### Homework 4: Statistics

Please complete this homework assignment in code cells in the iPython notebook. Include comments in your code when necessary. Please rename the notebook as SIS ID\_HW04.ipynb (your student ID number) and save the notebook once you have executed it as a PDF (note, that when saving as PDF you don't want to use the option with latex because it crashes, but rather the one to save it directly as a PDF).

For questions that ask you to make an interpretation, please write a sentence or two explaining your response. You can use "Markdown" language to create a cell (like this one) which is ordinary text instead of using code. To do this, create a new cell, and then look at the bar above which has a picture of a floppy disk. On the right side is a drop down menu that allows you change whether a cell is "Markdown" or "Code".

The homework should be submitted on bCourses under the Assignments tab (both the .ipynb and .pdf files). Please label it by your student ID number (SIS ID)

#### Problem 1: Central Limit Theorem

Here we will verify the Central Limit Theorem and reproduced a plot I showed in class (https://en.wikipedia.org/wiki/Central\_limit\_theorem#/media/File:Dice\_sum\_central\_limit\_theorem.svg)

- 1. Write a function that returns n integer random numbers, uniformly disributed between 1 and 6, inclusively. This represents n throws of a fair 6-sided die. The value that comes up at each throw will be called the "score".
- 2. Generate a distribution of 1000 dice throws and plot it as a histogram normalized to unit area. Compute the mean  $\mu_1$  and standard deviation  $\sigma_1$  of this distribution. Compare your numerical result to the analytical calculation.
- 3. Generate 1000 sets of throws of N=2,3,4,5,10 dice, computing the total sum of dice scores for each set. For each value of N, plot the distribution of total scores, and compute the mean  $\mu_N$  and standard deviation  $\sigma_N$  of each distribution. This should be similar to the plot at the link above.
- 4. Plot the standard deviation  $\sigma_N$  as a function of N. Does it follow the Central Limit Theorem?

```
import numpy as np
import matplotlib.pyplot as plt

def roll_dice(n):
    x = np.random.randint(low=1, high=7, size=n)
    y = []
    for val in x:
        y.append(int(val))
    return y

rolls = roll_dice(1000)

n, bins, patches = plt.hist(rolls, bins=6)
plt.show()

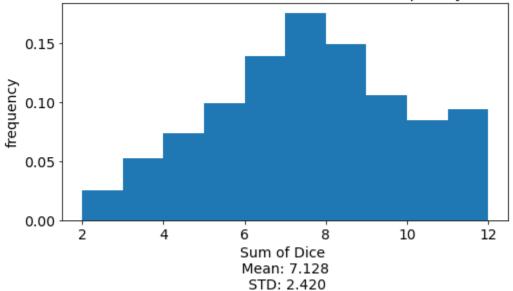
print('Mean = {0:.5f}'.format(np.mean(rolls)))
print('STD = {0:.5f}'.format(np.std(rolls)))
```

```
175 -
150 -
125 -
100 -
75 -
50 -
25 -
0 1 2 3 4 5 6

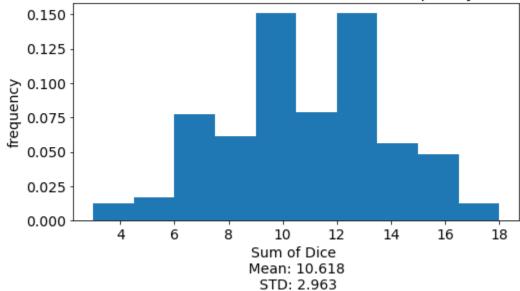
Mean = 3.54100
STD = 1.65297
```

```
N = [2, 3, 4, 5, 10]
In [25]:
          stds = []
          for val in N:
              rollTotal = []
              for i in range(1000):
                  rolls = roll_dice(val)
                  rollTotal.append(np.sum(rolls))
              n, bins, patches = plt.hist(rollTotal, density=True)
              plt.title('1000 Throws of {0} Fair Six-Side Dice Normalized Frequency vs Sum of Dice
              plt.xlabel('Sum of Dice \n Mean: {0:.3f} \n STD: {1:.3f}'.format(np.mean(rollTotal),
              plt.ylabel('frequency')
             plt.show()
              stds.append(np.std(rollTotal))
              rollTotal = []
          plt.scatter(N, stds)
          plt.title('1000 Throws of N fair six-side dice standard deviation of 1000 throw sums vs
          plt.ylabel('STD of sums')
         plt.xlabel('N number of dice thrown')
          plt.show()
```

