

Recap: Matched Filter

- If we have some template A that we want to find in our data, then our best-fit value for the amplitude is:
 $m = A^T N^{-1} d / A^T N^{-1} A$.
- If the noise is stationary and we want to shift the template, we can do so by taking the cross-correlation of A with $N^{-1} d$. Modulo edge effects, $A^T N^{-1} A$ is constant and so only needs to be computed once.

SNR

- Very often we want not just the amplitude, but to know if the matched filter has found something statistically significant.
- The error on m is just $\sqrt{(A^T N^{-1} A)^{-1}}$, so the matched filter SNR is just correlation of $(A, N^{-1} d) / \sqrt{A^T N^{-1} A}$.
- Of course, if your errors aren't right, maybe you want to get the error by looking at the RMS of the matched filter.

How about changing noise?

- Often we're searching for things where noise is varying - say, searching for sources in a map that is deeper in the middle.
- Top: $A^T N^{-1} d$ is easy - correlate A with $N^{-1} d$.
- Bottom: $A^T N^{-1} A$ we can correlate N^{-1} with A^2 .
- Amplitude stays ratio of the two (bottom is now a vector instead of a scalar), and SNR is $\text{top} / \sqrt{(\text{bottom})}$

Fourier Stuff

- Matched filters usually require Fourier transforms.
- Fourier transforms on the computer are discrete, since we don't have infinite computing power/RAM...
- Understanding how the computer does them is important.
- Usual discrete Fourier transform (DFT) is $F(k) = \sum f(x) \exp(-2\pi i k x / N)$ for N data points (not the noise matrix for now) where x goes from 0 to $N-1$, and k goes from 0 to $N-1$.

DFT Examples

- What is the DFT of $f(x)=1$ if $x=0$, otherwise 0?
- What is the DFT of $f(x)=1$ everywhere?
- If we have $\text{DFT}(f(x))$, what is $\text{DFT}(f(x+j))$?
- If $f(x)$ is real, what can we say about $F(k)$?

DFT Examples

- What is the DFT of $f(x)=1$ if $x=0$, otherwise 0?
 - $F(k)=1$ everywhere
- What is the DFT of $f(x)=1$ everywhere?
 - $F(0)=n$, otherwise 0
- If we have $\text{DFT}(f(x))$, what is $\text{DFT}(f(x+j))$?
 - $F_{\text{new}}(k) \rightarrow \exp(-2\pi i k j / N) F(k)$
- If $f(x)$ is real, what can we say about $F(k)$?
 - $F(-k)=F^*(k)$. NB - `numpy.fft.rfft/irfft` will do real-to-complex transforms.

DFT as Rotation

- if $F(k) = \sum f(x) \exp(-2\pi i k x / N)$
- For a single k value, this is just a dot product:
 $F(k_j) = f(x) \cdot \exp(-2\pi i k_j x / N)$
- Once again, we can write this down using linear algebra:
 $F = T f$ where $T_{mn} = \exp(-2\pi i m n / N)$ (switching from i, j to m, n to avoid confusion with $\sqrt{-1}$)

DFT as Rotation 2

- First, note that T is symmetric, since swapping n & m doesn't change anything.
- What is the dot product of columns of T ? Note that we want the dot product of a vector with itself to be the length, which implies we conjugate one copy.
- So, want $\sum \exp(2\pi i m n / N) \exp(-2\pi i m n' / N)$
 - This equals $N\delta(m-n)$
- What is T^{-1} then? just T^*/N , so $T^*T = NI$.
- Tells us that the inverse transform is just $f(x) = 1/N \sum F(k) \exp(2\pi i k x / N)$

Normalizations

- NB - a more natural mathematical definition would be $T = \exp(-2\pi i m n / N) / \sqrt{N}$. Then T^{-1} would be T^* .
- Not a good idea to put in extra square roots on a computer though...
- Becomes important when estimating noise in Fourier space.
- What is Parseval's theorem using usual DFT definitions?

