UT on day $0+0^{\rm h}.065$ 71 $d+1^{\rm h}.002$ 74 t

 $\label{eq:local_local} \mbox{Local mean sidereal time} = \mbox{GMST} \begin{tabular}{l} + \mbox{east} \\ - \mbox{west} \end{tabular} \mbox{longitude}$

https://aa.usno.navy.mil/publications/reports/ap19_for_web.pdf-

$$\omega_e = 7.292115 \times 10^{-5} \text{ rad/s (WGS84)}$$

$$\theta_{g0} = 100.4606184^{\circ} + 36,000.77005361T_{\rm UT1} + 0.00038793T_{\rm UT1}^2 - 2.6 \times 10^{-8}T_{\rm UT1}^3$$

 $T_{\rm UT1} = \text{number of Julian centuries from J2000.0}$

Julian Date

▶ Days since January 1, 4713 BCE, 12^h UT

$$JD = 367Y - int \left(\frac{7\left(Y + int\left(\frac{M+9}{12}\right)\right)}{4} \right) + int \left(\frac{275M}{9}\right) + D + 1721013.5 + \frac{UT}{24}$$
$$-\frac{1}{2}\operatorname{sgn}(100Y + M - 190002.5) + \frac{1}{2}$$
$$int(x) = \begin{cases} \lfloor x \rfloor & x \ge 0 \\ \lceil x \rceil & x < 0 \end{cases} \quad \operatorname{sgn}(x) = \begin{cases} 1 & x \ge 0 \\ -1 & x < 0 \end{cases}$$

- ▶ 1 Julian year is exactly 365.25 days, 1 Julian century is 100 Julian years
- ► Define: MJD = JD -2,400,000.5

IEEE 754: Standard for Floating-Point Arithmetic

- ► A floating point number is represented by two values:
 - 1. s: The significand (mantissa, coefficient)—fixed length (p) digit string in base b
 - 2. e: The exponent—a signed integer

$$f \approx \frac{s}{b^{p-1}} b^e$$

▶ The IEEE 754 double precision (binary64, default in MATLAB) data type has: b = 2, p = 52, and 11 exponent bits ($e \in [-1022, 1023]$).