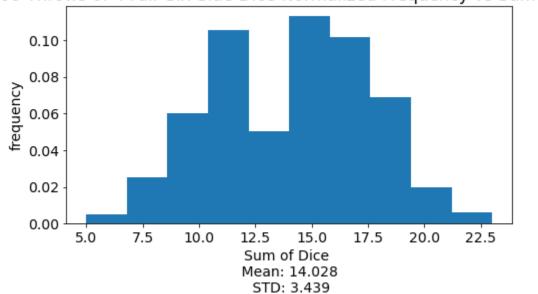
### 1000 Throws of 2 Fair Six-Side Dice Normalized Frequency vs Sum of Dice



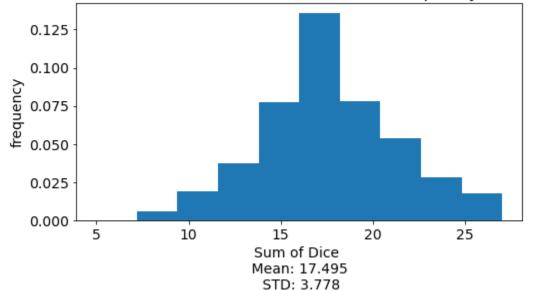
# 1000 Throws of 3 Fair Six-Side Dice Normalized Frequency vs Sum of Dice



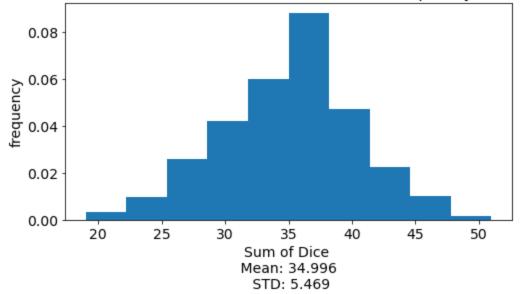
## 1000 Throws of 4 Fair Six-Side Dice Normalized Frequency vs Sum of Dice



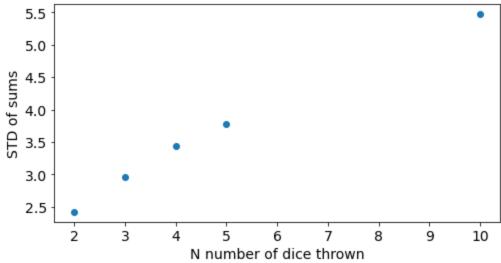
1000 Throws of 5 Fair Six-Side Dice Normalized Frequency vs Sum of Dice



#### 1000 Throws of 10 Fair Six-Side Dice Normalized Frequency vs Sum of Dice



1000 Throws of N fair six-side dice standard deviation of 1000 throw sums vs N



# Problem 2: Parity-violating asymmetry

The data sample for this problem comes from the E158 experiment at SLAC (a national lab near that Junior university across the Bay). E158 measured a parity-violating asymmetry in Møller (electron-electron) scattering. This was a fixed-target experiment, which scattered longitudinally-polarized electrons off atomic (unpolarized) electrons in the 1.5m liquid hydrogen target. The data below contains a snapshot of 10,000 "events" from this experiment (overall, the experiment collected almost 400 million such events over the course of about 4 months). Each event actually records a pair of pulses: one for the right-handed electron (spin pointing along momentum) and one for the left-handed electron. For each event, we record 4 variables:

- · Counter: event index
- Asymmetry: "raw" cross section asymmetry  $A_{raw}$  from one of the detector channels (there are 50 of these overall). The cross section asymmetry is defined as  $A_{raw} = \frac{\sigma_R \sigma_L}{\sigma_R + \sigma_L}$  The asymmetry is recorded in units of PPM (parts per million). It is called "raw" because corrections due to the difference in beam properties at the target are not yet applied.
- DeltaX: difference in beam position  $\Delta X = X_R X_L$  at the target in X direction in microns (with the convention that the beam is traveling along Z)