Periodicity

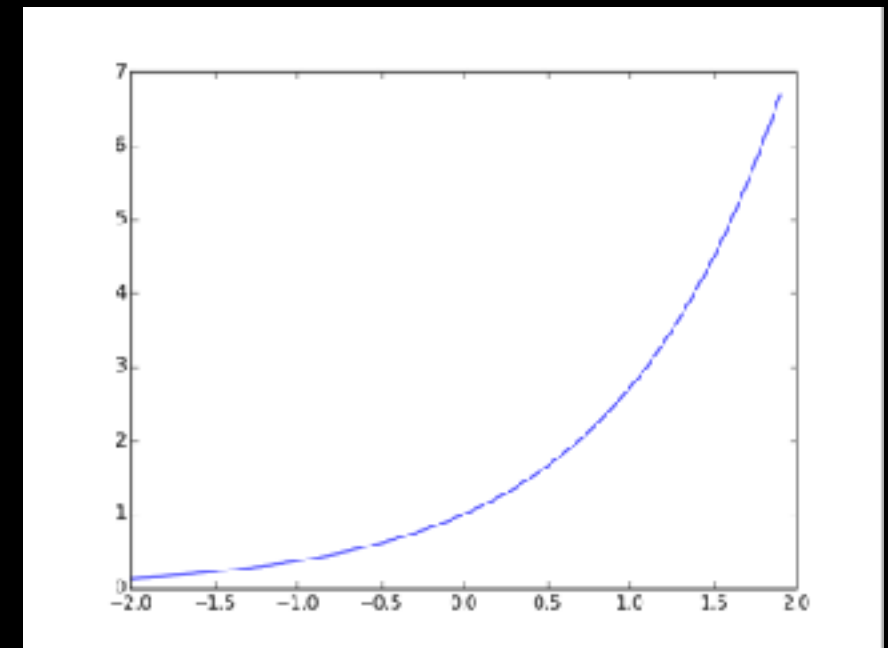
- $f(x) = \sum F(k) \exp(2\pi i k x / N)$
- What is $f(x+N)$? $\sum F(k) \exp(2\pi i k (x+N) / N)$
- $= \sum F(k) \exp(2\pi i k) \exp(2\pi i k x / N)$.
- $\exp(2\pi i k) = 1$ for integer k , so $f(x+N) = f(x)$
- DFT's are periodic - they just repeat themselves ad infinitum.

Periodicity

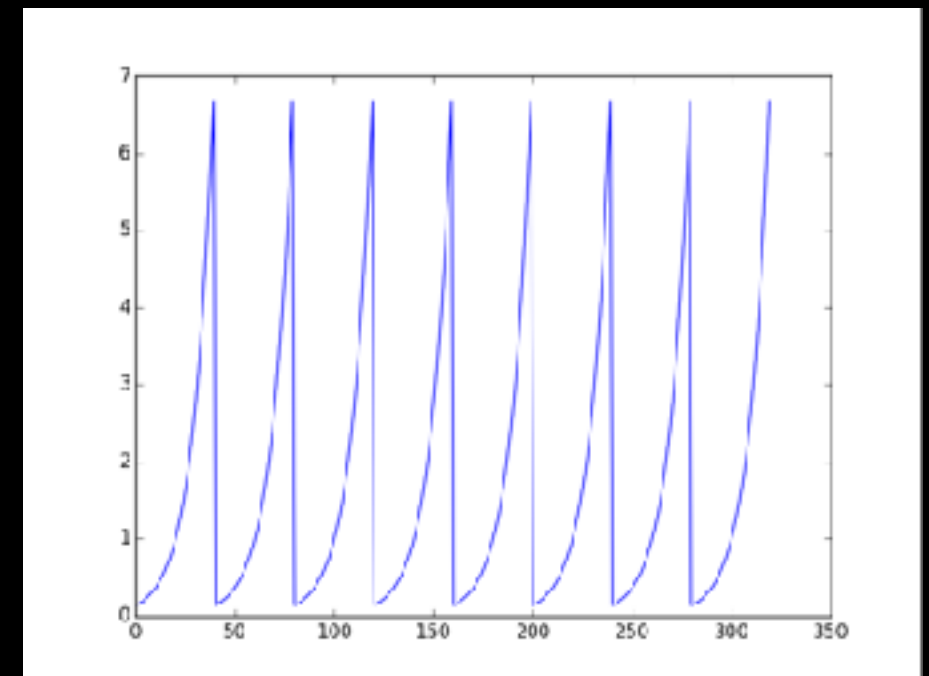
```
import numpy
from matplotlib import pyplot as plt

x=numpy.arange(-2,2,0.1)
y=numpy.exp(x)
plt.plot(x,y)
plt.savefig('fft_exp')
plt.show()

yy=numpy.concatenate((y,y))
yy=numpy.concatenate((yy,yy))
yy=numpy.concatenate((yy,yy))
plt.plot(yy)
plt.savefig('fft_exp_repeating')
plt.show()
```



- You may think you're taking top transform. You're not - you're taking the bottom one.
- In particular, jumps from right edge to left will strongly affect DFT



