

Phys 641 Lecture 4

Recap: Correlated Noise

- Correlated noise is common in wide variety of physical systems.
- Linear algebra formulation of χ^2 is accurate if $N_{ij} = \langle n_i n_j \rangle$ and noise is Gaussian
- Rotating into eigenvalues of noise matrix makes noise uncorrelated.

Stationary Noise

- For case of $N_{ij}=g(i-j)$, noise becomes uncorrelated in Fourier space, $\langle F_k^* F_{k'} \rangle = G_k \delta(k-k')$
- If T is (normalized) Fourier transform operator, then we can write $\chi^2 = r^T T^T N^{-1} T r$ where the residual $r = (\text{data} - \text{model})$ and N is now diagonal.
- As per problem set, this can still be extremely useful approximation even if noise isn't quite stationary.

Sample Problem

- We have seen error bars in LLS come from $(A^T N^{-1} A)^{-1}$. If condition number is bad then this may cause problems.
- Slightly more subtle - different columns of A could have different units, so condition number could change.
- Show that can rewrite $A^T N A$ as BCB where C has ones along the diagonal and B is diagonal. What are the elements of B ? What is inverse of BCB ?
- In this case, C is often called the correlation matrix. Its condition number is usually more relevant to instability. Diagonal elements of C^{-1} tell you the error bar hit from correlated parameters.
- (came up on Friday cosmo-ph...)

χ^2 Example - template fitting

- Let's fit a 1-parameter template to data, but possibly want to shift it (e.g. fitting a for a source amplitude at various positions).
- If A is n by 1 , then $A^T N^{-1} A$ is a scalar.
- now we have $m = A^T N^{-1} d / (A^T N^{-1} A)$.
- If N is stationary (i.e. $N_{ij} = f(i-j)$) and we shift the template $A_i \rightarrow A_{i+\delta}$, how does the denominator depend on δ ?
- Up to an overall constant, we can make $N^{-1} d$ and dot it against the various shifted A 's.

Template fitting ctd.

- What if we wanted to plot amplitude in terms of SNR?
- recall $\text{Var}(\mathbf{m}) = (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A})^{-1}$, so $\sigma(\mathbf{m}) = (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A})^{-1/2}$.
- So, $\text{SNR} = \text{signal/noise} = \mathbf{A}^T \mathbf{N}^{-1} \mathbf{d} / \mathbf{A}^T \mathbf{N}^{-1} \mathbf{A} / (\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A})^{-1/2} = \mathbf{A}^T \mathbf{N}^{-1} \mathbf{d} / \sqrt{\mathbf{A}^T \mathbf{N}^{-1} \mathbf{A}}$.

Code Example

```
import numpy
from matplotlib import pyplot as plt

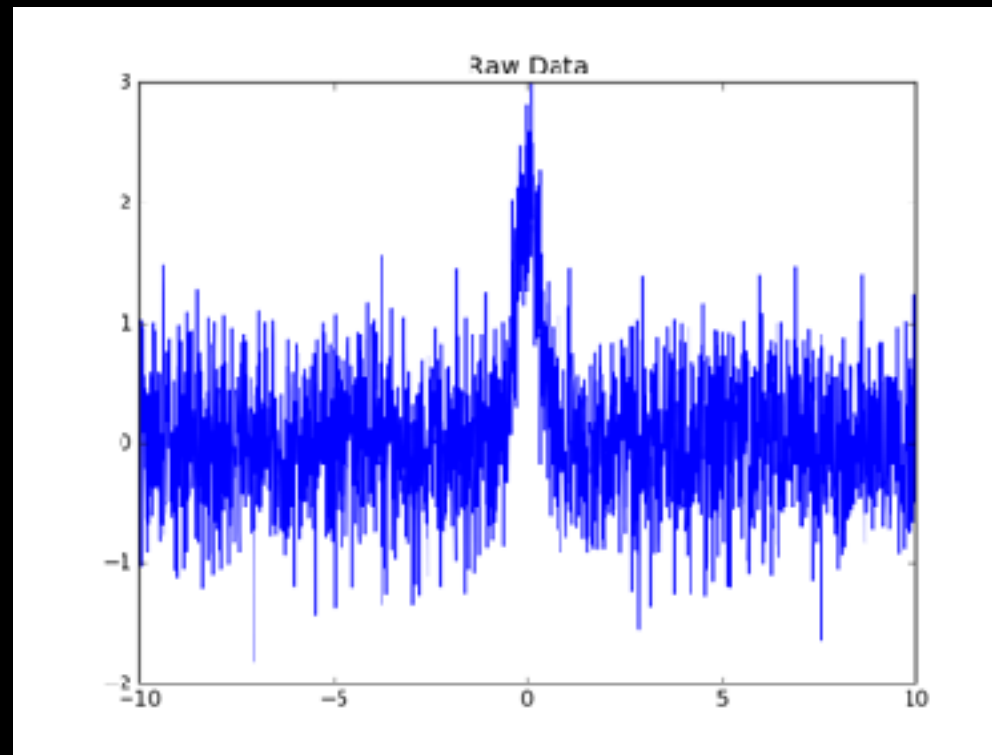
dx=0.01
noise=0.5
Ninv=1.0/noise**2
x=numpy.arange(-10,10,dx)
n=len(x)

x0=0
amp_true=2.0

sig=0.3
template=numpy.exp(-0.5*(x-x0)**2/sig**2)

dat=template*amp_true+numpy.random.randn(n)*noise

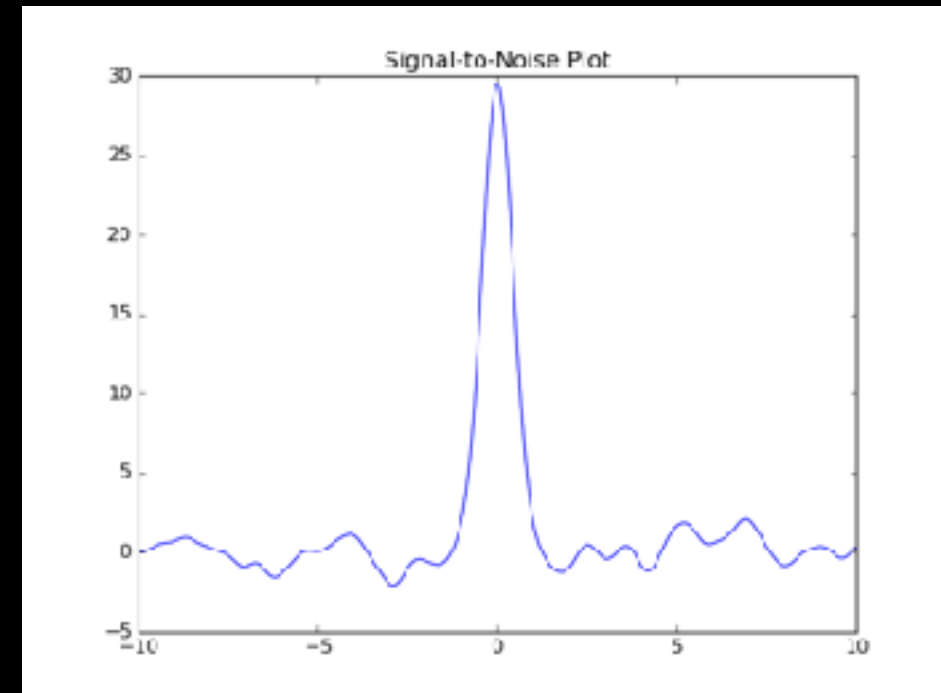
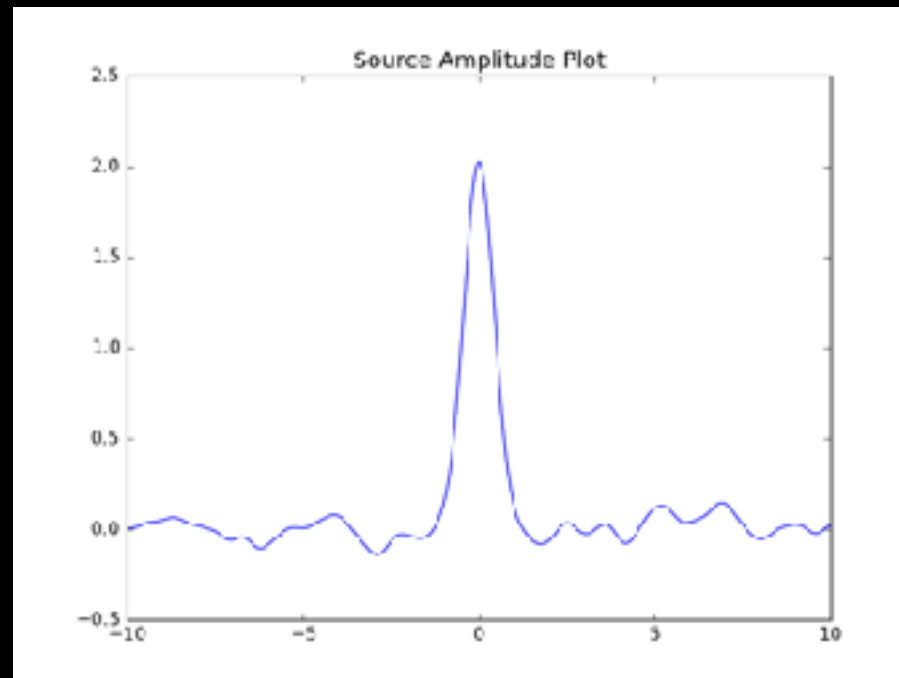
snr=numpy.zeros(n)
amp=numpy.zeros(n)
dat_filt=Ninv*dat
denom=(numpy.dot(template,Ninv*template))
rt_denom=numpy.sqrt(denom)
for i in range(n):
    template=numpy.exp(-0.5*(x-x[i])**2/sig**2)
    rhs=numpy.dot(template,dat_filt)
    snr[i]=rhs/rt_denom
    amp[i]=rhs/denom
```



```
plt.clf();
plt.plot(x,snr);
plt.title('Signal-to-Noise Plot')
plt.savefig('snr_plot.png')

plt.clf();
plt.plot(x,amp)
plt.title('Source Amplitude Plot')
plt.savefig('amp_plot.png')

plt.clf();
plt.plot(x,dat)
plt.title('Raw Data')
plt.savefig('dat_raw.png')
```



But wait!

- We took $\sum d(t)a(t-\tau)$. But, this is just the correlation of d with a . We can do this quickly using Fourier transforms.
- Alternatively, let $a^\diamond = a(-t)$. Then this is $\sum d(t)a^\diamond(\tau-t) = d \otimes a^\diamond$. By convolution theorem, this is $\text{IFT}(\text{FT}(d)\text{FT}(a^\diamond))$.
- However, $\text{FT} = \sum f(x)\exp(-2\pi i k x/N)$. $\text{FT}(f(-x)) = \sum f(-x)\exp(-2\pi i k x/N) = \sum f(x)\exp(2\pi i k x/N) = F^*(k)$. So, our output is just $\text{IFT}(\text{FT}(d)\text{FT}^*(a))$.
- NB - if a is symmetric, then $A(k)$ is real (why?) and we can skip the conjugates.

How about changing noise?

- Often we're searching for things where noise is varying - say, searching for sources in a map that is deeper in the middle.
- Top: $A^T N^{-1} d$ is easy - correlate A with $N^{-1} d$.
- Bottom: $A^T N^{-1} A$ we can correlate N^{-1} with A^2 .
- Amplitude stays ratio of the two (bottom is now a vector instead of a scalar), and SNR is $\text{top} / \sqrt{(\text{bottom})}$

Matched Filter

- We are beginning on a very widespread class of techniques called *matched filters*.
- In various forms, they show up in GW analysis, photometry in optical images, finding galaxy clusters in CMB maps, finding radar return echos...
- Extendable to non-independent noise, multiple (possibly correlated) simultaneous/multi-frequency datasets, many others.
- All arises from writing down χ^2 and minimizing. Make a habit of this.

Fourier Transforms

- On computer, discrete Fourier transform usually defined as $\sum f(x)\exp(-2\pi i kx/N)$, sum goes from 0 to N-1
- Can write this as a matrix multiply where ij^{th} element is $\exp(-2\pi i kx/N)$
- What is i^{th} column dotted with j^{th} column?
- what is the inverse of this matrix?
- Strongly encourage you to just memorize this. Numerical factors important in real life...

Some Fourier Theorems

- $x \rightarrow x + \delta, F(k) \rightarrow F(k) \exp(-2\pi i \delta k / N)$
- $x \rightarrow -x, F(k) \rightarrow F^*(k)$
- Convolution theorem: $FT(f \otimes g) = FT(f) * FT(g)$ where $f \otimes g = \int f(\tau) g(t - \tau) d\tau$.

Periodicity

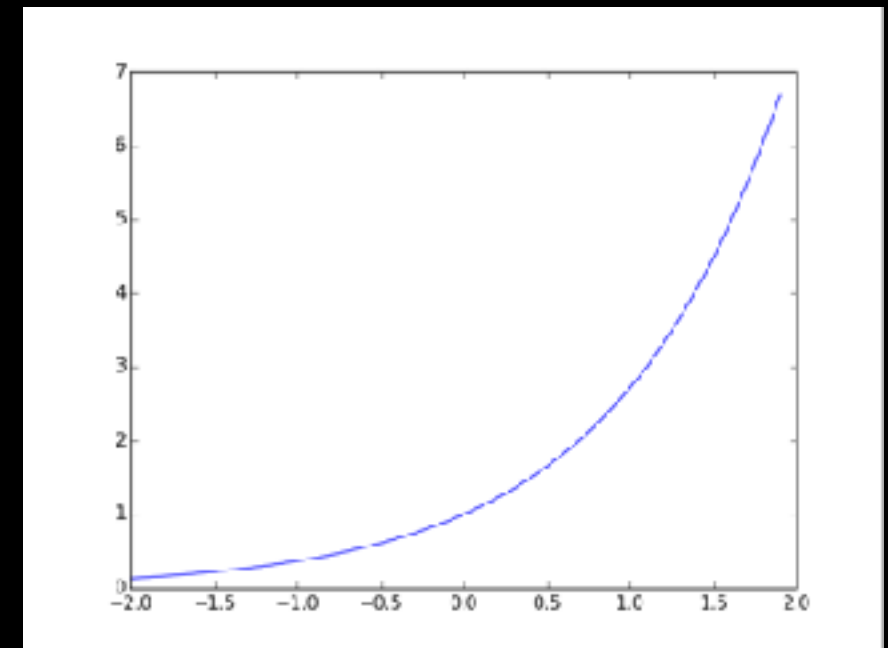
- $f(x) = \sum F(k) \exp(-2\pi i k x / N)$
- What is $f(x+N)$? $\sum F(k) \exp(-2\pi i k (x+N) / N)$
- $= \sum F(k) \exp(-2\pi i k) \exp(-2\pi i k x / N)$.
- $\exp(-2\pi i k) = 1$ for integer k , so $f(x+N) = f(x)$
- DFT's are periodic - they just repeat themselves ad infinitum.

Periodicity

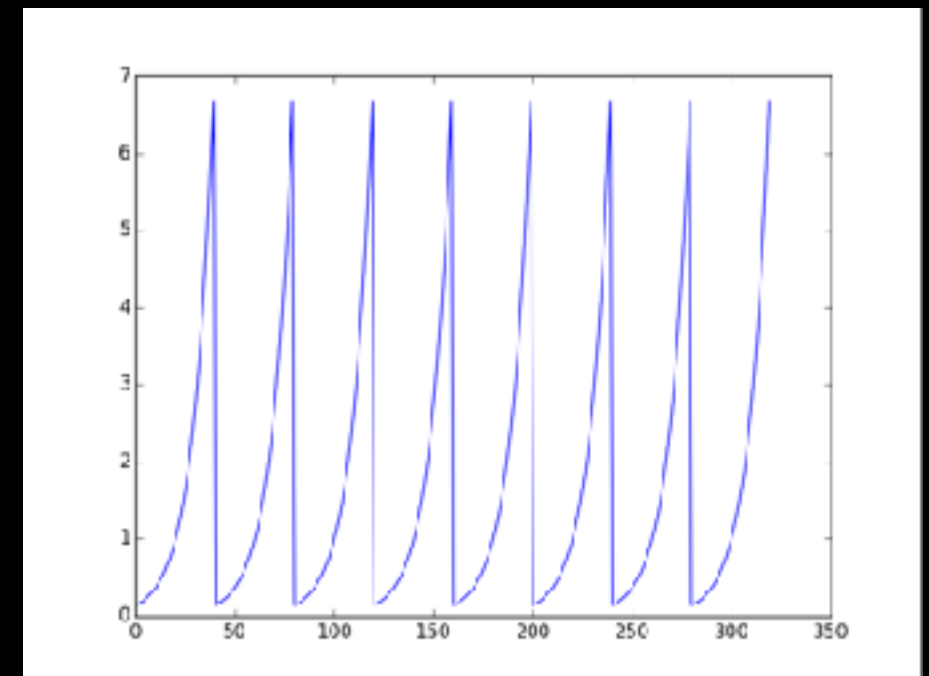
```
import numpy
from matplotlib import pyplot as plt

x=numpy.arange(-2,2,0.1)
y=numpy.exp(x)
plt.plot(x,y)
plt.savefig('fft_exp')
plt.show()

yy=numpy.concatenate((y,y))
yy=numpy.concatenate((yy,yy))
yy=numpy.concatenate((yy,yy))
plt.plot(yy)
plt.savefig('fft_exp_repeating')
plt.show()
```



- You may think you're taking top transform. You're not - you're taking the bottom one.
- In particular, jumps from right edge to left will strongly affect DFT



Aliasing

- $f(x) = \sum F(k) \exp(2\pi i k x / N)$
- What if I had higher frequency, $k > N$? let $k^* = k - N$ (i.e. k^* low freq.)
- $f(x) = \sum F(k) \exp(2\pi i (k^* + N) x / N) = \sum F(k) \exp(2\pi i x) \exp(2\pi i k^* x / N)$
- for x integer, middle term goes away: $\sum F(k^* + N) \exp(2\pi i k^* x / N)$
- High frequencies behave exactly like low frequencies - power has been *aliased* into main frequencies of DFT.
- Always keep this in mind! Make sure samples are fine enough to prevent aliasing.

Negative Frequencies

- All frequencies that are N apart behave identically
- DFT has frequencies up to $(N-1)$.
- Frequency $(N-1)$ equivalent to frequency (-1) . You will do better to think of DFT as giving frequencies $(-N/2, N/2)$ than frequencies $(0, N-1)$
- *Sampling (Nyquist) theorem*: if function is band-limited - highest frequency is ν - then I get full information if I sample *twice* per frequency, $dt = 1/(2\nu)$. Factor of 2 comes from aliasing.

Exercise

- What is the PDF of the sum of two independent random variables?
- Let's now add n independent identically distributed random variables together.
- Without (much) loss of generality can rescale so mean is 0 and variance is 1.

Continue...

- If I want to rescale sum so mean stays zero and variance stays 1, what factor should I rescale by?
- If $F(k)$ is FT of probability distribution, what is $F(0)$? What is $F(\delta)$ like if $\delta \ll 1/\sigma$?
- To leading order, $F(k) = 1 - ak^2 + \dots$
- What happens to $F(k)$ if $x_i \rightarrow x_i/a$?
- Leaves us with $\sum x_i / \sqrt{n}$ has FT $(1 - ak^2/n)^n$ as n goes to infinity. This turns into $\exp(-ak^2)$, so FT of distribution of sum is a Gaussian.
- If that's true, FT of a Gaussian is a Gaussian, so distribution of sum must also be a Gaussian. ergo central limit theorem.

Comin up..

- Getting very close to running on actual data.
- LIGO tutorial very handy - will be using things from <https://losc.ligo.org/tutorials/>
- Think about how to estimate power spectra...