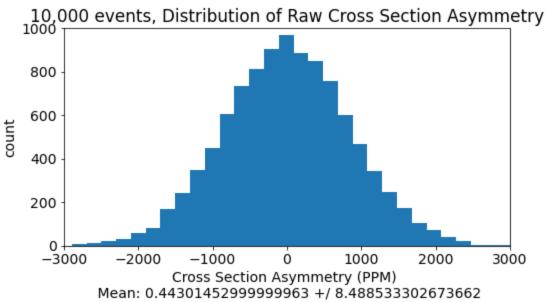
- DeltaY: difference in beam position  $\Delta Y = Y_R - Y_L$  at the target in Y direction in microns

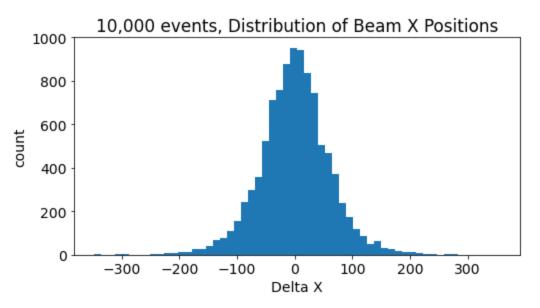
The data sample is provided in plain text format as the file asymdata.txt. Questions for this analysis:

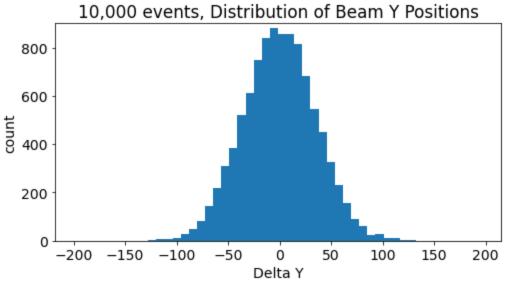
- 1. Read the data from the file, and plot distributions of  $A_{raw}$ ,  $\Delta X$ , and  $\Delta Y$ .
- 2. Compute the mean of the raw asymmetry distribution and its statistical uncertainty.
- 3. Plot  $A_{raw}$  vs  $\Delta X$ ,  $A_{raw}$  vs  $\Delta Y$ , and  $\Delta X$  vs  $\Delta Y$  as scatter plots.
- 4. Compute the correlation coefficients Corr(Asym, DeltaX), Corr(Asym, DeltaY), and Corr(DeltaX, DeltaY). See lecture notes, Workshop05.ipynb and Workshop05\_optional.ipynb or https://en.wikipedia.org/wiki/Pearson\_correlation\_coefficient for additional help understanding correlation coefficients. Which variables are approximately independent of each other?