Aliasing

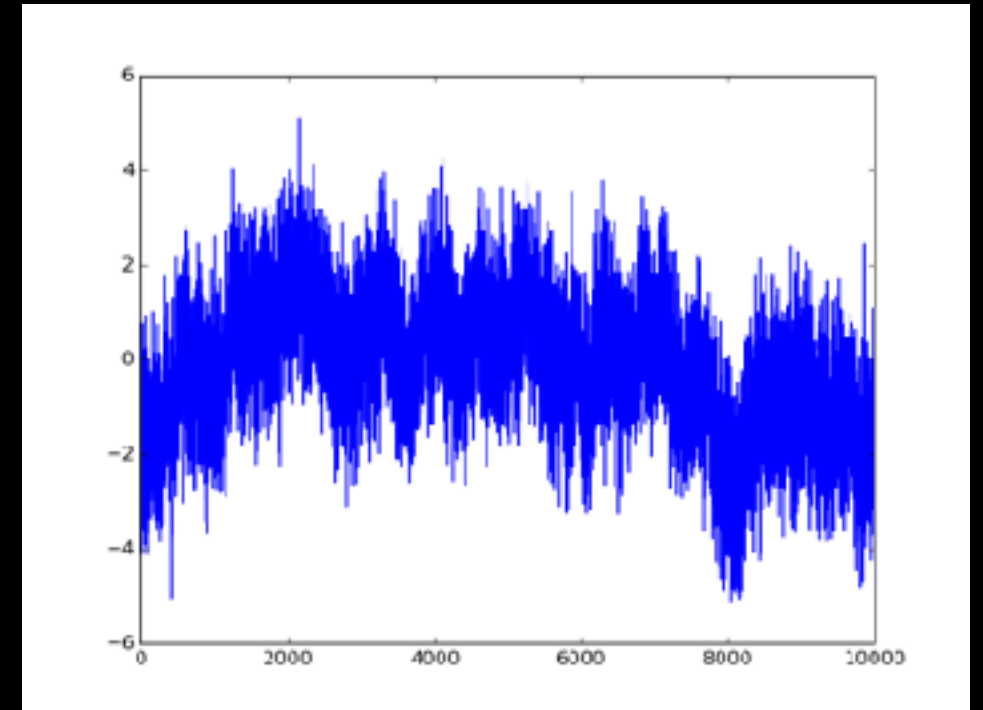
- $f(x) = \sum F(k) \exp(2\pi i k x / N)$
- What if I had higher frequency, $k > N$? let $k^* = k - N$ (i.e. k^* low freq.)
- $f(x) = \sum F(k) \exp(2\pi i (k^* + N) x / N) = \sum F(k) \exp(2\pi i x) \exp(2\pi i k^* x / N)$
- for x integer, middle term goes away: $\sum F(k^* + N) \exp(2\pi i k^* x / N)$
- High frequencies behave exactly like low frequencies - power has been *aliased* into main frequencies of DFT.
- Always keep this in mind! Make sure samples are fine enough to prevent aliasing.

Negative Frequencies

- All frequencies that are N apart behave identically
- DFT has frequencies up to $(N-1)$.
- Frequency $(N-1)$ equivalent to frequency (-1) . You will do better to think of DFT as giving frequencies $(-N/2, N/2)$ than frequencies $(0, N-1)$
- *Sampling (Nyquist) theorem*: if function is band-limited - highest frequency is ν - then I get full information if I sample *twice* per frequency, $dt = 1/(2\nu)$. Factor of 2 comes from aliasing.

Non-white noise

- Let's say we wanted to make some non-white noise. Typical case: variance is say $1/f$ plus a constant (white noise).
- Characterized by *knee frequency*, where $1/f$ part matches white noise, and index since not always exactly $1/f^{1.0}$.
- Often called pink noise, flicker noise, or just $1/f$ noise. In astro, $1/f$ usually doesn't mean index is -1 , just negative.



```
import numpy as np
from matplotlib import pyplot as plt

#specify what sort of noise/data we want
n=10000
alpha=-2.0
knee=100.0

#one easy thing to do is generate white noise, then scale its Fourier transform
dat=np.random.randn(n)
datft=np.fft.rfft(dat)
nuvec=np.arange(len(datft))+1 #the plus one is so the bottom frequency isn't zero

filtvec=np.sqrt(1+(nuvec/knee)**alpha)
datft=datft*filtvec
dat_pink=np.fft.irfft(datft)

plt.ion()
plt.clf()
plt.plot(dat_pink)
```

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