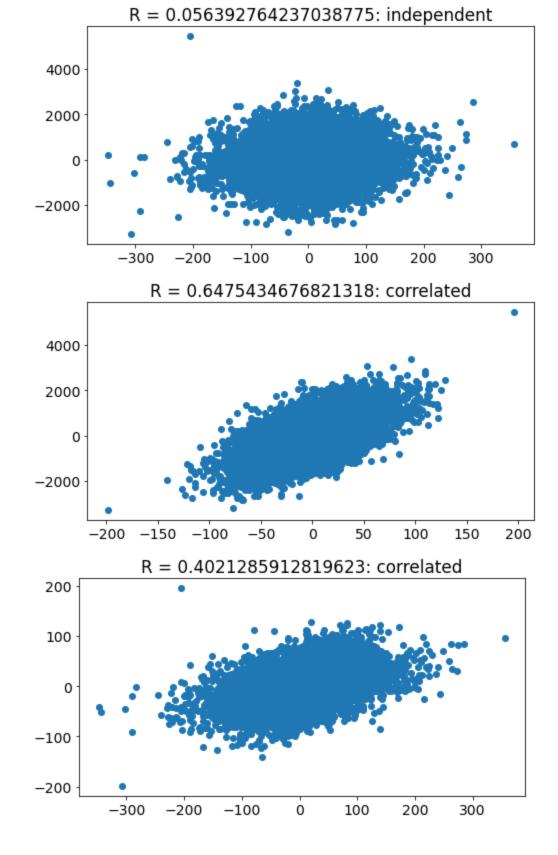
```
In [48]: f = np.loadtxt('asymdata.txt')
         event_index = []
         cross_section_asymmetry = []
         deltaX = []
         deltaY = []
         for value in f:
             event_index.append(value[0])
             cross_section_asymmetry.append(value[1])
             deltaX.append(value[2])
             deltaY.append(value[3])
         cross_section_asymmetry_mean_err = np.std(cross_section_asymmetry)/(len(cross_section_as
         plt.hist(cross_section_asymmetry, bins=44)
         plt.title('10,000 events, Distribution of Raw Cross Section Asymmetry')
         plt.ylabel('count')
         plt.xlabel('Cross Section Asymmetry (PPM) \n Mean: {0} +/ {1}'.format(np.mean(cross_sect
         plt.ylim([0, 1000])
         plt.xlim([-3000,3000])
         plt.show()
         plt.hist(deltaX, bins=58)
         plt.title('10,000 events, Distribution of Beam X Positions')
         plt.ylabel('count')
         plt.xlabel('Delta X')
         plt.ylim([0, 1000])
         plt.show()
         plt.hist(deltaY, bins=50)
         plt.title('10,000 events, Distribution of Beam Y Positions')
         plt.ylabel('count')
         plt.xlabel('Delta Y')
         plt.ylim([0, 900])
         plt.show()
         plt.scatter(deltaX, cross_section_asymmetry)
         plt.title('R = {}: independent'.format(np.corrcoef(deltaX, cross_section_asymmetry)[0][1
         plt.show()
         plt.scatter(deltaY, cross_section_asymmetry)
         plt.title('R = {}: correlated'.format(np.corrcoef(deltaY, cross_section_asymmetry)[0][1]
         plt.show()
         plt.scatter(deltaX, deltaY)
```

plt.title('R = {}: correlated'.format(np.corrcoef(deltaX, deltaY)[0][1])) plt.show()









# Problem 3: Gamma-ray peak

[Some of you may recognize this problem from Advanced Lab's Error Analysis Exercise. That's not an accident. You may also recognize this dataset from Homework04. That's not an accident either.]

You are given a dataset ( peak.dat ) from a gamma-ray experiment consisting of ~1000 hits. Each line in the file corresponds to one recorded gamma-ray event, and stores the the measured energy of the gamma-ray (in MeV). We will assume that the energies are randomly distributed about a common mean, and that each event is uncorrelated to others. Read the dataset from the enclosed file and:

- 1. Produce a histogram of the distribution of energies. Choose the number of bins wisely, i.e. so that the width of each bin is smaller than the width of the peak, and at the same time so that the number of entries in the most populated bin is relatively large. Since this plot represents randomly-collected data, plotting error bars would be appropriate.
- 2. Compute the mean and standard deviation of the distribution of energies and their statistical uncertainties. Assume the distribution is Gaussian and see the lecture notes for the formulas for the mean and variance of the sample and the formulas for the errors on these quantities.
- 3. Compute the fraction of events contained within  $\pm 1\sigma$  of the mean,  $\pm 2\sigma$  of the mean, and  $\pm 3\sigma$  of the mean (where  $\sigma$  is the standard deviation you computed in Part 2). Compare these fractions with the quantiles of the Gaussian distribution (see lecture notes)?
- 4. Fit the distribution to a Gaussian function using an unbinned fit (*Hint:* use scipi.stats.norm.fit() function), and compare the parameters of the fitted Gaussian with the mean and standard deviation computed in Part 2
- 5. Fit the distribution to a Gaussian function using a binned least-squares fit (*Hint:* use scipy.optimize.curve\_fit() function), and compare the parameters of the fitted Gaussian and their uncertainties to the parameters obtained in the unbinned fit above.
- 6. Re-make your histogram from (1) with twice as many bins, and repeat the binned least-squares fit from (3) on the new histogram. How sensitive are your results to binning?
- 7. How consistent is the distribution with a Gaussian? In other words, compare the histogram from (1) to the fitted curve, and compute a goodness-of-fit value, such as  $\chi^2/d$ .f.

```
import numpy as np
import matplotlib.pyplot as plt
from scipy.stats import norm
import scipy.optimize as fitter

# Once again, feel free to play around with the matplotlib parameters
plt.rcParams['figure.figsize'] = 8,4
plt.rcParams['font.size'] = 14

energies = np.loadtxt('peak.dat') # MeV
```

Recall plt.hist() isn't great when you need error bars, so it's better to first use np.histogram() --which returns the counts in each bin, along with the edges of the bins (there are n+1 edges for n bins). Once you find the bin centers and errors on the counts, you can make the actual plot with plt.bar(). Start with something close to bins = 25 as the second input parameter to np.histogram().

```
In [47]: # use numpy.histogram to get the counts and bin edges
    counts, bin_edges = np.histogram(energies, bins=25)

# bin_centers = 0.5*(bin_edges[1:]+bin_edges[:-1]) works for finding the bin centers

bin_centers = 0.5*(bin_edges[1:]+bin_edges[:-1])

# assume Poisson errors on the counts - errors go as the square root of the count
    counts_errors = np.sqrt(counts)

# now use plt.bar() to make the histogram with error bars (remember to label the plot)

width = (bin_edges[1]-bin_edges[0])*(0.8)

# Remember, curve_fit() will need a model function defined
```

```
def model(x, A, mu, sigma):
    '''Model function to use with curve_fit();
       it should take the form of a 1-D Gaussian'''
    exponentTerm = -(((x - mu)**2)/(2*sigma*sigma))
    return A * (np.e**(exponentTerm))
# Also make sure you define some starting parameters for curve_fit (we typically called
par0 = np.array([1,1, 1])
# You can use this to ensure the errors are greater than 0 to avoid division by 0 within
for i, err in enumerate(counts_errors):
   if err == 0:
        counts_errors[i] = 1
par, cov = fitter.curve_fit(model, bin_centers, counts, par0, counts_errors, absolute_si
stds_2 = []
stds_1 = []
stds_3 = []
mean = np.mean(bin_centers)
std = np.std(bin_centers)
for bins in bin_centers:
    if bins > (mean - (2 * std)):
        stds_2.append(bins)
    if bins > (mean - std) or bins < mean + std:</pre>
        stds_1.append(bins)
    if bins > (mean - (3 * std)) or bins < (mean + (3 * std)):
        stds_3.append(bins)
frac1 = len(bin_centers)/len(stds_1)
frac2 = len(bin_centers)/len(stds_2)
frac3 = len(bin_centers)/len(stds_3)
plt.bar(bin_centers, height=counts, width=width, yerr=counts_errors, color='r')
plt.title('Distribution of energies of 1000 gamma ray hits')
plt.ylabel('count')
plt.xlabel('Energy (MeV) \n Mean: {0:.5f} +/- {1:.5f} \n STD: {2:.5f} +/- {3:.5f} \n Fra
plt.show()
```

#### Distribution of energies of 1000 gamma ray hits 125 100 count 75 50 25 0 0.9 1.0 1.5 0.8 1.1 1.2 1.3 1.4 Energy (MeV) Mean: 1.17991 +/- 0.00332

STD: 0.20743 +/- 0.00239 Fraction within 1 std: 1.00000, 2 std: 1.00000, 3 std: 1.00000

You can use the list of energies directly as input to scipy.stats.norm.fit(); the returned values are the mean and standard deviation of a fit to the data.

```
In [10]: # Find the mean and standard deviation using scipy.stats.norm.fit()

par = norm.fit(bin_centers)

print ('mean = {0:4.2f}'.format(par[0]))
print ('sigma = {0:4.2f}'.format(par[1]))

# Compare these to those computed in the previous homework (or just find them again here
print('\n \n')

print(np.mean(bin_centers))
print(np.std(bin_centers))

mean = 1.18
sigma = 0.21
```

1.179913

0.2074299608743616

Now, using the binned values (found above with np.histogram()) and their errors use scipy.optimize.curve fit() to fit the data.

```
In [11]: # Remember, curve_fit() will need a model function defined

def model(x, A, mu, sigma):
    '''Model function to use with curve_fit();
    it should take the form of a 1-D Gaussian'''
    exponentTerm = -(((x - mu)**2)/(2*sigma*sigma))
    return A * (np.e**(exponentTerm))

# Also make sure you define some starting parameters for curve_fit (we typically called

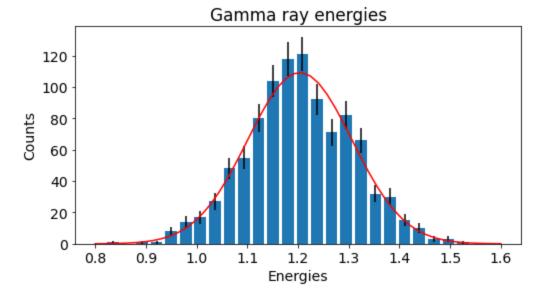
par0 = np.array([1,1, 1])

# You can use this to ensure the errors are greater than 0 to avoid division by 0 within

for i, err in enumerate(counts_errors):
```

```
if err == 0:
        counts_errors[i] = 1
par, cov = fitter.curve_fit(model, bin_centers, counts, par0, counts_errors, absolute_si
print(np.sqrt(cov[0,0]))
# Now use fitter.curve_fit() on the binned data and compare the best-fit parameters to t
# It's also useful to plot the fitted curve over the histogram you made in part 1 to che
# The best-fit value for a0 is the first element of the returned parameters
a0 = par[0]
# And the error estimate is the first diagonal element (row 0, column 0) of the covarian
error_a0 = np.sqrt(cov[0,0])
print('a0=\{0:6.3f\}+/-\{1:5.3f\}'.format(a0, error_a0))
# The best-fit value for a1 is the second element of the returned parameters
a1 = par[1]
# And the error estimate is the second diagonal element (row 1, column 1) of the covaria
error_a1 = np.sqrt(cov[1,1])
print('a1=\{0:6.3f\}+/-\{1:5.3f\}'.format(a1, error_a1))
a2 = par[2]
error_a2 = np.sqrt(cov[2, 2])
print('a2=\{0:6.3f\}+/-\{1:5.3f\}'.format(a2, error_a2))
chi_squared = np.sum(((model(bin_centers, *par)-counts)/counts_errors)**2)
print ('chi^2 = \{0:5.2f\}'.format(chi_squared))
reduced_chi_squared = (chi_squared)/(len(bin_centers)-len(par))
print ('chi^2/d.f.={0:5.2f}'.format(reduced_chi_squared))
# Plot best fit line over the data
counts_errors = np.sqrt(counts)
# now use plt.bar() to make the histogram with error bars (remember to label the plot)
width = (bin_edges[1]-bin_edges[0])*(0.8)
plt.bar(bin_centers, height=counts, width=width, yerr=counts_errors)
xfit = np.linspace(0.8, 1.6, 50)
plt.plot(xfit, model(xfit, *par), 'r-')
plt.title('Gamma ray energies')
plt.ylabel('Counts')
plt.xlabel('Energies')
plt.show()
# At this point, it's also useful to find the chi^2 and reduced chi^2 value of this binn
4.295742987893001
```

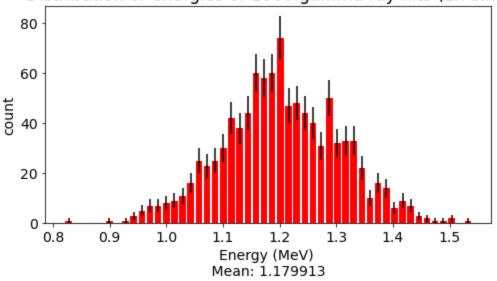
4.295/4298/893001 a0=109.163+/-4.296 a1= 1.204+/-0.003 a2=-0.103+/-0.002 chi^2 = 22.58 chi^2/d.f.= 1.03



```
In [12]: # use numpy.histogram to get the counts and bin edges
         counts, bin_edges = np.histogram(energies, bins=50)
         # bin_centers = 0.5*(bin_edges[1:]+bin_edges[:-1]) works for finding the bin centers
         bin_centers = 0.5*(bin_edges[1:]+bin_edges[:-1])
         # assume Poisson errors on the counts – errors go as the square root of the count
         counts_errors = np.sqrt(counts)
         # now use plt.bar() to make the histogram with error bars (remember to label the plot)
         width = (bin_edges[1]-bin_edges[0])*(0.8)
         plt.bar(bin_centers, height=counts, width=width, yerr=counts_errors, color='r')
         plt.title('Distribution of energies of 1000 gamma ray hits (2x bins)')
         plt.ylabel('count')
         plt.xlabel('Energy (MeV) \n Mean: {0} \n STD: {1}'.format(np.mean(bin_centers), np.std(b)
         plt.show()
         # Remember, curve_fit() will need a model function defined
         def model(x, A, mu, sigma):
             '''Model function to use with curve_fit();
                it should take the form of a 1-D Gaussian'''
             exponentTerm = -(((x - mu)**2)/(2*sigma*sigma))
             return A * (np.e**(exponentTerm))
         # Also make sure you define some starting parameters for curve_fit (we typically called
         par0 = np.array([1,1, 1])
         # You can use this to ensure the errors are greater than 0 to avoid division by 0 within
         for i, err in enumerate(counts_errors):
             if err == 0:
                 counts_errors[i] = 1
         par, cov = fitter.curve_fit(model, bin_centers, counts, par0, counts_errors, absolute_si
         # Now use fitter.curve_fit() on the binned data and compare the best-fit parameters to t
         # It's also useful to plot the fitted curve over the histogram you made in part 1 to che
         # The best-fit value for a0 is the first element of the returned parameters
```

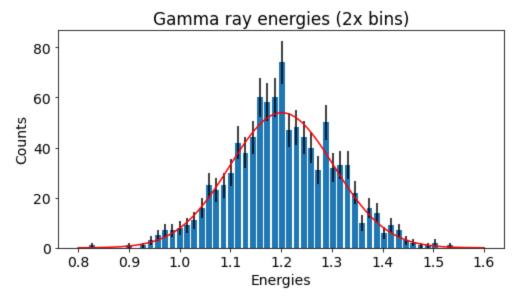
```
a0 = par[0]
# And the error estimate is the first diagonal element (row 0, column 0) of the covarian
error_a0 = np.sqrt(cov[0,0])
print('a0=\{0:6.3f\}+/-\{1:5.3f\}'.format(a0, error_a0))
# The best-fit value for a1 is the second element of the returned parameters
a1 = par[1]
# And the error estimate is the second diagonal element (row 1, column 1) of the covaria
error_a1 = np.sqrt(cov[1,1])
print('a1=\{0:6.3f\}+/-\{1:5.3f\}'.format(a1, error_a1))
a2 = par[2]
error_a2 = np.sqrt(cov[2, 2])
print('a2=\{0:6.3f\}+/-\{1:5.3f\}'.format(a2, error_a2))
chi_squared = np.sum(((model(bin_centers, *par)-counts)/counts_errors)**2)
print ('chi^2 = \{0:5.2f\}'.format(chi_squared))
reduced_chi_squared = (chi_squared)/(len(bin_centers)-len(par))
print ('chi^2/d.f.={0:5.2f}'.format(reduced_chi_squared))
# Plot best fit line over the data
counts_errors = np.sqrt(counts)
# now use plt.bar() to make the histogram with error bars (remember to label the plot)
width = (bin_edges[1]-bin_edges[0])*(0.8)
plt.bar(bin_centers, height=counts, width=width, yerr=counts_errors)
xfit = np.linspace(0.8, 1.6, 100)
plt.plot(xfit, model(xfit, *par), 'r-')
plt.title('Gamma ray energies (2x bins)')
plt.ylabel('Counts')
plt.xlabel('Energies')
plt.show()
# At this point, it's also useful to find the chi^2 and reduced chi^2 value of this binn
```

### Distribution of energies of 1000 gamma ray hits (2x bins)



STD: 0.20755458086810513

a0=53.895+/-2.164 a1= 1.201+/-0.003 a2=-0.103+/-0.003  $chi^2 = 39.91$  $chi^2/d.f. = 0.85$ 



Repeat this process with twice as many bins (i.e. now use bins = 50 in np.histogram(), or a similar value). Compute the  $\chi^2$  and reduced  $\chi^2$  and compare these values, along with the best-fit parameters between the two binned fits. Feel free to continue to play with the number of bins and see how it changes the fit.

In [ ]: