

# A VERSATILE THREE-DIMENSIONAL RAY TRACING COMPUTER PROGRAM FOR RADIO WAVES IN THE IONOSPHERE

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## PREFACE

This report documents the latest version of the three-dimensional ray tracing program originally described in "A Three-Dimensional Ray Tracing Computer Program," by R. M. Jones, ESSA Technical Report IER 17-ITSA 17, and later modified in "Modifications to the Three-Dimensional Ray Tracing Program Described in IER 17-ITSA 17," by R. M. Jones, ESSA Technical Memorandum ERLTM-ITS 134. This report replaces all of the material contained in the above two reports.



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R. Michael Jones\* and Judith J. Stephenson\*\*

This report describes an accurate, versatile FORTRAN computer program for tracing rays through an anisotropic medium whose index of refraction varies continuously in three dimensions. Although developed to calculate the propagation of radio waves in the ionosphere, the program can be easily modified to do other types of ray tracing because of its organization into subroutines.

The program can represent the refractive index by either the Appleton-Hartree or the Sen-Wyller formula, and has several ionospheric models for electron density, perturbations to the electron density (irregularities), the earth's magnetic field, and electron collision frequency.

For each path, the program can calculate group path length, phase path length, absorption, Doppler shift due to a time-varying ionosphere, and geometrical path length. In addition to printing these parameters and the direction of the wave normal at various points along the ray path, the program can plot the projection of the ray path on any vertical plane or on the ground and punch the main characteristics of each ray path on cards.

The documentation includes equations, flow charts, program listings with comments, definitions of program variables, deck set-ups, descriptions of input and output, and a sample case.

Key words: Ray tracing, computer program, radio waves, ionosphere, three-dimensional, Appleton-Hartree formula, Sen-Wyller formula.

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## 1. INTRODUCTION

This report describes a three-dimensional ray tracing program written in FORTRAN language for the CDC-3800 computer. Copies of the program deck are available from the Institute for Telecommunication Sciences.

Earlier versions of this program have been in use now for over nine years, both by us and by people scattered all over the world. During that time we have improved and modified the program to the extent that we now need to document these changes so that the present program will be easier to use. We have included the input parameter forms that we use to request ray path calculations because they give nearly all the necessary input data and describe the electron density, collision frequency, and magnetic field models.

## 2. GENERAL DESCRIPTION

This computer program traces the path of radio wave through a user-specified model of the ionosphere when given the transmitter location (longitude, latitude, and height above the ground), the frequency of the wave, the direction of transmission (both elevation and azimuth), the receiver height, and the maximum number of hops wanted.

## 3. RAY TRACING EQUATIONS

The program calculates ray paths by numerically integrating Hamilton's equations. Lighthill (1965) gives Hamilton's equations in four dimensions (three spatial and one time) for Cartesian coordinates. Haselgrove (1954) gives Hamilton's equations in three dimensions for spherical polar coordinates. Combining the two gives Hamilton's equations in four dimensions in which the three spatial coordinates are spherical polar (see Table 1 for a definition of the symbols):

$$\frac{dr}{d\tau} = \frac{\partial H}{\partial k_r}, \quad (1)$$

Table 1. List of the More Important Symbols

---

A	In section 3, absorption in decibels.
$B_o$	Magnetic induction of earth's magnetic field.
c	Speed of electromagnetic waves in free space.
C	Cosine of the angle of incidence on the ionosphere.
e	Charge of the electron (a negative number).
$F(w)$	$F(w) = w C_{3/2}(w) + i \frac{5}{2} C_{5/2}(w) = \frac{1}{3/2!} \int_0^{\infty} \frac{t^{3/2} \exp(-t) dt}{w - it}$ (Davies, 1965, p. 86)
f	Wave frequency.
$\Delta f$	Frequency shift of a wave due to a time varying ionosphere (sometimes called Doppler shift).
$f_H$	Gyro frequency for electrons, $ e B_o/2\pi m$ .
$f_N$	Plasma frequency, $(Ne^2/4\pi^2\epsilon_o m)^{1/2}$ .
$G(w)$	$wF(w)$ .
H	Hamiltonian.
$k_r, k_\theta, k_\varphi$	Components of the propagation vector in the r, $\theta$ , $\varphi$ directions--a vector perpendicular to the wave front having a magnitude $2\pi/\lambda = \omega/v$ .
m	Mass of electron.
N	Number of electrons per unit volume.
n	Phase refractive index (in general complex).
$n'$	Group refractive index (in general complex).
P	Phase path length, phase of wave divided by free space wave number $2\pi/\lambda_o$ .
$P'$	Group path length $ct$ .
$r, \theta, \varphi$	Coordinates of a point in spherical polar coordinates.
s	Geometric ray path length.
S	Sine of the angle of incidence on the ionosphere.

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Table 1. (Continued)

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t	Time, travel time of a wave packet.
U	$1 - iZ$ in the Appleton-Hartree formula or $Z/F(1/Z) = 1/G(1/Z)$ in the Sen-Wyller formula.
$V_r, V_\theta, V_\varphi$	Components of the wave normal direction in the $r, \theta,$ and $\varphi$ directions, normalized so that $V_r^2 + V_\theta^2 + V_\varphi^2 = \text{Real } \{n^2\}.$
v	Phase velocity.
X	$\omega_N^2/\omega^2 = f_N^2/f^2 = Ne^2/(\epsilon_0 m \omega^2).$
Y	$\omega_H/\omega = f_H/f.$
$Y_L$	$Y \cos \psi.$
$Y_T$	$Y \sin \psi.$
Z	$v/\omega$ or $v_m/\omega.$
$\epsilon_0$	Electric permittivity of free space.
$\theta$	Colatitude in spherical polar coordinates.
$\lambda$	Wavelength.
$\lambda_0$	Wavelength in free space.
$\nu$	Electron collision frequency.
$\nu_m$	Mean electron collision frequency.
$\rho$	Characteristic wave polarization (definition in Table 6).
$\rho_L$	Longitudinal polarization (definition in Table 6).
$\tau$	Independent variable in Hamilton's equations.
$\varphi$	Longitude in spherical polar coordinates.
$\psi$	Angle between wave normal and $-B_0.$
$\omega$	$2\pi f$ , angular wave frequency.
$\Delta\omega$	$2\pi\Delta f$ , angular frequency shift.
$\omega_H$	$2\pi f_H =  e B_0/m$ , angular gyrofrequency.
$\omega_N$	$2\pi f_N = (Ne^2/\epsilon_0 m)^{\frac{1}{2}}$ , angular plasma frequency.

---

$$\frac{d\theta}{d\tau} = \frac{1}{r} \frac{\partial H}{\partial k_\theta} , \quad (2)$$

$$\frac{d\varphi}{d\tau} = \frac{1}{r \sin\theta} \frac{\partial H}{\partial k_\varphi} , \quad (3)$$

$$\frac{dt}{d\tau} = - \frac{\partial H}{\partial \omega} , \quad (4)$$

$$\frac{dk_r}{d\tau} = - \frac{\partial H}{\partial r} + k_\theta \frac{d\theta}{d\tau} + k_\varphi \sin\theta \frac{d\varphi}{d\tau} , \quad (5)$$

$$\frac{dk_\theta}{d\tau} = \frac{1}{r} \left( - \frac{\partial H}{\partial \theta} - k_\theta \frac{dr}{d\tau} + k_\varphi r \cos\theta \frac{d\varphi}{d\tau} \right) , \quad (6)$$

$$\frac{dk_\varphi}{d\tau} = \frac{1}{r \sin\theta} \left( - \frac{\partial H}{\partial \varphi} - k_\varphi \sin\theta \frac{dr}{d\tau} - k_r r \cos\theta \frac{d\theta}{d\tau} \right) , \quad (7)$$

$$\frac{d\omega}{d\tau} = \frac{\partial H}{\partial t} , \quad (8)$$

The variables  $r$ ,  $\theta$ ,  $\varphi$  are the spherical polar coordinates of a point on the ray path;  $k_r$ ,  $k_\theta$ , and  $k_\varphi$  are the components of the propagation vector (wave normal direction normalized so that in free space

$$k_r^2 + k_\theta^2 + k_\varphi^2 = \frac{\omega^2}{c^2} , \quad (9)$$

where  $\omega = 2\pi f$  is the angular frequency of the wave and  $c$  is the speed of propagation of electromagnetic waves in free space);  $t$  is time, in (4) it is the propagation time of a wave packet, in (8) it expresses the variation with time of a time varying medium;  $\tau$  is a parameter whose value depends on the choice of the Hamiltonian  $H$ .

For actual calculation, the ray tracing program uses group path  $P' = ct$  as the independent variable because the derivatives with respect to  $P'$  are independent of the choice of Hamiltonian, allowing the program to switch Hamiltonians in the middle of a path. This choice automatically causes the program to take smaller steps in real path length near reflection where the calculations are more critical. The resulting equations obtained by dividing (1) through (8) by  $c$  times (4) are:

$$\frac{dr}{dP'} = -\frac{1}{c} \frac{\partial H/\partial k_r}{\partial H/\partial \omega} , \quad (9)$$

$$\frac{d\theta}{dP'} = -\frac{1}{rc} \frac{\partial H/\partial k_\theta}{\partial H/\partial \omega} , \quad (10)$$

$$\frac{d\varphi}{dP'} = -\frac{1}{rc \sin\theta} \frac{\partial H/\partial k_\varphi}{\partial H/\partial \omega} , \quad (11)$$

$$\frac{dk_r}{dP'} = \frac{1}{c} \frac{\partial H/\partial r}{\partial H/\partial \omega} + k_\theta \frac{d\theta}{dP'} + k_\varphi \sin\theta \frac{d\varphi}{dP'} , \quad (12)$$

$$\frac{dk_\theta}{dP'} = \frac{1}{r} \left( \frac{1}{c} \frac{\partial H/\partial \theta}{\partial H/\partial \omega} - k_\theta \frac{dr}{dP'} + k_\varphi r \cos\theta \frac{d\varphi}{dP'} \right) , \quad (13)$$

$$\frac{dk_\varphi}{dP'} = \frac{1}{r \sin\theta} \left( \frac{1}{c} \frac{\partial H/\partial \varphi}{\partial H/\partial \omega} - k_\varphi \sin\theta \frac{dr}{dP'} - k_\varphi r \cos\theta \frac{d\theta}{dP'} \right) , \quad (14)$$

$$\begin{aligned} \frac{d(\Delta f)}{dP'} &= \frac{1}{2\pi} \frac{d\Delta\omega}{dP'} = \frac{1}{2\pi} \frac{d\omega}{dP'} \\ &= -\frac{1}{2\pi} \frac{\partial H/\partial t}{\partial H/\partial \omega} . \end{aligned} \quad (15)$$

Equation (15) for the frequency shift of a wave propagating through a time varying medium follows directly from Hamilton's equations (4) and (8). An alternative derivation is given by Bennett (1967). For large frequency shifts, the frequency shift should be accumulated along the ray path and the shifted frequency used in calculations at each point on the ray path. Equations (1) through (8) imply that all eight dependent variables vary along the path, and that at each point on the path the instantaneous value of all parameters (including frequency) is used in further evaluations of the equations. However, the time variation of the ionosphere due to natural causes (such as solar flares) is so slow that the resulting frequency shifts are small enough (less than one part in  $10^5$ ) to have negligible effect on the propagation. For this reason, the program calculates frequency shift to compare with frequency shift measurements, but does not adjust the carrier frequency of the wave used in the propagation calculations.

The first six differential equations (9) through (14) are always integrated. The user can choose whether to have the program integrate (15) to calculate the frequency shift.

There are three other quantities that can be calculated by integration along the ray path. The phase path  $P$  (phase divided by the free space wavenumber  $2\pi/\lambda_0 = \omega/c$ ) is calculated by integrating

$$\begin{aligned} \frac{dP}{dP'} &= \frac{c}{\omega} \left( k_r \frac{dr}{dP'} + k_\theta r \frac{d\theta}{dP'} + k_\varphi r \sin\theta \frac{d\varphi}{dP'} \right) \\ &= -\frac{1}{\omega} \frac{k_r \frac{\partial H}{\partial k_r} + k_\theta \frac{\partial H}{\partial k_\theta} + k_\varphi \frac{\partial H}{\partial k_\varphi}}{\partial H / \partial \omega} . \end{aligned} \quad (16)$$

If the absorption per wavelength is small (as it must be for this type of ray tracing to be valid), then an approximate formula can be integrated

to give the absorption in decibels

$$\begin{aligned}\frac{dA}{dP'} &= -\frac{10}{\log_e 10} \frac{\omega}{c} \frac{\text{imag}(\frac{\omega^2}{c^2} n^2)}{k_r^2 + k_\theta^2 + k_\varphi^2} \frac{dP}{dP'} \\ &= \frac{10}{\log_e 10} \frac{\text{imag}(\frac{\omega^2}{c^2} n^2)}{k_r^2 + k_\theta^2 + k_\varphi^2} \frac{k_r \frac{\partial H}{\partial k_r} + k_\theta \frac{\partial H}{\partial k_\theta} + k_\varphi \frac{\partial H}{\partial k_\varphi}}{c \frac{\partial H}{\partial \omega}}, \quad (17)\end{aligned}$$

where  $n$  is the (complex) phase refractive index. The geometrical path length of the ray can be calculated by integrating

$$\begin{aligned}\frac{ds}{dP'} &= \sqrt{\left(\frac{dr}{dP'}\right)^2 + r^2 \left(\frac{d\theta}{dP'}\right)^2 + r^2 \sin^2 \theta \left(\frac{d\varphi}{dP'}\right)^2} \\ &= -\frac{\sqrt{\left(\frac{\partial H}{\partial k_r}\right)^2 + \left(\frac{\partial H}{\partial k_\theta}\right)^2 + \left(\frac{\partial H}{\partial k_\varphi}\right)^2}}{c \frac{\partial H}{\partial \omega}}. \quad (18)\end{aligned}$$

The user can choose to have frequency shift, phase path, absorption, or path length calculated using equations (15), (16), (17), or (18) and printed by setting the appropriate value in the input W array. (W59, W57, W58, W60 in Table 2.)

If the user wants to add differential equations to the program, he can do so by modifying subroutine HAMLTN, which evaluates Hamilton's equations.

The Hamiltonian and its derivatives are calculated by one of the versions of subroutine RINDEX, which also calculates the phase refractive index and its derivatives.

#### 4. CHOOSING AND CALCULATING THE HAMILTONIAN

Because Hamilton's equations guarantee that the Hamiltonian is constant along the ray path and because it is desirable to have the dispersion

relation satisfied at each point on the ray path, it is usual to write the dispersion relation in the form  $H = \text{constant}$  and choose that  $H$  as the Hamiltonian. Two problems arise. First, in a lossy medium the dispersion relation is complex, so that the resulting complex Hamiltonian gives ray paths having complex coordinates when used in Hamilton's equations. Second, in some cases some forms of the dispersion relation have computational advantages over others when used as a Hamiltonian.

Allowing the coordinates of the ray path to assume complex values is called ray tracing in complex space (Budden and Jull, 1964; Jones, 1970; Budden and Terry, 1971) which is the extension to three dimensions of the phase integral method (Budden, 1961). Ray tracing in complex space is necessary to calculate the propagation of LF radio waves in the D region of the ionosphere (Jones, 1970), and it may also be needed for some medium frequencies.

However, the effect of losses on the ray path of HF radio waves in the ionosphere is probably small, so that the only effect of losses is to attenuate the signal. For this case, then, it is desirable to find a prescription for calculating ray paths having real coordinates. Several methods exist for doing this, and except for computational difficulties, one is probably as good as another. One should recognize that along the ray path:

- (1) the dispersion cannot be exactly satisfied, or
- (2) Hamilton's equations cannot be satisfied, or
- (3) both of the above.

In our program, we have chosen to keep Hamilton's equations and require only the real part of the dispersion relation to be satisfied, neglecting the imaginary part. Another approach (Suchy, 1972) is to alter Hamilton's equations so that the full complex dispersion relation is still satisfied along a ray path having real coordinates. We are

reasonably certain that for any situation in which Suchy's method gives significantly different answers from ours, neither method is valid; ray tracing in complex space or an equivalent method would then be required.

Three choices for the Hamiltonian illustrate the computational difficulties involved. Haselgrove (1954) used the following Hamiltonian

$$H = \frac{c}{\omega} \frac{(k_r^2 + k_\theta^2 + k_\varphi^2)^{\frac{1}{2}}}{\text{real}(n)} - 1 , \quad (19)$$

which, except for the effects of errors in the numerical integration and the value of the independent variable, is equivalent to

$$\begin{aligned} H &= 1 - \frac{\omega}{c} \frac{\text{real}(n)}{(k_r^2 + k_\theta^2 + k_\varphi^2)^{\frac{1}{2}}} \\ &= \text{real} \left\{ 1 - \frac{\omega}{c} \frac{n}{(k_r^2 + k_\theta^2 + k_\varphi^2)^{\frac{1}{2}}} \right\} . \end{aligned} \quad (20)$$

There are eight versions of the subroutine RINDEX which calculate the Hamiltonian and its partial derivatives. (Eight versions allow the user to choose the Appleton-Hartree formula or the Sen-Wyller formula, and to include or ignore the earth's magnetic field and collisions.) Six of these versions (subroutines AHWFWC, AHWFNC, AHNFWC, AHNFNC, SWWF, and SWNF) use the following Hamiltonian:

$$\begin{aligned} H &= \frac{1}{2} \left( \frac{c^2}{\omega^2} (k_r^2 + k_\theta^2 + k_\varphi^2) - \text{real}(n^2) \right) \\ &= \text{real} \left\{ \frac{1}{2} \left( \frac{c^2}{\omega^2} (k_r^2 + k_\theta^2 + k_\varphi^2) - n^2 \right) \right\} . \end{aligned} \quad (21)$$

The other two versions (subroutines BQWFWC and BQWFNC) use as a Hamiltonian the real part of the quadratic equation whose solution is the Appleton-Hartree formula (Budden, 1961)

$$\begin{aligned}
H = \text{real} & \left\{ [(U - X) U^2 - Y^2 U] c^4 k^4 + X(k \cdot Y)^2 c^4 k^2 + \right. \\
& + [-2U(U - X)^2 + Y^2(2U - X)] c^2 k^2 \omega^2 - X(k \cdot Y)^2 c^2 \omega^2 + \\
& \left. + [(U - X)^2 - Y^2] (U - X) \omega^4 \right\} \quad (22)
\end{aligned}$$

except in or near free space (defined by  $X < 0.1$ ) where they also use (21) as the Hamiltonian. In (22),  $U = 1 - iZ$ , and  $X$ ,  $Y$ , and  $Z$  are the usual magnetoionic parameters.

In a lossy medium, the Hamiltonians in (20), (21), and (22) determine slightly different ray paths, but the differences are significant only when it is no longer valid to represent ray paths with coordinates that are real rather than complex. In fact, this is a weak criterion. The ray paths determined by these three Hamiltonians will become invalid before there are noticeable differences between the three ray paths. In a lossless medium, the above three Hamiltonians determine identical ray paths (except for integration errors).

For either a lossy or lossless medium, some of the above three Hamiltonians have computational difficulties. Special care must be taken in using (19) or (20) in an evanescent region (which is frequently necessary at or near vertical incidence because the numerical integration subroutine usually requires the evaluation of the differential equations not only on the ray path, but also at points near the ray path). For instance, in a lossless medium,  $\text{real}(n)$  is zero in an evanescent region, which leads to problems in (19) and (20). This problem will not arise in (21) because  $\text{real}(n^2)$  is well behaved in or at the boundary of an evanescent region, nor will it occur in using (22).

Neither (20) nor (21) (nor any other Hamiltonian based on the refractive index) will work for a ray passing through a spitze (Davies, 1965, p. 202) because the refractive index is indeterminate at a spitze, and

some of the derivatives of  $n$  diverge. So far, we have had no problems using (22) to calculate ray paths through a spizte with or without collisions.

However, the Hamiltonian in (22) will not work in or near free space because all of its derivatives are zero in free space. This problem is related to (22) not being able to distinguish between ordinary and extraordinary waves. To get started, the program uses (21) until the electron density is large enough that  $X$  is equal or greater than 1/10.

As far as we can tell, the AHWFWC (Appleton-Hartree, with field, with collisions) version of subroutine RINDEX has been made obsolete by the BQWFWC (Booker quartic, with field, with collisions) version. The latter will do everything the AHWFWC version will do and in addition it will calculate rays through spitzes. A few trial runs, however, indicate that AHWFWC runs about 30 percent faster than BQWFWC.

Similarly, the AHWFNC (Appleton-Hartree, with field, no collisions) version has been made obsolete by the BQWFNC version, which apparently runs just as fast as the AHWFNC version. We are continuing to include the AHWFNC version just in case there are undiscovered problems with the BQWFNC version.

In addition to the Appleton-Hartree formula, which is based on a constant collision frequency, the program also includes the generalized formula of Sen and Wyller (1960), which assumes a Maxwell-Boltzman distribution of electron energy and a collision frequency proportional to energy. Two versions of subroutine RINDEX use the Sen-Wyller formula for calculating the refractive index and the resulting Hamiltonian with its derivatives. These are SWWF, which includes the effects of the earth's magnetic field, and SWNF, which neglects the Earth's magnetic field. The SWWF version will probably not work for calculating rays through a spizte. It would be possible to make a version which used as

its Hamiltonian the quadratic equation whose solution is the Sen-Wyller formula for calculating rays through a spitze, but it is unlikely that we will ever do that.

The versions of subroutine RINDEX that use (21) for a Hamiltonian use the following formulas for calculating the derivatives of that Hamiltonian.

$$\frac{\partial H}{\partial t} = -n \frac{\partial n}{\partial t} , \quad (23)$$

$$\frac{\partial H}{\partial r} = -n \frac{\partial n}{\partial r} , \quad (24)$$

$$\frac{\partial H}{\partial \theta} = -n \frac{\partial n}{\partial \theta} , \quad (25)$$

$$\frac{\partial H}{\partial \varphi} = -n \frac{\partial n}{\partial \varphi} , \quad (26)$$

$$\frac{\partial H}{\partial \omega} = -\frac{n n^t}{\omega} , \quad (27)$$

$$\frac{\partial H}{\partial k_r} = \frac{c^2}{\omega^2} k_r - \frac{c}{\omega} n \frac{\partial n}{\partial V_r} , \quad (28)$$

$$\frac{\partial H}{\partial k_\theta} = \frac{c^2}{\omega^2} k_\theta - \frac{c}{\omega} n \frac{\partial n}{\partial V_\theta} , \quad (29)$$

$$\frac{\partial H}{\partial k_\varphi} = \frac{c^2}{\omega^2} k_\varphi - \frac{c}{\omega} n \frac{\partial n}{\partial V_\varphi} , \quad (30)$$

$$\vec{k} \cdot \frac{\partial H}{\partial \vec{k}} = k_r \frac{\partial H}{\partial k_r} + k_\theta \frac{\partial H}{\partial k_\theta} + k_\varphi \frac{\partial H}{\partial k_\varphi} = n^2 , \quad (31)$$

where  $n^t$  is the group refractive index defined by

$$n' = n + f \frac{dn}{df} = n + \omega \frac{dn}{d\omega} , \quad (32)$$

and  $V_r$ ,  $V_\theta$ , and  $V_\varphi$  are the components of the wave normal direction in the  $r$ ,  $\theta$ , and  $\varphi$  directions normalized so that

$$V_r^2 + V_\theta^2 + V_\varphi^2 = \text{Real } \{n^2\} . \quad (33)$$

The derivatives of the Hamiltonian in (22) are given in section 5.5.

## 5. REFRACTIVE INDEX EQUATIONS

The refractive index equations used in this ray tracing program are based either on the Appleton-Hartree formula (Budden, 1961) or on the generalized formula of Sen and Wyller (1960). There are eight versions of SUBROUTINE RINDEX, the subroutine that calculates the refractive index and its gradient:

- (1) Appleton-Hartree formula with field, with collisions.
- (2) Appleton-Hartree formula with field, no collisions.
- (3) Appleton-Hartree formula with collisions, no field.
- (4) Appleton-Hartree formula no field, no collisions.
- (5) Booker quartic with field, with collisions.
- (6) Booker quartic with field, no collisions.
- (7) Sen-Wyller formula with field.
- (8) Sen-Wyller formula, no field.

Each of these eight versions calculates  $n^2$ ,  $nn'$ ,  $n \partial n / \partial r$ ,  $n \partial n / \partial \theta$ ,  $n \partial n / \partial \varphi$ ,  $n \partial n / \partial V_r$ ,  $n \partial n / \partial V_\theta$ ,  $n \partial n / \partial V_\varphi$ ,  $n \partial n / \partial t$ , and the polarization, where  $n$  is the complex phase refractive index;  $n'$  is the complex group refractive index;  $r$ ,  $\theta$ , and  $\varphi$  are the spherical polar coordinates of a point on the ray path, and  $V_r$ ,  $V_\theta$ , and  $V_\varphi$  are the components of the wave normal direction in the  $r$ ,  $\theta$ , and  $\varphi$  directions. The quantities

$$n' = n + f \frac{dn}{df} = n + \omega \frac{dn}{d\omega} , \quad (32)$$

and  $V_r$ ,  $V_\theta$ , and  $V_\varphi$  are the components of the wave normal direction in the  $r$ ,  $\theta$ , and  $\varphi$  directions normalized so that

$$V_r^2 + V_\theta^2 + V_\varphi^2 = \text{Real } \{n^2\} . \quad (33)$$

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- (3) Appleton-Hartree formula with collisions, no field.
- (4) Appleton-Hartree formula no field, no collisions.
- (5) Booker quartic with field, with collisions.
- (6) Booker quartic with field, no collisions.
- (7) Sen-Wyller formula with field.
- (8) Sen-Wyller formula, no field.

Each of these eight versions calculates  $n^2$ ,  $nn'$ ,  $n \partial n / \partial r$ ,  $n \partial n / \partial \theta$ ,  $n \partial n / \partial \varphi$ ,  $n \partial n / \partial V_r$ ,  $n \partial n / \partial V_\theta$ ,  $n \partial n / \partial V_\varphi$ ,  $n \partial n / \partial t$ , and the polarization, where  $n$  is the complex phase refractive index;  $n'$  is the complex group refractive index;  $r$ ,  $\theta$ , and  $\varphi$  are the spherical polar coordinates of a point on the ray path, and  $V_r$ ,  $V_\theta$ , and  $V_\varphi$  are the components of the wave normal direction in the  $r$ ,  $\theta$ , and  $\varphi$  directions. The quantities

$X$ ,  $\partial X/\partial r$ ,  $\partial X/\partial \theta$ ,  $\partial X/\partial \varphi$ , and  $\partial X/\partial t$  are supplied by one of the versions of subroutine ELECTX which defines the electron density model. The quantities  $Y$ ,  $\partial Y/\partial r$ ,  $\partial Y/\partial \theta$ ,  $\partial Y/\partial \varphi$ ,  $Y_r$ ,  $\partial Y_r/\partial r$ ,  $\partial Y_r/\partial \theta$ ,  $\partial Y_r/\partial \varphi$ ,  $Y_\theta$ ,  $\partial Y_\theta/\partial r$ ,  $\partial Y_\theta/\partial \theta$ ,  $\partial Y_\theta/\partial \varphi$ ,  $Y_\varphi$ ,  $\partial Y_\varphi/\partial r$ ,  $\partial Y_\varphi/\partial \theta$ , and  $\partial Y_\varphi/\partial \varphi$  are supplied by one of the versions of subroutine MAGY which defines the magnetic field model. The quantities  $Z$ ,  $\partial Z/\partial r$ ,  $\partial Z/\partial \theta$ , and  $\partial Z/\partial \varphi$  are supplied by one of the versions of subroutine COLFRZ which defines the collision frequency model.

In our formulation, we have tried to avoid using multivalued functions, such as the square root or  $\cos^{-1}$ , wherever possible. Only twice do we use the square root. One instance is the square root in the Appleton-Hartree formula, unavoidable without adding more differential equations to the system. The second instance is a square root used to calculate polarization. This latter use is unimportant because the polarization is not used in the ray tracing equations.

It is desirable to avoid multivalued functions because, unless extreme care is used, the value of such a function can change discontinuously from one point on the ray path to the next. A particularly troublesome case occurs at reflection for vertical incidence. At that point, the real part of  $n^2$  goes through zero, and  $n$  changes from approximately purely real to approximately purely imaginary. Since the numerical integration subroutine usually requires the evaluation of the differential equations not only on the ray path, but also at points near the ray path, it is necessary to be able to evaluate the differential equations above the reflection height, that is, in an evanescent region.

We have found that it is possible to regroup the variables in the equations to avoid this problem: we calculate the real part of  $n^2$  and its derivatives instead of the real part of  $n$  and its derivatives. And we calculate  $nn'$  instead of  $n'$ .

It was not easy to avoid using multivalued functions, however. Many of the usual parameters used to compute the refractive index require the use of multivalued functions in their calculation. Thus, we couldn't calculate

$$V = \sqrt{V_r^2 + V_\theta^2 + V_\varphi^2}$$

nor  $\cos \psi$ , nor  $\sin \psi$ , where  $\psi$  is the angle between the wave normal direction and the earth's magnetic field. Thus, we also could not calculate

$$Y_L = Y \cos \psi$$

nor

$$Y_T = Y \sin \psi.$$

The most difficult part of avoiding the use of multivalued functions was in calculating the derivatives.

The following is a list of the equations calculated by the eight versions of subroutine RINDEX.

### 5.1 Appleton-Hartree Formula with Field, with Collisions

The square of the complex phase refractive index is given by

$$n^2 = 1 - 2X \frac{1 - iZ - X}{2(1 - iZ)(1 - iZ - X) - Y_T^2 \pm \sqrt{Y_T^4 + 4Y_L^2(1 - iZ - X)^2}} , \quad (34)$$

where

$$X = \frac{\frac{f^2}{N}}{f^2} , \quad (35)$$

$$Y = \frac{f_H}{f} , \quad (36)$$

$$Z = \frac{v}{2\pi f} , \quad (37)$$

$$Y_T = Y \sin \psi , \quad (38)$$

$$Y_L = Y \cos \psi , \quad (39)$$

$f_N$  is the plasma frequency,  $f_H$  is the electron gyrofrequency,  $\nu$  is the electron collision frequency,  $f$  is the wave frequency, and  $\psi$  is the angle between the wave normal direction and the earth's magnetic field.

The following equations parallel the formulas in this version of RINDEX.

$$V^2 = V_r^2 + V_\theta^2 + V_\varphi^2 \quad (40)$$

$$V \cdot Y = V_r Y_r + V_\theta Y_\theta + V_\varphi Y_\varphi \quad (41)$$

$$\frac{Y_L}{V} = \frac{V \cdot Y}{V^2} \quad (42)$$

$$Y_L^2 = \frac{(V \cdot Y)^2}{V^2} \quad (43)$$

$$Y_T^2 = Y^2 - Y_L^2 \quad (44)$$

$$Y_T^4 = (Y_T^2)^2 \quad (45)$$

$$U = 1 - iZ \quad (46)$$

$$RAD = \pm \sqrt{Y_T^4 + 4Y_L^2 (U-X)^2} \quad (47)$$

$$D = 2U(U-X) - Y_T^2 + RAD \quad (48)$$

$$n^2 = 1 - \frac{2X(U-X)}{D} \quad (49)$$

$$\frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} = \frac{2X(U-X) \left( -1 + \frac{Y_T^2 - 2(U-X)^2}{RAD} \right)}{D^2} \quad (50)$$

$$Y_T Y_L \frac{\partial \psi}{\partial r} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial r} - \left( V_r \frac{\partial Y_r}{\partial r} + V_\theta \frac{\partial Y_\theta}{\partial r} + V_\varphi \frac{\partial Y_\varphi}{\partial r} \right) \left( \frac{Y_L}{V} \right) \quad (51)$$

$$Y_T Y_L \frac{\partial \psi}{\partial \theta} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial \theta} - \left( V_r \frac{\partial Y_r}{\partial \theta} + V_\theta \frac{\partial Y_\theta}{\partial \theta} + V_\varphi \frac{\partial Y_\varphi}{\partial \theta} \right) \left( \frac{Y_L}{V} \right) \quad (52)$$

$$Y_T Y_L \frac{\partial \psi}{\partial \varphi} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial \varphi} - \left( V_r \frac{\partial Y_r}{\partial \varphi} + V_\theta \frac{\partial Y_\theta}{\partial \varphi} + V_\varphi \frac{\partial Y_\varphi}{\partial \varphi} \right) \left( \frac{Y_L}{V} \right) \quad (53)$$

$$n \frac{\partial n}{\partial X} = - \frac{\left( 2U(U-X)^2 - Y_T^2(U-2X) + \frac{Y_T^4(U-2X) + 4Y_L^2(U-X)^3}{RAD} \right)}{D^2} \quad (54)$$

$$n \frac{\partial n}{\partial Y} = \frac{2X(U-X)}{D^2 Y} \left( -Y_T^2 + \frac{Y_T^4 + 2Y_L^2(U-X)^2}{RAD} \right) \quad (55)$$

$$n \frac{\partial n}{\partial Z} = \frac{iX}{D^2} \left( -2(U-X)^2 - Y_T^2 + \frac{Y_T^4}{RAD} \right) \quad (56)$$

$$n \frac{\partial n}{\partial r} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial r} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial r} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial r} + \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} Y_L Y_T \frac{\partial \psi}{\partial r} \quad (57)$$

$$n \frac{\partial n}{\partial \theta} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \theta} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial \theta} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \theta} + \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} Y_T Y_L \frac{\partial \psi}{\partial \theta} \quad (58)$$

$$n \frac{\partial n}{\partial \varphi} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \varphi} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial \varphi} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \varphi} + \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} Y_T Y_L \frac{\partial \psi}{\partial \varphi} \quad (59)$$

$$n \frac{\partial n}{\partial V_r} = \frac{n}{Y_T Y_L} \frac{\partial n}{\partial \psi} \left( \frac{V_r Y_L^2}{V^2} - \left( \frac{Y_L}{V} \right) Y_r \right) \quad (60)$$

$$n \frac{\partial n}{\partial V_\theta} = \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \left( \frac{V_\theta Y_L^2}{V^2} - \left( \frac{Y_L}{V} \right) Y_\theta \right) \quad (61)$$

$$n \frac{\partial n}{\partial V_\varphi} = \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \left( \frac{V_\varphi Y_L^2}{V^2} - \left( \frac{Y_L}{V} \right) Y_\varphi \right) \quad (62)$$

$$nn' = n^2 - \left( 2Xn \frac{\partial n}{\partial X} + Y n \frac{\partial n}{\partial Y} + Z n \frac{\partial n}{\partial Z} \right) \quad (63)$$

$$\text{Polarization} = \rho = -i \frac{(-Y_T^2 + RAD) \sqrt{V^2}}{2(U-X) V \cdot Y} \quad (64a)$$

(Budden, 1961, page 49).

$$\text{Longitudinal polarization} = \frac{iX\sqrt{Y^2}}{(U-X)(U+i\frac{V \cdot Y}{\sqrt{V^2}} - \rho)} \quad (64b)$$

(Budden, 1961, page 54).

$$n \frac{\partial n}{\partial t} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial t} \quad (65)$$

## 5.2 Appleton-Hartree Formula with Field, no Collisions

The equations are the same as with field with collisions except for:

$$Z = \frac{\partial Z}{\partial r} = \frac{\partial Z}{\partial \theta} = \frac{\partial Z}{\partial \varphi} = 0, \quad (66)$$

$$U = 1. \quad (67)$$

5.3 Appleton-Hartree Formula no Field, With Collisions

$$U = 1 - iZ \quad (68)$$

$$n^2 = 1 - \frac{X}{U} \quad (69)$$

$$n \frac{\partial n}{\partial X} = - \frac{.5}{U} \quad (70)$$

$$n \frac{\partial n}{\partial Z} = - \frac{.5iX}{U^2} \quad (71)$$

$$nn' = n^2 - \left( 2X n \frac{\partial n}{\partial X} + Z n \frac{\partial n}{\partial Z} \right) \quad (72)$$

$$n \frac{\partial n}{\partial r} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial r} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial r} \quad (73)$$

$$n \frac{\partial n}{\partial \theta} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \theta} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \theta} \quad (74)$$

$$n \frac{\partial n}{\partial \varphi} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \varphi} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \varphi} \quad (75)$$

$$n \frac{\partial n}{\partial V_r} = 0 \quad (76)$$

$$n \frac{\partial n}{\partial V_\theta} = 0 \quad (77)$$

$$n \frac{\partial n}{\partial V_\varphi} = 0 \quad (78)$$

$$\text{Polarization} = i \quad (79a)$$

$$\text{Longitudinal polarization} = 0 \quad (79b)$$

$$n \frac{\partial n}{\partial t} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial t} \quad (80)$$

5.4 Appleton-Hartree Formula no Field, no Collisions

$$n^2 = 1 - X \quad (81)$$

$$n \frac{\partial n}{\partial X} = -.5 \quad (82)$$

$$n \frac{\partial n}{\partial r} = \left( n \frac{\partial n}{\partial X} \right) \left( \frac{\partial X}{\partial r} \right) \quad (83)$$

$$n \frac{\partial n}{\partial \theta} = \left( n \frac{\partial n}{\partial X} \right) \frac{\partial X}{\partial \theta} \quad (84)$$

$$n \frac{\partial n}{\partial \varphi} = \left( n \frac{\partial n}{\partial X} \right) \frac{\partial X}{\partial \varphi} \quad (85)$$

$$n \frac{\partial n}{\partial V_r} = 0 \quad (86)$$

$$n \frac{\partial n}{\partial V_\theta} = 0 \quad (87)$$

$$n \frac{\partial n}{\partial V_\varphi} = 0 \quad (88)$$

$$nn' = 1 \quad (89)$$

$$\text{Polarization} = i \quad (90a)$$

$$\text{Longitudinal polarization} = 0 \quad (90b)$$

$$n \frac{\partial n}{\partial t} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial t} \quad (91)$$

5.5 Booker Quartic with Field, with Collisions

The form of the dispersion relation used for the Hamiltonian in this version of subroutine RINDEX is the quadratic equation whose solution is the Appleton-Hartree formula. This Hamiltonian, given by

(22), is also a special case of the Booker quartic for  $S = 0$  and  $C = 1$  (Budden, 1961). This version uses the Hamiltonian in (22) only when the electron density is large enough that  $X$  is greater than or equal to  $1/10$ . For  $X$  less than  $1/10$ , the Hamiltonian in (21) is used.

Below are the equations for the derivatives of the Hamiltonian in (22). The equations for the derivatives of the Hamiltonian in (21) are the same as those in section 4.1.

$$k^2 = k_r^2 + k_\theta^2 + k_\varphi^2 , \quad (92)$$

$$k \cdot Y = k_r Y_r + k_\theta Y_\theta + k_\varphi Y_\varphi , \quad (93)$$

$$U = 1 - iZ , \quad (94)$$

$$A = (U - X) U^2 - U Y^2 , \quad (95)$$

$$B = -2U(U - X)^2 + Y^2(2U - X) , \quad (96)$$

$$\alpha = A c^4 k^4 + X(k \cdot Y)^2 c^4 k^2 , \quad (97)$$

$$\beta = B c^2 k^2 \omega^2 - X(k \cdot Y)^2 c^2 \omega^2 , \quad (98)$$

$$\gamma = ((U - X)^2 - Y^2)(U - X) \omega^4 , \quad (99)$$

$$H = \alpha + \beta + \gamma , \quad (100)$$

$$\begin{aligned} \frac{\partial H}{\partial X} = & -U^2 c^4 k^4 + (k \cdot Y)^2 c^4 k^2 + (4U(U - X) - Y^2)c^2 k^2 \omega^2 + \\ & -(k \cdot Y)^2 c^2 \omega^2 + (-3(U - X)^2 + Y^2) \omega^4 , \end{aligned} \quad (101)$$

$$\frac{\partial H}{\partial(Y^2)} = -U c^4 k^4 + (2U - X) c^2 k^2 \omega^2 - (U - X) \omega^4 , \quad (102)$$

$$\frac{\partial H}{\partial((k \cdot Y)^2)} = X c^2 (c^2 k^2 - \omega^2) , \quad (103)$$

$$\begin{aligned} \frac{\partial H}{\partial U} &= (2 U(U - X) + U^2 - Y^2) c^4 k^4 + \\ &+ 2(Y^2 - (U - X)^2 - 2U(U - X)) c^2 k^2 \omega^2 + \\ &+ (3(U - X)^2 - Y^2) \omega^4 , \end{aligned} \quad (104)$$

$$\frac{\partial H}{\partial Z} = -i \frac{\partial H}{\partial U} , \quad (105)$$

$$\frac{\partial H}{\partial(k^2)} = 2 A c^4 k^2 + X(k \cdot Y)^2 c^4 + B c^2 \omega^2 , \quad (106)$$

$$\frac{\partial H}{\partial t} = \frac{\partial H}{\partial X} \frac{\partial X}{\partial t} , \quad (107)$$

$$\begin{aligned} \frac{\partial H}{\partial r} &= \frac{\partial H}{\partial X} \frac{\partial X}{\partial r} + 2 \frac{\partial H}{\partial(Y^2)} Y \frac{\partial Y}{\partial r} + \frac{\partial H}{\partial Z} \frac{\partial Z}{\partial r} + \\ &+ 2 \frac{\partial H}{\partial((k \cdot Y)^2)} (k \cdot Y) \left( k_r \frac{\partial Y_r}{\partial r} + k_\theta \frac{\partial Y_\theta}{\partial r} + k_\varphi \frac{\partial Y_\varphi}{\partial r} \right) , \end{aligned} \quad (108)$$

$$\begin{aligned} \frac{\partial H}{\partial \theta} &= \frac{\partial H}{\partial X} \frac{\partial X}{\partial \theta} + 2 \frac{\partial H}{\partial(Y^2)} Y \frac{\partial Y}{\partial \theta} + \frac{\partial H}{\partial Z} \frac{\partial Z}{\partial \theta} + \\ &+ 2 \frac{\partial H}{\partial((k \cdot Y)^2)} (k \cdot Y) \left( k_r \frac{\partial Y_r}{\partial \theta} + k_\theta \frac{\partial Y_\theta}{\partial \theta} + k_\varphi \frac{\partial Y_\varphi}{\partial \theta} \right) , \end{aligned} \quad (109)$$

$$\begin{aligned}\frac{\partial H}{\partial \varphi} = & \frac{\partial H}{\partial X} \frac{\partial X}{\partial \varphi} + 2 \frac{\partial H}{\partial(Y^2)} Y \frac{\partial Y}{\partial \varphi} + \frac{\partial H}{\partial Z} \frac{\partial Z}{\partial \varphi} + \\ & + 2 \frac{\partial H}{\partial((k \cdot Y)^2)} (k \cdot Y) \left( k_r \frac{\partial Y_r}{\partial \varphi} + k_\theta \frac{\partial Y_\theta}{\partial \varphi} + k_\varphi \frac{\partial Y_\varphi}{\partial \varphi} \right) , \quad (110)\end{aligned}$$

$$\begin{aligned}\frac{\partial H}{\partial \omega} = & (2\beta + 4\gamma)/\omega - 2 \frac{\partial H}{\partial X} \frac{X}{\omega} - 2 \frac{\partial H}{\partial(Y^2)} \frac{Y^2}{\omega} + \\ & - 2 \frac{\partial H}{\partial((k \cdot Y)^2)} \frac{(k \cdot Y)^2}{\omega} - \frac{\partial H}{\partial Z} \frac{Z}{\omega} , \quad (111)\end{aligned}$$

$$\frac{\partial H}{\partial k_r} = 2 \frac{\partial H}{\partial(k^2)} k_r + 2(k \cdot Y) \frac{\partial H}{\partial((k \cdot Y)^2)} Y_r , \quad (112)$$

$$\frac{\partial H}{\partial k_\theta} = 2 \frac{\partial H}{\partial(k^2)} k_\theta + 2(k \cdot Y) \frac{\partial H}{\partial((k \cdot Y)^2)} Y_\theta , \quad (113)$$

$$\frac{\partial H}{\partial k_\varphi} = 2 \frac{\partial H}{\partial(k^2)} k_\varphi + 2(k \cdot Y) \frac{\partial H}{\partial((k \cdot Y)^2)} Y_\varphi , \quad (114)$$

$$k^2(\text{calculated}) = k^2 \frac{-\beta \pm \sqrt{\beta^2 - 4\alpha\gamma}}{2\alpha} , \quad (115)$$

$$k \cdot \frac{\partial H}{\partial k} = k_r \frac{\partial H}{\partial k_r} + k_\theta \frac{\partial H}{\partial k_\theta} + k_\varphi \frac{\partial H}{\partial k_\varphi} = 4\alpha + 2\beta , \quad (116)$$

$$\text{Polarization} = i \frac{\sqrt{k^2}}{k \cdot Y} \left( U + \frac{X\omega^2}{c^2 k^2 \text{calculated} - \omega^2} \right) , \quad (117)$$

$$\begin{aligned}\text{Longitudinal} \\ \text{Polarization} = i \frac{\sqrt{Y^2 - \frac{(k \cdot Y)^2}{k^2}}}{U - X} \left( 1 - \frac{c^2 k^2 \text{calculated}}{\omega^2} \right) . \quad (118)\end{aligned}$$

### 5.6 Booker Quartic with Field, no Collisions

All the equations here are the same as for the Booker quartic version with collisions (section 5.5) except for:

$$Z = \frac{\partial Z}{\partial r} = \frac{\partial Z}{\partial \theta} = \frac{\partial Z}{\partial \varphi} = 0 , \quad (119)$$

$$U = 1. \quad (120)$$

All the variables except polarization are real; polarization is pure imaginary.

### 5.7 Sen-Wyller Formula with Field

It is possible to write the generalized Appleton-Hartree formula of Sen and Wyller (1960) in the following form:

$$n^2 = 1 - \frac{2X(U-X) + 2 AUX \sin^2 \psi}{2U(U-X)(1+A) + 2 AUX \sin^2 \psi - U((1-BC)U + A(U+X)) \sin^2 \psi + RAD} , \quad (121)$$

where

$$RAD = \pm \sqrt{U^2((1-BC)U + A(U+X))^2 \sin^4 \psi + U^2(U-X)^2(C-B)^2 \cos^2 \psi} , \quad (122)$$

$$A = \frac{C + B}{2} - 1 , \quad (123)$$

$$B = \frac{F\left(\frac{1}{Z}\right)}{F\left(\frac{1-Y}{Z}\right)} , \quad (124)$$

$$C = \frac{F\left(\frac{1}{Z}\right)}{F\left(\frac{1+Y}{Z}\right)} , \quad (125)$$

$$U = \frac{Z}{F\left(\frac{1}{Z}\right)} \quad (126)$$

$$F(w) = w C_{3/2}(w) + i \frac{5}{2} C_{5/2}(w) = \frac{1}{\frac{3}{2}!} \int_0^\infty \frac{t^{3/2} e^{-t} dt}{w - it} , \quad (127)$$

$$X = \frac{f^2 N}{f^2} , \quad (128)$$

$$Y = \frac{f_H}{f} , \quad (129)$$

$$Z = \frac{v_m}{2\pi f} , \quad (130)$$

$f_N$  is the plasma frequency,  $f_H$  is the electron gyrofrequency,  $v_m$  is the mean electron collision frequency,  $f$  is the wave frequency, and  $\psi$  is the angle between the wave normal direction and the earth's magnetic field.  
(Note that if we would use

$$F(w) = \frac{1}{w-i} , \quad (131)$$

then (121) would reduce to the usual Appleton-Hartree formula.)

This version of RINDEX calls subroutine FSW. The first argument in the calling sequence is the argument  $w$  of  $F(w)$ . The second argument in the calling sequence is the value of the function  $F(w)$  calculated by FSW. The third argument in the calling sequence is the derivative of the function,  $F'(w)$ .

The following equations parallel the formulas in this version of RINDEX.

$$\alpha = F\left(\frac{1}{Z}\right) \quad (132)$$

$$\alpha' = F'\left(\frac{1}{Z}\right) \quad (133)$$

$$\beta = F\left(\frac{1-Y}{Z}\right) \quad (134)$$

$$\beta' = F'\left(\frac{1-Y}{Z}\right) \quad (135)$$

$$\gamma = F\left(\frac{1+Y}{Z}\right) \quad (136)$$

$$\gamma' = F'\left(\frac{1+Y}{Z}\right) \quad (137)$$

$$U = \frac{Z}{\alpha} \quad (138)$$

$$\frac{\partial U}{\partial Z} = \left(1 + \frac{\alpha'}{\alpha Z}\right) / \alpha \quad (139)$$

$$B = \frac{\alpha}{\beta} \quad (140)$$

$$\frac{\partial B}{\partial Y} = \frac{B}{Z} - \frac{\beta'}{\beta} \quad (141)$$

$$\frac{\partial B}{\partial Z} = -\frac{B}{Z^2} \left( \frac{\alpha'}{\alpha} - (1-Y) \frac{\beta'}{\beta} \right) \quad (142)$$

$$C = \frac{\alpha}{Y} \quad (143)$$

$$\frac{\partial C}{\partial Y} = - \frac{C}{Z} \frac{Y'}{Y} \quad (144)$$

$$\frac{\partial C}{\partial Z} = - \frac{C}{Z^2} \left( \frac{\alpha'}{\alpha} - (1+Y) \frac{Y'}{Y} \right) \quad (145)$$

$$A = .5(B+C) - 1 \quad (146)$$

$$\frac{\partial A}{\partial Y} = .5 \left( \frac{\partial B}{\partial Y} + \frac{\partial C}{\partial Y} \right) \quad (147)$$

$$\frac{\partial A}{\partial Z} = .5 \left( \frac{\partial B}{\partial Z} + \frac{\partial C}{\partial Z} \right) \quad (148)$$

$$V^2 = V_r^2 + V_\theta^2 + V_\varphi^2 \quad (149)$$

$$V \cdot Y = V_r Y_r + V_\theta Y_\theta + V_\varphi Y_\varphi \quad (150)$$

$$Y_L^2 = \frac{(V \cdot Y)^2}{V^2} \quad (151)$$

$$Y_T^2 = Y^2 - Y_L^2 \quad (152)$$

$$\sin^2 \psi = \frac{Y_T^2}{Y^2} \quad (153)$$

$$\cos^2 \psi = \frac{Y_L^2}{Y^2} \quad (154)$$

$$T_1 = [(1-BC) U^2 + A U(U+X)] \sin^2 \psi \quad (155)$$

$$\frac{\partial T_1}{\partial X} = +AU \sin^2 \psi \quad (156)$$

$$\frac{\partial T_1}{\partial Y} = \left( U(U+X) \frac{\partial A}{\partial Y} - U^2 \left( B \frac{\partial C}{\partial Y} + C \frac{\partial B}{\partial Y} \right) \right) \sin^2 \psi \quad (157)$$

$$\begin{aligned} \frac{\partial T_1}{\partial Z} &= \left( 2U \frac{\partial U}{\partial Z} (1-BC+A) + AX \frac{\partial U}{\partial Z} - U^2 \left( B \frac{\partial C}{\partial Z} + C \frac{\partial B}{\partial Z} \right) \right. \\ &\quad \left. + U(U+X) \frac{\partial A}{\partial Z} \right) \sin^2 \psi \end{aligned} \quad (158)$$

$$\frac{1}{Y_L Y_T} \frac{\partial T_1}{\partial \psi} = \frac{2T_1}{Y_T^2} \quad (159)$$

$$T_2 = U^2 (C-B)^2 (U-X)^2 \cos^2 \psi \quad (160)$$

$$\frac{\partial T_2}{\partial X} = -2(U-X) U^2 (C-B)^2 \cos^2 \psi \quad (161)$$

$$\frac{\partial T_2}{\partial Y} = 2U^2 (U-X)^2 \cos^2 \psi (C-B) \left( \frac{\partial C}{\partial Y} - \frac{\partial B}{\partial Y} \right) \quad (162)$$

$$\frac{\partial T_2}{\partial Z} = 2U^2 (U-X)^2 (C-B) \left( \frac{\partial C}{\partial Z} - \frac{\partial B}{\partial Z} \right) \cos^2 \psi + 2T_2 \left( \frac{1}{U} + \frac{1}{U-X} \frac{\partial U}{\partial Z} \right) \quad (163)$$

$$\frac{1}{Y_L Y_T} \frac{\partial T_2}{\partial \psi} = - \frac{2T_2}{Y_L^2} \quad (164)$$

$$RAD = \pm \sqrt{T_1^2 + T_2^2} \quad (165)$$

$$\frac{\partial \text{RAD}}{\partial X} = \frac{T_1 \frac{\partial T_1}{\partial X} + \frac{1}{2} \frac{\partial T_2}{\partial X}}{\text{RAD}} \quad (166)$$

$$\frac{\partial \text{RAD}}{\partial Y} = \frac{T_1 \frac{\partial T_1}{\partial Y} + \frac{1}{2} \frac{\partial T_2}{\partial Y}}{\text{RAD}} \quad (167)$$

$$\frac{\partial \text{RAD}}{\partial Z} = \frac{T_1 \frac{\partial T_1}{\partial Z} + \frac{1}{2} \frac{\partial T_2}{\partial Z}}{\text{RAD}} \quad (168)$$

$$\frac{1}{Y_L Y_T} \frac{\partial \text{RAD}}{\partial \psi} = \frac{T_1 \left( \frac{1}{Y_L Y_T} \frac{\partial T_1}{\partial \psi} \right) + \frac{1}{2} \left( \frac{1}{Y_L Y_T} \frac{\partial T_2}{\partial \psi} \right)}{\text{RAD}} \quad (169)$$

$$D = 2U(U-X)(1+A) - T_1 + \text{RAD} + 2AU \sin^2 \psi \quad (170)$$

$$\frac{\partial D}{\partial X} = -2U - \frac{\partial T_1}{\partial X} + \frac{\partial \text{RAD}}{\partial X} + 2AU \sin^2 \psi \quad (171)$$

$$\frac{\partial D}{\partial Y} = 2U(U-X) \frac{\partial A}{\partial Y} - \frac{\partial T_1}{\partial Y} + \frac{\partial \text{RAD}}{\partial Y} + 2U \sin^2 \psi \frac{\partial A}{\partial Y} \quad (172)$$

$$\begin{aligned} \frac{\partial D}{\partial Z} &= 2(1+A) \frac{\partial U}{\partial Z} (2U-X) + 2U(U-X) \frac{\partial A}{\partial Z} - \frac{\partial T_1}{\partial Z} + \frac{\partial \text{RAD}}{\partial Z} + \\ &\quad + 2AX \sin^2 \psi \frac{\partial U}{\partial Z} + 2UX \sin^2 \psi \frac{\partial A}{\partial Z} \end{aligned} \quad (173)$$

$$\frac{1}{Y_L Y_T} \frac{\partial D}{\partial \psi} = - \left( \frac{1}{Y_L Y_T} \frac{\partial T_1}{\partial \psi} \right) + \left( \frac{1}{Y_L Y_T} \frac{\partial \text{RAD}}{\partial \psi} \right) + 2AUX/Y^2 \quad (174)$$

$$n^2 - 1 = \frac{-2X}{D} \left( (U-X) + UA \sin^2 \psi \right) \quad (175)$$

$$n^2 = 1 + (n^2 - 1) \quad (176)$$

$$n \frac{\partial n}{\partial X} = \frac{1}{2} (n^2 - 1) \left( \frac{1}{X} - \frac{1}{D} \frac{\partial D}{\partial X} \right) + \frac{X}{D} \quad (177)$$

$$n \frac{\partial n}{\partial Y} = - \frac{X U \sin^2 \psi}{D} \frac{\partial A}{\partial Y} - \frac{(n^2 - 1)}{2D} \frac{\partial D}{\partial Y} \quad (178)$$

$$n \frac{\partial n}{\partial Z} = - \frac{X}{D} (1 + A \sin^2 \psi) \frac{\partial U}{\partial Z} - \frac{XU}{D} \sin^2 \psi \frac{\partial A}{\partial Z} - \frac{(n^2 - 1)}{2D} \frac{\partial D}{\partial Z} \quad (179)$$

$$\frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} = - \frac{XUA}{Y^2 D} - \frac{(n^2 - 1)}{2D} \left( \frac{1}{Y_L Y_T} \frac{\partial D}{\partial \psi} \right) \quad (180)$$

$$Y_L Y_T \frac{\partial \psi}{\partial r} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial r} - \left( V_r \frac{\partial Y_r}{\partial r} + V_\theta \frac{\partial Y_\theta}{\partial r} + V_\varphi \frac{\partial Y_\varphi}{\partial r} \right) \frac{V \cdot Y}{V^2} \quad (181)$$

$$Y_L Y_T \frac{\partial \psi}{\partial \theta} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial \theta} - \left( V_r \frac{\partial Y_r}{\partial \theta} + V_\theta \frac{\partial Y_\theta}{\partial \theta} + V_\varphi \frac{\partial Y_\varphi}{\partial \theta} \right) \frac{V \cdot Y}{V^2} \quad (182)$$

$$Y_L Y_T \frac{\partial \psi}{\partial \varphi} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial \varphi} - \left( V_r \frac{\partial Y_r}{\partial \varphi} + V_\theta \frac{\partial Y_\theta}{\partial \varphi} + V_\varphi \frac{\partial Y_\varphi}{\partial \varphi} \right) \frac{V \cdot Y}{V^2} \quad (183)$$

$$n \frac{\partial n}{\partial r} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial r} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial r} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial r} + \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( Y_L Y_T \frac{\partial \psi}{\partial r} \right) \quad (184)$$

$$n \frac{\partial n}{\partial \theta} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \theta} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial \theta} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \theta} + \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( Y_L Y_T \frac{\partial \psi}{\partial \theta} \right) \quad (185)$$

$$n \frac{\partial n}{\partial \varphi} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \varphi} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial \varphi} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \varphi} + \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( Y_L Y_T \frac{\partial \psi}{\partial \varphi} \right) \quad (186)$$

$$n \frac{\partial n}{\partial V_r} = \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( V_r Y_L^2 - (V \cdot Y) Y_r \right) / V^2 \quad (187)$$

$$n \frac{\partial n}{\partial V_\theta} = \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( V_\theta Y_L^2 - (V \cdot Y) Y_\theta \right) / V^2 \quad (188)$$

$$n \frac{\partial n}{\partial V_\varphi} = \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( V_\varphi Y_L^2 - (V \cdot Y) Y_\varphi \right) / V^2 \quad (189)$$

$$nn' = n^2 - \left( 2X \left( n \frac{\partial n}{\partial X} \right) + Y \left( n \frac{\partial n}{\partial Y} \right) + Z \left( n \frac{\partial n}{\partial Z} \right) \right) \quad (190)$$

$$\text{polarization} = \rho = \frac{i(T_1 - \text{RAD}) Y \sqrt{V^2}}{U(U-X)(C-B)(V \cdot Y)} \quad (191a)$$

$$\text{longitudinal polarization} = \frac{X(.5i(C-B)\rho + A \cos \psi) \sqrt{\sin^2 \psi}}{\rho((U-X)(1+.5i(C-B)\rho \cos \psi) + A(U-X \cos \psi))} \quad (191b)$$

where

$$\cos \psi = \frac{V \cdot Y}{Y \sqrt{V^2}} \quad (192)$$

$$n \frac{\partial n}{\partial t} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial t} \quad (193)$$

### 5.8 Sen-Wyller Formula no Field

This subroutine uses the function G which is related to the function F (defined in the previous section) as follows:

$$G(w) = wF(w) . \quad (194)$$

The following equations parallel the formulas used in this version of RINDEX.

$$F_1 = F \left( \frac{1}{Z} \right) \quad (195)$$

$$F'_1 = F' \left( \frac{1}{Z} \right) \quad (196)$$

$$G_1 = G \left( \frac{1}{Z} \right) \quad (197)$$

$$G'_1 = G' \left( \frac{1}{Z} \right) \quad (198)$$

$$n^2 = 1 - X G_1 \quad (199)$$

$$n \frac{\partial n}{\partial X} = -\frac{1}{2} G_1 \quad (200)$$

$$n \frac{\partial n}{\partial Z} = \frac{X}{2} \frac{G'_1}{Z^2} \quad (201)$$

$$n \frac{\partial n}{\partial r} = \left( n \frac{\partial n}{\partial X} \right) \frac{\partial X}{\partial r} + \left( n \frac{\partial n}{\partial Z} \right) \frac{\partial Z}{\partial r} \quad (202)$$

$$n \frac{\partial n}{\partial \theta} = \left( n \frac{\partial n}{\partial X} \right) \frac{\partial X}{\partial \theta} + \left( n \frac{\partial n}{\partial Z} \right) \frac{\partial Z}{\partial \theta} \quad (203)$$

$$n \frac{\partial n}{\partial \varphi} = \left( n \frac{\partial n}{\partial X} \right) \frac{\partial X}{\partial \varphi} + \left( n \frac{\partial n}{\partial Z} \right) \frac{\partial Z}{\partial \varphi} \quad (204)$$

$$nn' = n^2 - \left( 2X \left( n \frac{\partial n}{\partial X} \right) + Z \left( n \frac{\partial n}{\partial Z} \right) \right) \quad (205)$$

$$\text{polarization} = i \quad (206a)$$

$$\text{longitudinal polarization} = 0 \quad (206b)$$

$$n \frac{\partial n}{\partial t} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial t} \quad (207)$$

## 6. IONOSPHERIC MODELS

When using the program, one must specify ionospheric models which define electron density, collision frequency (if the effect of collisions is being considered), and the earth's magnetic field (if its effects are being taken into account) as a function of position in space. Each of these three characteristics of the ionosphere is defined by a separate subroutine.

Appendices 3, 4, 5, and 6 contain descriptions, input parameter forms and listings of ionospheric models that now exist. These ionospheric models are not likely to cover the needs of everyone who wants to use the program. Anticipating this when we wrote the program, we made it possible to add models easily. The user may make up his own ionospheric models by simply writing subroutines to define electron density, collision frequency, and the earth's magnetic field (and their gradients) as a function of position in space in spherical polar coordinates, following the form of the subroutines in appendices 3, 4, 5, and 6.

Appendix 3 contains electron density models; appendix 4 contains models of irregularities which may be applied as perturbations to any of the electron density models; appendix 5 contains models of the earth's magnetic field; and appendix 6 contains collision frequency models.

Having several versions of the subroutines for refractive index, electron density, collision frequency, and the earth's magnetic field gives the user not only a wide choice among ionospheric models, but also a variety of compromises between cost and an accurate description of the ionosphere, while still keeping the program simple.

## 7. FINDING THE RAY PATHS THAT CONNECT A TRANSMITTER AND RECEIVER

The reason for using a ray tracing program is to find all important ray paths that connect a given transmitter and receiver (either or both of which may be on a satellite) on a particular frequency and such properties of these ray paths as group time delay, phase time delay, and absorption of the wave. "All important" ray paths include those that reflect from the various ionospheric layers (including multiple reflections) and that propagate off the great circle path.

Since basically all that a ray tracing program can do is to calculate the path of a ray when given the transmitter location, frequency, and direction of transmission, it cannot directly calculate those ray paths that arrive at a specified receiver. The problem is to know, before tracing the ray, in which directions to transmit the ray so that it will arrive at the receiver. Since there are no general solutions to this problem, the user of a ray tracing program must rely on some sort of trial and error technique to find those ray paths that connect the transmitter and receiver. This involves varying the direction of transmission until a ray is found that reaches the receiver. If a ray tracing program does this automatically, we say that it has a homing feature. This program does not have such a feature. To find all the paths connecting the transmitter with the receiver requires a very elaborate homing routine because "homing in" on a receiver takes more judgment and common sense than speed in performing massive calculations. Therefore, the person using the program is more fitted to this task than is the computer program itself.

As an aid, however, the program allows the user to specify the receiver height, the number of hops, and a range of azimuth and elevation angles-of-transmission that he thinks will include those rays that

will arrive at the receiver. The program then calculates a ray path for each of the azimuth and elevation angles-of-transmission specified in the range. Usually only in the case of ionospheres with large horizontal gradients will the azimuth angle-of-transmission have to be varied. The program will calculate each ray path far enough to intersect or make a closest approach to the receiver height for the requested number of hops. The user can then interpolate between those rays which surround the receiver.

We define the point of "closest approach" as the point on the ray path where the wave normal direction is horizontal. It approximates an apogee if the receiver is above the apogee height and it approximates a perigee if the receiver is below the perigee height. The approximation is good for oblique propagation. When the earth's magnetic field is neglected, a point of "closest approach" is exactly an apogee or perigee.

We count one hop every time the ray crosses the receiver height. If the receiver is on the ground, a ground reflection counts as one hop for the downcoming ray before the ground reflection, and another hop for the upgoing ray after the ground reflection. We count two hops every time the ray passes through a point of "closest approach" to the receiver height. This procedure helps make rays that have the same hop number have a ground range that is a continuous function of the direction of transmission.

## 8. OUTPUT

### 8.1 Printout

Periodically and at selected points during a ray trace, the program will print information giving the position of the current ray path point, the direction of the wave normal, and the cumulative values of quantities being integrated along the ray path such as group path, phase path, absorption, and Doppler shift. Appendix 8c contains a sample of the printout.

## 8.2 Punched Cards

The program will punch a card at the beginning of each ray, a card at each ground reflection, a card at each crossing of the receiver height, two cards for each closest approach to the receiver height, and a card at the end of each ray to summarize the main results of the ray path calculations. These cards are explained in figures 1 and 2.

These cards are very useful as input data to other computer programs and for plotting the results of the ray tracing. In fact, these cards represent the most useful form of output for production ray path calculations. This method, called the rayset information-storage technique, was developed by Dr. T. A. Croft (Croft and Gregory, 1963) of Stanford University.

## 8.3 Plots of the Ray Path

A plot of the actual ray path, especially for very irregular ionospheres, can be helpful in understanding what sometimes seems like strange results in light of the input data. Thus, the program has an option for plotting, providing, of course, that the user has a plotter and plotting subroutines such as those described in appendix 7. The program can plot the projection of the ray path on any vertical plane or on the ground. The input parameter forms for plots of the ray path (appendix 1i) give more details. Appendix 8e contains sample plots of the raypath.

## 9. DECK SET UP

The versatility gained by having several versions of some of the subroutines is somewhat offset because the user must learn the deck set up in order to make necessary substitutions. Figure 3 shows the deck setup, including the subroutines that make up the main deck and those which are frequently exchanged with alternate versions. The order of the subroutines is unimportant.

Col. 80	"T" indicates ray at transmitter
Col. 79	Maximum number of hops
Col. 74-78	Imaginary part
Col. 69-73	Real part
Col. 54-63	Elevation angle of transmission, deg
Col. 44-53	Azimuth angle of transmission, deg clockwise of north
Col. 35-43	Frequency, MHz
Col. 26-34	Height of receiver above ground, km
Col. 20-25	Longitude of transmitter, deg
Col. 14-19	Latitude of transmitter, deg
Col. 5-13	Height of transmitter above ground, km
Col. 4	O, X, or N indicates ordinary, extraordinary, or no field
Col. 1-3	Ionospheric identification

▲ indicates implied decimal

Figure 1. Sample transmitter rayset.

Col. 80	Alphabetic character indicating type of rayset*
Col. 79	Hop number
Col. 74-78	Imaginary part of wave polarization
Col. 69-73	Real part of wave polarization
Col. 63-68	Doppler shift, Hz
Col. 57-62	Absorption, db
Col. 51-56	Phase path minus straight line distance between transmitter and ray point, km
Col. 45-50	Group path minus straight line distance between transmitter and ray point, km
Col. 37-44	Straight line distance between transmitter and ray point, km
Col. 31-36	Elevation angle of arrival, deg
Col. 25-30	At ray point
Col. 19-24	At transmitter
Col. 10-18	Ground range between transmitter and receiver, km
Col. 1-9	Height of ray, km

OL00E 506125

▲ indicates implied decimal

- \* G ray ground reflected. The height punched out is the apogee height since the last ground reflect.
- M ray made a closest approach to the receiver height
- P ray penetrated
- R ray at the receiver height. The height punched out is the height of the ray farthest from receiver height since last crossing of receiver height
- S program reached maximum number of steps

Figure 2. Sample minimum distance rayset.

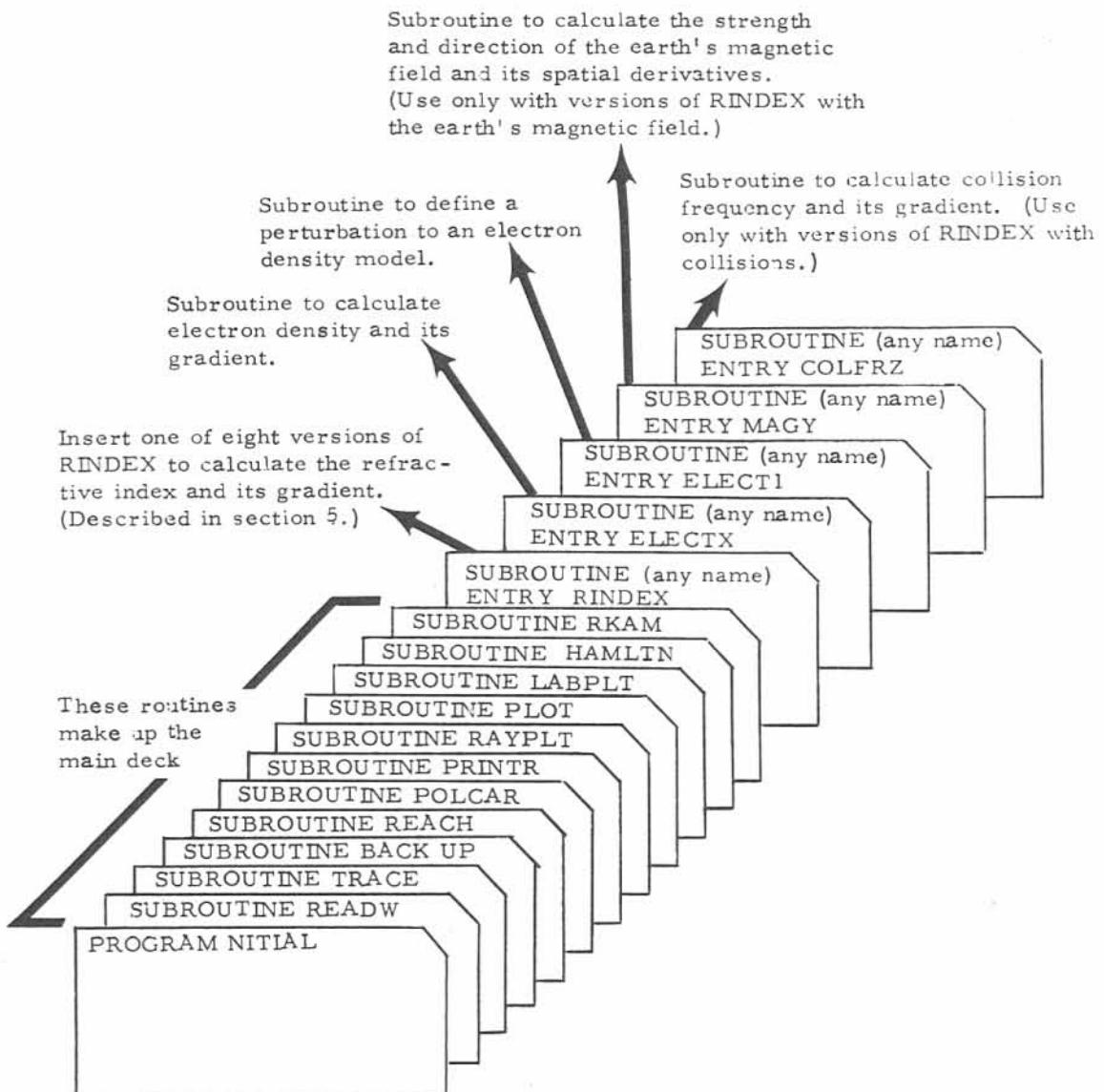


Figure 3. Program deck set-up

## 10. INPUT

The input data for a ray tracing program divide themselves naturally into two groups:

First, data that control the type of ray trace requested, such as the transmitter location and frequency, plus parameters describing analytic models of the ionosphere. Since there are few of these, efficiency in packing such data can be exchanged for versatility and ease of data handling. Therefore, by putting only one piece of data on each card, we gain the conveniences of reading in these data in any order and of having the program read in only those data that are different from those of the previous case. A number in the first three columns of each card identifies the data being read in. Table 2 defines the identifying numbers that are subscripts for a linear array, W. The last 56 columns of the card are available for comments.

We have also provided a method for conversion of units for input. The computer program needs angles in radians, whereas people usually like to use angles in degrees. The program is set up for angles in radians, but putting a "1" in column 18 allows the user to enter the angle in degrees and have the program make the conversion. A "1" in column 19 allows the user to enter central earth angles as the great circle distance along the ground in kilometers. (The program will calculate the latitude of a transmitter which is 500 km north of the equator, for instance.) The program expects distances in kilometers. A "1" in column 20 indicates a distance in nautical miles, and a "1" in column 21 indicates a distance in feet.

Appendix 8b contains a sample of how the cards are to be punched. If two or more cards have the same identifying number, the last one dominates. A card with the first three columns blank indicates the end of this type of data cards.

Table 2. Description of the Input Data for the W Array

W1	= 1. for ordinary ray = -1. for extraordinary ray
W2*	Radius of the earth in km
W3	Height of transmitter above the earth in km
W4	North geographic latitude of the transmitter
W5	East geographic longitude of the transmitter
W7	Initial frequency in MHz
W8	Final frequency in MHz
W9	Step in frequency in MHz (zero for a fixed frequency)
W11	Initial azimuth angle of transmission
W12	Final azimuth angle of transmission
W13	Step in azimuth angle of transmission (zero for a fixed azimuth)
W15	Initial elevation angle of transmission
W16	Final elevation angle of transmission
W17	Step in elevation angle of transmission (zero for a fixed elevation)
W20	Receiver height above the earth in km
W21	Nonzero to skip to the next frequency after the ray has penetrated the ionosphere
W22*	Maximum number of hops
W23*	Maximum number of steps per hop
W24*	North geographic latitude of the north geomagnetic pole
W25*	East geographic longitude of north geomagnetic pole
W41*	=1. for Runge-Kutta integration =2. for Adams-Moulton integration without error checking =3. for Adams-Moulton integration with relative error check =4. for Adams-Moulton integration with absolute error check
W42*	Maximum allowable single step error
W43*	Ratio of maximum single step error to minimum single step error
W44*	Initial integration step size in km (step in group path)
W45*	Maximum step length in km
W46*	Minimum step length in km
W47*	Factor by which to increase or decrease step length
W57	=1. to integrate, =2. to integrate and print phase path
W58	=1. to integrate, =2. to integrate and print absorption
W59	=1. to integrate, =2. to integrate and print doppler shift
W60	=1. to integrate, =2. to integrate and print path length
W71	Number of steps between periodic printout
W72	Nonzero to punch raysets on cards
W81	=0. to not plot ray path =1. to plot projection of ray path on a vertical plane =2. to plot projection of ray path on the ground
W82-88	Parameters used when plotting
W100-149	Parameters for analytic electron density models
W150-199	Parameters for perturbations to electron density models
W200-249	Parameters for analytic magnetic field models
W250-299	Parameters for analytic collision frequency models

\*These values have been initialized in the main program but may be reset by reading them in. See Appendix 1b for the initial values.

A second group of input data cards are necessary if nonanalytic ionospheric models such as the electron density profile defined by subroutine TABLEX or the collision frequency profile defined by subroutine TABLEZ are used. Each subroutine defining a nonanalytic ionospheric model reads in data cards according to a format defined in that subroutine. An element in the W array controls the reading of these cards. (See table 2.) Figure 4 shows the order in which these data cards should be arranged.

## 11. ACCURACY

The numerical integration subroutine has a built-in mechanism to check errors and adjust the integration step length accordingly. If the errors get larger than a maximum specified by the user, the routine will decrease the step length in order to maintain the accuracy. On the other hand, if the accuracy is greater than that required by the user, the routine will increase the step length in order to reduce the computing cost. The user specifies the desired accuracy in W42 (see table 2). W42 is the maximum allowable relative error in any single step for any of the equations being integrated. To get a very accurate (but expensive) ray trace, one can use a small W42 (about  $10^{-5}$  or  $10^{-6}$ ). For a cheap, approximate ray trace, one should use a large W42 ( $10^{-3}$  or even  $10^{-2}$ ). For cases in which all of the variables being integrated increase monotonically, the total relative error can be guaranteed to be less than W42. Otherwise, the total relative error cannot be easily estimated.

The far left column of the printout from the ray path calculation gives an indication of the integration error in the magnitude of the vector which points in the wave normal direction. Although the calculation of this error is made independently of the error calculation in the numerical integration routine, we have found that except near reflection for vertical or near vertical incidence this error is usually of the same order

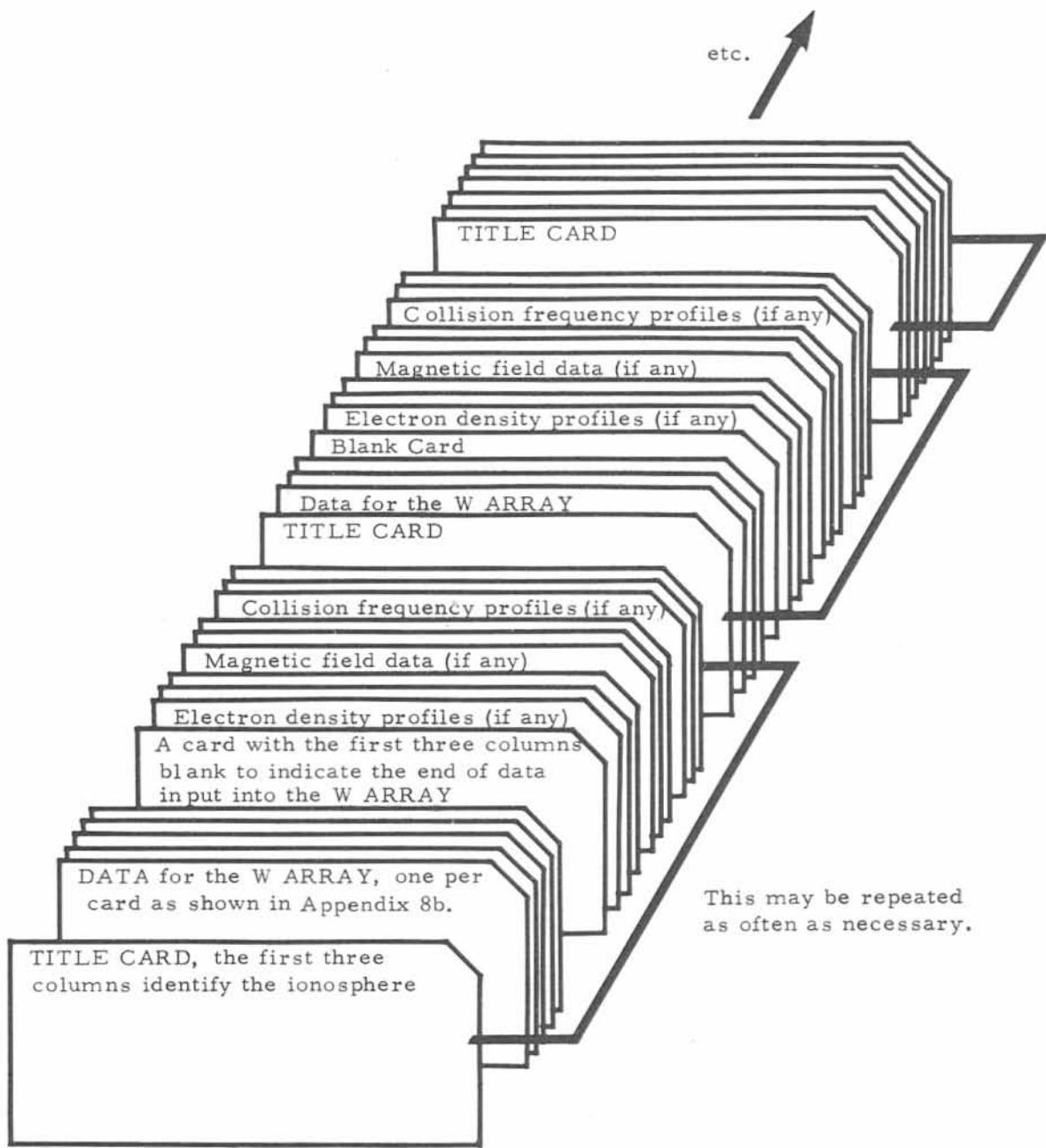


Figure 4. Data deck set-up.

of magnitude as that specified in W42. We have found that whenever this error has exceeded W42 by several orders of magnitude, the electron density subroutine we had written was calculating a gradient of electron density inconsistent with the spatial variation of electron density being calculated. See the general description of electron density models in Appendix 3a for more information.

## 12. COORDINATE SYSTEMS

The program uses two different spherical polar coordinate systems, namely, a geographic and a computational coordinate system. Input data for the coordinates of the transmitter (W4 and W5) and input data for the coordinates of the north pole of the computational coordinate system (W24 and W25) are entered in geographic coordinates. (Putting W25 equal to  $0^\circ$  and W24 equal to  $90^\circ$  would superimpose the two north poles and equate the two coordinate systems.)

When the two coordinate systems do not coincide, the three types of ionospheric models calculate electron density, the earth's magnetic field, and collision frequency in terms of the computational coordinate system. In particular, the dipole model of the earth's magnetic field uses the axis of the computational coordinate system as the axis for the dipole field. Thus, when using this dipole model, the computational coordinate system is a geomagnetic coordinate system, and both electron density and collision frequency must be defined in geomagnetic coordinates. Dudziak (1961) describes the transformations between these coordinate systems.

### 13. HOW THE PROGRAM WORKS

This ray tracing program consists of various subroutines that perform specific tasks in calculating ray paths. This division of labor facilitates modifying the program to solve specific problems. Often it may be necessary to change only one or two subroutines to convert the program to a different use.

The main program (NITIAL) sets up the initial conditions (transmitter location, wave frequency, and direction of transmission) for each ray trace. In setting up the initial conditions for each ray trace, the main program (NITIAL) steps frequency, azimuth angle of transmission, and elevation angle of transmission. The details of the workings of NITIAL can be found in the flow chart in figure 5. Then subroutine TRACE calculates one ray path for the requested number of crossings of the specified receiver height. Subroutine TRACE is the heart of the ray tracing program. It is the most complicated subroutine included, but also the most important to understand. The flow chart in figure 6 should help to explain TRACE.

Subroutine RKAM integrates the differential equations numerically using an Adams-Moulton predictor-corrector method with a Runge-Kutta starter. Subroutine HAMLTN evaluates the differential equations to be integrated. Subroutine RINDEX calculates the phase refractive index and its gradients, the group refractive index, and the polarization. (Eight versions of subroutine RINDEX are included.) Subroutines ELECTX, ELECT1, MAGY, and COLFRZ calculate the ionospheric electron density, perturbations to the electron density (irregularities), the earth's magnetic field, and the electron collision frequency, respectively. Several versions of these four subroutines are included and it is easy to add more. Subroutine REACH calculates a straight-line segment of a ray path in free space between the earth and the

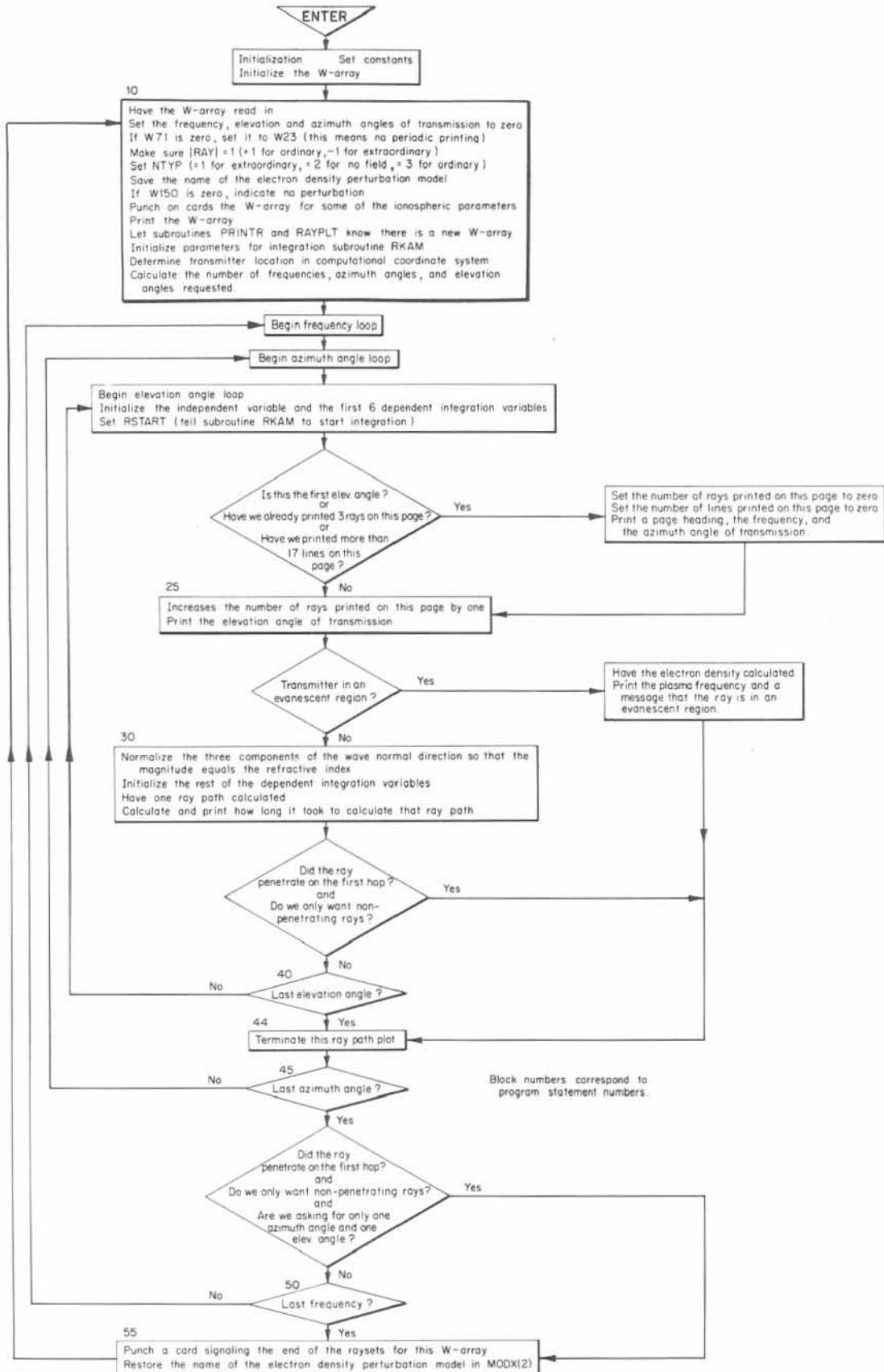


Figure 5. Flow chart for program NITIAL.

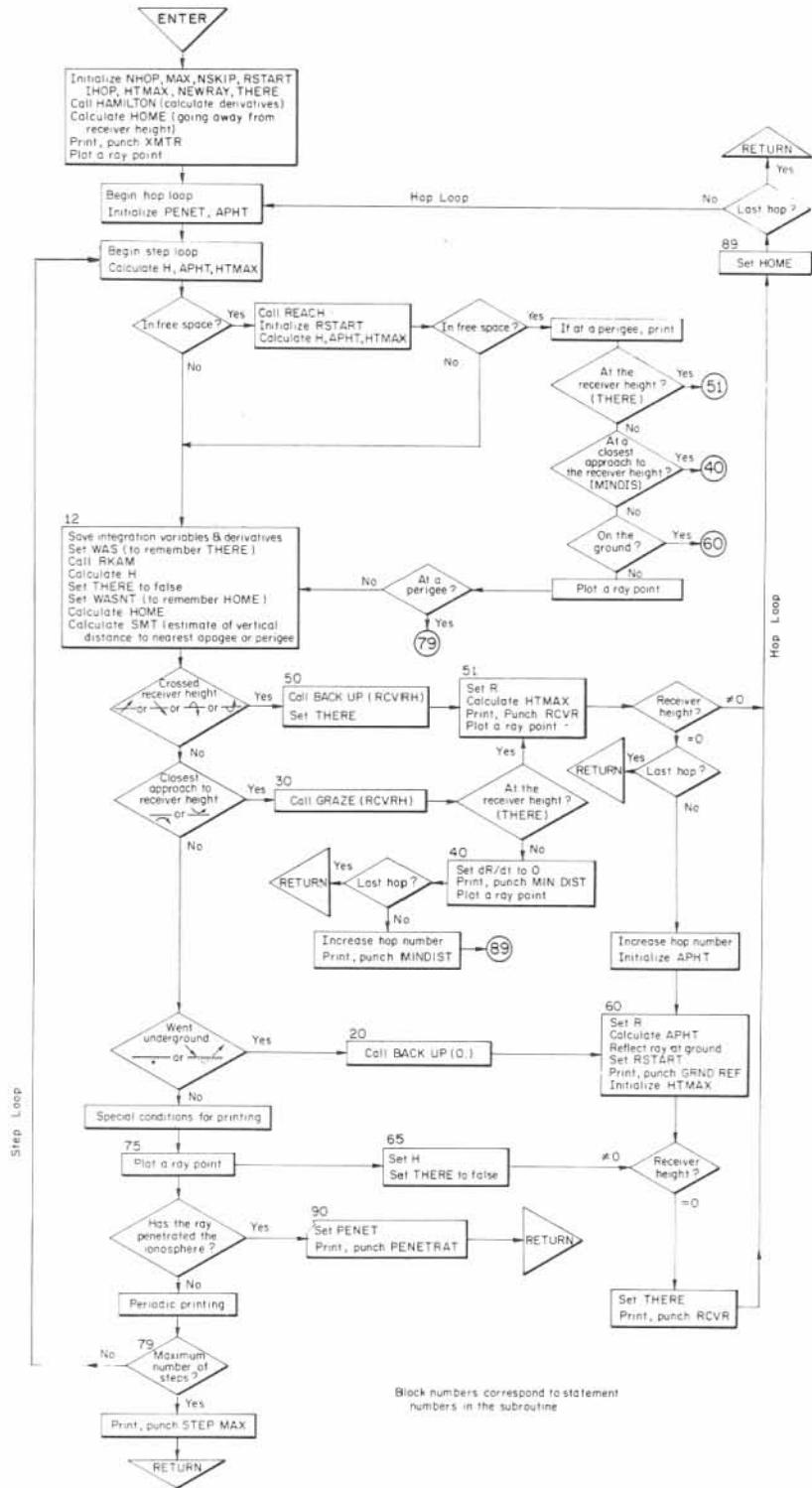


Figure 6. Flow chart for subroutine TRACE.

The ray graphics illustrate the path of a ray during a single step.

ionosphere or between ionospheric layers. Subroutine BACK UP finds an intersection of the ray with the receiver height or with the ground. Subroutine PRINTR prints information describing the ray path and punches the results on cards (raysets). Subroutine RAYPLT plots the ray path. The block diagram in figure 7 shows the relationship among these (and other) subroutines.

The listings of most of the subroutines have comments that should help in understanding how they work. In addition, Tables 3 through 14 define the variables in the common blocks.

#### 14. ACKNOWLEDGMENTS

Part of the organization of this program into subroutines follows that of the program of Dudziak (1961), in particular for subroutines RKAM, HAMLTN , RINDEX, ELECTX, MAGY, and COLFRZ. Also, the coordinate transformation in subroutine PRINTR and the method for data input via the W array are taken from the program of Dudziak (1961). The term "rayset," the idea of punching results of each hop for each ray trace onto cards, and the idea of automatically plotting ray paths come from the program of Croft and Gregory (1963). The quasi-parabolic layer electron density model QPARAB is taken from the paper by Croft and Hoogasian (1968). Notice that the quasi-parabolic layer that is now in the program is slightly different from the one in the program of Jones (1966). Subroutine RKAM is a modification of subroutine RKAMSUB, which was written by G. J. Lastman and is available through the CDC CO-OP library (the CO-OP identification is D2 UTEX RKAMSUB). Subroutine GAUSEL was written by L. David Lewis, Space Environment Laboratory, National Oceanic and Atmospheric Administration. Subroutine FSW was written in conjunction with Helmut Kopka of the Max-Planck-Institut für Aeronomie, Lindau/Harz, Germany.

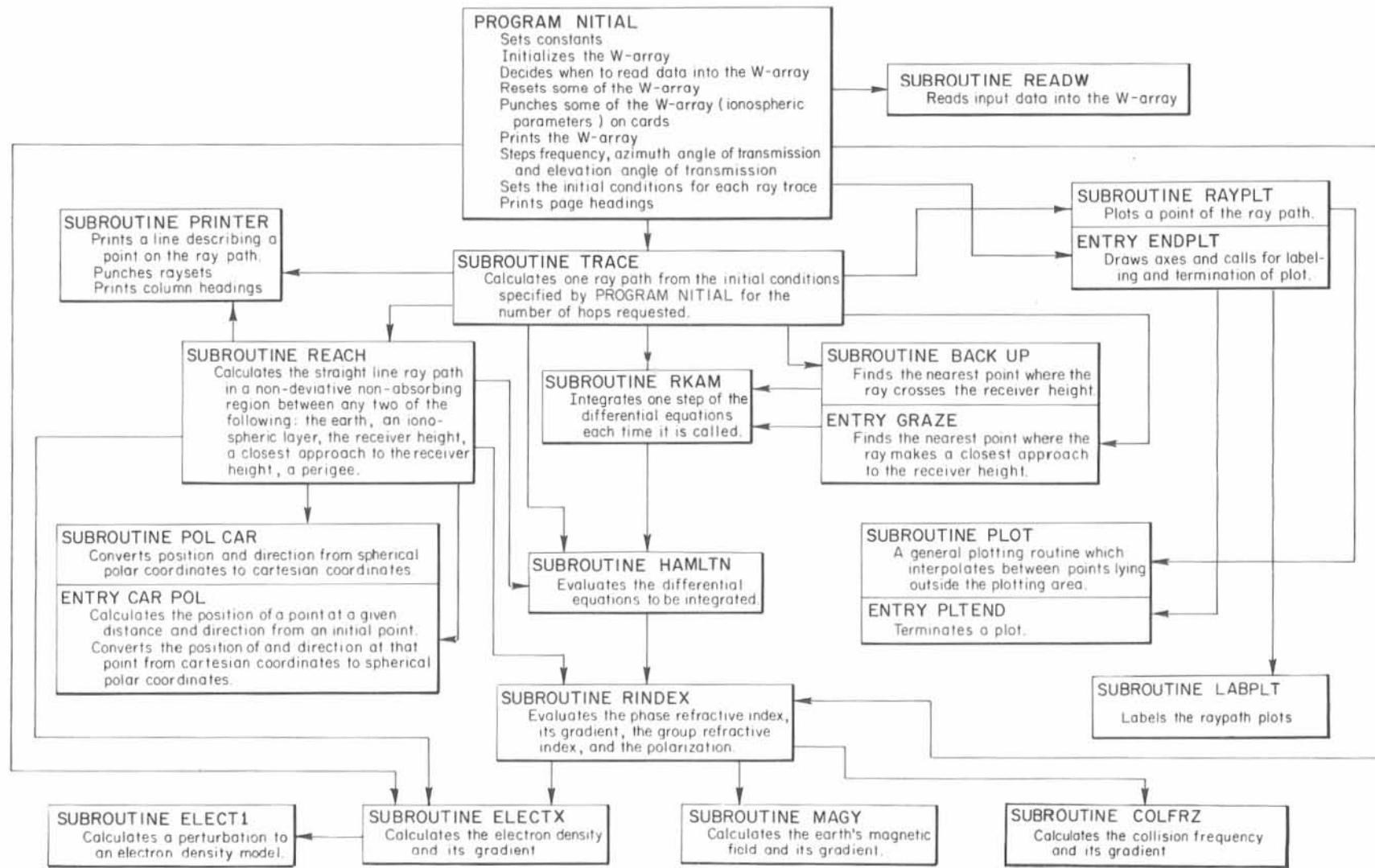


Figure 7. Block diagram for the ray tracing program.

Table 3. Definitions of the Parameters in Blank Common

Position in Common	Variable Name	Definition
1-20	R	The dependent variables in the differential equations being integrated--the definitions of the first six are fixed, but the others may be varied by the program user.
1	R(1)	$r$
2	R(2)	$\theta$
3	R(3)	$\varphi$
4	R(4)	$k_r$
5	R(5)	$k_\theta$
6	R(6)	$k_\varphi$
7-12	R(7)-R(12)	Those variables the user has chosen to integrate, taken in the following order: P -phase path in kilometers A -absorption in decibels $\Delta f$ -Doppler shift in hertz s -geometrical path length in kilometers
13-20	R(13)-R(20)	Reserved for future expansion.
21	T	Group path in kilometers (the independent variable in the differential equations).
22	STP	Step length in group path.
23-42	DRDT	The derivatives of the dependent variables with respect to the independent variable T.

R and T are initialized in program NITIAL and changed in subroutines RKAM, REACH, and BACK UP.

STP is calculated in subroutine RKAM.

DRDT is calculated in subroutine HAMILTN and used in subroutine RKAM.

Table 4. Definitions of the Parameters in Common Block /CONST/

Position in Common	Variable Name	Definition
1	PI	$\pi$
2	PIT2	$2\pi$
3	PID2	$\pi/2$
4	DEGS	$180.0/\pi$
5	RAD	$\pi/180.0$
6	K	Ratio of the square of the plasma frequency to the electron density in $\text{MHz}^2 \text{cm}^3 = r_e c^2 / \pi = e^2 / (4\pi^2 \epsilon_0 m)$ , where $r_e$ is the classical electron radius, $c$ is the free space speed of light, $e$ is the charge on the electron, $m$ is the mass of the electron, and $\epsilon_0$ is the capacitativity of a vacuum.
7	C	Free space speed of light in km/sec.
8	LOGTEN	$\log_e 10$

These parameters are set in program NITIAL.

Table 5. Definitions of the Parameters in Common Block /RK/

Position in Common	Variable Name	Definition
1	N	The number of equations being integrated.
2	STEP	The initial step in group path in kilometers.
3	MODE	Defines type of integration used (same as W41), see Table 2.
4	E1MAX	Maximum allowable single step error (same as W42).
5	E1MIN	Minimum allowable single step error (= W42/W43).
6	E2MAX	Maximum step length (same as W45).
7	E2MIN	Minimum step length (same as W46).
8	FACT	Factor by which to increase or decrease step length (same as W47).
9	RSTART	Nonzero to initialize numerical integration, zero to continue integration.

These parameters are calculated in program NITIAL (some are temporarily reset in subroutine BACK UP) and are used in subroutine RKAM.

Table 6. Definition of the Parameters in Common Block /RIN/

Position in Common	Variable Name	Definition
1, 2, 3	MODRIN	Description of version of RINDEX in BCD.
4	COLL	= 1 if this version of RINDEX includes collisions, = 0 otherwise.
5	FIELD	= 1 if this version of RINDEX includes the earth's magnetic field, = 0 otherwise.
6	SPACE	TRUE, if the ray is in a nondeviative, nonabsorbing medium.
7, 8	KAY2	$k^2$ , square of the complex phase refractive index times $\omega^2/c^2$ .
9, 10	H	Hamiltonian (complex)
11, 12	PHPT	$\partial H/\partial t$ (complex)
13, 14	PHPH	$\partial H/\partial r$ (complex)
15, 16	PHPHTH	$\partial H/\partial \theta$ (complex)
17, 18	PHPH	$\partial H/\partial \varphi$ (complex)
19, 20	PHPOM	$\partial H/\partial \omega$ (complex)
21, 22	PHPKR	$\partial H/\partial k_r$ (complex)
23, 24	PHPKTH	$\partial H/\partial k_\theta$ (complex)
25, 26	PHPKPH	$\partial H/\partial k_\varphi$ (complex)
27, 28	KPHPK	$\vec{k} \cdot \vec{\partial H}/\partial \vec{k}$ (complex)
29, 30	POLAR	$= k_r \partial H/\partial k_r + k_\theta \partial H/\partial k_\theta + k_\varphi \partial H/\partial k_\varphi$ Characteristic polarization of the wave; equal to the ratio of the component of the electric field perpendicular with the earth's magnetic field to the transverse component of the electric field parallel with the earth's magnetic field (complex) (Budden, 1961, p. 49, eq. (5.13)).

Table 6. (Continued)

Position in Common	Variable Name	Definition
31, 32	LPOLAR	Characteristic longitudinal polarization of the wave; equal to the ratio of the longitudinal component of the electric field to the component of the electric field perpendicular with the earth's magnetic field. (complex) Budden, 1961, p. 54, eq. (5.38)).
33	SGN	= +1 or -1; used for ray tracing in complex space.

These parameters are calculated in subroutine RINDEX and used in subroutine HAMLTN.

Note: In some subroutines, the real and imaginary parts of the complex variables have separate names.

Table 7. Definitions of the Parameters in Common Block /FLG/  
(See Subroutine TRACE)

Position in Common	Variable Name	Definition
1	NTYP	= 1 for extraordinary, = 2 for no field, = 3 for ordinary
2	NEWWR	Set equal to . TRUE. to tell subroutine RAYPLT there is a new W array.
3	NEWWP	Set equal to . TRUE. to tell subroutine PRINTR there is a new W array.
4	PENET	Set equal to . TRUE. if the ray just penetrated.
5	LINES	Number of lines printed on the current page.
6	IHOP	Hop number (at the beginning of each ray, subroutine TRACE sets this parameter to zero so that subroutine RAYPLT will begin a new line in plotting the ray path and subroutine PRINTR will print column headings and punch a transmitter rayset).
7	HPUNCH	The height to be punched on the raysets.

Table 8. Definitions of the Parameters in Common Block /XX/

Position in Common	Variable Name	Definition
1	MODX(1)	BCD name of the electron density subroutine.
2	MODX(2)	BCD name of the subroutine defining a perturbation to the electron density model.
3	X	X in Appleton-Hartree formula, square of the ratio of the plasma frequency to the wave frequency.
4	PXPR	$\frac{\partial X}{\partial r}$
5	PXPTH	$\frac{\partial X}{\partial \theta}$
6	PXPPH	$\frac{\partial X}{\partial \varphi}$
7	PXPT	$\frac{\partial X}{\partial t}$ , where t is time; used for calculating Doppler shifts.
8	HMAX	Height of maximum electron density.

These parameters are calculated in subroutine ELECTX, possibly modified in subroutine ELECT1, and are mainly used in subroutine RINDEX.

Table 9. Definitions of the Parameters in Common Block /YY/

Position in Common	Variable Name	Definition
1	MODY	BCD name of the subroutine defining the earth's magnetic field.
2	Y	Y in the Appleton-Hartree formula, ratio of the electron gyrofrequency to the wave frequency.
3	PYPR	$\frac{\partial Y}{\partial r}$
4	PYPTH	$\frac{\partial Y}{\partial \theta}$
5	PYPFH	$\frac{\partial Y}{\partial \phi}$
6	YR	$Y_r$ , proportional to the component of the earth's magnetic field in the $r$ direction.
7	PYRPR	$\frac{\partial Y_r}{\partial r}$
8	PYRPT	$\frac{\partial Y_r}{\partial \theta}$
9	PYRPP	$\frac{\partial Y_r}{\partial \phi}$
10	YTH	$Y_\theta$
11	PYTPT	$\frac{\partial Y_\theta}{\partial r}$
12	PYTPT	$\frac{\partial Y_\theta}{\partial \theta}$

Table 9. (Continued)

Position in Common	Variable Name	Definition
13	PYTPP	$\frac{\partial Y}{\partial \varphi}$
14	YPH	$Y_{\varphi}$
15	PYPPR	$\frac{\partial Y}{\partial r}$
16	PYPPT	$\frac{\partial Y}{\partial \theta}$
17	PYPBP	$\frac{\partial Y}{\partial \varphi}$

These parameters are calculated in subroutine MAGY and are mainly used in subroutine RINDEX.

Table 10. Definitions of the Parameters in Common Block /ZZ/

Position in Common	Variable Name	Definition
1	MODZ	BCD name of the collision frequency subroutine.
2	Z	Z in the Appleton-Hartree formula, ratio of the electron-neutral collision frequency to the angular wave frequency.
3	PZPR	$\frac{\partial Z}{\partial r}$
4	PZPTH	$\frac{\partial Z}{\partial \theta}$
5	PZPPH	$\frac{\partial Z}{\partial \varphi}$

These parameters are calculated in subroutine COLFRZ and are mainly used in subroutine RINDEX.

Table 11. Definitions of the Parameters in Common Block /TRAC/

Position in Common	Variable Name	Definition
1	GROUND	. TRUE. if the ray is on the surface of the earth.
2	PERIGE	. TRUE. if the ray has just made a perigee.
3	THERE	. TRUE. if the ray is at the receiver height.
4	MINDIS	. TRUE. if the ray has just made a closest approach to the receiver height.
5	NEWRAY	Set equal to .TRUE. to tell subroutine REACH that this is a new ray.
6	SMT	An estimation of the vertical distance to an apogee or perigee of the ray.

These parameters are used for communication between subroutine TRACE and subroutines REACH and BACK UP.

Table 12. Definition of the Parameter in Common Block /COORD/

Position in Common	Variable Name	Definition
1	S	The straight line distance along the ray from the position of the ray where REACH was called to the present position.

This parameter is used for communication between subroutine REACH and subroutine POL CAR.

Table 13. Definitions of the Parameters in Common Block /PLT/

Position in Common	Variable Name	Definition
1	XMIN0, XL	The x coordinate of the left side of the plotting area in kilometers.
2	XMAX0, XR	The x coordinate of the right side of the plotting area in kilometers.
3	XMIN0, YB	The y coordinate of the bottom of the plotting area in kilometers.
4	YMAX0, YT	The y coordinate of the top of the plotting area in kilometers.
5	RESET	Set equal to one whenever the plotting area is changed.

These parameters are used for communication between subroutine RAYPLT and subroutine PLOT.

Table 14. Definitions of the Parameters in Common Block /DD/

Position in Common	Variable Name	Definition
1	IN	Intensity. IN = 0 specifies normal intensity. IN = 1 specifies high intensity.
2	IOR	Orientation. IOR = 0 specifies upright orientation. IOR = 1 specifies rotated orientation (90° counterclockwise).
3	IT	Italics (Font). IT = 0 specifies non-Italic (Roman) symbols. IT = 1 specifies Italic symbols.
4	IS	Symbol size. IS = 0 specifies miniature size. IS = 1 specifies small size. IS = 2 specifies medium size. IS = 3 specifies large size.
5	IC	Symbol case. IC = 0 specifies upper case. IC = 1 specifies lower case.
6	ICC	Character code, 0-63 (R1 format). ICC and IC together specify the symbol plotted.
7	IX	X -coordinate, 0-1023.
8	IY	Y -coordinate, 0-1023.

We also want to thank those who have used our program and have pointed out errors or made suggestions. In particular, we are grateful to Dr. T. M. Georges of the Wave Propagation Laboratory, National Oceanic and Atmospheric Administration, for his suggestions resulting from extensive use of the program, for development of some of the ionospheric models (DCHAPT, DTORUS, WAVE, WAVE2), and for financing part of the development of ray tracing through a spitzer.

Examples of use of the ray tracing program are shown in the reports by Stephenson and Georges (1969) and Georges (1971).

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## APPENDIX 1. LISTINGS OF THE MAIN PROGRAM AND SUBROUTINES IN THE MAIN DECK

This appendix contains listings of the main program and those subroutines that have only one version (with the exception of subroutine RAYPLT, which has a do-nothing version for users lacking a plotter to plot ray paths). Thus, the routines which form the contents of this appendix will be used in all ray path calculations.

Additionally, this appendix contains the main input parameter form for ray tracing and the input parameter forms for plotting. These forms are very useful when using the program because they indicate the input parameters needed for ray path calculations.

The contents of this appendix are arranged as follows:

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INPUT PARAMETER FORM  
FOR THREE-DIMENSIONAL RAY PATHS

Name \_\_\_\_\_ Project No. \_\_\_\_\_ Date \_\_\_\_\_

Ionospheric ID (3 characters) \_\_\_\_\_

Title (75 characters) \_\_\_\_\_

Models:	Electron density	_____
	Perturbation	_____
	Magnetic field	_____
	Ordinary	_____ (W1 = + 1.)
	Extraordinary	_____ (W1 = - 1.)
	Collision frequency	_____
Transmitter:	Height	_____ km, nautical miles, feet (W3)
	Latitude	_____ rad, deg, km (W4)
	Longitude	_____ rad, deg, km (W5)
	Frequency, initial	_____ MHz (W7)
	final	_____ (W8)
	step	_____ (W9)
	Azimuth angle, initial	_____ rad, deg clockwise of north (W11)
	final	_____ (W12)
	step	_____ (W13)
	Elevation angle, initial	_____ rad, deg (W15)
	final	_____ (W16)
	step	_____ (W17)
Receiver:	Height	_____ km, nautical miles, feet (W20)
Penetrating rays:	Wanted	_____ (W21 = 0.)
	Not wanted	_____ (W21 = 1.)
Maximum number of hops		_____ (W22)
Maximum number of steps per hop		_____ (W23)
Maximum allowable error per step		_____ (W42)
Additional calculations:		= 1. to integrate = 2. to integrate and print
Phase path		_____ (W57)
Absorption		_____ (W58)
Doppler shift		_____ (W59)
Path length		_____ (W60)
Other		_____
		_____
		_____

Printout: Every \_\_\_\_\_ steps of the ray trace (W71)

Punched cards (raysets): \_\_\_\_\_ (W72 = 1.)

```

C PROGRAM NITIAL                                NITI001
      SETS THE INITIAL CONDITIONS FOR EACH RAY AND CALLS TRACE   NITI002
DIMENSION MFLD(2)                                NITI003
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,C,LOGTEN    NITI004
COMMON /FLG/ NTYP,NEWWR,NEWWP,PENET,LINES,IHOP,HPUNCH  NITI005
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,N2,N2I,PNP(10),POLAR,  NITI006
1          LPOLAR,SGN                           NITI007
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART  NITI008
COMMON /XX/ HOOK(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX        NITI009
COMMON /YY/ HODY,Y(16),ZZ/MODZ,Z(4)                  NITI010
COMMON R(20),T,STP,DRDT(20)/WW/I)(10),W0,W(400)       NITI011
EQUIVALENCE (RAY,W(1)),(EARTH,R,W(2)),(XMTRH,W(3)),(TLAT,W(4)),  NITI012
1 (TLON,W(5)),(F,W(6)),(FBEG,W(7)),(FEND,W(8)),(FSTEP,W(9)),  NITI013
2 (AZ1,W(10)),(AZBEG,W(11)),(AZEND,W(12)),(AZSTEP,W(13)),  NITI014
3 (BETA,W(14)),(ELBEG,W(15)),(ELEND,W(16)),(ELSTEP,W(17)),  NITI015
4 (ONLY,W(21)),(HOP,W(22)),(MAXSTP,W(23)),(PLAT,W(24)),(PLON,W(25)) NITI016
5 (INTYP,W(41)),(MAXERR,W(42)),(ERATIO,W(43)),(STEP1,W(44)),  NITI017
6 (STPMAX,W(45)),(STPMIN,W(46)),(FACTR,W(47)),(SKIP,W(71)),  NITI018
7 (RAYSET,W(72)),(PLT,W(81)),(PERT,W(150))           NITI019
LOGICAL SPACE,NEWWR,NEWWP,PENET                  NITI020
REAL N2,N2I,LOGTEN,K,MAXSTP,INTYP,MAXERR,MU        NITI021
COMPLEX PNP,POLAR,LPOLAR                         NITI022
NDATE=IDATE(0)                                    NITI025
SECOND=KLOCK(0)*.001                            NITI026
COLL=4H NO                                     NITI027
IF (COLL.NE.0.) KOLL=4HWITH                   NITI028
C***** CONSTANTS                               NITI029
PI=3.1415926536                                  NITI030
PIT2=2.*PI                                       NITI031
PID2=PI/2.                                       NITI032
DEGS=180./PI                                     NITI033
RAD=PI/180.                                      NITI034
C=2.997925E5                                    NITI035
K=2.81785E-15*C**2/PI                          NITI036
LOGTEN=ALOG(10.)                                 NITI037
C***** INITIALIZE SOME VARIABLES IN THE W ARRAY  NITI038
DO 5 NW=1,400                                     NITI039
5 W(NW)=0.                                       NITI040
PLON=0.                                         NITI041
PLAT=PID2                                     NITI042
EARTH=6370.                                     NITI043
INTYP=3.                                         NITI044
MAXERR=1.E-4                                     NITI045
ERATIO=50.                                        NITI046
STEP1=1.                                         NITI047
STPMAX=100.                                       NITI048
STPMIN=1.E-8                                     NITI049
FACTR=0.5                                         NITI050
MAXSTP=1000.                                      NITI051
HOP=1.                                           NITI052
C***** READ W ARRAY AND PRINT NON-ZERO VALUES  NITI053
10 CALL READ W                                    NITI054
F=BETA=AZ1=0.                                     NITI055
IF (SKIP.EQ.0.) SKIP=MAXSTP                      NITI056
12 RAY=SIGN(1.,RAY)                             NITI057
NTYP=2.+FIELD*RAY                                NITI058
GO TO (13,14,15), NTYP                          NITI059
13 MFLD(1)=8HEXTRAORD                           NITI060
MFLD(2)=5HINARY                                 NITI061
GO TO 16                                         NITI062
14 MFLD(1)=8HNO FIELD                           NITI063
MFLD(2)=1H                                       NITI064
GO TO 16                                         NITI065

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15 MFLD(1)=8HORDINARY NITI066
MFLD(2)=8H NITI067
16 MODSAV=MODX(2) NITI068
IF (PERT.EQ.0.) MODX(2)=6H NITI069
IF (RAYSET.NE.0.) PUNCH 2000, ID,MODX(1),(W(NW),NW=101,107), NITI070
1 MODX(2),(W(NW),NW=151+157),MODY,(W(NW),NW=201,207), NITI071
2 MODZ,(W(NW),NW=251,257) NITI072
2000 FORMAT (10A8,4(/A8,2X,7E10.3)) NITI073
PRINT 1000, ID,NUATE,MODX,MODY,MODZ,MODRIN,MFLD,KOLL NIIIU4
1000 FORMAT (1H1,10A8,25X,A8/4(1X+A8),24X,3A8,1X+A8,A8,A8,1X,A4+ NITI075
1 11H COLLISIONS/) NITI076
PRINT 1050 NITI077
1050 FORMAT (85H INITIAL VALUES FOR THE W ARRAY -- ALL ANGLES IN RADIANNIIIU79
1S, ONLY NONZERO VALUES PRINTED/) NIIIU79
DO 17 NW=1,400 NIIIU80
IF (W(NW).NE.0.) PRINT 1700, NW,W(NW) NITI081
1700 FORMAT (I4,E19.11) NITI082
17 CONTINUE NITI083
***** LET SUBROUTINES PRINTR AND RAYPLT KNOW THERE IS A NEW W ARRAYNITI084
NEWWP=.TRUE. NITI085
NEWWR=.TRUE. NITI086
***** INITIALIZE PARAMETERS FOR INTEGRATION SUBROUTINE RKAM NITI087
N=6 NITI088
DO 20 NR=7,20 NITI089
IF (W(50+NR).NE.0.) N=N+1 NITI090
20 CONTINUE NIIIU91
MODE=INTYP NITI092
STEP=STEP1 NIIIU93
E1MAX=MAXERR NIIIU94
E1MIN=MAXERR/ERATIO NITI095
E2MAX=STPMAX NITI096
E2MIN=STPMIN NITI097
FACT=FACTR NITI098
***** DETERMINE TRANSMITTER LOCATION IN COMPUTATIONAL COORDINATE NITI099
***** SYSTEM (GEOMAGNETIC COORDINATES IF DIPOLE FIELD IS USED) NIIIU100
R0=EARTH+XMTRH NITI101
SP=SIN (PLAT) NITI102
CP=SIN (PID2-PLAT) NITI103
SDPH=SIN (TLON-PLON) NITI104
CDPH=SIN (PID2-(TLON-PLON)) NITI105
SL=SIN (TLAT) NITI106
CL=SIN (PID2-TLAT) NITI107
ALPHA=ATAN2(-SDPH*CP,-CDPH*CP*SL+SP*CL) NITI108
TH0=ACOS (CDPH*CP*CL+SP*SL) NITI109
PH0=ATAN2(SDPH*CL,CDPH*SP*CL-CP*SL) NITI110
***** LOOP ON FREQUENCY, AZIMUTH ANGLE, AND ELEVATION ANGLE NITI111
NFREQ=1 NIIIU112
IF (FSTEP.NE.0.) NFREQ=(FEND-FBEG)/FSIEP+1.5 NIIIU113
NAZ=1 NIIIU114
IF (AZSTEP.NE.0.) NAZ=(AZEND-AZBEG)/AZSIEP+1.5 NITI115
NBETA=1 NITI116
IF (ELSTEP.NE.0.) NBETA=(ELEND-ELBEG)/ELSTEP+1.5 NITI117
DO 50 NF=1,NFREQ NITI118
FBEG+(NF-1)*FSTEP NITI119
DO 45 J=1,NAZ NITI120
AZ1=AZBEG+(J-1)*AZSTEP NITI121
AZA=AZ1*DEGS NIIIU122
GAMMA=PI-AZ1+ALPHA NITI123
SGAMMA=SIN (GAMMA) NITI124
CGAMMA=SIN (PID2-GAMMA) NITI125
DO 40 I=1,NBETA NITI126
BETA=ELBEG+(I-1)*ELSTEP NITI127

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      EL=BETA*DFGS          NITI128
      CBETA=SIN (PID2-BETA) NITI129
      R(1)=R0                NITI130
      R(2)=TH0               NITI131
      R(3)=PH0               NITI132
      R(4)=SIN (BETA)        NITI133
      R(5)=CBETA*CGAMMA     NITI134
      R(6)=CBETA*SGAMMA     NITI135
      T=0.                  NITI136
      RSTART=1.              NITI137
C      SGN=1.   (NEED FOR RAY TRACING IN COMPLEX SPACE.) NITI138
C***** ALLOW IONOSPHERIC MODEL SUBROUTINES TO READ AND PRINT DATA NITI139
      CALL RINDEX
      IF (I.NE.1.AND.NPAGE.LT.3.AND.LINES.LE.17) GO TO 25
      NPAGE=LINES=0
      PRINT 1000, ID,NDATE,MODX,MODY,MODZ,MUDRIN,MFLD,KULL
      PRINT 2400, F,AZA
2400 FORMAT (18X,11HFREQUENCY =,F12.6,37H MHZ, AZIMUTH ANGLE OF TRANSMINITI145
      ISSION =,F12.6,4H DEG) NITI146
25  NPAGE=NPAGE+1          NITI147
      PRINT 2500, EL          NITI148
2500 FORMAT (/31X,33HELEVATION ANGLE OF TRANSMISSION =,F12.6,4H DEG/) NITI149
      IF (N2.GT.0.) GO TO 30
      CALL ELECTX
      FN=SIGN (SQRT (ABS (X))*F,X)
      PRINT 2900, FN          NITI153
2900 FORMAT (58HOTRANSMITTER IN EVANESCENT REGION, TRANSMISSION IMPOSSINITI154
      1BLE/20H0PLASMA FREQUENCY = ,E17.10) NITI155
      GO TO 44
30  MU=SQRT (N2/(R(4)**2+R(5)**2+R(6)**2)) NITI157
      DO 34 NN=4,6            NITI158
34  R(NN)=R(NN)*MU        NITI159
      DO 35 NN=7,N            NITI160
35  R(NN)=0.              NITI161
      CALL TRACE
      OSEC=SECOND
      SECOND=KLOCK(n)*.001    NITI163
      DIFF=SECOND-OSEC        NITI164
      PRINT 3500, DIFF        NITI165
      PRINT 3500, DIFF        NITI166
3500 FORMAT (36X,26HTHIS RAY CALCULATION TOOK ,F8.3,4H SEC) NITI167
      IF (PENET.AND.ONLY.NE.0..AND.IHOP.EQ.1) GO TO 44
40  CONTINUE               NITI168
44  IF (PLT.NE.0.) CALL ENDPLT NITI169
45  CONTINUE               NITI170
      IF(PENET.AND.ONLY.NE.0..AND.IHOP.EQ.1.AND.NAZ.EQ.1.AND.NBETA.EQ.1)NITI172
      1 GO TO 55
50  CONTINUE               NITI173
55  IF (RAYSET.NE.0.) PUNCH 5000 NITI174
5000 FORMAT (78X,1H-)
      MODX(2)=MODSAV         NITI176
      GO TO 10
      END                     NITI178
                                         NITI179-

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SUBROUTINE READ W                               READ001
C           READS W ARRAY                      READ002
C           A 1 IN THE FOLLOWING COLUMNS WILL MAKE THE DESCRIBED CONVERSIONSREAD003
COL 18  DEGREES TO RADIANS                   READ004
COL 19  GREAT CIRCLE DISTANCE IN KM TO RADIANREAD005
COL 20  NAUTICAL MILES TO KM                  READ006
COL 21  FEET TO KM                         READ007
C
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,DUM(3)    READ008
COMMON /WW/ ID(10),W0,W(400)                   READ009
EQUIVALENCE (EARTH, W(2))                     READ010
INTEGER DEG,FEET                           READ011
READ 1000, ID                      READ012
1000 FORMAT (10A8)                      READ013
IF (EOF,60) 3,4                         READ014
3 CALL EXIT                            READ015
4 READ 1100, NW,W(NW),DEG,KM,NM,FEET      READ016
1100 FORMAT (I3,E14.7,5I1)                 READ017
IF (NW.EQ.0) GO TO 10                    READ018
IF (NW.GT.0.AND.NW.LE.400) GO TO 5       READ019
PRINT 4000, NW                         READ020
PRINT 4000, NW                         READ021
4000 FORMAT(15H1THE SUBSCRIPT ,I3,77H ON THE W-ARRAY INPUT IS OUT OF BOREAD022
1UNDS. ALLOWABLE VALUES ARE 1 THROUGH 400.)   READ023
CALL EXIT                                READ024
5 IF (DEG.NE.0.) W(NW)=W(NW)*RAD          READ025
IF (KM.NE.0) W(NW)=W(NW)/EARTH           READ026
IF (NM.NE.0) W(NW)=W(NW)*1.852          READ027
IF (FEET.NE.0) W(NW)=W(NW)*3.048006096E-4  READ028
GO TO 4                                     READ029
10 RETURN                                 READ030
END                                      READ031-

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SUBROUTINE TRACE                               TRAC001
C           CALCULATES THE RAY PATH             TRAC002
DIMENSION ROLD(20),DROLD(20)                  TRAC003
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART   TRAC004
COMMON /FLG/ NTYP,NEWWR,NEWWP,PENET,LINES,IHOP,HPUNCH        TRAC005
COMMON /TRAC/ GROUND,PERIGE,THERE,MINDIS,NEWRAY,SMT         TRAC006
COMMON /RIN/ MODIN(3),COLL,FIELD,SPACE,N2,PNP(10),POLAR,LPOLAR TRAC007
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX            TRAC008
COMMON R(20),T,STP,DRDT(20) /WW/ ID(10),W0,W(400)           TRAC009
LOGICAL SPACE,HOME,WASN1,UNDRGD,GROUND,PERIGE,THERE,MINDIS,NEWWR,  TRAC010
1           NEWWP,PENET,NEWRAY,WAS              TRAC011
REAL MAXSTP                                TRAC012
COMPLEX N2,PNP,POLAR,LPOLAR                 TRAC013
EQUIVALENCE (EARTH, W(2)),(RCVRH,W(20)),(HOP,W(22)),(MAXSTP,W(23)) TRAC014
1,(SKIP,W(71)),(RAYSET,W(72)),(PLT,W(81))          TRAC015
N>HOP=HOP                                  TRAC016
MAX=MAXSTP                                TRAC017
NSKIP=SKIP                                TRAC018
RSTART=1.                                    TRAC019
CALL HAMLTN                                TRAC020
HOME=DRDT(1)*(R(1)-EARTH-RCVRH).GE.0.          TRAC021
***** IHOP=0 TELLS PRINTR TO PRINT HEADING AND PUNCH A TRANSMITTER TRAC022
***** RAYSET AND TELLS RAYPLT TO START A NEW RAY          TRAC023
IHOP=0                                     TRAC024
CALL PRINTR (8HXHTR ,0.)                   TRAC025
IF (PLT.NE.0.) CALL RAYPLT                TRAC026
HTMAX=0.                                    TRAC027
NEWRAY=.TRUE.                                TRAC028
THERE=R(1)-EARTH.R.EQ.RCVR1                TRAC029

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C***** LOOP ON NUMBER OF HOPS
10 IHOP=IHOP+1
   IF (IHOP.GT.NHOP) RETURN
   PENET=.FALSE.
   APHT=RCVRH
C***** LOOP ON MAXIMUM NUMBER OF STEPS PER HOP
   DO 79 J=1,MAX
   H=R(1)-EARTH
   IF (ABS(H-RCVRH).GT.ABS(APHT-RCVRH)) APHT=H
   HTMAX=AMAX1(H,HTMAX)
   IF (.NOT.SPACE) GO TO 12
   CALL REACH
   RSTART=1.
   H=R(1)-EARTH
   IF (ABS(H-RCVRH).GT.ABS(APHT-RCVRH)) APHT=H
   HTMAX=AMAX1(H,HTMAX)
   IF (.NOT.SPACE) GO TO 12
   IF (PERIGE) CALL PRINTR (8HPERIGEE ,0.)
   IF (THERE) GO TO 51
   IF (MINDIS) GO TO 40
   IF (GROUND) GO TO 60
   IF (PLT.NE.0.) CALL RAYPLT
   IF (PERIGE) GO TO 79
12 DO 13 L=1,N
   ROLD(L)=R(L)
13 DROLD(L)=DRDT(L)
   TOLD=T
   HAS=THERE
   CALL RKAM
   H=R(1)-EARTH
   THERE=.FALSE.
   HASNT=.NOT.HOME
   HOME=DROLD(1)*(H-RCVRH).GE.0.
   TMP=(DRDT(1)-DROLD(1))*(T-TOLD)
   SMT=0.
   IF (TMP.NE.0.) SMT=0.5*(R(1)-ROLD(1)+0.5*TMP)**2/ABS(TMP)
   IF (((H-RCVRH)*(ROLD(1)-EARTH-RCVRH).LT.0..AND..NOT.HAS).OR.
1 (HAS.AND.DRDT(1)*DROLD(1).LT.0..AND.HOME)) GO TO 50
   IF (HOME.AND.HASNT) GO TO 30
   IF (H.LT.0..OR.DRDT(1).GT.0..AND.DROLD(1).LT.0..AND.SMT.GT.H)
1 GO TO 20
   IF (DROLD(1).LT.0..AND.DRDT(1).GT.0.) CALL PRINTR(8HPERIGEE ,0.)
   IF (DROLD(1).GT.0..AND.DRDT(1).LT.0.) CALL PRINTR(8HAPOGEE ,0.)
   IF (DROLD(2)*DRDT(2).LT.0.) CALL PRINTR(8HMAX LAT ,0.)
   IF (DROLD(3)*DRDT(3).LT.0.) CALL PRINTR(8HMAX LONG,0.)
   DO 14 I=4,6
   IF (ROLD(I)*R(I).LT.0.) CALL PRINTR(8HHAVE REV,0.)
14 CONTINUE
   GO TO 75
C***** RAY WENT UNDERGROUND
20 CALL BACK UP(0.)
   GO TO 60
C***** RAY MAY HAVE MADE A CLOSEST APPROACH
30 CALL GRAZE(RCVRH)
   IF (THERE) GO TO 51
40 DRDT(1)=0.
   HPUNCH=R(1)-EARTH
   CALL PRINTR(8HMIN DIST,RAYSET)
   IF (PLT.NE.0.) CALL RAYPLT
   IF (IHOP.GE.NHOP) RETURN
   IHOP=IHOP+1
   CALL PRINTR (8HMIN DIST,RAYSET)
   GO TO 89

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C***** RAY CROSSED RECEIVER HEIGHT          TRAC093
 50 CALL BACK UP(RCVRH)                      TRAC094
    THERE=.TRUE.                               TRAC095
 51 R(1)=EARTH+RCVRH                         TRAC096
    HTMAX=AMAX1(RCVRH,HTMAX)                  TRAC097
    HPUNCH=APHT                                TRAC098
    CALL PRINTR(8HRCVR      ,RAYSET)          TRAC099
    IF (PLT.NE.0.) CALL RAYPLT                TRAC100
    IF (RCVRH.NE.0.) GC TO 89                 TRAC101
    IF (IHOP.GE.NHOP) RETURN                  TRAC102
    IHOP=IHOP+1                                TRAC103
    APHT=RCVRH                                TRAC104
    APHT=RCVRH                                TRAC105
C***** GROUND REFLECT                      TRAC106
 60 R(1)=EARTH+R
    IF (ABS(RCVRH).GT.ABS(APHT-RCVRH)) APHT=0.   TRAC107
    R(4)=ABS (R(4))
    DRDT(1)=ABS (DRDT(1))
    RSTART=1.
    HPUNCH=HTMAX
    CALL PRINTR(8HGRND REF,RAYSET)
    HTMAX=0.
    IF (RCVRH.NE.0.) GO TO 65
    THERE=.TRUE.
    HPUNCH=APHT
    CALL PRINTR (8HRCVR      ,RAYSET)
    GO TO 89
 55 H=0.
    THERE=.FALSE.
C***** EXCEEDED MAXIMUM NUMBER OF STEPS     TRAC119
 75 IF (PLT.NE.0.) CALL RAYPLT               TRAC122
    IF (H.GT.HMAX.AND.H.GT.RCVRH.AND.DRDT(1).GT.0.) GO TO 90   TRAC123
    IF (MOD(J,NSKIP).EQ.0) CALL PRINTR(8H      ,0.)
 79 CONTINUE                                 TRAC125
C***** EXCEEDED MAXIMUM NUMBER OF STEPS     TRAC126
 75 IF (PLT.NE.0.) CALL RAYPLT               TRAC127
    IF (H.GT.HMAX.AND.H.GT.RCVRH.AND.DRDT(1).GT.0.) GO TO 90   TRAC128
    IF (MOD(J,NSKIP).EQ.0) CALL PRINTR(8H      ,0.)
 79 CONTINUE                                 TRAC129
C***** EXCEEDED MAXIMUM NUMBER OF STEPS     TRAC130
 75 IF (PLT.NE.0.) CALL RAYPLT               TRAC131
    IF (H.GT.HMAX.AND.H.GT.RCVRH.AND.DRDT(1).GT.0.) GO TO 90   TRAC132
    IF (MOD(J,NSKIP).EQ.0) CALL PRINTR(8H      ,0.)
 79 CONTINUE                                 TRAC133
 90 PENET=.TRUE.
    HPUNCH=H
    CALL PRINTR(8HPENETRAT,RAYSET)
    RETURN
  END                                         TRAC138

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SUBROUTINE BACK UP(HS)
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART
COMMON /TRAC/ GROUND,PERIGE,THERE,MINDIS,NEWRAY,SMT
COMMON R(20),T,STP,DRDT(20) /WW/ ID(10),W0,W(400)
EQUIVALENCE (EARTH,W(2)),(INTYP,W(41)),(STEP1,W(44))
REAL INTYP
LOGICAL GROUND,PERIGE,THERE,MINDIS,NEWRAY,HOME
C***** DIAGNOSTIC PRINTOUT
C      CALL PRINTR (8HBACK UP0,0.)
C***** GOING AWAY FROM THE HEIGHT HS
HOME=DRDT(1)*(R(1)-EARTH-HS).GE.0.
IF (HS.GT.0..AND..NOT.HOME.OR.HS.EQ.0..AND.DRDT(1).GT.0.) GO TO 30BACK012

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***** FIND NEAREST INTERSECTION OF RAY WITH THE HEIGHT HS          BACK013
DO 10 I=1,10
STEP=-(R(1)-EARTH-R-HS)/DRDT(1)                                     BACK014
STEP=SIGN(AMIN1(ABS(STP),ABS(STEP)),STEP)
IF (ABS(R(1)-EARTH-R-HS).LT..5E-4.AND.ABS(STEP).LT.1.) GO TO 60    BACK015
***** DIAGNOSTIC PRINTOUT                                         BACK016
C      CALL PRINTR(8HBACK UP1,0.)
MODE=1
RSTART=1.
CALL RKAM
10 RSTART=1.

C ***** FIND NEAREST CLOSEST APPROACH OF RAY TO THE HEIGHT HS      BACK017
ENTRY GRAZE
THERE=.FALSE.
***** DIAGNOSTIC PRINTOUT                                         BACK018
C      CALL PRINTR (8HGRAZE 0 +0.)
IF (SMT.GT.ABS(R(1)-EARTH-R-HS)) GO TO 30                         BACK019
DO 20 I=1,10
STEP=-R(4)/DRDT(4)
STEP=SIGN(AMIN1(ABS(STP),ABS(STEP)),STEP)
IF (ABS(R(4)).LE.1.E-6.AND.ABS(STEP).LT.1.) GO TO 60                BACK020
***** DIAGNOSTIC PRINTOUT                                         BACK021
C      CALL PRINTR (8HGRAZE 1 +0.)
MODE=1
RSTART=1.
CALL RKAM
RSTART=1.
IF (DRDT(4)*(R(1)-EARTH-R-HS).LT.0.) GO TO 30                      BACK022
IF(R(5).EQ.0..AND.R(6).EQ.0.) GO TO 60                                BACK023
20 CONTINUE
***** IF A CLOSEST APPROACH COULD NOT BE FOUND IN 10 STEPS, IT      BACK024
***** PROBABLY MEANS THAT THE RAY INTERSECTS THE HEIGHT HS        BACK025
30 CONTINUE
***** DIAGNOSTIC PRINTOUT                                         BACK026
C      CALL PRINTR (8HBACK UP2,0.)
MODE=1
***** ESTIMATE DISTANCE TO NEAREST INTERSECTION OF RAY WITH HEIGHT BACK027
***** HS BEHIND THE PRESENT RAY POINT                           BACK028
STEP=(-R(4)-SQRT(R(4)**2-2.*((R(1)-EARTH-R-HS)*DRDT(4)))/DRDT(4) BACK029
RSTART=1.
CALL RKAM
RSTART=1.
***** FIND NEAREST INTERSECTION OF RAY WITH HEIGHT HS            BACK030
DO 40 I=1,10
STEP=-(R(1)-EARTH-R-HS)/DRDT(1)                                     BACK031
STEP=SIGN(AMIN1(ABS(STP),ABS(STEP)),STEP)
IF (ABS(R(1)-EARTH-R-HS).LT..5E-4.AND.ABS(STEP).LT.1.) GO TO 60    BACK032
***** DIAGNOSTIC PRINTOUT                                         BACK033
C      CALL PRINTR (8HBACK UP3,0.)
MODE=1
RSTART=1.
CALL RKAM
40 RSTART=1.
50 THERE=.TRUE.
***** RESET STANDARD MODE AND INTEGRATION TYPE                  BACK034
60 MODE=INTYP
STEP=STEP1
RETURN
END

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SUBROUTINE REACH          REAC001
C      CALCULATES THE STRAIGHT LINE RAY PATH BETWEEN THE EARTH    REAC002
C      AND THE IONOSPHERE OR BETWEEN IONOSPHERIC LAYERS           REAC003
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART      REAC004
COMMON /TRAC/ GROUND,PERIGE,THERE,MINDIS,NEWRAY,SMT            REAC005
COMMON /COORD/ S
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,N2,N2I,PNP(10),POLAR,   REAC006
1      LPOLAR
COMMON /XXX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX             REAC008
COMMON R(20),T,STP,DRDT(20) /WW/ ID(10),W0,W(400)              REAC009
EQUIVALENCE (EARTH,R(2)),(XMTRH,W(3)),(RCVRH,W(20))           REAC010
LOGICAL CROSS,CROSSG,CROSSR,SPACE,GROUND,PERIGE,THERE,MINDIS,   REAC011
1      NEWRAY, RSPACE
REAL N2,N2I
COMPLEX PNP,POLAR,LPOLAR
DATA (NSTEP=500)
CALL HAMLTN
H=R(1)-EARTH
IF (.NOT.NEWRAY.AND..NOT.RSPACE) CALL PRINTR(8HEXIT ION,0.)
NEWRAY=.FALSE.
V=SQRT(R(4)**2+R(5)**2+R(6)**2)
C***** NORMALIZE THE WAVE NORMAL DIRECTION TO ONE
R(4)=R(4)/V
R(5)=R(5)/V
R(6)=R(6)/V
C***** NEGATIVE OF DISTANCE ALONG RAY TO CLOSEST APPROACH TO CENTER REAC26
C***** OF EARTH
UP=R(1)*R(4)
RADG=EARTH**2-R(1)**2*(R(5)**2+R(6)**2)
DISTG=SQRT (AMAX1(0.,RADG))
C***** DISTANCE ALONG RAY TO FIRST INTERSECTION WITH OR CLOSEST REAC31
C***** APPROACH TO THE EARTH
SG=-UP-DISTG
C***** CROSSG IS TRUE IF THE RAY WILL INTERSECT OR TOUCH THE EARTH REAC34
CROSSG=UP.LT.0..AND.RADG.GE.0.
RADR=(EARTH+RCVRH)**2-R(1)**2*(R(5)**2+R(6)**2)
DISTR=SQRT (AMAX1(0.,RADR))
C***** DISTANCE ALONG RAY TO THE FIRST INTERSECTION WITH OR CLOSEST REAC38
C***** APPROACH TO THE RECEIVER HEIGHT
SR=DISTR-UP
IF (UP.LT.0..AND.DISTR.LT.-UP.AND.R(1).NE.EARTH+RCVRH) SR=-DISTR
1 -UP
C***** CROSSR IS TRUE IF THE RAY WILL INTERSECT WITH OR MAKE A REAC43
C***** CLOSEST APPROACH TO THE RECEIVER HEIGHT
CROSSR=R(4).LT.0..OR.R(1).LT.(EARTH+RCVRH)
CROSS=CROSSG.OR.CROSSR
C***** MAXIMUM DISTANCE IN WHICH TO LOOK FOR THE IONOSPHERE
S1=AMIN1(SR,SG)
IF(.NOT.CROSSG) S1=SR
IF (UP.GE.0.) GO TO 15
CROSS=.TRUE.
C***** IF RAY IS GOING DOWN, S1 IS AT MOST THE DISTANCE TO A PERIGEE REAC52
S1=AMIN1(S1,-UP)
C***** CONVERT THE POSITION AND DIRECTION OF THE RAY TO CARTESIAN REAC54
C***** COORDINATES
15 CALL POL CAR
SSTEP=100.
S=SSTEP
DO 20 I=1,NSTEP
  IF ((S-SSTEP).GT.S1.AND.CROSS) GO TO 25
C***** CONVERT POSITION AND DIRECTION TO SPHERICAL POLAR COORDINATES REAC61
C***** AT A DISTANCE S ALONG THE RAY
  CALL CAR POL
  CALL ELECTX
C***** FREE SPACE

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        IF (X.EQ.0.) GO TO 20                                REAC 66
        CALL RINDEX                                         REAC 67
C***** EFFECTIVELY FREE SPACE                           REAC 68
        IF (SPACE) GO TO 20                                 REAC 69
        IF (SSTEP.LT.0.5E-4) GO TO 25                      REAC 70
C***** RAY IN THE IONOSPHERE. STEP BACK OUT           REAC 71
        S=S-SSTEP                                         REAC 72
C***** DECREASE STEP SIZE                            REAC 73
        SSTEP=SSTEP/10.                                    REAC 74
20      S=S+SSTEP                                     REAC 75
        PRINT 2000, NSTEP                                REAC 76
2000    FORMAT (9H EXCEEDED,I5,26H STEPS IN SUBROUTINE REACH) REAC 77
        CALL EXIT                                         REAC 78
25      IF(CROSS) S=AMIN1(S,S1)                         REAC 79
C***** CONVERT POSITION AND DIRECTION TO SPHERICAL POLAR COORDINATES REAC 80
C***** AT A DISTANCE S ALONG THE RAY                  REAC 81
        CALL CAR POL                                      REAC 82
C***** AVOID THE RAY BEING SLIGHTLY UNDERGROUND       REAC 83
        R(1)=AMAX1(R(1),EARTH)
C***** ONE STEP INTEGRATION                          REAC 84
        IF (N.LT.7) GO TO 31                            REAC 85
        DO 30 NN=7,N                                     REAC 86
30      R(NN)=P(NN)+S*DROT(NN)                        REAC 87
31      T=T+S                                         REAC 88
        CALL RINDEX                                     REAC 89
C***** AT A PERIGEE                                REAC 90
        PERIGE=S.EQ.(-UP)                             REAC 91
C***** CORRECT MINOR ERRORS                         REAC 92
        IF (PERIGE) R(4)=0.                            REAC 93
C***** KEEP CONSISTENCY AFTER CORRECTING MINOR ERRORS REAC 94
        DROT(1)=R(4)                                    REAC 95
C***** ON THE GROUND                               REAC 96
        GROUND=S.EQ.SG.AND.CROSSG                     REAC 97
C***** AT THE RECEIVER HEIGHT                      REAC 98
        THERE=S.EQ.SR.AND.CROSSR.AND..NOT.PERIGE     REAC 99
C***** AT A CLOSEST APPROACH TO THE RECEIVER HEIGHT REAC100
        MINDIS=PERIGE.AND.S.EQ.SR.AND.CROSSR          REAC101
        RSPACE=SPACE                                     REAC102
        V=SQRT(N2/(R(4)**2+R(5)**2+R(6)**2))         REAC103
C***** RENORMALIZE THE WAVE NORMAL DIRECTION TO = SQRT(REAL(N**2)) REAC104
        R(4)=R(4)*V                                     REAC105
        R(5)=R(5)*V                                     REAC106
        R(6)=R(6)*V                                     REAC107
        RSTART=1.                                       REAC108
        IF (.NOT.SPACE) CALL PRINTR (8HENTR ION,0.)
        RETURN                                         REAC109
        END                                             REAC110
                                                               REAC111
                                                               REAC112-

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SUBROUTINE POL CAR                                     POLC001
DIMENSION X0(6),X(6),R0(4)                         POLC002
COMMON R(6) /COORD/ S                               POLC003
COMMON /CONST/ PI,PIT2,PID2,DUM(5)                 POLC004
C
C      CONVERTS SPHERICAL COORDINATES TO CARTESIAN   POLC005
IF (R(5).EQ.0...AND.R(6).EQ.0..) GO TO 1          POLC006
VERT=0.
SINA=SIN(R(2))
COSA=SIN(PID2-R(2))
SINP=SIN(R(3))
COSP=SIN(PID2-R(3))
X0(1)=R(1)*SINA*COSP
X0(2)=R(1)*SINA*SINP
X0(3)=R(1)*COSA

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X(4)=R(4)*SINA*COSP+R(5)*COSA*COSP-R(6)*SINP          POLC016
X(5)=R(4)*SINA*SINP+R(5)*COSA*SINP+R(6)*COSP         POLC017
X(6)=R(4)*COSA-R(5)*SINA                                POLC018
RETURN
C      VERTICAL INCIDENCE
1 VERT=1.
R0(1)=R(1)
R0(2)=R(2)
R0(3)=R(3)
R0(4)=SIGN (1.,R(4))
RETURN
C      STEPS THE RAY A DISTANCE S, AND THEN
C      CONVERTS CARTESIAN COORDINATES TO SPHERICAL COORDINATES
ENTRY CAR POL
IF (VERT.NE.0.) GO TO 2
X(1)=X0(1)+S*X(4)                                      POLC030
X(2)=X0(2)+S*X(5)                                      POLC031
X(3)=X0(3)+S*X(6)                                      POLC032
TEMP=SQRT(X(1)**2+X(2)**2)                               POLC033
R(1)=SQRT(X(1)**2+X(2)**2+X(3)**2)                      POLC034
R(2)=ATAN2(TEMP,X(3))                                    POLC035
R(3)=ATAN2(X(2),X(1))                                    POLC036
R(4)=(X(1)*X(4)+X(2)*X(5)+X(3)*X(6))/R(1)           POLC037
R(5)=(X(3)*(X(1)*X(4)+X(2)*X(5))-(X(1)**2+X(2)**2)*X(6))/1
R(6)=(X(1)*X(5)-X(2)*X(4))/TEMP                       POLC040
1 (R(1)*TEMP)
R(6)=(X(1)*X(5)-X(2)*X(4))/TEMP                       POLC041
R(6)=(X(1)*X(5)-X(2)*X(4))/TEMP                       POLC042
RETURN
C      VERTICAL INCIDENCE
2 R(1)=R0(1)+R0(4)*S                                     POLC043
R(2)=R0(2)
R(3)=R0(3)
R(4)=R0(4)
R(5)=0.
R(6)=0.
RETURN
END
PRIN 52-
C      SUBROUTINE PRINTR(NWHY,CARD)
PRINTS OUTPUT AND PUNCHES RAYSETS WHEN REQUESTED          PRIN001
DIMENSION G(3,3),G1(3,3),TYPE(3),HEADR1(20),HEADR2(20),UNITS(20),     PRIN002
1     HEAD1(20),HEAD2(20),UNIT(20),RPRINT(20),NPR(20)          PRIN003
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,DUM(3)                PRIN004
COMMON /FLG/ NTYP,NEWWR,NEWWP,PENET,LINES,IHOP,HPUNCH        PRIN005
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,N2,N2I,PNP(10),POLAR(2),    PRIN006
1     LPOLAR(2)                                              PRIN007
COMMON R(20),T /WW/ ID(10),WD,W(400)                      PRIN008
EQUIVALENCE (THETA,R(2)),(PHI,R(3))                      PRIN009
EQUIVALENCE (EARTH,R(2)),(XMTRH,W(3)),(TLAT,W(4)),(TLON,W(5)),    PRIN010
1 (F,W(6)),(AZ1,W(10)),(BETA,W(14)),(RCVRH,W(20)),(HOP,W(22)),   PRIN011
2 (PLAT,W(24)),(PLON,W(25)),(RAYSET,W(72))              PRIN012
LOGICAL SPACE,NEWWR,NEWWP,PENET                           PRIN013
REAL N2,N2I,LPOLAR                                         PRIN014
COMPLEX PNP                                              PRIN015
DATA (TYPE=1HK,1HN,1HO)                                    PRIN016
2,(HEADR1(7)=5H PHAS),(HEADR2(7)=6HE PATH),(UNITS(7)=6H   KM ), PRIN018
3,(HEADR1(8)=5H ABSO),(HEADR2(8)=6HRPTION),(UNITS(8)=6H   DB ), PRIN019
4,(HEADR1(9)=5H DOP),(HEADR2(9)=6HPLER ),(UNITS(9)=6H C/S ), PRIN020
5,(HEADR1(10)=5H PATH ),(HEADR2(10)=6HLENGTH),(UNITS(10)=6H   KM ) PRIN023
CALL RINDEX                                              PRIN024
IF (.NOT.NEWWP) GO TO 10                                 PRIN025

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***** NEW W ARRAY -- REINITIALIZE PRIN026
NEWWP=.FALSE. PRIN027
SPL=SIN (PLON-TLON) PRIN028
CPL=SIN (PID2-(PLON-TLON)) PRIN029
SP=SIN (PLAT) PRIN030
CP=SIN (PID2-PLAT) PRIN031
SL=SIN (TLAT) PRIN032
CL=SIN (PID2-TLAT) PRIN033
***** MATRIX TO ROTATE COORDINATES PRIN034
G(1,1)=CPL*SP*CL-CP*SL PRIN035
G(1,2)=SPL*SP PRIN036
G(1,3)=-SL*SP*CPL-CL*CP PRIN037
G(2,1)=-SPL*CL PRIN038
G(2,2)=CPL PRIN039
G(2,3)=SL*SPL PRIN040
G(3,1)=CL*CP*CPL+SP*SL PRIN041
G(3,2)=CP*SPL PRIN042
G(3,3)=-SL*CP*CPL+SP*CL PRIN043
DENM=G(1,1)*G(2,2)*G(3,3)+G(1,2)*G(3,1)*G(2,3)+G(2,1)*G(3,2)*G(1,3) PRIN044
1)-G(2,2)*G(3,1)*G(1,3)-G(1,2)*G(2,1)*G(3,3)-G(1,1)*G(3,2)*G(2,3) PRIN045
***** THE MATRIX G1 IS THE INVERSE OF THE MATRIX G PRIN046
G1(1,1)=(G(2,2)*G(3,3)-G(3,2)*G(2,3))/DENM PRIN047
G1(1,2)=(G(3,2)*G(1,3)-G(1,2)*G(3,3))/DENM PRIN048
G1(1,3)=(G(1,2)*G(2,3)-G(2,2)*G(1,3))/DENM PRIN049
G1(2,1)=(G(3,1)*G(2,3)-G(2,1)*G(3,3))/DENM PRIN050
G1(2,2)=(G(1,1)*G(3,3)-G(3,1)*G(1,3))/DENM PRIN051
G1(2,3)=(G(2,1)*G(1,3)-G(1,1)*G(2,3))/DENM PRIN052
G1(3,1)=(G(2,1)*G(3,2)-G(3,1)*G(2,2))/DENM PRIN053
G1(3,2)=(G(3,1)*G(1,2)-G(1,1)*G(3,2))/DENM PRIN054
G1(3,3)=(G(1,1)*G(2,2)-G(2,1)*G(1,2))/DENM PRIN055
R0=EARTH*XMTRH PRIN056
***** CARTESIAN COORDINATES OF TRANSMITTER PRIN057
XR=R0*G(1,1) PRIN058
YR=R0*G(2,1) PRIN059
ZR=R0*G(3,1) PRIN060
CTHR=G(3,1) PRIN061
STHR=SIN (ACOS (CTHR)) PRIN062
PHIR=ATAN2(YR,XR) PRIN063
ALPH=ATAN2(G(3,2),G(3,3)) PRIN064
***** PRIN065
NR=6 PRIN066
NP=0 PRIN067
DO 7 NN=7,20 PRIN068
IF (W(NN+50).EQ.0.) GO TO 7 PRIN069
***** DEPENDENT VARIABLE NUMBER NN IS BEING INTEGRATED PRIN070
***** NR IS THE NUMBER OF DEPENDENT VARIABLES BEING INTEGRATED PRIN071
NR=NR+1 PRIN072
IF (W(NN+50).NE.2.) GO TO 7 PRIN073
***** DEPENDENT VARIABLE NUMBER NN IS BEING INTEGRATED AND PRINTED. PRIN074
***** NP IS THE NUMBER OF DEPENDENT VARIABLES BEING INTEGRATED AND PRIN075
***** PRINTE) PRIN076
NP=NP+1 PRIN077
***** SAVE THE INDEX OF THE DEPENDENT VARIABLE TO PRINT PRIN078
NPR(NP)=NR PRIN079
HEAD1(NP)=HEADR1(NN) PRIN080
HEAD2(NP)=HEADR2(NN) PRIN081
UNIT(NP)=UNITS(NN) PRIN082
7 CONTINUE PRIN083
VP1=MIN0(NP,3) PRIN084
PDEV=ABSORB=DOPP=0. PRIN085

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***** PRINT COLUMN HEADINGS AT THE BEGINNING OF EACH RAY PRIN086
10 IF (IHOP.NE.0) GO TO 12 PRIN087
PRINT 1100, (HEAD1(NN),HEAD2(NN),NN=1,NP1) PRIN088
1100 FORMAT (44X,7HAZIMUTH/43X,9HDEVIATION,8X,9HELEVATION/ PRIN089
1 19X,16HHEIGHT RANGE,1X,2(5X,12HXMTR LOCAL),5X,26HPOLARIZATIPRIN090
20N GROUP PATH,5A6,A5) PRIN091
PRINT 1150, (JNIT(NN),NN=1,NP1) PRIN092
1150 FORMAT (13X,2(8X,2HKM),2X,2(6X,3HDEG,5X,3HDEG),6X,12HREAL IMAG,PRIN093
1 7X,2HKM,4X,3(4X,A6,2X)) PRIN094
IF (RAYSET.EQ.0.) GO TO 12 PRIN095
C***** PUNCH A TRANSMITTER RAYSET PRIN096
TLOND=TLON*DEGS PRIN097
IF (TLOND.LT.0.) TLOND=TLOND+360. PRIN098
TLATO=TLAT*DEGS PRIN099
IF (TLATO.LT.0.) TLATO=TLATO+360. PRIN100
AZ=AZ1*DEGS PRIN101
EL=BETA*DEGS PRIN102
NHOP=HOP PRIN103
PUNCH 1200, ID(1),TYPE(NTYP),XMTRH,TLATO,TLOND,RCVRH,F,AZ,EL,POLARPRIN104
1,NHOP,1HT PRIN105
1200 FORMAT (A3,A1,4PF9.0,3P2F6.0,4P2F9.0,5P2F10.0,5X,2P2F5.0,I1,A1) PRIN106
C*****
12 V=0. PRIN107
IF (N2.NE.0.) V=(R(4)**2+R(5)**2+R(6)**2)/N2-1. PRIN108
H=R(1)-EARTH R PRIN109
STH=SIN (THETA) PRIN110
CTH=SIN (PID2-THETA) PRIN111
C***** CARTESIAN COORDINATES OF RAY POINT, ORIGIN AT TRANSMITTER PRIN112
XP=R(1)*STH*SIN (PID2-PHI)-XR PRIN113
YP=R(1)*STH*SIN (PHI)-YR PRIN114
ZP=R(1)*CTH-ZR PRIN115
C***** CARTESIAN COORDINATES OF RAY POINT, ORIGIN AT TRANSMITTER ANDPRIN116
C***** ROTATED PRIN117
EPS=XP*G1(1,1)+YP*G1(1,2)+ZP*G1(1,3) PRIN118
ETA=XP*G1(2,1)+YP*G1(2,2)+ZP*G1(2,3) PRIN119
ZETA=XP*G1(3,1)+YP*G1(3,2)+ZP*G1(3,3) PRIN120
RCE2=ETA**2+ZETA**2 PRIN121
RCE=SQRT (RCE2) PRIN122
RCE=SQRT (RCE2) PRIN123
C***** GROUND RANGE PRIN124
RANGE=EARTH R*ATAN2(RCE,EARTH R+EPS) PRIN125
C***** ANGLE OF WAVE NORMAL WITH LOCAL HORIZONTAL PRIN126
ELL=ATAN2(R(4),SQRT (R(5)**2+R(6)**2))*DEGS PRIN127
C***** STRAIGHT LINE DISTANCE FROM TRANSMITTER TO RAY POINT PRIN128
SR=SQRT (RCE2+EPS**2) PRIN129
IF (NP.LT.1) GO TO 16 PRIN130
DO 15 I=1,NP PRIN131
NN=NPR(I) PRIN132
15 RPRINT(I)=R(NN) PRIN133
16 IF (SR.GE.1.E-6) GO TO 20 PRIN134
C***** TOO CLOSE TO TRANSMITTER TO CALCULATE DIRECTION FROM PRIN135
C***** TRANSMITTER PRIN136
PRINT 1500, V,NWHY,H,RANGE,ELL,POLAR,T,(RPRINT(NN),NN=1,NP1) PRIN137
1500 FORMAT (1X,E6.0,1X,A8,F10.4,F11.4,26X,F8.3,F9.3,F8.3,4F12.4) PRIN138
GO TO 40 PRIN139
C***** ELEVATION ANGLE OF RAY POINT FROM TRANSMITTER PRIN140
20 EL=ATAN2(EPS,RCE)*DEGS PRIN141
IF (RCE.GE.1.E-6) GO TO 30 PRIN142
C***** NEARLY DIRECTLY ABOVE OR BELOW TRANSMITTER. CAN NOT CALCULATEPRIN143
C***** AZIMUTH DIRECTION FROM TRANSMITTER ACCURATELY PRIN144
PRINT 2500, V,NWHY,H,RANGE,EL,ELL,POLAR,T,(RPRINT(NN),NN=1,NP1) PRIN145
2500 FORMAT (1X,E6.0,1X,A8,F10.4,F11.4,17X,F9.3,F8.3,F9.3,F8.3, PRIN146
1 4F12.4) PRIN147
GO TO 40 PRIN148

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***** AZIMUTH ANGLE OF RAY POINT FROM TRANSMITTER PRIN149
 30 ANGA=ATAN2(ETA,ZETA) PRIN150
   AZDEV=180.-AMOD(540.-(AZ1-ANGA)*DEGS,360.) PRIN151
   IF (R(5).NE.0..OR.R(6).NE.0.) GO TO 34 PRIN152
C***** WAVE NORMAL IS VERTICAL, SO AZIMUTH DIRECTION CANNOT BE PRIN153
C***** CALCULATED PRIN154
   PRINT 3000, V,NWHT,H,RANGE,AZDEV,EL,ELL,POLAR,T,(RPRINT(NN),NN=1, PRIN155
   1 NP1) PRIN156
 3000 FORMAT (1X,E6.0,1X,A8,F10.4,F11.4,F9.3,8X,F9.3,F8.3, PRIN157
   1 4F12.4) PRIN158
   GO TO 40 PRIN159
 34 ANA=ANGA-ALPH PRIN160
   SANA=SIN (ANA) PRIN161
   SPHI=SANA*STHR/STH PRIN162
   CPHI=-SIN (PID2-ANA)*SIN (PID2-(PHI-PHIR))+SANA*SIN (PHI-PHIR) PRIN163
   1 *CTHR PRIN164
   AZA=180.-AMOD (540.-(ATAN2(SPHI,CPHI)-ATAN2(R(6),R(5)))*DEGS,360.) PRIN165
   PRINT 3500, V,NWHT,H,RANGE,AZDEV,AZA,EL,ELL,POLAR,T,(RPRINT(NN),NNPRIN166
   1 =1,NP1) PRIN167
 3500 FORMAT (1X,E6.0,1X,A8,F10.4,F11.4,2(F9.3,F8.3),F9.3,F8.3, PRIN168
   1 4F12.4) PRIN169
C*****
 40 LINES=LINES+1 PRIN170
   IF (NP.LE.3) GO TO 45 PRIN171
C***** ADDITIONAL LINE TO PRINT REMAINING DEPENDENT INTEGRATION PRIN173
C***** VARIABLES PRIN174
   PRINT 4000, (RPRINT(NN),NN=4,NP) PRIN175
 4000 FORMAT (99X,3F12.4) PRIN176
   LINES=LINES+1 PRIN177
   45 IF (CARD.EQ.0.) RETURN PRIN178
C
C***** PUNCH A RAYSET PRIN179
   IF (AZDEV.LT.-90.) AZDEV=AZDEV+360. PRIN180
   IF (AZA.LT.-90.) AZA=AZA+360. PRIN181
   TDEV=T-SR PRIN182
   NR=6 PRIN183
   IF (W(57).EQ.0.) GO TO 47 PRIN184
C***** PHASE PATH PRIN185
   NR=NR+1 PRIN186
   PDEV=R(NR)-SR PRIN187
   47 IF (W(58).EQ.0.) GO TO 48 PRIN188
C***** ABSORPTION PRIN189
   NR=NR+1 PRIN190
   ABSORB=R(NR) PRIN191
C***** DOPPLER SHIFT PRIN192
   48 IF (W(59).NE.0.) DOPP=R(NR+1) PRIN193
   PUNCH 4500, HPUNCH,RANGE,AZDEV,AZA,ELL,SR,TDEV,PDEV,ABSORB,DOPP, PRIN194
   1 POLAR,IHOP,NWHT PRIN195
 4500 FORMAT (4P2F9.0,3P3F6.0,3PF8.0,3P4F6.0,2P2F5.0,I1,A1) PRIN196
   RETURN PRIN197
   END PRIN198
                                         PRIN199-

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INPUT PARAMETER FORM FOR PLOTTING THE PROJECTION  
OF THE RAY PATH ON A VERTICAL PLANE

Coordinates of the left edge of the graph:

Longitude = \_\_\_\_\_

Coordinates of the right edge of the graph:

Latitude = \_\_\_\_\_

Longitude = \_\_\_\_\_

Height above the ground of the bottom of the graph = km (W88)

Distance between tic marks = \_\_\_\_\_ rad  
\_\_\_\_\_ deg (W87)  
\_\_\_\_\_ km

(W81 = 1.)

## INPUT PARAMETER FORM FOR PLOTTING THE PROJECTION OF THE RAY PATH ON THE GROUND

Coordinates of the left edge of the graph:

Latitude = \_\_\_\_\_

Longitude = \_\_\_\_\_

Coordinates of the right edge of the graph:

Latitude = \_\_\_\_\_ rad deg north (W85)  
km

Factor to expand lateral deviation scale by = (W82)

Distance between tic marks on range scale = \_\_\_\_\_ rad  
deg (W87)  
km

(W81 = 2.)

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C SUBROUTINE RAYPLT          YPLT001
    REPLACES SUBROUTINES RAYPLT, PLOT, AND LABPLT IF PLOTS ARE
    NOT WANTED OR IF A PLOTTER IS NOT AVAILABLE      YPLT002
COMMON /WW/ ID(10),W0,W(400)      YPLT003
EQUIVALENCE (PLT,W(81))          YPLT004
PLT=0.                           YPLT005
ENTRY ENDPLT                     YPLT006
RETURN                          YPLT007
END                            YPLT008
                                YPLT  9-

C SUBROUTINE RAYPLT          RAYP001
    W(81)=1. PLOTS PROJECTION OF RAYPATH ON VERTICAL PLANE   RAYP002
    =2. PLOTS PROJECTION OF RAYPATH ON GROUND             RAYP003
COMMON /PLT/ XL,XR,YB,YT,RESET      RAYP004
COMMON /CONST/ PI,PIT2,PID2,DUM(5)      RAYP005
COMMON /FLG/ NTYP,NEWWR,NEWWP,PENET,LINES,IHOP,HPUNCH      RAYP006
COMMON R(6) /WW/ ID(10),W0,W(400)      RAYP007
EQUIVALENCE (TH,R(2)),(PH,R(3))      RAYP008
EQUIVALENCE (EARTH,R(2)),(PLAT,W(24)),(PLON,W(25)),(PLT,W(81)), RAYP009
1 (FACTR,W(82)),(LLAT,W(83)),(LLON,W(84)),(RLAT,W(85)),(RLON,W(86)) RAYP010
2*(TIC,W(87)),(HB*W(88))           RAYP011
REAL LLAT,LLON,LTIC                 RAYP012
LOGICAL NEWWR,NEWWP,PENET          RAYP013
IF (.NOT.NEWWR) GO TO 5            RAYP014
C
C     NEW W ARRAY -- REINITIALIZE          RAYP015
NEWWR=.FALSE.                      RAYP016
RESET=1.                           RAYP017
RAYP018
CONVERT COORDINATES OF VERTICAL PLANE FROM GEOGRAPHIC TO GEOMAGNETIC RAYP019
SW=SIN (PLAT)                      RAYP020
CW=SIN (PID2-PLAT)                 RAYP021
SLM=SIN (LLAT)                     RAYP022
CLM=SIN (PID2-LLAT)                RAYP023
SRM=SIN (RLAT)                     RAYP024
CRM=SIN (PID2-RLAT)                RAYP025
CDPHI=SIN (PID2-(LLON-PLON))       RAYP026
PHL=ATAN2(SIN (LLON-PLON)*CLM,CDPHI*SW*CLM-CW*SLM)      RAYP027
CTHL=CDPHI*CW*CLM+SW*SLM          RAYP028
STHL=SIN (ACOS (CTHL))            RAYP029
CDPHI=SIN (PID2-(RLON-PLON))       RAYP030
PHR=ATAN2(SIN (RLON-PLON)*CRM,CDPHI*SW*CRM-CW*SRM)      RAYP031
CTHR=CDPHI*CW*CRM+SW*SRM          RAYP032
STHR=SIN (ACOS (CTHR))            RAYP033
CLR=CTHL*CTHR+STHL*STHR*SIN (PID2-(PHL-PHR))      RAYP034
SLR=SQRT (1.-CLR**2)              RAYP035
IF (PLT.EQ.2.) GO TO 3            RAYP036
FACTR=1.                           RAYP037
R0=EARTH,R+HB                      RAYP038
ALPHA=.5*ACOS (CLR)                RAYP039
XR=R0*SIN (ALPHA)                 RAYP040
XL=-XR                            RAYP041
YB=R0*SIN (PID2-ALPHA)            RAYP042
YT=YB+2.*XR                         RAYP043
GO TO 5                            RAYP044
3 IF (FACTR.EQ.0.) FACTR=1.        RAYP045
ALPH1=ATAN2(STHR*SIN (PHR-PHL),(CTHR-CTHL*CLR)/STHL)      RAYP046
XL=0.                             RAYP047
XR=EARTH,R*ACOS (CLR)              RAYP048
YT=0.5*XR/FACTR                   RAYP049
YB=-YT                            RAYP050

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C
      5 STH=SIN (TH) RAYP051
      CTH=SIN (PID2-TH) RAYP052
      CR=CTHR*CTH+STHR*STH*SIN (PID2-(PHR-PH)) RAYP053
      CL=CTHL*CTH+STHL*STH*SIN (PID2-(PHL-PH)) RAYP054
      CEA=ATAN2(CR-CL*CLR,CL*SLR) RAYP055
      NEW=1 RAYP056
      IF (IHOP.NE.0) NEW=0 RAYP057
      IF (PLT.EQ.2.) GO TO 10 RAYP058
      CALL PLOT (R(1)*SIN(CEA-ALPHA),R(1)*SIN(PID2-(CEA-ALPHA))+NEW) RAYP059
      RETURN RAYP060
10 SL=SQRT (1.-CL**2) RAYP061
      TMP1=STH*SIN (PH-PHL) RAYP062
      TMP2=(CTH-CTHL*CL)/STHL RAYP063
      ALPH2=0. RAYP064
      IF (TMP1.NE.0..OR.TMP2.NE.0..) ALPH2=ATAN2(TMP1,TMP2) RAYP065
      CALL PLOT (EARTH*CEA+EARTH*ASIN(SL*SIN (ALPH1-ALPH2)),NEW) RAYP066
      RETURN RAYP067
RAYP068
RAYP069
C          DRAW AXES AND CALL FOR LABELING AND TERMINATION OF THIS PLOT RAYP070
ENTRY ENDPLT RAYP071
TICKX=0.01*(YT-YB) RAYP072
IF (PLT.EQ.2.) GO TO 25 RAYP073
R1=EARTH-TICKX RAYP074
X=XL RAYP075
Y=YB RAYP076
CALL PLOT (X,Y,1) RAYP077
NTIC=2 RAYP078
IF (TIC.NE.0..) NTIC=NTIC+2.*ALPHA/TIC RAYP079
NLINE=MAX0 (1,100/NTIC) RAYP080
DO 20 I=1,NTIC RAYP081
ANG=-ALPHA+(I-1)*TIC RAYP082
CALL PLOT (R1*SIN (ANG),R1*SIN (PID2-ANG),0) RAYP083
CALL PLOT (X,Y,0) RAYP084
DO 20 J=1,NLINE RAYP085
ANG=ANG+TIC/NLINE RAYP086
X=EARTH*SIN (ANG) RAYP087
Y=EARTH*SIN (PID2-ANG) RAYP088
20 CALL PLOT (X,Y,0) RAYP089
CALL PLOT (XR,YB,0) RAYP090
GO TO 50 RAYP091
25 DTIC=TIC*EARTH RAYP092
LTIC=DTIC/FACTR RAYP093
TICY=XL+0.01*(XR-XL) RAYP094
NTIC=YT/LTIC RAYP095
TIC1=-LTIC*NTIC RAYP096
CALL PLOT (XL,YB,1) RAYP097
NTIC=2*NTIC+1 RAYP098
DO 30 I=1,NTIC RAYP099
Y=TIC1+(I-1)*LTIC RAYP100
CALL PLOT (XL,Y,0) RAYP101
CALL PLOT (TICY,Y,0) RAYP102
30 CALL PLOT (XL,Y,0) RAYP103
CALL PLOT (XL,YT,0) RAYP104
CALL PLOT (XL,0.,1) RAYP105
NTIC=(XR-XL)/DTIC RAYP106
DO 40 I=1,NTIC RAYP107
X=I*DTIC RAYP108
CALL PLOT (X,0.,0) RAYP109
CALL PLOT (X,TICKX,0) RAYP110
40 CALL PLOT (X,0.,0) RAYP111
CALL PLOT (XR,0.,0) RAYP112
50 CALL LABPLT RAYP113
CALL PLTEND RAYP114
RETURN RAYP115
END RAYP116-

```

```

SUBROUTINE PLOT (X,Y,NEW)
COMMON /PLT/ XMINO,XMAXO,YMINO,YMAXO,RESET          PLOT001
COMMON /DD/ INT,IOR,IT,IS,IC,ICC,IX,IY             PLOT002
DATA (INITAL=1),(MINX=0),(MINY=0),(MAXX=1023),(MAXY=1023),
1 (MINX0=23),(MINY0=23),(MAXX0=1023),(MAXY0=1023)      PLOT003
PLOT004
PLOT005
PLOT006
PLOT007
PLOT008
PLOT009
PLOT010
PLOT011
PLOT012
PLOT013
PLOT014
PLOT015
PLOT016
PLOT017
PLOT018
PLOT019
PLOT020
PLOT021
PLOT022
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PLOT041
PLOT042
PLOT043
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PLOT046
PLOT047
PLOT048
PLOT049
PLOT050
PLOT051
PLOT052
PLOT053
PLOT054
PLOT055
PLOT056
PLOT057
PLOT058
PLOT059
PLOT060
PLOT061
PLOT062
PLOT063
PLOT064
PLOT065

C           INITIALIZE LIBRARY PLOTTING ROUTINES
C           IF (INITAL.EQ.0) GO TO 1
C           INITIAL=0
C           CALL DDINIT (1+1H)

C           COMPUTE SCALE FACTORS
1  IF (RESET.EQ.0.) GO TO 5
RESET=0.
XSCALE=(MAXX0-MINX0)/(XMAX0-XMIN0)
YSCALE=(MAXY0-MINY0)/(YMAX0-YMIN0)
XMIN=XMIN0-(MINX0-MINX)/XSCALE
YMIN=YMIN0-(MINY0-MINY)/YSCALE
XMAX=XMAX0+(MAXX-MAXX0)/XSCALE
YMAX=YMAX0+(MAXY-MAXY0)/YSCALE

C           START A NEW LINE
5  IF (NEW.EQ.0) GO TO 10
IX=MINX0+(X-XMIN0)*XSCALE
IY=MINY0+(Y-YMIN0)*YSCALE
IF (IX.GE.MINX AND IX.LE.MAXX AND IY.GE.MINY AND IY.LE.MAXY)
1 CALL DDBP
GO TO 50

C           HORIZONTAL DISPLACEMENT
10 XS=X-XOLD
YS=Y-YOLD
IF (XS) 11,12,16
C           NEGATIVE
11 X1=XMAX
X2=XMIN
GO TO 20
C           ZERO
12 IF (YS) 13,50,14
13 S1=(YMAX-YOLD)/YS
S2=(YMIN-YOLD)/YS
GO TO 40
14 S1=(YMIN-YOLD)/YS
S2=(YMAX-YOLD)/YS
GO TO 40
C           POSITIVE
16 X1=XMIN
X2=XMAX

C           VERTICAL DISPLACEMENT
20 IF (YS) 21,22,26
C           NEGATIVE
21 Y1=YMAX
Y2=YMIN
GO TO 30
C           ZERO
22 S1=(X1-XOLD)/XS
S2=(X2-XOLD)/XS
GO TO 40
C           POSITIVE
26 Y1=YMIN
Y2=YMAX

C           S1=AMAX1((X1-XOLD)/XS,(Y1-YOLD)/YS)
S2=AMIN1((X2-XOLD)/XS,(Y2-YOLD)/YS)

```

```

C          PLOT LINE -- CHECKING FOR BORDER CROSSINGS          PLOT066
C          S=SQRT(XS**2+YS**2)          PLOT067
40 IF (S2.LT.0..OR.S*S1-S.GT.0.) GO TO 50          PLOT068
IF (S1.LT.0.) GO TO 42          PLOT069
C          PREVIOUS POINT OFF GRAPH          PLOT070
IX=MINX0+(XOLD+XS*S1-XMIN0)*XSCALE+0.5          PLOT071
IY=MINY0+(YOLD+YS*S1-YMIN0)*YSCALE+0.5          PLOT072
CALL DDDB          PLOT073
42 IF (S*S2-S.GT.0.) GO TO 44          PLOT074
C          CURRENT POINT OFF GRAPH          PLOT075
IX=MINX0+(XOLD+XS*S2-XMIN0)*XSCALE+0.5          PLOT076
IY=MINY0+(YOLD+YS*S2-YMIN0)*YSCALE+0.5          PLOT077
CALL DDVC          PLOT078
GO TO 50          PLOT079
C          CURRENT POINT ON GRAPH          PLOT080
44 IX=MINX0+(X-XMIN0)*XSCALE+0.5          PLOT081
IY=MINY0+(Y-YMIN0)*YSCALE+0.5          PLOT082
CALL DDVC          PLOT083
C          EXIT ROUTINE          PLOT084
C          PLOT085
50 XOLD=X          PLOT086
YOLD=Y          PLOT087
RETURN          PLOT088
C          PLOT089
C          TERMINATE THE CURRENT PLOT          PLOT090
ENTRY PLTEND          PLOT091
CALL DDFR          PLOT092
RETURN          PLOT093
END          PLOT 95-

```

```

SUBROUTINE LABPLT          LABP001
C          LABEL THE CURRENT PLOT          LABP002
DIMENSION LABEL(9),TYPE(3)          LABP003
COMMON /DD/ INT,IOR,IT,IS,IC,ICC,IX,IY          LABP004
COMMON /CONST/ PI,PIT2,PID2,DEGS,DUM(4)          LABP005
COMMON /FLG/ NTYP,NEWWR,NEWWP,PENET,LINES,IHOP,HPUNCH          LABP006
COMMON /WW/ ID(10),WD,W(400)          LABP007
EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(AZ1,W(10)),(PLT,W(81)),          LABP008
1 (FACTR,W(82)),(TIC,W(87))          LABP009
LOGICAL NEWWR,NEWWP,PENET          LABP010
REAL LTIC          LABP011
DATA (TYPE=8HEXTRAORD,8HNO FIELD,8HORDINARY)          LABP012
IOR=IT=0          LABP013
IS=2          LABP014
IX=0      $ IY=1023 $ CALL DOTAB $ CALL DOTEXT (7,ID)          LABP015
NDATE=IDATE(0)          LABP016
CALL DOTEXT (1,NDATE)          LABP017
AZA=AZ1*DEGS          LABP018
DTIC=TIC*EARTH          LABP019
ENCODE (72,1000,LABEL) F,AZA,TYPE(NTYP),DTIC          LABP020
1000 FORMAT (3HF =,F7.3,6H, AZ =,F7.2,2H, ,A8,2H, ,F7.2,24H KM BETWEEN          LABP021
1TICK MARKS#.)
IX=0      $ IY=991 $ CALL DOTAB $ CALL DOTEXT (9,LABEL)          LABP022
IF (PLT.EQ.1.) RETURN          LABP023
LTIC=DTIC/FACTR          LABP024
ENCODE (32,2000,LABEL) LTIC          LABP025
2000 FORMAT (F7.2,24H KM BETWEEN TICK MARKS#.)
IOR=1          LABP026
IX=0      $ IY=0      $ CALL DOTAB $ CALL DOTEXT (4,LABEL)          LABP027
IOR=0          LABP028
RETURN          LABP029
END          LABP030
LABP 32-

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C      SUBROUTINE RKAM          RKAM001
      NUMERICAL INTEGRATION OF DIFFERENTIAL EQUATIONS   RKAM002
COMMON /RK/ NN,SPACE,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART   RKAM003
COMMON Y(20),T,STEP,DYDT(20)   RKAM004
DIMENSION DELY(4,20),BET(4),XV(5),FV(4,20),YU(5,20)   RKAM005
TYPE DOUBLE YU               RKAM006
IF (RSTART.EQ.0.) GO TO 1000   RKAM007
LL=MM=1                      RKAM008
IF (MODE.EQ.1) MM=4           RKAM009
ALPHA=T                        RKAM010
EPM=0.0                         RKAM011
BET(1)=BET(2)=0.5              RKAM012
BET(3)=1.0                      RKAM013
BET(4)=0.0                      RKAM014
STEP=SPACE                      RKAM015
R=19.0/270.0                     RKAM016
XV(MM)=T                        RKAM017
IF (E1MIN.LE.0.) E1MIN=E1MAX/55.   RKAM018
IF (FACT.LE.0.) FACT=0.5          RKAM019
CALL HAMLTN                     RKAM020
DO 320 I=1,NN                   RKAM021
FV(MM,I)=DYDT(I)                RKAM022
320 YU(MM,I)=Y(I)                RKAM023
RSTART=0.                         RKAM024
GO TO 1001                       RKAM025
1000 IF (MODE.NE.1) GO TO 2000   RKAM026
C
C      RUNGE-KUTTA             RKAM027
1031 DO 1034 K=1,4               RKAM028
DO 1350 I=1,NN                   RKAM029
DELY(K,I)=STEP*FV(MM,I)          RKAM030
Z=YU(MM,I)                      RKAM031
1350 Y(I)=Z+BET(K)*DELY(K,I)    RKAM032
T=BET(K)*STEP+XV(MM)            RKAM033
CALL HAMLTN                     RKAM034
DO 1034 I=1,NN                   RKAM035
FV(MM,I)=DYDT(I)                RKAM036
1034 DO 1039 I=1,NN               RKAM037
DEL=(DELY(1,I)+2.0*DELY(2,I)+2.0*DELY(3,I)+DELY(4,I))/6.0   RKAM038
1039 YU(MM+1,I)=YU(MM,I)+DEL   RKAM039
MM=MM+1                         RKAM040
XV(MM)=XV(MM-1)+STEP            RKAM041
DO 1400 I=1,NN                   RKAM042
1400 Y(I)=YU(MM,I)              RKAM043
T=XV(MM)                         RKAM044
CALL HAMLTN                     RKAM045
IF (MODE.EQ.1) GO TO 42          RKAM046
DO 150 I=1,NN                   RKAM047
150 FV(MM,I)=DYDT(I)            RKAM048
IF (MM.LE.3) GO TO 1001          RKAM049
C
C      ADAMS-MOULTON           RKAM050
2000 DO 2048 I=1,NN               RKAM051
DEL=STEP*(55.*FV(4,I)-59.*FV(3,I)+37.*FV(2,I)-9.*FV(1,I))/24.   RKAM052
Y(I)=YU(4,I)+DEL                RKAM053
2048 DELY(1,I)=Y(I)              RKAM054
T=XV(4)+STEP                     RKAM055
CALL HAMLTN                     RKAM056
XV(5)=T                          RKAM057
DO 2051 I=1,NN                   RKAM058
DEL=STEP*(9.*DYDT(I)+19.*FV(4,I)-5.*FV(3,I)+FV(2,I))/24.        RKAM059
YU(5,I)=YU(4,I)+DEL              RKAM060
2051 Y(I)=YU(5,I)                RKAM061
CALL HAMLTN                     RKAM062
IF (MODE.LE.2) GO TO 42          RKAM063

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C          ERROR ANALYSIS                         RKAM066
C          SSE=0.0                                RKAM067
DO 3033 I=1,NN                           RKAM068
EPSIL=R*ABS(Y(I)-DELY(1,I))           RKAM069
IF (MODE.EQ.3.AND.Y(I).NE.0.) EPSIL=EPSIL/ABS(Y(I))
IF (SSE.LT.EPSIL) SSE=EPSIL            RKAM071
3033 CONTINUE                            RKAM072
IF (E1MAX.GT.SSE) GO TO 3035          RKAM073
IF (ABS(STEP).LE.E2MIN) GO TO 42      RKAM074
LL=MM=1                                RKAM075
STEP=STEP*FACT                          RKAM076
GO TO 1001                             RKAM077
3035 IF (LL.LE.1.OR.SSE.GE.E1MIN.OR.E2MAX.LE.ABS(STEP)) GO TO 42
LL=2                                    RKAM079
MM=3                                    RKAM080
XV(2)=XV(3)                            RKAM081
XV(3)=XV(5)                            RKAM082
DO 5363 I=1,NN                           RKAM083
FV(2,I)=FV(3,I)                        RKAM084
FV(3,I)=DYDT(I)                        RKAM085
YU(2,I)=YJ(3,I)                        RKAM086
5363 YU(3,I)=YJ(5,I)                  RKAM087
STEP=2.0*STEP                           RKAM088
GO TO 1001                            RKAM089
RKAM090
C          EXIT ROUTINE                      RKAM091
C          EXIT ROUTINE                      RKAM092
42 LL=2                                RKAM093
MM=4                                    RKAM094
DO 12 K=1,3                            RKAM095
XV(K)=XV(K+1)                          RKAM096
DO 12 I=1,NN                           RKAM097
FV(K,I)=FV(K+1,I)                      RKAM098
12 YU(K,I)=YU(K+1,I)                  RKAM099
XV(4)=XV(5)                            RKAM100
DO 52 I=1,NN                           RKAM101
FV(4,I)=DYDT(I)                        RKAM102
52 YU(4,I)=YU(5,I)                    RKAM103
IF (MODE.LE.2) RETURN                  RKAM104
E=ABS(XV(4)-ALPHA)                     RKAM105
IF (E.LE.EPM) GO TO 2000              RKAM106
EPM=E                                RKAM107
RETURN                               RKAM108
END                                  RKAM109-

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SUBROUTINE HAMLTN                                         HAML001
C***** CALCULATES HAMILTONS EQUATIONS FOR RAY TRACING   HAML002
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,C,LOGTEN          HAML003
COMMON /RIN/  MOORIN(3),COLL,FIELD,SPACE,KAY2,KAY2I,        HAML004
1      H,HI,PHPT,PHPTI,PHPH,PHPRI,PHPHT,PHPHTI,PHPPH,PHPPHI HAML005
2 , PHPOM,PHPOMI,PHPKR,PHPKRI,PHPKTH,PHPKTI, PHPKPH,PHPKPI HAML006
3 , KPHPK,KPHPKI,POLAR,POLARI,LPOLAR,LPOLRI             HAML007
COMMON R(20),T,STP,DRDT(20) /WW/ ID(10),W0,W(400)         HAML008
EQUIVALENCE (TH,R(2)),(PH,R(3)),(KR,R(4)),(KTH,R(5)),(KPH,R(6)), HAML009
1 (DTHDT,DRDT(2)),(DPHDT,DRDT(3)),(DKRDT,DRDT(4)),(DKTHDT,DRDT(5)), HAML010
2 (DKPHOT,DRDT(6)),(F,W(6))                            HAML011
REAL KR,KTH,KPH,KPHPK,KPHPKI,LPOLAR,LPOLRI,LOGTEN,K,KAY2,KAY2I HAML012
OM=PIT2*1.E6+F                                         HAML013
STH=SIN(TH)                                           HAML014
CTH=SIN(PID2-TH)                                     HAML015
RSTH=R(1)*STH                                       HAML016
RCTH=R(1)*CTH                                     HAML017
CALL RINDEX                                         HAML018
DRDT=-PHPKR/(PHPOM*C)                                HAML019
DTHDT=-PHPKTH/(PHPOM*R(1)*C)                         HAML020
DPHDT=-PHPKPH/(PHPOM*RSTH*C)                         HAML021
DKRDT=PHPH/(PHPOM*C)+KTH*DTHDT+KPH*STH*DPHDT       HAML022
DKTHDT=(PHPTH/(PHPOM*C)-KTH*DRDT+KPH*RCTH*DPHDT)/R(1) HAML023
DKPHOT=(PHPPH/(PHPOM*C)-KPH*STH*DRDT-KPH*RCTH*DTHDT)/RSTH HAML024
NR=6                                                 HAML025
C***** PHASE PATH                                     HAML026
IF (W(57).EQ.0.) GO TO 10                           HAML027
NR=NR+1                                              HAML028
DRDT(NR)=-KPHPK/PHPOM/OM                           HAML029
C***** ABSORPTION                                     HAML030
10 IF (W(58).EQ.0.) GO TO 15                         HAML031
NR=NR+1                                              HAML032
DRDT(NR)= 10./LOGTEN*KPHPK*KAY2I/(KR*KR+KTH*KTH+KPH*KPH)/PHPOM/C HAML033
C***** DOPPLER SHIFT                                 HAML034
15 IF (W(59).EQ.0.) GO TO 20                         HAML035
NR=NR+1                                              HAML036
DRDT(NR)=-PHPT/PHPOM/C/PIT2                         HAML037
C***** GEOMETRICAL PATH LENGTH                     HAML038
20 IF (W(60).EQ.0.) GO TO 25                         HAML039
NR=NR+1                                              HAML040
DRDT(NR)=-SQRT(PHPKR**2+PHPKTH**2+PHPKPH**2)/PHPOM /C HAML041
C***** OTHER CALCULATIONS                         HAML042
25 CONTINUE                                         HAML043
RETURN                                              HAML044
END                                                 HAML045-

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## APPENDIX 2. VERSIONS OF THE REFRACTIVE INDEX SUBROUTINE (RINDEX)

This ray tracing program gains versatility without sacrificing speed by having several versions of some of the subroutines. For example, the 8 versions of the refractive index subroutine allow the user to decide for each ray path calculation whether to include or ignore various aspects of the propagation medium such as the earth's magnetic field or collisions between electrons and neutral air molecules.

If collisions are included, the user has the option of using the Appleton-Hartree formula (which assumes a constant collision frequency) or the Sen-Wyller formula (which assumes a Maxwell distribution of electron energies and a collision frequency proportional to energy). The Sen-Wyller formula is generally assumed to be more accurate, especially in the lower ionosphere, but the Appleton-Hartree formula can often be used with an effective collision frequency profile to save computer time.

When the effect of the earth's magnetic field is included and ray paths are calculated near vertical incidence, a spitz (Davies, 1965, p. 202) often occurs in the ray path. (At a spitz, the usual formulas for refractive index become indeterminate because the wave normal is parallel with the earth's magnetic field and the wave frequency equals the local plasma frequency.) Two versions of the refractive index subroutine have been developed to calculate ray paths through a spitz. These two versions will also work in the absence of a spitz, but the standard versions are much faster.

The input to the refractive index subroutines is through blank common and common blocks /XX/, /YY/, and /ZZ/. Output is through common block /RIN/. The refractive index subroutine is called through the entry RINDEX. The subroutine names are used only for user identification. The following 8 versions of the refractive index subroutine are

listed in this appendix:

a.	Subroutine AHWFWC (Appleton-Hartree formula with field, with collisions)	93
b.	Subroutine AHWFNC (Appleton-Hartree formula with field, no collisions)	94
c.	Subroutine AHNFWC (Appleton-Hartree formula no field, with collisions)	96
d.	Subroutine AHNFNC (Appleton-Hartree formula no field, no collisions)	97
e.	Subroutine BQWFWC (Booker Quartic with field, with collisions)	98
f.	Subroutine BQWFNC (Booker Quartic with field, no collisions)	100
g.	Subroutine SWWF (Sen-Wyller formula with field)	102
h.	Subroutine SWNF (Sen-Wyller no field) Subroutine FGSW Subroutine FSW Fresnel integral function C Fresnel integral function S	105 106 106 108 108

```

C SUBROUTINE A4WFC          WFWC001
      CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE
      APPLETON-HARTREE FORMULA WITH FIELD, WITH COLLISIONS        WFWC002
C COMMON /CONST/ PI,PIT2,PI02,DEGS,RADIAN,K,C,LOGTEN           WFWC003
C COMMON /RIN/ MODRIN(3),CO_L,FIELD,SPACE,KAY2,H,PHPT,PHPH,PHPTH,   WFWC004
1     PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,KPHPK,POLAR,LPOLAR       WFWC006
C COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX            WFWC007
C COMMON /YY/ MOODY,Y,PYPR,PYPTH,PYPH,PYRPR,PYRPT,PYRPP,YTH,PYTPT,WFWC008
1     ,PYTPT,PYTPP,YPH,PYPPR,PYPPT,PYPPP                         WFWC009
C COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH                         WFWC010
C COMMON R,TH,PH,KR,KTH,KPH /WW/ ID(10),W0,W(400)             WFWC011
C COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART    WFWC012
EQUIVALENCE (RAY,W(1)),(F,W(6))                                WFWC013
LOGICAL SPACE                                                 WFWC014
REAL KR,KTH,KPH,K2                                         WFWC015
COMPLEX N2,PNPR,PNPTH,PNPPH,PNPVR,PNPVTH,PNPVPH,NNP,PNPT,   WFWC016
1     POLAR,LPOLAR,I,U,RAD,D,PNPPS,PNPX,PNPY,PNPZ,UX,UX2,D2, WFWC017
2     KAY2,1,PHPT,PHPH,PHPTH,PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH, WFWC018
3     KPHPK                                         WFWC019
DATA (MODRIN=8HAPPLETON,8H-HARTREE,8H FORMULA),(COLL=1.),      WFWC020
1     (FIELD=1.),                                         WFWC021
2     (X=0.), (PXPR=0.), (PXPTH=0.), (PXPPH=0.), (PXPT=0.), WFWC022
3     (Y=0.), (PYPR=0.), (PYPTH=0.), (PYPPH=0.), (YR=0.), (PYRPR=0.), WFWC023
4     (PYRPT=0.), (PYRPP=0.), (YTH=0.), (PYTPT=0.), (PYTPT=0.), WFWC024
5     (PYTPP=0.), (YPH=0.), (PYPPR=0.), (PYPPT=0.), (PYPPP=0.) WFWC025
5     ,(Z=0.), (PZPR=0.), (PZPTH=0.), (PZPPH=0.), WFWC026
7     (I=(0.,1.)), (ABSLIM=1.E-5)                         WFWC027
ENTRY RINDEX                                                 WFWC028
OM=PIT2*1.E6*F                                         WFWC029
C2=C*C                                                 WFWC030
K2=KR*KR+KTH*KTH+KPH*KPH                               WFWC031
OM2=OM*OM                                                 WFWC032
VR =C/OM*KR                                              WFWC033
VTH=C/OM*KTH                                             WFWC034
VPH=C/OM*KPH                                             WFWC035
CALL ELECTX                                              WFWC036
CALL MAGY                                                 WFWC037
V2=VR**2+VTH**2+VPH**2                                  WFWC038
VDOTY=VR*YR+VTH*YTH+VPH*YPH                            WFWC039
YLV=VDOTY/V2                                              WFWC040
YL2=VDOTY**2/V2                                         WFWC041
YT2=Y**2-YL2                                              WFWC042
YT4=YT2*YT2                                              WFWC043
CALL COLFRZ                                              WFWC044
U=CMPLX(1.,-Z)                                           WFWC045
UX=U-X                                                 WFWC046
UX2=UX*UX                                              WFWC047
RAD=RAY*CSQRT(YT4+4.*YL2*UX2)                           WFWC048
D=2.*U*UX-YT2+RAD                                         WFWC049
D2=D*D                                                 WFWC050
N2=1.-2.*X*UX/D                                         WFWC051
PNPPS=2.*X*UX*(-1.+YT2-2.*UX2)/RAD)/D2                WFWC052
PPSPR =YL2/Y*PYPR -(VR*PYRPR+VTH*PYTPT+VPH*PYPPR)*YLV  WFWC053
PPSPTH=YL2/Y*PYPTH -(VR*PYRPT+VTH*PYTPT+VPH*PYPPT)*YLV WFWC054
PPSPPH=YL2/Y*PYPPH -(VR*PYRPP+VTH*PYTPP+VPH*PYPPP)*YLV WFWC055
PNPX=-(2.*U*UX2-YT2*(U-2.*X)+(YT4*(U-2.*X)+4.*YL2*UX*UX2)/RAD)/D2 WFWC056
PNPY=2.*X*UX*(-YT2+(YT4+2.*YL2*UX2)/RAD)/(D2*Y)        WFWC057
PNPZ=I*X*(-2.*UX2-YT2+YT4/RAD)/D2                      WFWC058
PNPR =PNPX*PXPR +PNPY*PYPR +PNPZ*PZPR +PNPPS*PPSPR    WFWC059
PNPTH=PNPX*PXPTH+PNPY*PYPTH+PNPZ*PZPTH+PNPPS*PPSPTH  WFWC060
PNPPH=PNPX*PXPPH+PNPY*PYPPH+PNPZ*PZPPH+PNPPS*PPSPPH  WFWC061
PNPVR =PNPPS*(VR *YL2/V2-YLV*YR )                      WFWC062
PNPVTH=PNPPS*(VTH*YL2/V2-YLV*YTH)                      WFWC063
PNPVPH=PNPPS*(VPH*YL2/V2-YLV*YPH)                      WFWC064
NNP=N2-(2.*X*PNPX+Y*PNPY+Z*PNPZ)                      WFWC065

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PNPT=PNPX*PXPT WFWC066
SPACE=REAL(N2).EQ.1..AND.ABS(AIMAG(N2)).LT.ABSLIM WFWC067
POLAR=-I*SQRT(V2)*(-YT2+RAD)/(2.*VDOOTY*UX) WFWC068
GAM=(-YT2+RAD)/(2.*UX) WFWC069
LPOLAR=I*X*SQRT(YT2)/(UX*(U+GAM)) WFWC070
KAY2=OM2/C2*N2 WFWC071
IF(RSTART.EQ.0.) GO TO 1 WFWC072
SCALE=SQRT(REAL(KAY2)/K2) WFWC073
KR=SCALE*KR WFWC074
KTH=SCALE*KTH WFWC075
KPH=SCALE*KPH WFWC076
1 CONTINUE WFWC077
C***** CALCULATES A HAMILTONIAN H WFWC078
H=.5*(C2*K2/OM2-N2) WFWC079
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO WFWC080
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI. WFWC081
    PHPT=-PNPT WFWC082
    PHPR=-PNPR WFWC083
    PHPTH=-PNPTH WFWC084
    PHPH=-PNPPH WFWC085
    PHPOM=-NNP/OM WFWC086
    PHPKR=C2/OM2*KR -C/OM*PNPVR WFWC087
    PHPKTH=C2/OM2*KTH-C/OM*PNPVTH WFWC088
    PHPKPH=C2/OM2*KPH-C/OM*PNPVPH WFWC089
    KPHPK=N2 WFWC090
    RETURN WFWC091
    END WFWC092

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SUBROUTINE AHWFNC WFNC001
C   CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE WFNC002
C   APPLETON-HARTREE FORMULA WITH FIELD, NO COLLISIONS WFNC003
COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,C,LOGTN WFNC004
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,KAY2I, WFNC005
1           H,HI,PHPT,PHPTI,PHPR,PHPRI,PHPTH,PHPTHI,PHPPH,PHPHIWFNC006
2           ,PHPOM,PHPOMI,PHPKR,PHPKRI,PHPKTH,PHPKTI,PHPKPH,PHPKPIWFNC007
3           ,KPHPK,KPHPKI,POLAR,POLARI,LPOLAR,LPOLRI,SGN WFNC008
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX WFNC009
COMMON /YY/ MODY,Y,PYPR,PYPHTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPRWFNC010
1           ,PYTPT,PYTTPP,YPH,PYPPR,PYPPT,PYPPP WFNC011
COMMON /ZZ/ MODZ,Z(4) WFNC012
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART WFNC013
COMMON R,TH,PH,KR,KTH,KPH /WW/ ID(10),W0,W(400) WFNC014
EQUIVALENCE (RAY,W(1)),(F,W(6)) WFNC015
LOGICAL SPACE WFNC016
REAL KR,KTH,KPH,K2,KPHPK,KPHPKI,KAY2,KAY2I,N2,NNP,LPOLAR,LPOLRI WFNC017
DATA (MODRIN=8HAPPLETON,8H-HARTREE,8H FORMULA),(COLL=0.), WFNC018
1           (FIELD=1.),(KAY2I=0.),(HI=0.),(PHPTI=0.),(PHPRI=0.),
2           (PHPTHI=0.),(PHPHI=0.),(PHPOMI=0.),(PHPKRI=0.),(PHPKTI=0.),
3           (PHPKPI=0.),(KPHPKI=0.),(POLAR=0.),(LPOLAR=0.),
4           (X=0.),(PXPR=0.),(PXPTH=0.),(PXPPH=0.),(PXPT=0.),
5           (Y=0.),(PYPR=0.),(PYPHTH=0.),(PYPPH=0.),(YR=0.),(PYRPR=0.),
6           (PYRPT=0.),(PYRPP=0.),(YTH=0.),(PYTPR=0.),(PYTPT=0.),
7           (PYTTPP=0.),(YPH=0.),(PYPPR=0.),(PYPPT=0.),(PYPPP=0.),
8           (MODZ=1H),(U=1.)) WFNC019
ENTRY RINDEX WFNC020
OM=PIT2*1.E6*f WFNC021
C2=C*C WFNC022
K2=KR*KR+KTH*KTH+KPH*KPH WFNC023
OM2=OM*OM WFNC024
VR =C/OM*KR WFNC025
VTH=C/OM*KTH WFNC026
VPH=C/OM*KPH WFNC027

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CALL ELECTX                               WFNC035
CALL MAGY                                 WFNC036
V2=VR**2+VTH**2+VPH**2                  WFNC037
VDOTY=VR*YR+VTH*YTH+VPH*YPH            WFNC038
YLV=VDOTY/V2                            WFNC039
YL2=VDOTY**2/V2                         WFNC040
YT2=Y**2-YL2                            WFNC041
YT4=YT2*YT2                            WFNC042
UX=U-X                                 WFNC043
UX2=UX*UX                             WFNC044
RAD=RAY*SQRT(YT4+4.*YL2*UX2)           WFNC045
D=2.*UX-YT2+RAD                         WFNC046
D2=D*D                                WFNC047
N2=1.-2.*X*UX/D                         WFNC048
PNPPS=2.*X*UX*(-1.+((YT2-2.*UX)/RAD)/D2) WFNC049
PPSPR=YL2/Y*PYPR -(VR*PYRPR+VTH*PYTPR+VPH*PYPPR)*YLV WFNC050
PPSPTH=YL2/Y*PYPTH -(VR*PYRPT+VTH*PYTPT+VPH*PYPPT)*YLV WFNC051
PPSPPH=YL2/Y*PYPPH -(VR*PYRPP+VTH*PYTPP+VPH*PYPPP)*YLV WFNC052
PNPX=-(2.*UX2-YT2*(U-2.*X)+(YT4*(U-2.*X)+4.*YL2*UX*UX2)/RAD)/D2 WFNC053
PNPY=2.*X*UX*(-YT2+(YT4+2.*YL2*UX2)/RAD)/(D2*Y) WFNC054
NP=N2-(2.*X*NPX+Y*PNPY)                WFNC055
PNPR =PNPX*PXPR +PNPY*PYPR +PNPPS*PPSPR WFNC056
PNPTH=PNPX*PXPTH+PNPY*PYPTH+PNPPS*PPSPTH WFNC057
PNPPH=PNPX*PXPPH+PNPY*PYPPH+PNPPS*PPSPPH WFNC058
PNPVR =PNPPS*(VR *YL2-VDOTY*YR )/V2 WFNC059
PNPVTH=PNPPS*(VTH*YL2-VDOTY*YTH)/V2 WFNC060
PNPVPH=PNPPS*(VPH*YL2-VDOTY*YPH)/V2 WFNC061
PNPT=PNPX*PXPT                          WFNC062
SPACE=N2.EQ.1.                           WFNC063
POLARI=SQRT(V2)*(YT2-RAD)/(2.*VDOTY*UX) WFNC064
GAM=(-YT2+RAD)/(2.*UX)                  WFNC065
LPOLRI=X*SQRT(YT2)/(UX*(U+GAM))        WFNC066
KAY2=OM2/C2*N2                           WFNC067
IF(RSTART.EQ.0.) GO TO 1                 WFNC068
SCALE=SQRT(KAY2/K2)                      WFNC069
KR =SCALE*KR                            WFNC070
KTH=SCALE*KTH                           WFNC071
KPH=SCALE*KPH                           WFNC072
1  CONTINUE                               WFNC073
C***** CALCULATES A HAMILTONIAN H          WFNC074
H=.5*(C2*K2/OM2-N2)                      WFNC075
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO WFNC076
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI. WFNC077
PHPT =-PNPT                               WFNC078
PHPR =-PNPR                               WFNC079
PHPTH=-PNPTH                            WFNC080
PHPH=-PNPPH                            WFNC081
PHPOM=-NNP/OM                           WFNC082
PHPKR =C2/OM2*KR -C/OM*PNPVR             WFNC083
PHPKTH=C2/OM2*KTH-C/OM*PNPVTH            WFNC084
PHPKPH=C2/OM2*KPH-C/OM*PNPVPH            WFNC085
KPHPK=N2                                WFNC086
RETURN                                  WFNC087
END                                     WFNC088-

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C      SUBROUTINE A4NFWC          NFWC010
      CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE
      APPLETON-HARTREE FORMULA -- NO FIELD, WITH COLLISIONS      NFWC011
      COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,C,LOGTEN        NFWC012
      COMMON /RIN/ MODIN(3),COLL,FIELD,SPACE,KAY2,H,PHPT,PHPH,PHPTH,
1           PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,KPHPK,POLAR,LPOLAR, NFWC013
2           SGN                                              NFWC014
      COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX          NFWC015
      COMMON /YY/ MODY,Y(16)                                     NFWC016
      COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH                      NFWC017
      COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART NFWC018
      COMMON R,TH,PH,KR,KTH,KPH      /WW/ ID(10),W0,W(400)       NFWC019
      EQUIVALENCE (RAY,W(1)),(F,W(6))                         NFWC020
      LOGICAL SPACE                                         NFWC021
      REAL KR,KTH,KPH,K2                                     NFWC022
      COMPLEX KAY2,H,PHPT,PHPH,PHPOM,PHPKR,PHPKTH,PHPKPH,      NFWC023
1           KPHPK,POLAR,LPOLAR,U,I,PNPX,PNPZ,                 NFWC024
2           N2,PNPR,PNPTH,PNPPH,PNPVR,PNPVTH,PNPVPH,NNP,PNPT   NFWC025
      DATA (MODIN=8HAPPLETON,8H-HARTREE,8H FORMULA),(COLL=1.),
1           (FIELD=0.),(POLAR=(0.,1.)),(LPOLAR=(0.,0.)),        NFWC026
2           (X=0.),(PXPR=0.),(PXPTH=0.),(PXPPH=0.),(PXPT=0.),
3           (MODY=1H),                                         NFWC027
4           (Z=0.),(PZPR=0.),(PZPTH=0.),(PZPPH=0.),           NFWC028
5           (I=(0.,1.)),(ABSLIM=1.E-5),(PNPVR=0.),(PNPVTH=0.),(PNPVPH=0.) NFWC029
      ENTRY RINDEX                                         NFWC030
      OM=PIT2*1.E6*F                                      NFWC031
      C2=C*C
      K2=KR*KR+KTH*KTH+KPH*KPH                         NFWC032
      OM2=OM*OM                                         NFWC033
      VR =C/OM*KR                                         NFWC034
      VTH=C/OM*KTH                                        NFWC035
      VPH=C/OM*KPH                                        NFWC036
      CALL ELECTX                                         NFWC037
      CALL COLFRZ                                         NFWC038
      U=1.-I*Z                                           NFWC039
      N2=1.-X/U                                         NFWC040
      PNPX=-1./(2.*U)                                     NFWC041
      PNPZ=-I*X/(2.*U**2)                                NFWC042
      NNP=N2-(2.*X*PNPX+Z*PNPZ)                          NFWC043
      PNPR =PNPX*PXPR +PNPZ*PZPR                         NFWC044
      PNPTH=PNPX*PKPTH+PNPZ*PZPTH                        NFWC045
      PNPPH=PNPX*PXPPH+PNPZ*PZPH                         NFWC046
      PNPT=PNPX*PKPT                                       NFWC047
      SPACE=REAL(N2).EQ.1..AND.ABS(AIMAG(N2)).LT.ABSLIM    NFWC048
      KAY2=OM2/C2*N2                                       NFWC049
      IF(RSTART.EQ.0.) GO TO 1                           NFWC050
      SCALE=SQRT(REAL(KAY2)/K2)                          NFWC051
      KR =SCALE*KR                                         NFWC052
      KTH=SCALE*KTH                                        NFWC053
      KPH=SCALE*KPH                                        NFWC054
      1  CONTINUE                                         NFWC055
C***** CALCULATES A HAMILTONIAN H
      H=.5*(C2*K2/OM2-N2)                                NFWC056
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI. NFWC057
      PHPT ==PNPT                                         NFWC058
      PHPR ==PNPR                                         NFWC059
      PHPTH==PNPTH                                       NFWC060
      PHPPH==PNPPH                                       NFWC061
      PHPOM==NNP/OM                                       NFWC062
      PHPKR =C2/OM2*KR                                    NFWC063
      PHPKTH=C2/OM2*KTH                                  NFWC064
      PHPKPH=C2/OM2*KPH                                  NFWC065
      KPHPK=N2                                           NFWC066
      RETURN
      ENO

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SUBROUTINE AHNFNC
C      CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE      NFNC001
C      APPLETON-HARTREE FORMULA -- NO FIELD, NO COLLISIONS      NFNC002
C      COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,C,LOGTEN      NFNC003
C      COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,KAY2I,      NFNC004
C      1          H,HI,PHPT,PHPTI,PHPH,PHPRI,PHPTH,PHPTHI,PHPPH,PHPHI,NFNC005
C      2          PHPOM,PHPOMI,PHPKR,PHPKRI,PHPKTH,PHPKTI,PHPKPH,PHPKPINFNC006
C      3          , $\langle$ PHPK,KPHPK,POLAR,POLARI,LPOLAR,LPOLRI,SGN      NFNC007
C      COMMON /XXX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX      NFNC008
C      COMMON /YY/ MODY,Y(16) /ZZ/ MODZ,Z(4)      NFNC009
C      COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART      NFNC010
C      COMMON R,TH,PH,KR,KTH,KPH /WW/ ID(10),W0,W(400)      NFNC011
C      EQUIVALENCE (RAY,W(1)),(F,W(6))      NFNC012
C      LOGICAL SPACE      NFNC013
C      REAL N2,NNP,KR,KTH,KPH,K2,KPHPK,KPHPKI,KAY2,KAY2I,LPOLAR,LPOLRI      NFNC014
C      DATA (MODRIN=8HAPPLETON,8H-HARTREE,8H FORMULA),(COLL=0.),      NFNC015
C      1          (FIELD=0.), (KAY2I=0.), (HI=0.), (PHPTI=0.), (PHPRI=0.),      NFNC016
C      2          (PHPTHI=0.), (PHPHI=0.), (PHPOMI=0.), (PHPKRI=0.), (PHPKTI=0.),      NFNC017
C      3          (PHPKPI=0.), (KPHPK=0.), (POLAR=0.), (POLARI=1.), (LPOLAR=0.),      NFNC018
C      4          (LPOLRI=1.),      NFNC019
C      4          (X=0.), (PXPR=0.), (PXPTH=0.), (PXPPH=0.), (PXPT=0.),      NFNC020
C      5          (MODY=1H ), (MODZ=1H ),      NFNC021
C      5          (NNP=1.), (PNPX=-0.5), (PNPVR=0.), (PNPVTH=0.), (PNPVPH=0.)      NFNC022
C      ENTRY RINDEX      NFNC023
C      OM=PIT2*1.E6*F      NFNC024
C      C2=C*C      NFNC025
C      K2=KR*KR+KTH*KTH+KPH*KPH      NFNC026
C      OM2=OM*OM      NFNC027
C      VR =C/OM*KR      NFNC028
C      VTH=C/OM*KTH      NFNC029
C      VPH=C/OM*KPH      NFNC030
C      CALL ELECTX      NFNC031
C      PNPR =PNPX*PXPR      NFNC032
C      PNPTH=PNPX*PXPTH      NFNC033
C      PNPPH=PNPX*PXPPH      NFNC034
C      PNPT=PNPX*PXPT      NFNC035
C      N2=1.-X      NFNC036
C      SPACE=N2.EQ.1.      NFNC037
C      KAY2=OM2/C2*N2      NFNC038
C      IF(RSTART.EQ.0.) GO TO 1      NFNC039
C      SCALE=SQRT(KAY2/K2)      NFNC040
C      KR =SCALE*KR      NFNC041
C      KTH=SCALE*KTH      NFNC042
C      KPH=SCALE*KPH      NFNC043
C      1 CONTINUE      NFNC044
C***** CALCULATES A HAMILTONIAN H      NFNC045
C      H=.5*(C2*K2/OM2-N2)      NFNC046
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO      NFNC047
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI.      NFNC048
C      PHPT =-PNPT      NFNC049
C      PHPH =-PNPR      NFNC050
C      PHPTH=-PNPTH      NFNC051
C      PHPPH=-PNPPH      NFNC052
C      PHPOM=-NNP/OM      NFNC053
C      PHPKR =C2/OM2*KR      NFNC054
C      PHPKTH=C2/OM2*KTH      NFNC055
C      PHPKPH=C2/OM2*KPH      NFNC056
C      KPHPK=N2      NFNC057
C      RETURN      NFNC058
C      END      NFNC059
C      NFNC060

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SUBROUTINE BQWFWC                                BQWC001
C***** CALCULATES A HAMILTONIAN H                BQWC002
C***** (= BOOKER QUARTIC FOR VERTICAL INCIDENCE, S=0, C=1)    BQWC003
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO      BQWC004
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI.     BQWC005
C***** WITH FIELD, WITH COLLISIONS                 BQWC006
COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,C,LOGTEN      BQWC007
COMMON /RIN/ MODRIN(3),CJLL,FIELD,SPACE,KAY2,H,PHPT,PHPR,PHPTH,   BQWC008
1          PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,KPHPK,POLAR,LPOLAR,   BQWC009
2          SGN                                              BQWC010
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX      BQWC011
COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPFH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPTBQWC012
1          PYTPT,PYTPP,PYH,PYPPR,PYPPT,PYPPP               BQWC013
COMMON /ZZ/ M0DZ,Z,PZPR,PZPTH,PZPPH                  BQWC014
COMMON R,TH,PH,KR,KTH,KPH /WW/ IO(10),W0,W(400)        BQWC015
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART  BQWC016
COMMON /FLG/ NTYP,NEWWR,NEWWP,PENET,LINES,IHOP,HPUNCH     BQWC017
EQUIVALENCE (RAY,W(1)),(F,W(6))                      BQWC018
LOGICAL SPACE                                         BQWC019
REAL KR,KTH,KPH,K2,KDOTY,<4>,KDOTY2                BQWC020
COMPLEX KAY2,-1,PHPT,PHPR,PHPTH,PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,   BQWC021
1          POLAR,LPOLAR,I,U,RAD,D,PNPPS,PNPX,PNPY,PNPZ,UX,UX2,D2,   BQWC022
2          KPHPK,U2,A,B,ALPHA,BETA,GAMMA,PHPX,PHPY2,PHPK2,PHPU,PHPZ,   BQWC023
3          N2,PNPR,PNPTH,PNPPH,PNPVR,PNPVTH,PNPVPH>NNP,PNPT           BQWC024
DATA (MODRIN=8HBOOKER Q,8HUARTIC, ,8HS=0, C=1), (COLL=1.),   BQWC025
1          (FIELD=1.),                                         BQWC026
2          (X=0.), (PXPR=0.), (PXPTH=0.), (PXPPH=0.), (PXPT=0.),   BQWC027
3          (Y=0.), (PYPR=0.), (PYPTH=0.), (PYPFH=0.), (YR=0.), (PYRPR=0.),   BQWC028
4          (PYRPT=0.), (PYRPP=0.), (YTH=0.), (PYTPR=0.), (PYTPT=0.),   BQWC029
5          (PYTPP=0.), (PYH=0.), (PYPPR=0.), (PYPTT=0.), (PYPPP=0.)   BQWC030
6          ,(Z=0.), (PZPR=0.), (PZPTH=0.), (PZPPH=0.),   BQWC031
7          (I=(0.,1.)), (ABSLIM=1.E-5), (SGN=1.)             BQWC032
ENTRY RINDEX                                         BQWC033
DM=PIT2*1.E6*F                                      BQWC034
C2=C*C
K2=KR*KR+KTH*KTH+KPH*KPH                          BQWC035
OM2=OM*OM                                         BQWC036
CALL ELECTX                                         BQWC037
IF(X.LT..1) GO TO 2                                BQWC039
K4=K2*K2                                         BQWC040
OM4=OM2*OM2                                         BQWC041
C4=C2*C2                                         BQWC042
CALL MAGY                                           BQWC043
Y2=Y*Y                                         BQWC044
KDOTY=KR*YR+KTH*YTH+KPH*YPH                         BQWC045
KDOTY2=KDOTY*KDOTY                               BQWC046
CALL COLFRZ                                         BQWC047
U=CMPLX(1.,-Z)                                     BQWC048
U2=U*U                                         BQWC049
UX=U-X                                         BQWC050
UX2=UX*UX                                         BQWC051
A=UX*U2-U*Y2                                       BQWC052
B=-2.*U*UX2+Y2*(2.*U-X)                           BQWC053
ALPHA=A*C4*K4+X*KDOTY2*C4*K2                     BQWC054
BETA=B*C2*OM2*K2-X*KDOTY2*C2*OM2                 BQWC055
GAMMA=(UX2-Y2)*UX*OM4                            BQWC056
H=ALPHA+BETA+GAMMA                                BQWC057
PHPX=-U2*C4*K4+KDOTY2*C4*K2+(4.*U*UX-Y2)*C2*OM2*K2-KDOTY2*C2*OM2+  BQWC058
1 (-3.*UX2+Y2)*OM4                                BQWC059
PHPY2=-U*C4*K4+(2.*U-X)*C2*OM2*K2-UX*OM4       BQWC060
PHPY2=X*C2*(C2*K2-OM2)                           BQWC061
PHPU=(2.*U*UX+U2-Y2)*C4*K4+2.* (Y2-UX2-2.*U*UX)*C2*K2*OM2+(3.*UX2  BQWC062
1 -Y2)*OM4                                         BQWC063
PHZ=-I*PHPU                                         BQWC064
PHPK2=2.*A*C4*K2+X*KDOTY2*C4+B*C2*OM2           BQWC065

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PHPT=PHPX*PXPT          BQWC066
PHPR =PHPK*PXPR +PHPY2*2.*Y*PYPR +PHPKY2 *2.*KDOTY* BQWC067
1 (KR*PYRPR+KTH*PYTPR+KPH*PYPPR) +PHPZ*PZPR          BQWC068
PHPTH=PHPX*PXPTH+PHPY2*2.*Y*PYPTH+PHPKY2 *2.*KDOTY* BQWC069
1 (KR*PYRPT+KT4*PYTPT+KPH*PYPPT) +PHPZ*PZPTH          BQWC070
PHPPH=PHPX*PXPPH+PHPY2*2.*Y*PYPPH+PHPKY2 *2.*KDOTY* BQWC071
1 (KR*PYRPP+KTH*PYTPP+KPH*PYPPP) +PHPZ*PZPPH          BQWC072
PHPOM=(2.*BETA+4.*GAMMA)/OM          BQWC073
1 -2.*PHPX*X/OM-2.*PHPY2*Y2/OM-2.*PHPKY2 *KDOTY2/OM -PHPZ*Z/OM BQWC074
PHPKR= 2.*PHPK2*KR +2.*KDOTY*PHPKY2 *YR          BQWC075
PHPKTH=2.*PHPK2*KTH+2.*KDOTY*PHPKY2 *YTH          BQWC076
PHPKPH=2.*PHPK2*KPH+2.*KDOTY*PHPKY2 *YPH          BQWC077
<AY2=K2*(-BETA+SGN*RAY*CSQRT(BETA**2-4.*ALPHA*GAMMA))/(2.*ALPHA) BQWC078
C
C IF(RSTART.EQ.0.) GO TO 1          BQWC079
SCALE=SQRT((-REAL(BETA)+SGN*RAY*SQRT(REAL(BETA)**2 BQWC080
1 -4.*REAL(ALPHA)*REAL(GAMMA)))/(2.*REAL(ALPHA)))          BQWC081
KR =SCALE*KR          BQWC082
KTH=SCALE*KTH          BQWC083
KPH=SCALE*KPH          BQWC084
1 CONTINUE          BQWC085
C***** THE FOLLOWING 3 CARDS USED FOR RAY TRACING IN COMPLEX SPACE BQWC086
C IF(CABS((-BETA-SGN*RAY*CSQRT(BETA**2-4.*ALPHA*GAMMA))/ALPHA-2.). BQWC087
C 1LT.CABS((-BETA+SGN*RAY*CSQRT(BETA**2-4.*ALPHA*GAMMA))/ALPHA-2.) BQWC088
C 2 .AND.RSTART.EQ.0.) SGN=-SGN          BQWC089
KPHPK=4.*ALPHA+2.*BETA          BQWC090
SPACE=CABS(C2*KAY2/OM2-1.) .LT.ABSLIM          BQWC091
POLAR =SQRT(K2)*(U+X*OM2/(C2*KAY2-OM2))/KDOTY*I          BQWC092
LPOLR = SQRT(Y2-KDOTY2/K2)/UX*(1.-C2*KAY2/OM2)*I          BQWC093
RETURN          BQWC094
C      CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE BQWC095
C      APPLETON-HARTREE FORMULA WITH FIELD, WITH COLLISIONS          BQWC096
C
2 CONTINUE          BQWC097
VR =C/OM*KR          BQWC098
VTH=C/OM*KTH          BQWC099
VPH=C/OM*KPH          BQWC100
CALL MAGY          BQWC101
V2=VR**2+VTH**2+VPH**2          BQWC102
VDOTY=VR*YR+VTH*YTH+VPH*YPH          BQWC103
YLV=VDOTY/V2          BQWC104
YL2=VDOTY**2/V2          BQWC105
YT2=Y**2-YL2          BQWC106
YT4=YT2*YT2          BQWC107
CALL COLFRZ          BQWC108
U=CMPLX(1.,-Z)          BQWC109
UX=U-X          BQWC110
UX2=UX*UX          BQWC111
RAD=SGN*RAY*CSQRT(YT4+4.*YL2*UX2)          BQWC112
D=2.*U*UX-YT2+RA0          BQWC113
D2=D*D          BQWC114
N2=1.-2.*X*UX/D          BQWC115
PNPPS=2.*X*UX*(-1.+YT2-2.*UX2/RAD)/D2          BQWC116
PPSPR= YL2/Y*PYPR -(VR*PYRPR+VTH*PYTPR+VPH*PYPPR)*YLV          BQWC117
PPSPTH=YL2/Y*PYPTH-(VR*PYRPT+VTH*PYTPT+VPH*PYPPT)*YLV          BQWC118
PPSPPH=YL2/Y*PYPPH-(VR*PYRPP+VTH*PYTPP+VPH*PYPPP)*YLV          BQWC119
PNPX=-(2.*U*UX2-YT2*(U-2.*X)+(YT4+4.*YL2*UX2)/RAD)/(D2*Y)          BQWC120
PNPY=2.*X*UX*(-YT2+(YT4+2.*YL2*UX2)/RAD)/(D2*Y)          BQWC121
PNPZ=I*X*(-2.*UX2-YT2+YT4/RAD)/D2          BQWC122
NNP=N2-(2.*X*PNPX+Y*PNPY+Z*PNPZ)          BQWC123
PNPR =PNPX*PXPR +PNPY*PYPR +PNPZ*PZPR +PNPPS*PPSPR          BQWC124
PNPTH=PNPX*PXPTH+PNPY*PYPTH+PNPZ*PZPTH+PNPPS*PPSPTH          BQWC125
PNPPH=PNPX*PXPPH+PNPY*PYPPH+PNPZ*PZPPH+PNPPS*PPSPPH          BQWC126
PNPVR =PNPPS*(VR *YL2-VDOTY*YR )/V2          BQWC127
PNPVTH=PNPPS*(VTH*YL2-VDOTY*YTH)/V2          BQWC128
PNPVPH=PNPPS*(VPH*YL2-VDOTY*YPH)/V2          BQWC129

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PNPT=PNPX*PXPT                                BQWC131
SPACE=REAL(N2).EQ.1..AND.ABS(AIMAG(N2)).LT.ABSLIM   BQWC132
POLAR=-I*SQRT(V2)*(-YT2+RAD)/(2.*VDOOTY*UX)      BQWC133
GAM=(-YT2+RAD)/(2.*UX)                          BQWC134
LPOLAR=I*X*SQRT(YT2)/(UX*(U+GAM))            BQWC135
KAY2=OM2/C2*N2                                 BQWC136
IF(RSTART.EQ.0.) GO TO 3                      BQWC137
SCALE=SQRT(REAL(KAY2)/K2)                      BQWC138
KR =SCALE*KR                                    BQWC139
KTH=SCALE*KTH                                  BQWC140
KPH=SCALE*KPH                                  BQWC141
CONTINUE                                         BQWC142
H=.5*(C2*K2/OM2-N2)                           BQWC143
PHPT =-PNPT                                     BQWC144
PHPR =-PNPR                                     BQWC145
PHPTH=-PNPTH                                    BQWC146
PHPPH=-PNPPH                                    BQWC147
PHPOM=-NNP/OM                                   BQWC148
PHPKR =C2/OM2*KR -C/OM*PNPVR                  BQWC149
PHPKTH=C2/OM2*KTH-C/OM*PNPVTH                 BQWC150
PHPKPH=C2/OM2*KPH-C/OM*PNPVPH                 BQWC151
KPHPK=N2                                       BQWC152
RETURN                                           BQWC153
END                                              BQWC154

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SUBROUTINE BQHFCN                                BQNC001
C***** CALCULATES A HAMILTONIAN H                BQNC002
C***** (= BOOKER QUARTIC FOR VERTICAL INCIDENCE, S=0, C=1) BQNC003
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO BQNC004
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI. BQNC005
C***** WITH FIELD, NO COLLISIONS                 BQNC006
COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,C,LOGTEN BQNC007
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,KAY2I, BQNC008
1          ,HI,PHPT,PHPTI,PHPR,PHPRI,PHPTH,PHPPH,PHPHI,BQNC009
2          ,PHPOMI,PHPKR,PHPKRI,PHPKTH,PHPKTI,PHPKPH,PHPKPI,BQNC010
3          ,KPHPK,KPHPKI,POLAR,POLARI,LPOLEAR,LPOLEI,SGN BQNC011
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX    BQNC012
COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPR,BQNC013
1          ,PYTPT,PYTPP,YPH,PYPPR,PYPPT,PYPPP           BQNC014
COMMON /ZZ/ MODZ,Z(4)                            BQNC015
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART BQNC016
COMMON R,TH,P4,KR,KTH,KP4 /WW/ ID(10),W0,W(400)       BQNC017
EQUIVALENCE (RAY,W(1)),(F,W(6))                 BQNC018
LOGICAL SPACE                                     BQNC019
REAL N2,NNP,LPOLEAR,LPOLEI,KR,KTH,KPH,K2,KDOTY,K4,KDOTY2, BQNC020
1          KPHPK,KPHPKI,KAY2,KAY2I                 BQNC021
DATA (MODRIN=8HBOOKER Q,8HUARTIC, ,8HS=0, C=1),(COLL=0.), BQNC022
1          (FIELD=1.),(KAY2I=0.),(HI=0.),(PHPTI=0.),(PHPRI=0.), BQNC023
2          (PHPTHI=0.),(PHPPH=0.),(PHPOMI=0.),(PHPKR=0.),(PHPKRI=0.), BQNC024
3          (PHPKPI=0.),(KPHPKI=0.),(POLAR=0.),(LPOLEAR=0.), BQNC025
4          (X=0.),(PXPR=0.),(PXPTH=0.),(PXPPH=0.),(PXPT=0.), BQNC026
5          (Y=0.),(PYPR=0.),(PYPTH=0.),(PYPPH=0.),(YR=0.),(PYRPR=0.), BQNC027
6          (PYRPT=0.),(PYRPP=0.),(YTH=0.),(PYTPR=0.),(PYTPT=0.), BQNC028
7          (PYTPP=0.),(YPH=0.),(PYPPR=0.),(PYPPT=0.),(PYPPP=0.), BQNC029
8          (MODZ=1H ),(U=1.),(U2=1.)                 BQNC030
ENTRY RINDEX                                     BQNC031
OM=PIT2*1.E6*= BQNC032
C2=C*C C                                         BQNC033
K2=KR*KR+KTH*KTH+KPH*KPH                         BQNC034
OM2=OM*OM                                         BQNC035

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CALL ELECTX                                BQNC036
IF(X.LT..1) GO TO 2                         BQNC037
<4=K2*K2                                     BQNC038
OM4=OM2*OM2                                    BQNC039
C4=C2*C2                                     BQNC040
CALL MAGY                                     BQNC041
Y2=Y*Y                                       BQNC042
KDOTY=KR*YR+KTH*YTH+KPH*YPH                BQNC043
KDOTY2=KDOTY*KDOTY                           BQNC044
UX=U-X                                       BQNC045
UX2=UX*UX                                     BQNC046
A=UX*U2-U*Y2                                 BQNC047
B=-2.*U*UX2+Y2*(2.*U-X)                     BQNC048
ALPHA=A*C4*K4*X*KDOTY2*C4*K2               BQNC049
BETA=B*C2*OM2*K2-X*KDOTY2*C2*OM2           BQNC050
GAMMA=(UX2-Y2)*UX*OM4                        BQNC051
H=ALPHA+BETA+GAMMA                          BQNC052
PHPX=-U2*C4*<4+KDOTY2*C4*K2+(4.*U*UX-Y2)*C2*OM2*K2-KDOTY2*C2*OM2+
1 (-3.*UX2+Y2)*OM4                         BQNC053
BQNC054
PHPY2=-U*C4*K4+(2.*U-X)*C2*OM2*K2-UX*OM4 BQNC055
PHPKY2 =X*C2*(C2*K2-OM2)                    BQNC056
PHPK2=2.*A*C4*K2+X*KDOTY2*C4+B*C2*OM2    BQNC057
PHPT=PHPX*PXPT                               BQNC058
PHPR =PHPX*PXPR +PHPY2*2.*Y*PYPR +PHPKY2 *2.*KDOTY*
1 (KR*PYRPR+KTH*PYTPR+KPH*PYPPR)           BQNC059
PHPTH=PHPX*PXPT+PHPY2*2.*Y*PYPTH+PHPKY2 *2.*KDOTY*
1 (KR*PYRPT+KTH*PYTP+KPH*PYPP)              BQNC061
BQNC062
PHPPH=PHPX*PXPH+PHPY2*2.*Y*PYPPH+PHPKY2 *2.*KDOTY*
1 (KR*PYRPP+KTH*PYTPP+KPH*PYPPP)            BQNC063
BQNC064
PHPM=(2.*BETA+4.*GAMMA)/OM                  BQNC065
1 -2.*PHPX*X/OM-2.*PHPY2*Y2/OM-2.*PHPKY2 *KDOTY2/OM BQNC066
BQNC067
PHPKR= 2.*PHPK2*KR +2.*KDOTY*PHPKY2 *YR
PHPKTH=2.*PHPK2*KTH+2.*KDOTY*PHPKY2 *YTH
BQNC068
PHPKPH=2.*PHPK2*KPH+2.*KDOTY*PHPKY2 *YPH
BQNC069
KAY2 = K2 *(-BETA+RAY*SQRT(BETA**2-4.*ALPHA*GAMMA))/(2.*ALPHA)
BQNC070
IF(RSTART.EQ.0.) GO TO 1
BQNC071
SCALE=SQRT(KAY2/K2)
BQNC072
KR =SCALE*KR
BQNC073
KTH=SCALE*KTH
BQNC074
KPH=SCALE*KPH
BQNC075
CONTINUE                                     BQNC076
BQNC077
KPHPK=4.*ALPHA+2.*BETA
SPACE=KAY2.EQ.OM2/C2
BQNC078
POLARI=SQRT(K2)*(U+X*OM2/(C2*KAY2-OM2))/KDOTY
BQNC079
LPOLRI= SQRT(Y2-KDOTY2/K2)/UX*(1.-C2*KAY2/OM2)
BQNC080
RETURN                                         BQNC081
C      CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE
C      APPLETON-HARTREE FORMULA WITH FIELD, NO COLLISIONS          BQNC082
BQNC083
2  CONTINUE                                     BQNC084
VR =C/OM*KR                                     BQNC085
VTH=C/OM*KTH                                    BQNC086
VPH=C/OM*KPH                                    BQNC087
CALL MAGY                                     BQNC088
V2=VR**2+VTH**2+VPH**2                         BQNC089
VDOTY=VR*YR+VTH*YTH+VPH*YPH                   BQNC090
YLV=VDOTY/V2                                    BQNC091
YL2=VDOTY**2/V2                                BQNC092
YT2=Y**2-YL2                                    BQNC093
YT4=YT2*YT2                                    BQNC094
JK=U-X                                       BQNC095
UX2=UX*UX                                     BQNC096
RAD=RAY*SQRT(YT4+4.*YL2*JK2)                  BQNC097
D=2.*UX-YT2+RAD                                BQNC098
D2=D*D                                       BQNC099
N2=1.-2.*X*UX/D                                BQNC100

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PNPPS=2.*X*UX*(-1.+YT2-2.*UX2)/RAD)/D2          BQNC101
PPSPR= YL2/Y*PYPR -(VR*PYRPR+VTH*PYTPR+VPH*PYPPR)*YLV   BQNC102
PPSPTH=YL2/Y*PYPTH-(VR*PYRPT+VTH*PYTPT+VPH*PYPTP)*YLV   BQNC103
PPSPPH=YL2/Y*PYPPH-(VR*PYRPP+VTH*PYTPP+VPH*PYPPP)*YLV   BQNC104
PNPX=- (2.*UX2-YT2*(U-2.*X)+(YT4*(U-2.*X)+4.*YL2*UX*UX2)/RAD)/D2
PNPY=2.*X*UX*(-YT2+(YT4+2.*YL2*UX2)/RAD)/(D2*Y)        BQNC105
NNP=N2-(2.*X*PNPX+Y*PNPY)                         BQNC106
PNPR =PNPX*PX>R +PNPY*PYPR +PNPPS*PPSPR           BQNC107
PNPTH=PNPX*PXPTH+PNPY*PYPTH+PNPPS*PPSPTH         BQNC108
PNPPH=PNPX*PXPPH+PNPY*PYPPH+PNPPS*PPSPPH         BQNC109
PNPVR =PNPPS*(VR *YL2-VDOTY*YR )/V2              BQNC110
PNPVTH=PNPPS*(VTH*YL2-VDOTY*YTH)/V2             BQNC111
PNPVPH=PNPPS*(VPH*YL2-VDOTY*YPH)/V2             BQNC112
PNPT=PNPX*PXPT                                     BQNC113
SPACE=N2.EQ.1.                                     BQNC114
POLARI=SQRT(V2)*(YT2-RAD)/(2.*VDOTY*UX)          BQNC115
GAM=(-YT2+RAD)/(2.*UX)                            BQNC116
LPOLRI=K*SQRT(YT2)/(UX*(J+GAM))                 BQNC117
KAY2=OM2/C2*N2                                     BQNC118
IF(RSTART.EQ.0.) GO TO 3                          BQNC119
SCALE=SQRT(KAY2/K2)                                BQNC120
KR =SCALE*KR                                       BQNC121
KTH=SCALE*KTH                                      BQNC122
KPH=SCALE*KPH                                      BQNC123
KPH=SCALE*KPH                                      BQNC124
3 CONTINUE                                         BQNC125
H=.5*(C2*K2/OM2-N2)                                BQNC126
PHPT =-PNPT                                         BQNC127
PHPR =-PNPR                                         BQNC128
PHPTH=-PNPTH                                       BQNC129
PHPPH=-PNPPH                                       BQNC130
PHPOM=-NNP/OM                                      BQNC131
PHPKR =C2/OM2*KR -C/OM*PNPVR                      BQNC132
PHPKTH=C2/OM2*KTH-C/OM*PNPVTH                     BQNC133
PHPKPH=C2/OM2*KPH-C/OM*PNPVPH                     BQNC134
KPHPK=N2                                           BQNC135
RETURN                                            BQNC136
END                                              BQNC137-

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C SUBROUTINE SHWF                               SHWF001
C CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE SHWF002
C SEN-WYLLER FORMULA -- WITH FIELD               SHWF003
C NEEDS SUBROUTINE FSW AND FUNCTIONS C AND S.    SHWF004
COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,SEA,LOGTN SHWF005
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,H,PHPT,PHPR,PHPTH, SHWF006
1      PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,KPHPK,POLAR,LPOLAR, SHWF007
2      SGN                                         SHWF008
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX     SHWF009
COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPRSHWF010
1      ,PYPTP,PYTPP,YPH,PYPPR,PYPPT,PYPPP                SHWF011
COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH                  SHWF012
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART SHWF013
COMMON R,TH,P1,KR,KTH,KPH /WW/ ID(10),W0,W(400)       SHWF014
EQUIVALENCE (RAY,W(1)),(F,W(6))                   SHWF015
LOGICAL SPACE                                      SHWF016
REAL KR,KTH,KPH,K2                                 SHWF017
COMPLEX KAY2,H,PHPT,PHPR,PHPTH,PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH, SHWF018
1      KPHPK,POLAR,LPOLAR,I,U,RAD,D,PNPPS,PNPX,PNPY,PNPZ,UX,UX2, SHWF019
2      ALPHA,BETA,GAMMA,A,B,C,TEMP1,TEMP2,TEMP3,ALPOAL,BEPOBE, SHWF020
3      GAPOGA,CB2,N2M1,J2,D2GA,DAL,DBET,DGAM,DADY,DAOZ,DBDY,DBDZ,SHWF021
4      DCDY,DCDZ,DUDZ,DT1DX,DT1DY,DT1DZ,DT1DPS,DT2DX,DT2DY,DT2DZ,SHWF022
5      DT2DPS,DRAADX,DRAADDY,DRAADDZ,DROOPS,DDDX,DDDY,DDDZ,DDDPX, SHWF023

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6      UPX,N2,PNPR,PNPT1,PNPPH,PNPVR,PNPVTH,PNPVPH,NNP,PNPT      SWWF024
DATA (MODRIN=BH      SE,84N-WYLLER,BH FORMULA),(COLL=1.),      SWWF025
1      (FIELD=1.),,(LPOLAR=(0.,0.)),,                           SWWF026
2      (X=0.),,(PXPR=0.),,(PKPTH=0.),,(PXPPH=0.),,(PXPT=0.),, SWWF027
3      (Y=0.),,(PYPR=0.),,(PYPTH=0.),,(PYPPH=0.),,(YR=0.),,(PYRPR=0.),, SWWF028
4      (PYRPT=0.),,(PYRPP=0.),,(YTH=0.),,(PYTPR=0.),,(PYTPT=0.),, SWWF029
5      (PYTPP=0.),,(YPH=0.),,(PYPPR=0.),,(PYPPT=0.),,(PYPPP=0.),, SWWF030
6      (Z=0.),,(PZPR=0.),,(PZPTH=0.),,(PZPPH=0.),,              SWWF031
7      (I=(0.,1.)),,(ABSLIM=1.E-5)                                SWWF032
ENTRY RINDEX
OM=PIT2*1.E6*F
C2=SEA*SEA
K2=KR*KR+KTH*KTH+KPH*KPH
OM2=OM*OM
VR =SEA/OM*KR
VTH=SEA/OM*KTH
VPH=SEA/OM*KP4
CALL ELECTX
CALL MAGY
OPY=1.+Y
OMY=1.-Y
CALL COLFRZ
Z2=Z*Z
CALL FSW(1./Z,ALPHA,DAL)
ALPOAL=DAL/ALPHA
CALL FSW(OMY/Z,BETA,DBET)
BEPLOBE=DBET/BETA
CALL FSW(OPY/Z,GAMMA,DGAM)
GAPOGA=DGAM/GAMMA
U=Z/ALPHA
DUDZ=(1.+ALPOAL/Z)/ALPHA
U2=U*U
UX=U-X
UPX=U+X
B=ALPHA/BETA
DBDY=B*BEPLOBE/Z
DBDZ=-B*(ALPOAL-OMY*BEPLOBE)/Z2
C=ALPHA/GAMMA
DCDY=-C*GAPOGA/Z
DCDZ=-C*(ALPOAL-OPY*GAPOGA)/Z2
A=.5*(B+C)-1.
DADY=.5*(DBDY+DCDY)
DADZ=.5*(DBDZ+DCDZ)
TEMP3=(1.-B*C)*U2+A*U*UPX
V2=    VR**2+VTH**2+VPH**2
VDOTY=VR*YR+VTH*YTH+VPH*YPH
YL2=VDOTY**2/V2
YT2=Y**2-YL2
Y2=Y*Y
S2PSI=YT2/Y2
C2PSI=YL2/Y2
UX2=UX*UX
CB2=(C-B)**2
TEMP1=TEMP3*S2PSI
DT1DX=A*U*S2PSI
DT1DY=(U*UPX*DADY-U2*(B*DCDY+C*DBDY))*S2PSI
DT1DZ=(2.*U*DUDZ*(1.-B*C+A)+A*X*DUDZ-U2*(B*DCDZ+C*DBDZ)+U*UPX*DADZ)SWWF080
1)*S2PSI
C      (1/YLYT) D/DPSI(TEMP1)                                SWWF081
DT1DPS=2.*TEMP1/YT2                                         SWWF082
TEMP2=U2*CB2*X2*C2PSI                                         SWWF083
DT2DX=-2.*UX*U2*CB2*C2PSI                                     SWWF084
DT2DY=2.*U2*UX2*C2PSI*(C-B)*(DCDY-DBDY)                      SWWF085
DT2DZ=2.*U2*UX2*C2PSI*(C-B)*(DCDZ-DBDZ)+2.*TEMP2*(1./U+1./UX)*DUDZSWWF087
C      (1/YLYT) D/DPSI(TEMP2)                                SWWF088
DT2DPS=-2.*TEMP2/YL2                                         SWWF089

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RAD=RAY*CSQRT(TEMP1**2+TEMP2) SWWF090
DRA0DX=(TEMP1*DT1DX +.5 *DT2DX ) /RAD SWWF091
DRA0DY=(TEMP1*DT1DY + .5*DT2DY ) /RAD SWWF092
DRA0DZ=(TEMP1*DT1DZ + .5*DT2DZ ) /RAD SWWF093
C (1/YLYT) D/DPSI(RAD) SWWF094
DRDOPS=(TEMP1*DT1DPS+ .5*DT2DPS) /RAD SWWF095
D=2.*U*UX*(1.+A)-TEMP1+RAD+2.*A*U*X*S2PSI SWWF096
DDDX=-2.*U-DT1DX*DRA0DX+2.*A*U*S2PSI SWWF097
DDDY= 2.*U*UX*DADY-DT1DY+DRA0DY+2.*U*S2PSI*DADY SWWF098
DDDZ=2.*{1.+A}*DUDZ*(U+UX)+2.*U*UX*DAOZ-DT1DZ+DRA0DZ+2.*X*S2PSI* SWWF099
1 (A*DUDZ+J*DADZ) SWWF100
C (1/YLYT) D/DPSI(D) SWWF101
DD0PS=-DT1DPS+DRDOPS+2.*A*U*X/Y2 SWWF102
N2M1=-2.*X*(UX+U*A*S2PSI)/D SWWF103
N2=1.+N2M1 SWWF104
C N D/DX(N) SWWF105
PNPX=-(JX+U*A*S2PSI)*(1.-X*DD0X/D)/D+X/D SWWF106
C N D/DY(N) SWWF107
PNPY=-X*U*S2PSI/D*DADY-.5*N2M1/D*DD0Y SWWF108
C N D/DZ(N) SWWF109
PNPZ=-X*(1.+A*S2PSI)/D*DUDZ-X*U*S2PSI/D*DADZ-.5*N2M1/D*DD0Z SWWF110
C (N/YLYT) D/DPSI(N) SWWF111
PNPPS=-X*U/A/(D*Y2) -.5*N2M1/D*DD0PS SWWF112
YLV=VOOTY/V2 SWWF113
C (YLYT) D/DR(PSI) SWWF114
PPSPR=YL2/Y*PYPR-(VR*PYRPR+VTH*PYTPR+VPH*PYPPR)*YLV SWWF115
C (YLYT) D/DTHETA(PSI) SWWF116
PPSPTH=YL2/Y*PYPTH-(VR*PYRPT+VTH*PYTPT+VPH*PYPPT)*YLV SWWF117
C (YLYT) D/DPHI(PSI) SWWF118
PPSPPH=YL2/Y*PYPPH-(VR*PYRPP+VTH*PYTPP+VPH*PYPPP)*YLV SWWF119
PNPR=PNPX*PXPR+PNPY*PYPR+PNPZ*PZPR+PNPPS*PPSPR SWWF120
PNPTH=PNPX*PXPTH+PNPY*PYPTH+PNPZ*PZPTH+PNPPS*PPSPTH SWWF121
PNPPH=PNPK*PXPPH+PNPY*PYPPH+PNPZ*PZPPH+PNPPS*PPSPHH SWWF122
PNPVR=PNPPS*(VR*YL2/V2-YLV*YR) SWWF123
PNPVTH=PNPPS*(VTH*YL2/V2-YLV*YTH) SWWF124
PNPVPH=PNPPS*(VPH*YL2/V2-YLV*YPH) SWWF125
NNP=N2-(2.*X*PNPX+Y*PNPY+Z*PNPZ) SWWF126
PNPT=PNPX*PXPT SWWF127
POLAR=I*(TEMP1-RAD)*Y*SQRT(V2)/(U*UX*(C-B)*VDOOTY) SWWF128
COSPSI=VDOOTY/(Y*SQRT(V2)) SWWF129
LPOLAR=.5*I*(C-B)*POLAR+A*COSPSI)*SQRT(S2PSI)/ SWWF130
1 (POLAR*(JX*(1.+.5*I*(C-B)*COSPSI*POLAR)+A*(U-X*C2PSI))) SWWF131
SPACE=REAL(N2).EQ.1..AND.ABS(AIMAG(N2)).LT.ABSLIM SWWF132
KAY2=OM2/C2*N2 SWWF133
IF(RSTART.EQ.0.) GO TO 1 SWWF134
SCALE=SQRT(REAL(KAY2)/K2) SWWF135
KR =SCALE*KR SWWF136
KTH=SCALE*KTH SWWF137
KPH=SCALE*KPH SWWF138
1 CONTINUE SWWF139
C***** CALCULATES A HAMILTONIAN H SWWF140
H=.5*(C2*K2/OM2-N2) SWWF141
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO SWWF142
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI. SWWF143
PHPT =-PNPT SWWF144
PHPR =-PNPR SWWF145
PHPTH=-PNPTH SWWF146
PHPPH=-PNPPH SWWF147
PHPOM=-NP/OM SWWF148
PHPKR =C2/OM2*KR -SEA/OM*PNPVR SWWF149
PHPKTH=C2/OM2*KTH-SEA/OM*PNPVTH SWWF150
PHPKPH=C2/OM2*KPH-SEA/OM*PNPVPH SWWF151
KPHPK=N2 SWWF152
RETURN SWWF153
END SWWF154-

```

```

C SUBROUTINE SWNF          SWNF001
C   CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE    SWNF002
C   SEN-WYLLER FORMULA -- NO FIELD                                SWNF003
C   NEEDS SUBROUTINES FGSW AND FSW AND FUNCTIONS C AND S.          SWNF004
COMMON /CONST/ PI,PIT2,PIJ2,DEGS,RADIAN,K,C,LOGTEN           SWNF005
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,H,PHPT,PHPM,PHPTH,   SWNF006
1      PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,KPHPK,POLAR,LPOLAR,     SWNF007
2      SGN                                         SWNF008
COMMON /XXX/ MOOX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX           SWNF009
COMMON /YY/ MOODY,Y(16)                                     SWNF010
COMMON /ZZ/ MOOZ,Z,PZPR,PZPTH,PZPPH                         SWNF011
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART    SWNF012
COMMON R,TH,PH,KR,KTH,KPH /HW/ ID(10),W0,W(400)             SWNF013
EQUIVALENCE (RAY,H(1)),(F,W(6))                           SWNF014
LOGICAL SPACE                                         SWNF015
REAL KR,KTH,KPH,K2                                     SWNF016
COMPLEX KAY2,I,PHPT,PHPM,PHPTH,PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,   SWNF017
1      KPHPK,POLAR,LPOLAR,PNPX,PNPZ,F1,DF,G1,DG1,           SWNF018
2      N2,PNPR,PNPTH,PNPPH,PNPVR,PNPTH,PNPVPH,NNP,PNP           SWNF019
DATA (MODRIN=8H      SE,8H N-WYLLER,8H FORMULA),(COLL=1.),       SWNF020
1      (FIELD=0.),(POLAR=0.,1.),,(LPOLAR=0.,0.),            SWNF021
2      (X=0.),(PXPR=0.),(PXPTH=0.),(PXPPH=0.),(PXPT=0.),        SWNF022
3      (MOODY=1H ),                                         SWNF023
4      (Z=0.),(PZPR=0.),(PZPTH=0.),(PZPPH=0.),            SWNF024
5      (ABSLIM=1.E-5),(PNPVR=0.),(PNPTH=0.),(PNPVPH=0.)        SWNF025
ENTRY RINDEX                                         SWNF026
OM=PIT2*1.E6*F                                     SWNF027
C2=C*C                                         SWNF028
K2=KR*KR+KTH*KTH+KPH*KPH                         SWNF029
OM2=OM*OM                                         SWNF030
VR =C/OM*KR                                         SWNF031
VTH=C/OM*KTH                                         SWNF032
VPH=C/OM*KPH                                         SWNF033
CALL ELECTX                                         SWNF034
CALL COLFRZ                                         SWNF035
CALL FGSW(1./Z,F1,DF1,G1,DG1)                      SWNF036
N2=1.-X*G1                                         SWNF037
PNPX=-.5*G1                                         SWNF038
PNPZ=.5*X*DG1/Z**2                                 SWNF039
PNPR=PNPX*PXPR+PNPZ*PZPR                         SWNF040
PNPTH=PNPX*PXPTH+PNPZ*PZPTH                       SWNF041
PNPPH=PNPX*PXPPH+PNPZ*PZPPH                     SWNF042
NNP=N2-(Z.*X*PNPX+Z*PNPZ)                         SWNF043
PNPT=PNPX*PXPT                                         SWNF044
SPACE=REAL(N2).EQ.1..AND.ABS(AIMAG(N2)).LT.ABSLIM   SWNF045
KAY2=OM2/C2*N2                                         SWNF046
IF(RSTART.EQ.0.) GO TO 1                           SWNF047
SCALE=SQRT(REAL(KAY2)/K2)                          SWNF048
KR =SCALE*KR                                         SWNF049
KTH=SCALE*KTH                                         SWNF050
KPH=SCALE*KPH                                         SWNF051
1  CONTINUE                                         SWNF052
***** CALCULATES A HAMILTONIAN H                   SWNF053
H=.5*(C2*K2/OM2-N2)                               SWNF054
***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO   SWNF055
***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI. SWNF056
PHPT =-PNPT                                         SWNF057
PHPM =-PNPR                                         SWNF058
PHPTH=-PNPTH                                       SWNF059
PHPPH=-PNPPH                                       SWNF060
PHPOM=-NNP/OM                                       SWNF061
PHPKR =C2/OM2*KR                                    SWNF062
PHPKTH=C2/OM2*KTH                                  SWNF063
PHPKPH=C2/OM2*KPH                                  SWNF064
KPHPK=N2                                         SWNF065
RETURN                                         SWNF066
END                                         SWNF067-

```

```

SUBROUTINE FGSW (X,F,DF,G,DG)          FGSW001
COMPLEX F,DF,G,DG                     FGSW002
CALL FSW (X,F,DF)                      FGSW003
IF(ABS(X).GT.50.) GO TO 1              FGSW004
G=X*F                                  FGSW005
DG=F+X*DF                             FGSW006
RETURN                                 FGSW007
1 X2=X*X                               FGSW008
X3=X2*X                               FGSW009
T2=2.*X2                              FGSW010
T3=3.*X2                              FGSW011
T4=4.*X2                              FGSW012
T8=8.*X2                              FGSW013
T12=12.*X2                            FGSW014
T16=16.*X2                            FGSW015
G=CMPLX(1.-35./T4*(1.-99./T4*(1.-195./T4*(1.-323./T4)))/T4,
12.5*(1.-63./T4*(1.-143./T4*(1.-255./T4*(1.-399./T4))))/X)   FGSW016
DG=.5*CMPLX(35.*(1.-99./T2*(1.-585./T8*(1.-323./T3*(1.-2415./T16))) FGSW017
1))/X3,                                FGSW018
2-5.*(1.-189./T4*(1.-715./T12*(1.-357./T4*(1.-513./T4))))/X2)   FGSW019
RETURN                                 FGSW020
END                                    FGSW021
                                         FGSW 22-

```

```

SUBROUTINE FSW (Z,F,DF)                FSW 001
C      F(Z) = Z*C3/2(Z) + 2.5*I*C5/2(Z) AND DF(Z) = DF/DZ        FSW 002
C      WHERE THE INPUT Z IS REAL AND THE OUTPUT F AND DF ARE COMPLEX. FSW 003
C      NEEDS THE SUBPROGRAMS FOR THE FRESNEL INTEGRAL FUNCTIONS S AND CFSW 004
DIMENSION A(10),B(10),D(10)           FSW 005
COMPLEX F,DF,C1,C2,C3,C8,W,TEMP,I    FSW 006
DATA (I=(0.,1.)), (PI=3.1415926536), (A3=1.3333333333) FSW 007
DATA (C2=(1.,1.)), (C3=(1.,-1.)), (C4=.79788456 1),(C6=1.3333333333) FSW 008
C      C4=SQRT (2./PI)                         FSW 009
DATA (A=.36230845E-02,.29579186E+00,.23193588E+01,.91355870E+01,   FSW 010
1.25856287E+02,.60488560E+02,.12562218E+03,.24214980E+03,   FSW 011
2.44918106E+03,.84244774E+03),   FSW 012
3(B=.16747479E-02,.84796280E-01,.25285001E+00,.22665857E+00,   FSW 013
4.83871933E-01,.13811875E-01,.98017417E-03,.26299148E-04,   FSW 014
5.19761006E-06,.18781476E-09),   FSW 015
5(D=.10080653E-03,.46117941E-01,.38507643E+00,.68507885E+00,   FSW 016
7.42648105E+00,.10742102E+00,.10985920E-01,.40924533E-03,   FSW 017
8.41881263E-05,.54513142E-08),(G=1.5045055)   FSW 018
C1=2./3.*I                           FSW 019
C8=C2*A3*SQRT(PI/2.)                 FSW 020
X=Z                                  FSW 021
X2=X*X                               FSW 022
X3=X2*X                               FSW 023
IF(ABS(X).GT.50.) GO TO 500          FSW 024
IF(ABS(X).GT.6.) GO TO 1             FSW 025
IF(ABS(X).LT..05) GO TO 200          FSW 026
C      FRESNEL
IF(X.GT.0.) GO TO 300               FSW 027
100 Y=C4*SQRT(-X)                  FSW 028
K2=X*X                               FSW 029
H=(COS(X)+I*SIN(X))*(1.-C3*(C(Y)+I*S(Y)))   FSW 030
F =C1+C6*(X+C3*X*Y/W)               FSW 031
DF=A3*CMPLX(1.,X)+CMPLX(1.5,X)*A3*C3*X/Y*W   FSW 032
      RETURN                           FSW 033
300 Y=C4*SQRT(X)                  FSW 034
X2=X*X                               FSW 035
H=(COS(X)+I*SIN(X))*(1.-C2*(C(Y)-I*S(Y)))   FSW 036
F =C1+C6*(X-C2*X*Y/W)               FSW 037
DF=A3*CMPLX(1.,X)-CMPLX(1.5,X)*A3*C2*X/Y*W   FSW 038
      RETURN                           FSW 039
                                         FSW 040

```

```

C          POWER SERIES
200 X=ABS(Z)          FSW 041
X2=X*X          FSW 042
X3=X2*X          FSW 043
X4=X*X3          FSW 044
X5=X*X4          FSW 045
TEMP=-C8* SQRT(X)*CEXP(I*X)          FSW 046
F=CMPLX(4./3.*X-16./9.*X3+64./315.*X5,2./3.+8./3.*X2-32./45.*X4)          FSW 047
1   +TEMP*X          FSW 048
DF=CMPLX(4./3.-16./3.*X2+64./63.*X4,16./3.*X-128./45.*X3          FSW 049
1   +256./945.*X5)          FSW 050
2   +TEMP*CMPLX(1.5,X)          FSW 051
IF(Z.GE.0.) RETURN          FSW 052
F=-CONJG(F)          FSW 053
DF=CONJG(DF)          FSW 054
RETURN          FSW 055
FSW 056
C          HERMITE
1 XQ = X**2          FSW 057
X2=XQ          FSW 058
FR = 0.          FSW 059
FI = 0.          FSW 060
DFR = 0.          FSW 061
DFI = 0.          FSW 062
DO 2 J = 1,10          FSW 063
SS = A(J) + XQ          FSW 064
SB = B(J)/SS          FSW 065
SD = D(J)/SS          FSW 066
FR = FR + SB          FSW 067
FI = FI + SD          FSW 068
DFR = DFR + SB/SS          FSW 069
2 DFI = DFI + SD/SS          FSW 070
F = CMPLX(X*FR,FI)*G          FSW 071
DF = G*(FR - 2.*X*CMPLX(X*DFR,DFI))          FSW 072
RETURN          FSW 073
FSW 074
C          ASYMPTOTIC
500 X2=X*X          FSW 075
X3=X2*X          FSW 076
X4=X3*X          FSW 077
X5=X4*X          FSW 078
T2=2.*X2          FSW 079
T3=3.*X2          FSW 080
T4=4.*X2          FSW 081
T8=8.*X2          FSW 082
T16=16.*X2          FSW 083
T28=28.*X2          FSW 084
F=CMPLX((1.-35./T4*(1.-99./T4*(1.-195./T4*(1.-323./T4)))))/X          FSW 085
1,5.*(1.-63./T4*(1.-143./T4*(1.-255./T4*(1.-399./T4))))/T2          FSW 086
DF=-CMPLX((1.-105./T4*(1.-165./T4*(1.-273./T4*(1.-2907./T28)))))/X2          FSW 087
1,5.*(1.-63./T2*(1.-429./T8*(1.-255./T3*(1.-1995./T16))))/X3          FSW 088
RETURN          FSW 089
END          FSW 090
FSW 91-

```

```

FUNCTION C(X)                               C  001
DOUBLEPRECISION   PIH, XD, Y, V, A, QZ, QN, Q, Z      C  002
DATA (A1=0.3183099),(A2=0.10132),(B1=0.0968),(B2=0.154) C  003
PIH = 1.570796326794897                  C  004
XA = ABS(X)                                C  005
IF (XA.GT.4.)    GOTO 20                   C  006
C                                         C  007
XD = X                                     C  008
Y = PIH*XD*XD                            C  009
V = Y*Y                                    C  010
A = 1.D0                                    C  011
Z = A                                      C  012
M = 15.*(XA + 1.)                         C  013
DO 10   I = 1, M                           C  014
KZ=2*(I-1)                                 C  015
KV=4*(I-1)                                 C  016
QZ = KV + 1                               C  017
QN = (KZ + 1)*(KZ + 2)*(KV + 5)           C  018
Q = QZ/QN                                  C  019
A = -A*Q*V                                 C  020
Z = Z + A                                  C  021
Z = Z*XD                                    C  022
C = Z                                      C  023
RETURN                                     C  024
C                                         C  025
20   W = PIH*X**X                         C  026
XV=XA**4                                  C  027
C=0.5+(A1-B1/XV)*SIN(W)/XA-(A2-B2/XV)*COS(W)/XA**3 C  028
IF (X.LT.0.)      C = -C                 C  029
RETURN                                     C  030
END                                         C  31-

```

```

FUNCTION S(X)                               S  001
DOUBLEPRECISION   PIH, XD, Y, V, A, QZ, QN, Q, Z      S  002
DATA (A1=0.3183099),(A2=0.10132),(B1=0.0968),(B2=0.154) S  003
PIH = 1.570796326794897                  S  004
C                                         S  005
XA = ABS(X)                                S  006
IF (XA.GT.4.)    GOTO 20                   S  007
C                                         S  008
XD = X                                     S  009
Y = PIH*XD*XD                            S  010
V = Y*Y                                    S  011
A = Y/3.D0                                 S  012
Z = A                                      S  013
M = 15.*(XA + 1.)                         S  014
DO 10   I = 1, M                           S  015
KZ=2*(I-1)                                 S  016
KV=4*(I-1)                                 S  017
QZ = KV + 3                               S  018
QN = (KZ + 2)*(KZ + 3)*(KV + 7)           S  019
Q = QZ/QN                                  S  020
A = -A*Q*V                                 S  021
10  Z = Z + A                             S  022
Z = Z*XD                                    S  023
S = Z                                      S  024
RETURN                                     S  025
C                                         S  026
20   W = PIH*X**X                         S  027
XV=XA**4                                  S  028
S=0.5-(A1-B1/XV)*COS(W)/XA-(A2-B2/XV)*SIN(W)/XA**3 S  029
IF (X.LT.0.)      S = -S                 S  030
RETURN                                     S  031
END                                         S  32-

```

### APPENDIX 3. ELECTRON DENSITY SUBROUTINES WITH INPUT PARAMETER FORMS

The following electron density models are available. The input parameter forms, which describe the model, and the subroutine listings are given on the pages shown.

a.	Tabular profiles (TABLEX)	111
b.	Subroutine GAUSEL	113
c.	Chapman layer with tilts, ripples, and gradients (CHAPX)	115
d.	Chapman layer with variable scale height (VCHAPX)	117
e.	Double, tilted $\alpha$ -Chapman layer (DCHAPT)	118
f.	Linear Layer (LINEAR)	120
g.	Plain or quasi-parabolic layer (QPARAB)	121
h.	Analytic equatorial model (BULGE)	122
i.	Exponential profile (EXPX)	124

A further source of versatility in this ray tracing program is the ease with which specific ionospheric models, suited to the users needs, may be introduced. To add electron density models not included in the program, the user must write a subroutine that calculates the normalized electron density ( $X$ ) and its gradient ( $\partial X / \partial r$ ,  $\partial X / \partial \theta$ ,  $\partial X / \partial \phi$ ) as a function of position in spherical coordinates ( $r$ ,  $\theta$ ,  $\phi$ ). ( $X = 80.5 \times 10^{-6} N/f^2$ , where  $N$  is the electron density in  $\text{cm}^{-3}$  and  $f$  is the wave frequency in MHz.)

Both  $X$  and its gradient must be continuous functions of position. The formulas for  $\partial X / \partial r$ ,  $\partial X / \partial \theta$ , and  $\partial X / \partial \phi$  must be consistent with the variation of  $X$  with  $r$ ,  $\theta$ , and  $\phi$ . Otherwise, the program will run slowly and give incorrect results.

The coordinates  $r$ ,  $\theta$ ,  $\phi$  refer to the computational coordinate system, which may not be the same as geographic coordinates. In particular, they are geomagnetic coordinates when the earth-centered dipole model of the earth's magnetic field is used.

The input to the subroutine ( $r$ ,  $\theta$ ,  $\phi$ ) is through blank common. (See

Table 3.) The output is through common block /XX/. (See Table 8.) It is useful if the name of the subroutine suggests the model to which it corresponds. The subroutine should have an entry point ELECTX so that other subroutines in the program can call it. Any parameters needed by the subroutine should be input into W101 through W149 of the W array. (See Table 2.) If the model needs massive amounts of data, these should be read in by the subroutine following the example of TABLEX. As in the already existing electron density subroutines, provision should be made for perturbations to the electron density model (irregularities) by having the statement

```
IF(PERT.NE.0.) CALL ELECT1
```

before the RETURN statement at the end of the subroutine.

INPUT PARAMETER FORM FOR SUBROUTINE TABLEX

## IONOSPHERIC ELECTRON DENSITY PROFILE

First card tells how many profile points in I4 format. The cards following the first card give the height and electron density of the profile points one point per card in F8.2, E12.4 format. The heights must be in increasing order. Set W100 = 1.0 to read in a new profile. After the cards are read, TABLEX will reset W100 = 0.0. This subroutine makes an exponential extrapolation down using the bottom 2 points in the profile.

```

SUBROUTINE TABLEX TABX001
C      CALCULATES ELECTRON DENSITY AND GRADIENT FROM PROFILES HAVING TABX002
C      THE SAME FORM AS THOSE USED BY CROFTS RAY TRACING PROGRAM TABX003
C MAKES AN EXPONENTIAL EXTRAPOLATION DOWN USING THE BOTTOM TWO POINTS TABX004
C      NEEDS SUBROUTINE GAUSEL TABX005
      DIMENSION HPC(250),FN2C(250),ALPHA(250),BETA(250),GAMMA(250), TABX006
1      DFLTA(250),SLOPE(250),MAT(4,5) TABX007
      COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2) TABX008
      COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX TABX009
      COMMON R(6) /WW/ ID(10),W0,W(400) TABX010
      EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(READFN,W(100)),(PERT,W(150)) TABX011
      REAL MAT,K TABX012
      DATA (MODX(I)=6HTABLEX) TABX013
      ENTRY ELECTX TABX014
      IF (READFN.EQ.0.) GO TO 10 TABX015
      READFN=0.
      READ 1000, NOC,(HPC(I),FN2C(I),I=1,NOC) TABX016
1000 FORMAT (I4/(F8.2,E12.4)) TABX017
      PRINT 1200, (HPC(I),FN2C(I),I=1,NOC) TABX018
1200 FORMAT(1H1,14X,6HHEIGHT,4X,16HELECTRON DENSITY/(1X,F20.10,E20.10))TABX020
      A=0.
      IF (FN2C(1).NE.0.) A= ALOG (FN2C(2)/FN2C(1))/(HPC(2)-HPC(1)) TABX021
      FN2C(1)=K*FN2C(1) TABX022
      FN2C(2)=K*FN2C(2) TABX023
      SLOPE(1)=A*FN2C(1) TABX024
      SLOPE(NOC)=0. TABX025
      NMAX=1 TABX026
      DO 6 I=2,NOC TABX027
      IF (FN2C(I).GT.FN2C(NMAX)) NMAX=I TABX028
      IF (I.EQ.NOC) GO TO 4 TABX029
      FN2C(I+1)=K*FN2C(I+1) TABX030
      DO 3 J=1,3 TABX031
      M=I+J-2 TABX032
      MAT(J,1)=1. TABX033
      MAT(J,2)=HPC(M) TABX034
      MAT(J,3)=HPC(M)**2 TABX035
      MAT(J,4)=FN2C(M) TABX036
3      CALL GAUSEL (MAT,4,3,4,NRANK) TABX037
      IF (NRANK.LT.3) GO TO 60 TABX038
      SLOPE(I)=MAT(2,4)+2.*MAT(3,4)*HPC(I) TABX039
4      DO 5 J=1,2 TABX040
      M=I+J-2 TABX041
      MAT(J,1)=1. TABX042
      MAT(J,2)=HPC(M) TABX043
      MAT(J,3)=HPC(M)**2 TABX044
      MAT(J,4)=HPC(M)**3 TABX045
      MAT(J,5)=FN2C(M) TABX046
      L=J+2 TABX047
      MAT(L,1)=0. TABX048
      MAT(L,2)=1. TABX049
      MAT(L,3)=2.*HPC(M) TABX050
      MAT(L,4)=3.*HPC(M)**2 TABX051
5      MAT(L,5)=SLOPE(M) TABX052
      CALL GAUSEL (MAT,4,4,5,NRANK) TABX053
      IF (NRANK.LT.4) GO TO 60 TABX054
      ALPHA(I)=MAT(1,5) TABX055
      BFTA(I)=MAT(2,5) TABX056
      GAMMA(I)=MAT(3,5) TABX057
6      DELTA(I)=MAT(4,5) TABX058
      HMAX=HPC(NMAX) TABX059
      NH=2 TABX060
      NH=2 TABX061
10     H=R(1)-EARTH TABX062
      F2=F**F TABX063
      PXPR=0. TABX064
      IF (H.GE.HPC(1)) GO TO 12 TABX065
11     NH=2 TABX066
      X=0. TABX067

```

```

IF(FN2C(1).EQ.0.) GO TO 50 TABX068
X=FN2C(1)*EXP(A*(H-HPC(1)))/F2 TABX069
PXPR=A*X TABX070
GO TO 50 TABX071
12 IF (H.GE.HPC(NOC)) GO TO 18 TABX072
NSTEP=1 TABX073
IF (H.LT.HPC(NH-1)) NSTEP=-1 TABX074
15 IF (HPC(NH-1).LE.H.AND.H.LT.HPC(NH)) GO TO 16 TABX075
NH=NH+NSTEP TABX076
GO TO 15 TABX077
16 X=(ALPHA(NH)+H*(BETA(NH)+H*(GAMMA(NH)+H*DELTA(NH))))/F2 TABX078
PXPR=(BETA(NH)+H*(2.*GAMMA(NH)+H*3.*DELTA(NH)))/F2 TABX079
GO TO 50 TABX080
18 X=FN2C(NOC)/F2 TABX081
50 IF (PERT.NE.0.) CALL ELECT1 TABX082
RETURN TABX083
60 PRINT 6000, I,HPC(I) TABX084
6000 FORMAT(4H THE,I4,55HTH POINT IN THE ELECTRON DENSITY PROFILE HAS TTABX085
1HE HEIGHT,F8.2,40H KM, WHICH IS THE SAME AS ANOTHER POINT.)
CALL EXIT TABX086
END TABX087

```

```

SUBROUTINE GAUSEL (C,NRD,NRR,NCC,NSF) GAUS001
C***** SAME AS SUBROUTINE GAUSSEL WRITTEN BY L. DAVID LEWIS ****GAUS002
DIMENSION C(NRD,NCC),L(128,2) GAUS003
C BITS = 2.**-18 GAUS004
DATA (BITS=3.8146972656E-6) GAUS005
NR=NRP GAUS006
NC=NCC GAUS007
IF(NC.LT.NR.OR.NR.GT.128.OR.NR.LE.0) CALL EXIT GAUS008
C INITIALIZE. GAUS009
NSF=0 GAUS010
NRM=NR-1 GAUS011
NRP=NR+1 GAUS012
D=1. GAUS013
LSD=1 GAUS014
DO 1 KR=1,NR GAUS015
L(KR,1)=KR GAUS016
1 L(KR,2)=0 GAUS017
IF(NR.EQ.1) GO TO 42 GAUS018
C ELIMINATION PHASE. GAUS019
DO 41 KP=1,NRM GAUS020
KPP=KP+1 GAUS021
PM=0. GAUS022
MPN=0 GAUS023
C SEARCH COLUMN KP FROM DIAGONAL DOWN FOR MAX PIVOT. GAUS024
DO 2 KR=KP,NR GAUS025
LKR=L(KR,1) GAUS026
PT=ABS(C(LKR,KP)) GAUS027
IF(PT.LE.PM) GO TO 2 GAUS028
PM=PT GAUS029
MPN=KR GAUS030
LMP=LKR GAUS031
2 CONTINUE GAUS032
C IF MAX PIVOT IS ZERO, MATRIX IS SINGULAR. GAUS033
IF(MPN.EQ.0) GO TO 9 GAUS034
NSF=NSF+1 GAUS035
IF(MPN.EQ.KP) GO TO 3 GAUS036

```

```

C      NEW ROW NUMBER KP HAS MAX PIVOT.          GAUS041
C      LSD=-LSD                                     GAUS042
C      L(KP,2)=L(MPN,1)=L(KP,1)                   GAUS043
C      L(KP+1)=LMP                                 GAUS044
C
C      ROW OPERATIONS TO ZERO COLUMN KP BELOW DIAGONAL.   GAUS045
3      MKP=L(KP,1)                                GAUS046
      P=C(MKP,KP)                                GAUS047
      D=D*P                                      GAUS048
      DO 41 KR=KPP,NR                           GAUS049
      MKR=L(KR,1)                                GAUS050
      Q=C(MKR,KP)/P                            GAUS051
      IF(Q.EQ.0.) GO TO 41                      GAUS052
C
C      SUBTRACT Q * PIVOT ROW FROM ROW KR.        GAUS053
      DO 4 LC=KPP,NC                           GAUS054
      R=Q*C(MKP,LC)                            GAUS055
      C(MKR,LC)=C(MKR,LC)-R                  GAUS056
4      IF(ABS(C(MKR,LC)).LT.ABS(R)*BITS) C(MKR,LC)=0.  GAUS057
41     CONTINUE                                  GAUS058
C
C      LOWER RIGHT HAND CORNER.                 GAUS059
42     LNR=L(NR,1)                                GAUS060
      P=C(LNR,NR)                                GAUS061
      IF(P.EQ.0.) GO TO 9                        GAUS062
      NSF=NSF+1                                 GAUS063
      D=D*P*LSD                                GAUS064
      IF(NR.EQ.NC) GO TO 8                      GAUS065
C
C      BACK SOLUTION PHASE.                     GAUS066
      DO 61 MC=NRP,NC                           GAUS067
      C(LNR,MC)=C(LNR,MC)/P                  GAUS068
      IF(NR.EQ.1) GO TO 61                      GAUS069
      DO 6 LL=1,NRM                           GAUS070
      KR=NR-LL                                 GAUS071
      MR=L(KR,1)                                GAUS072
      KRP=KR+1                                 GAUS073
      DO 5 MS=KRP,NR                           GAUS074
      LMS=L(MS,1)                                GAUS075
      R=C(MR,MS)*C(LMS,MC)                  GAUS076
      C(MR,MC)=C(MR,MC)-R                  GAUS077
5      IF(ABS(C(MR,MC)).LT.ABS(R)*BITS) C(MR,MC)=0.  GAUS078
6      C(MR,MC)=C(MR,MC)/C(MR,KR)            GAUS079
61     CONTINUE                                  GAUS080
C
C      SHUFFLE SOLUTION ROWS BACK TO NATURAL ORDER.    GAUS081
      DO 71 LL=1,NRM                           GAUS082
      KR=NR-LL                                 GAUS083
      MKR=L(KR,2)                                GAUS084
      IF(MKR.EQ.0) GO TO 71                      GAUS085
      MKP=L(KR,1)                                GAUS086
      DO 7 LC=NRP,NC                           GAUS087
      Q=C(MKR,LC)                                GAUS088
      C(MKR,LC)=C(MKP,LC)                  GAUS089
      C(MKP,LC)=Q                                GAUS090
7      CONTINUE                                  GAUS091
C
C      NORMAL AND SINGULAR RETURNS. GOOD SOLUTION COULD HAVE D=0.  GAUS092
8      C(1,1)=D                                GAUS093
      GO TO 91                                 GAUS094
9      C(1,1)=0.                                GAUS095
91     RETURN                                 GAUS096
      END                                     GAUS097
                                         GAUS098
                                         GAUS099
                                         GAUS100
                                         GAUS101
                                         GAUS102
                                         GAUS103
                                         GAUS104-

```

## INPUT PARAMETER FORM FOR SUBROUTINE CHAPX

An ionospheric electron density model consisting of a Chapman layer with tilts, ripples, and gradients

$$\begin{aligned}
 f_N^2 &= f_c^2 \exp(-\alpha(1-z-e^{-z})) \\
 z &= \frac{h - h_{max}}{H} \\
 f_c^2 &= f_{co}^2 \left( 1 + A \sin\left(2\pi\left(\theta - \frac{\pi}{2}\right)/B\right) + C\left(\theta - \frac{\pi}{2}\right) \right) \\
 h_{max} &= h_{max_0} + E\left(\theta - \frac{\pi}{2}\right) R_0
 \end{aligned}$$

$f_N$  is the plasma frequency

$h$  is the height above the ground

$R_0$  is the radius of the earth in km

and  $\theta$  is the colatitude in radians.

Specify:

Critical frequency at the equator,  $f_{co} = \underline{\hspace{2cm}}$  MHz (W101)

Height of the maximum electron density at the equator,  $h_{max_0} = \underline{\hspace{2cm}}$  km (W102)

Scale height,  $H = \underline{\hspace{2cm}}$  km (W103)

$\alpha = \underline{\hspace{2cm}}$  (W104, 0.5 for an  $\alpha$  Chapman layer, 1.0 for a  
 $\beta$  Chapman layer)

Amplitude of periodic variation of  $f_c^2$  with latitude,  $A = \underline{\hspace{2cm}}$  (W105)

Period of variation of  $f_c^2$  with latitude,  $B = \underline{\hspace{2cm}}$  rad  
 $\deg$  (W106)  
 $\text{km}$

Coefficient of linear variation of  $f_c^2$  with latitude,  $C = \underline{\hspace{2cm}}$   $\text{rad}^{-1}$  (W107)

Tilt of the layer,  $E = \underline{\hspace{2cm}}$  rad (W108)  
 $\deg$

```

C SUBROUTINE CHAPX          CHAP001
C   CHAPMAN LAYER WITH TILTS, RIPPLES, AND GRADIENTS    CHAP002
C   W(104) = 0.5 FOR AN ALPHA-CHAPMAN LAYER             CHAP003
C   = 1.0 FOR A BETA-CHAPMAN LAYER                      CHAP004
C COMMON /CONST/ PI,PIT2,PID2,DUM(5)                   CHAP005
C COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX    CHAP006
C COMMON R(6) /WW/ ID(10),W0,W(400)                  CHAP007
C EQUIVALENCE (THETA,R(2))                           CHAP008
C EQUIVALENCE (EARTH,R,W(2)),(F+W(6)),(FC,W(101)),(HM,W(102)),    CHAP009
C 1 (SH,W(103)),(ALPHA,W(104)),(A,W(105)),(B,W(106)),(C,W(107)),    CHAP010
C 2 (E,W(108)),(PERT,W(150))                         CHAP011
C DATA (MODX(1)=6H CHAPX)                          CHAP012
C ENTRY ELECTX                                     CHAP013
C THETA2=THETA-PID2                                CHAP014
C HMAX=HM+EARTH*R*E*THETA2                         CHAP015
C H=R(1)-EARTH           CHAP016
C Z=(H-HMAX)/SH                                     CHAP017
C D=0.                                         CHAP018
C IF (B.NE.0.) D=PIT2/B                            CHAP019
C TEMP=1.+A*SIN(D*THETA2)+C*THETA2                CHAP020
C EXZ=1.-EXP(-Z)                                    CHAP021
C X=(FC/F)**2*TEMP*EXP(ALPHA*(EXZ-Z))            CHAP022
C PXPR=-ALPHA*X*EXZ/SH                            CHAP023
C PXPTH=XX*(D*A*SIN(PID2-D*THETA2)+C)/TEMP-PXPR*EARTH*R*E    CHAP024
C IF (PERT.NE.0.) CALL ELECT1                     CHAP025
C RETURN                                         CHAP026
C END                                           CHAP 27-

```

## INPUT PARAMETER FORM FOR SUBROUTINE VCHAPX

An ionospheric electron density model consisting of a Chapman layer with variable scale height

$$f_N^2 = f_c^2 \tau^{\frac{1}{2}} e^{\frac{1}{2}(1-\tau)}$$

$$\tau = \left( \frac{h_{\max}}{h} \right)^x$$

$h$  is the height above the ground.

Specify:

critical frequency,  $f_c$  = \_\_\_\_\_ MHz (W101)

height of maximum electron density,  $h_{\max}$  = \_\_\_\_\_ km (W102)

$x$  = \_\_\_\_\_ (W103)

SUBROUTINE VCHAPX	VCHA001
CHAPMAN LAYER WITH VARIABLE SCALE HEIGHT	VCHA002
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	VCHA003
COMMON R(6) /WW/ ID(10),W0,W(400)	VCHA004
EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(FC,W(101)),(HM,W(102)),	VCHA005
1 (CHI,W(103)),(PERT,W(150))	VCHA006
DATA (MODX(1)=6HVCHAPX)	VCHA007
ENTRY ELECTX	VCHA008
HMAX=HM	VCHA009
X=PXPR=0.	VCHA010
H=R(1)-EARTH	VCHA011
IF (H.LE.0.) GO TO 50	VCHA012
TAU=(HM/H)**CHI	VCHA013
X=(FC/F)**2*SQRT(TAU)*EXP(0.5*(1.-TAU))	VCHA014
PXPR=.5*X*(TAU-1.)*CHI/H	VCHA015
50 IF (PERT.NE.0.) CALL ELECT1	VCHA016
RETURN	VCHA017
END	VCHA018-

INPUT PARAMETER FORM FOR SUBROUTINE DCHAPT

An ionospheric electron density model consisting of a double, tilted  
a-Chapman layer

$$f_N^2 = f_{c1}^2 \exp \frac{1}{2} (1 - z_1 - e^{-z_1}) + f_{c2}^2 \exp \frac{1}{2} (1 - z_2 - e^{-z_2})$$

$$z_1 = \frac{h-h_{m1}}{H_1} ; z_2 = \frac{h-h_{m2}}{H_2}$$

$$f_{c1}^2 = f_{c10}^2 C(\theta - \pi/2)$$

$$f_{c2}^2 = f_{c20}^2 C(\theta - \pi/2)$$

$$h_{m1} = h_{m10} + R_o E \left( \frac{\pi}{180} \right) (\theta - \frac{\pi}{2})$$

$$h_{m2} = h_{m20} + R_o E \left( \frac{\pi}{180} \right) (\theta - \frac{\pi}{2})$$

Specify:

$$f_{c10} = \text{MHz } (f_{c1} \text{ at equator}) \quad (W101)$$

$$h_{m10} = \text{Km } (h_{m1} \text{ at equator}) \quad (W102)$$

$$H_1 = \text{Km} \quad (W103)$$

$$f_{c20} = \text{MHz } (f_{c2} \text{ at equator}) \quad (W104)$$

$$h_{m20} = \text{Km } (h_{m2} \text{ at equator}) \quad (W105)$$

$$H_2 = \text{Km} \quad (W106)$$

$$C = \text{rad}^{-1} \text{ (fractional change in } f_{c1}, f_{c2}, \text{ position for increases southward)} \quad (W107)$$

$$E = \text{deg } \text{(positive for upward tilt to the south)} \quad (W108)$$

```

      SUBROUTINE DCHAPT          DCHA001
      TWO CHAPMAN LAYERS WITH TILTS   DCHA002
      COMMON /CONST/ PI,PIT2,PID2,DUM(5) DCHA003
      COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX DCHA004
      COMMON R(6) /WW/ ID(10),W0,W(400) DCHA005
      EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(FC1,W(101)),(HM1,W(102)), DCHA006
      1 (SH1,W(103)),(FC2,W(104)),(HM2,W(105)),(SH2,W(106)),(C,W(107)), DCHA007
      2 (E,W(108)),(PERT,W(150)) DCHA008
      DATA (MODX(1)=6HDCHAPT) DCHA009
      ENTRY ELECTX DCHA010
      EARTH=EARTH*E DCHA011
      THETA2=R(2)-PID2 DCHA012
      HMAX=HM1+EARTH*THETA2 DCHA013
      X=PXPR=PXPTH=0. DCHA014
      H=R(1)-EARTH R DCHA015
      IF (H.LT.0.) GO TO 50 DCHA016
      Z1=(H-HMAX)/SH1 DCHA017
      EXPZ1=1.-EXP(-Z1) DCHA018
      TEMP=1.+C*THETA2 DCHA019
      X=(FC1/F)**2*TEMP*EXP(.5*(EXPZ1-Z1)) DCHA020
      PXPR=-0.5*X*EXPZ1/SH1 DCHA021
      PXPTH=X*C/TEMP-PXPR*EARTH DCHA022
      IF (FC2.EQ.0.) GO TO 50 DCHA023
      Z2=(H-HM2-EARTH*THETA2)/SH2 DCHA024
      EXPZ2=1.-EXP(-Z2) DCHA025
      X2=(FC2/F)**2*TEMP*EXP(.5*(EXPZ2-Z2)) DCHA026
      X=X+X2 DCHA027
      PXPR2=-0.5*X2*EXPZ2/SH2 DCHA028
      PXPR=PXPR+PXPR2 DCHA029
      PXPTH=PXPTH+X2*C/TEMP-PXPR2*EARTH DCHA030
      50 IF (PERT.NE.0.) CALL ELECT1 DCHA031
      RETURN DCHA032
      END DCHA033-

```

## INPUT PARAMETER FORM FOR SUBROUTINE LINEAR

An ionospheric electron density model consisting of a linear layer

$N = 0$  for  $h \leq h_{\min}$   
 $N = A(h - h_{\min})$  for  $h > h_{\min}$

The ray will penetrate if  $h > h_{\max}$ .

Specify:

$A = \underline{\hspace{2cm}}$  electrons/cm<sup>3</sup>/ km (W101)

$h_{\max} = \underline{\hspace{2cm}}$  km (W102)

$h_{\min} = \underline{\hspace{2cm}}$  km (W103)

```
C      SUBROUTINE LINEAR                               LINE001
          LINEAR ELECTRON DENSITY MODEL                LINE002
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2)    LINE003
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX   LINE004
COMMON R(6) /WW/ ID(10),W0,W(400)                 LINE005
EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(FACT,W(101)),(HM,W(102)),
1 (HMIN,W(103)),(PERT,W(150))                   LINE006
REAL K                                         LINE007
DATA (MODX(1)=6HLINEAR)                         LINE008
ENTRY ELECTX                                     LINE009
H=R(1)-EARTH          LINE010
HMAX=HM                                     LINE011
X=PXPR=0.                                     LINE012
IF (H.LE.HMIN) GO TO 50                         LINE013
PXPR=K*FACT/F**2                                LINE014
X=PXPR*(H-HMIN)                                LINE015
50 IF (PERT.NE.0.) CALL ELECT1                  LINE016
RETURN                                         LINE017
END                                            LINE018
                                                LINE019-
```

## INPUT PARAMETER FORM FOR SUBROUTINE QPARAB

An ionospheric electron density model consisting of a parabolic or a quasi-parabolic layer (concentric)

$$f_N^2 = f_c^2 \left[ 1 - \frac{h-h_{\max}}{Y_m} \cdot C^2 \right] \quad \text{if } f_N^2 > 0.$$

$$f_N^2 = 0, \text{ otherwise.}$$

$C = 1$ , for a parabolic layer

$$C = \frac{R_o + h_{\max} - Y_m}{R_o + h} \quad \text{for a quasi-parabolic layer}$$

where  $R_o$  is the radius of the earth.

Specify:

Critical frequency,  $f_c = \underline{\hspace{2cm}}$  Mc/s (W101)

Height of maximum electron density,  $h_{\max} = \underline{\hspace{2cm}}$  km. (W102)

Semi-thickness,  $Y_m = \underline{\hspace{2cm}}$  km. (W103)

Type of profile:

Plain parabolic  $\underline{\hspace{2cm}}$  (W104 = 0.)

Quasi-parabolic  $\underline{\hspace{2cm}}$  (W104 = 1.)

```

C      SUBROUTINE QPARAB          PARA001
C      PLAIN PARABOLIC OR QUASI-PARABOLIC PROFILE    PARA002
C      W(104) = 0. FOR A PLAIN PARABOLIC PROFILE    PARA003
C              = 1. FOR A QUASI-PARABOLIC PROFILE    PARA004
C      COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX    PARA005
C      COMMON R(6) /WW/ ID(10),W0,W(400)           PARA006
C      EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(FC,W(101)),(HM,W(102)),
1     (YM,W(103)),(QUASI,W(104)),(PERT,W(150))    PARA007
      DATA (MODX(1)=6HQPARAB)                      PARA008
      ENTRY ELECTX                                     PARA009
      HMAX=HM                                         PARA010
      PXPR=0.                                         PARA011
      H=R(1)-EARTH          PARA012
      FCF2=(FC/F)**2          PARA013
      CONST=1.                                         PARA014
      IF (QUASI.EQ.1.) CONST=(EARTH+HM-YM)/R(1)    PARA015
      Z=(H-HM)/YM*CONST          PARA016
      X=MAX1(F(0.),FCF2*(1.-Z*Z))    PARA017
      IF (X.EQ.0.) GO TO 50          PARA018
      IF (QUASI.EQ.1.) CONST=(EARTH+HM)*(EARTH+HM-YM)/R(1)**2
      PXPR=-2.*Z*FCF2/YM*CONST          PARA019
50   IF (PERT.NE.0.) CALL ELECT1          PARA020
      RETURN          PARA021
      END          PARA022
                                         PARA023
                                         PARA024-

```

## INPUT PARAMETER FORM FOR SUBROUTINE BULGE

An analytic ionospheric electron density model which represents the general latitude variation of the equatorial ionosphere (afternoon, equinox, sunspot maximum) - see the center panel of figure 3.18b, page 133 of Davies (1965).

The model is an alpha Chapman layer with parameters which vary with geomagnetic latitude.

$$\frac{f_N^2}{N} = \frac{f_c^2}{e} e^{\frac{1}{2}(1-z-e^{-z})}$$

$$\text{where } z = \frac{h - h_{\max}}{H}$$

$f_N$  is the plasma frequency

$f_c$  is the critical frequency

$h_{\max}$  is the height of the maximum electron density

H is the scale height

h is height

---

$f_c$ ,  $h_{\max}$ , H vary with geomagnetic latitude in the following way:

if  $h < 100$  km,  $h_{\max} = 350$  km,  $f_c = 15$  Mc/s

---

For  $h \geq 100$  km,

$h_{\max} = 350$  if  $\lambda \geq 24^\circ$

$h_{\max} = 430 + 80 \cos\left(\frac{180}{24}\lambda\right)$  if  $\lambda < 24^\circ$

$\lambda$  is the geomagnetic latitude in degrees

---

$$f_c = \sqrt{50\left(\frac{\lambda}{8}\right)^2 \exp\left(2 - \left|\frac{\lambda}{8}\right|\right) + 40}$$

---

In all cases H is determined by the constraint that

$$\frac{f_N^2}{N} = 2 \text{ Mc/s at } 100 \text{ km.}$$

---

```

C SUBROUTINE BULGE                                BULG001
      ANALYTICAL MODEL OF THE VARIATION OF THE EQUATORIAL F2 LAYER    BULG002
C IN GEOMAGNETIC LATITUDE (EQUATORIAL BULGE AND ANOMALY)          BULG003
C SEE FIGURE 3-18B, PAGE 133 IN DAVIES (1965).                  BULG004
C THIS MODEL HAS NO VARIATION IN GEOMAGNETIC LONGITUDE.           BULG005
COMMON /CONST/ PI,PIT2,PID2,DUM(5)                 BULG006
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX        BULG007
COMMON R(6)/WW/ ID(10),W0,W(400)                  BULG008
EQUIVALENCE (EARTH,R(2)),(F,W(6)),(PERT,W(150))       BULG009
DATA (MODX(1)=6H BULGE)                           BULG010
ENTRY ELECTX                                     BULG011
H=R(1)-EARTH
PHMPTH=PFC2PTH=0.
HMAX=350.
FC2=225.
IF(H<LT+100.) GO TO 2                         BULG016
C   EQUATORIAL BULGE                               BULG017
BULLAT=7.5*(PID2-R(2))
IF(ABS(BUL LAT).GE.PI) GO TO 1                 BULG018
HMAX=430.+80.*COS(BULLAT)                      BULG020
PHMPTH=600.*SIN(BUL LAT)                        BULG021
C   EQUATORIAL ANOMALY                            BULG022
1  ANMLAT=22.5*(PID2-R(2))/PI                   BULG023
POW=2.-ABS(ANM LAT)
FC2=50.*ANM LAT**2*EXP( POW ) + 40.
PFC2PTH=-1125./PI*POW*ANMLAT*EXP(POW)         BULG025
C   FORCING PLASMA FREQ AT 100 KM TO BE 2 MHZ IN ORDER TO CALCULATE SH BULG027
2  ALPHA=2.* ALOG(FC2/4.)+1.                     BULG028
Z100=-ALOG(ALPHA)
DO 3 I=1,5                                       BULG029
3  Z100=-ALOG(ALPHA-Z100)                       BULG030
SH=(100.-HMAX)/Z100
Z=(H-HMAX)/SH
EXZ=1.-EXP(-Z)
X=FC2*EXP(.5*(EXZ-Z))/F**2
PXPR=-0.5*X*EXZ/SH
PXPTH=-PXPR*(1.-Z/Z100)*PHMPTH+(1.-Z*EXZ/(Z100*(1.-EXP(-Z100)))) BULG036
1  *X/FC2*PFC2PTH                               BULG037
IF (PERT.NE.0.) CALL ELECT1
RETURN
END                                              BULG040
                                                BULG041-

```

## INPUT PARAMETER FORM FOR SUBROUTINE EXPX

An exponential electron density profile

$$N = N_0 e^{a(h-h_0)}$$

$h$  is the height above the ground.

Specify:

the electron density at the height  $h_0$ ,  $N_0 = \underline{\hspace{2cm}}$  cm<sup>-3</sup> (W101)

the reference height,  $h_0 = \underline{\hspace{2cm}}$  km (W102)

the exponential increase of  $N$  with height,  $a = \underline{\hspace{2cm}}$  km<sup>-1</sup> (W103)

```

C      SUBROUTINE EXPX          EXPX001
      EXPONENTIAL ELECTRON DENSITY MODEL    EXPX002
      COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2)    EXPX003
      COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX    EXPX004
      COMMON R(6) /WW/ ID(10),W0,W(400)    EXPX005
      EQUIVALENCE (EARTH, W(2)), (F, W(6)),    EXPX006
1  (NO, W(101)), (HO, W(102)), (A, W(103)), (PERT, W(150))    EXPX007
      REAL N, NO, K    EXPX008
      DATA (MODX(1)=4HEXPX), (HMAX=350.)    EXPX009
      ENTRY ELECTX    EXPX010
      H=R(1)-EARTH    EXPX011
      N=NO * EXP(A*(H-H0))    EXPX012
      X=K*N/F**2    EXPX013
      PXPR=A*X    EXPX014
      IF (PERT.NE.0.) CALL ELECT1    EXPX015
      RETURN    EXPX016
      END    EXPX017-

```

#### APPENDIX 4. PERTURBATIONS TO ELECTRON DENSITY MODELS WITH INPUT PARAMETER FORMS

The following perturbations to electron density models (irregularities) are available. The input parameter forms, which describe the perturbation, and the subroutine listings are given on the pages shown.

a. Do-nothing perturbation (ELECT1)	126
b. East-west irregularity with an elliptical cross-section above the equator (TORUS)	127
c. Two east-west irregularities with elliptical cross-sections above the equator (DTORUS)	129
d. Increase in electron density at any latitude (TROUGH)	131
e. Increase in electron density produced by a shock wave (SHOCK)	132
f. "Gravity-wave" irregularity (WAVE)	134
g. "Gravity-wave" irregularity (WAVE2)	136
h. Height profile of time derivative of electron density for calculating Doppler shift (DOPPLER)	138

To add other perturbations to electron density models the user must write a subroutine to modify the normalized electron density ( $X$ ) and its gradient ( $\partial X / \partial r$ ,  $\partial X / \partial \theta$ ,  $\partial X / \partial \psi$ ) as a function of position in spherical polar coordinates ( $r$ ,  $\theta$ ,  $\psi$ ).

The restrictions on electron density models also apply to perturbations. Again, the coordinates  $r$ ,  $\theta$ ,  $\psi$  refer to the computational coordinate system, which may not be the same as geographic coordinates. In particular, they are geomagnetic coordinates when the earth-centered dipole model of the earth's magnetic field is used.

The input to the subroutine is through blank common (see Table 3) for the position ( $r$ ,  $\theta$ ,  $\psi$ ) and through common block /XX/ (see Table 8) for the unperturbed electron density and its gradient. The output is through common block /XX/. It is useful if the name of the subroutine suggests the perturbation model to which it corresponds. It should have an entry point ELECT1 so that it may be called by an electron density subroutine. Any parameters needed by the subroutine should be input into W151 through W199 of the W array. (See Table 2.)

If no perturbation is wanted, the following subroutine should be used.

```
C      SUBROUTINE ELECT1          ELEC001
      USE WHEN AN ELECTRON DENSITY PERTURBATION IS NOT WANTED   ELEC002
COMMON /XXX/ MODX(2),X(6)           ELEC003
COMMON /WW/  ID(10),W0,W(400)       ELEC004
EQUIVALENCE (PERT,W(150))         ELEC005
DATA (MODX(2)=6H NONE )          ELEC006
PERT=0.
RETURN                           ELEC007
END                             ELEC008
                                ELEC009-
```

## INPUT PARAMETER FORM FOR SUBROUTINE TORUS

A perturbation to an ionospheric electron density model consisting of an East-West irregularity with an elliptical cross section above the equator

$$N = N_0 (1 + \Delta)$$

$$\Delta = C_0 \exp \left\{ - \left[ \frac{(R_0 + H_0)(\theta - \pi/2) \cos \beta + (R - R_0 - H_0) \sin \beta}{A} \right]^2 \right\}$$

$$- \left[ \frac{(R - R_0 - H_0) \cos \beta - (R_0 + H_0)(\theta - \pi/2) \sin \beta}{B} \right]^2 \}$$

$R_0$  is the radius of the earth.

$R, \theta, \varphi$  give the position in spherical polar coordinates.

$N_0(R, \theta, \varphi)$  is any ionospheric electron density model.

Specify:

$$C_0 = \underline{\hspace{2cm}}. \quad (W151)$$

$$\text{Semi-major axis of ellipse, } A = \underline{\hspace{2cm}} \text{ km } (W152)$$

$$\text{Semi-minor axis of ellipse, } B = \underline{\hspace{2cm}} \text{ km } (W153)$$

$$\text{Tilt of ellipse, } \beta = \underline{\hspace{2cm}} \text{ degrees } (W154)$$

$$\text{Height of torus from ground, } H_0 = \underline{\hspace{2cm}} \text{ km } (W155)$$

(W150: = 1. to use perturbation, = 0. to ignore perturbation)

```

SUBROUTINE TORUS                               TOR 001
COMMON /CONST/ PI,PIT2,PID2,DUM(5)           TOR 002
COMMON /XXX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX   TOR 003
COMMON R(6) /WW/ ID(10),W0,W(400)          TOR 004
EQUIVALENCE (EARTH,R,W(2)),(C0,W(151)),(A,W(152)),(B,W(153)),
1 (BETA,W(154)),(HO,W(155))               TOR 005
REAL LAMBDA                                TOR 006
DATA (PDPP=0.),(MODX(2)=6H TORUS)          TOR 007
ENTRY ELECT1                                 TOR 008
IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN    TOR 010
IF (C0.EQ.0.) RETURN                         TOR 011
R0=EARTH+HO                                  TOR 012
Z=R(1)-R0                                    TOR 013
LAMBDA=R0*(R(2)-PID2)                      TOR 014
SINBET=SIN(BETA)                            TOR 015
COSBET=COS(BETA)                           TOR 016
P=LAMBDA*COSBET+Z*SINBET                  TOR 017
Y=Z*COSBET-LAMBDA*SINBET                  TOR 018
DELTA=C0*EXP(-(P/A)**2-(Y/B)**2)          TOR 019
DEL1=DELTA+1.                                TOR 020
PDPR=-2.*DELTA*(P*SINBET/A**2+Y*COSBET/B**2)  TOR 021
PDPT=-2.*DELTA*(P*R0*COSBET/A**2-Y*R0*SINBET/B**2)  TOR 022
PXPR=PXPR*DEL1+X*PDPR                     TOR 023
PXPTH=PXPTH*DEL1+X*PDPT                   TOR 024
PXPPH=PXPPH*DEL1+X*PDPP                  TOR 025
X=X*DEL1                                    TOR 026
RETURN                                     TOR 027
END                                         TOR 028-

```

## INPUT PARAMETER FORM FOR SUBROUTINE DTORUS

A perturbation to an ionospheric electron density model consisting of two east-west irregularities with elliptical cross sections above the equator. Since the model is expressed in spherical coordinates and does not depend on longitude, the perturbation is actually a torus circling the earth above the equator.

$$N = N_0 (1 + \Delta)$$

$$\Delta = C_1 \exp \left\{ - \left[ \frac{(r_0 + H_1)(\theta - \pi/2) \cos \beta + (r - r_0 - H_1) \sin \beta}{A_1} \right]^2 \right\}$$

$$- \left[ \frac{(r - r_0 - H_1) \cos \beta - (r_0 + H_1)(\theta - \pi/2) \sin \beta}{B_1} \right]^2 \}$$

$$+ C_2 \exp \left\{ - \left[ \frac{(r_0 + H_2)(\theta - \pi/2 + \delta\theta) \cos \beta + (r - r_0 - H_2) \sin \beta}{A_2} \right] \right\}$$

$$- \left[ \frac{(r - r_0 - H_2) \cos \beta - (r_0 + H_2)(\theta - \pi/2 + \delta\theta) \sin \beta}{B_2} \right]^2 \}$$

$\delta\theta$  = Northward angular displacement of the lower blob from the upper one

$$= \frac{H_1 - H_2}{\tan \beta (r_0 + H_2)}$$

$r_0$  is the radius of the earth.

$r, \theta, \psi$  are spherical (earth-centered) polar coordinates.

$N_0(r, \theta, \psi)$  is any electron density model.

Specify:

use perturbation \_\_\_\_\_ (W150 = 1.)

ignore perturbation \_\_\_\_\_ (W150 = 0.)

Fractional perturbation electron density at the center of the upper blob,  $C_1 =$  \_\_\_\_\_ (W151).

That of the lower blob,  $C_2 =$  \_\_\_\_\_ (W156).

Height (above ground) of the center of the upper blob,  
 $H_1 =$  \_\_\_\_\_ km (W155).

That of the lower blob,  $H_2 = \underline{\hspace{2cm}}$  km (W159).

Angle (with a horizontal southward vector) of the line joining the blob centers,  $\theta = \underline{\hspace{2cm}} \text{rad}$  ( $\underline{\hspace{2cm}} \text{deg}$ ) (W154).

Semi-axis of the upper blob, to the  $1/e$  perturbation contour, in the direction of the line joining the blobs,  $A_1 = \underline{\hspace{2cm}}$  km (W152).

That of the lower blob,  $A_2 = \underline{\hspace{2cm}}$  km (W157).

Semi-axis of the upper blob in the direction normal to the line joining the blobs,  $B_1 = \underline{\hspace{2cm}}$  km (W153).

That of the lower blob,  $B_2 = \underline{\hspace{2cm}}$  km (W158).

SUBROUTINE DTORUS	DTOR001
COMMON /CONST/ PI,PIT2,PID2,DUM(5)	DTOR002
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT	DTOR003
COMMON R(6) /WW/ ID(10),W0,W(400)	DTOR004
EQUIVALENCE (EARTH,R,W(2)),(C1,W(151)),(A1,W(152)),(B1,W(153)),	DTOR005
1 (BETA,W(154)),(H1,W(155)),(C2,W(156)),(A2,W(157)),(B2,W(158)),	DTOR006
2 (H2,W(159))	DTOR007
REAL LAMBDA1,LAMBDA2	DTOR008
DATA (MODX(2)=6HDTORUS),(PDPP=0.)	DTOR009
ENTRY ELECT1	DTOR010
IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0..) RETURN	DTOR011
IF (C1.EQ.0.) RETURN	DTOR012
R1=EARTH+H1	DTOR013
R2=EARTH+H2	DTOR014
Z1=R(1)-R1	DTOR015
Z2=R(1)-R2	DTOR016
LAMBDA1=R1*(R(2)-PID2)	DTOR017
LAMBDA2=R2*(R(2)-PID2+(H1-H2)/R2/TANF(BETA))	DTOR018
SINBET=SIN(BETA)	DTOR019
COSBET=COS(BETA)	DTOR020
P1=LAMBDA1*COSBET+Z1*SINBET	DTOR021
P2=LAMBDA2*COSBET+Z2*SINBET	DTOR022
Y1=Z1*COSBET-LAMBDA1*SINBET	DTOR023
Y2=Z2*COSBET-LAMBDA2*SINBET	DTOR024
DELT1=C1*EXP(-(P1/A1)**2-(Y1/B1)**2)	DTOR025
DELT2=C2*EXP(-(P2/A2)**2-(Y2/B2)**2)	DTOR026
DEL1=1.+DELT1+DELT2	DTOR027
PDPR1=-2.*DELT1*(P1*SINBET/A1**2+Y1*COSBET/B1**2)	DTOR028
PDPR2=-2.*DELT2*(P2*SINBET/A2**2+Y2*COSBET/B2**2)	DTOR029
PDPT1=-2.*DELT1*(P1*R1*COSBET/A1**2-Y1*R1*SINBET/B1**2)	DTOR030
PDPT2=-2.*DELT2*(P2*R2*COSBET/A2**2-Y2*R2*SINBET/B2**2)	DTOR031
PXPR=PXPR*DEL1+X*(PDPR1+PDPR2)	DTOR032
PXPTH=PXPTH*DEL1+X*(PDPT1+PDPT2)	DTOR033
PXPPH=PXPPH*DEL1*PDPP	DTOR034
X=X*DEL1	DTOR035
RETURN	DTOR036
END	DTOR037-

## INPUT PARAMETER FORM FOR SUBROUTINE TROUGH

A perturbation to an ionospheric electron density model consisting of an increase in electron density near any latitude

$$N = (1 + \Delta) N_0 (R, \theta, \varphi)$$

$$W = B \text{ for } \frac{\pi}{2} - \theta - \lambda \geq 0$$

$$\Delta = A \exp\left(-\left(\frac{\pi/2 - \theta - \lambda}{W}\right)^2\right)$$

$$W = B \times C \text{ for } \frac{\pi}{2} - \theta - \lambda < 0$$

$N_0 (R, \theta, \varphi)$  is any ionospheric electron density model.

$R, \theta, \varphi$  give the position in spherical polar coordinates.

Specify:

Amplitude of the perturbation,  $A = \underline{\hspace{2cm}}$  (W151)

half width of the perturbation,  $B = \underline{\hspace{2cm}}$  degrees (W152)

latitude of the perturbation,  $\lambda = \underline{\hspace{2cm}}$  degrees (W153)

width factor for South of trough,  $C = \underline{\hspace{2cm}}$  (W154)

(W150: = 1. to use perturbation, = 0. to ignore perturbation)

```

C      SUBROUTINE TROUGH                                TROU001
          A PERTURBATION TO AN ELECTRON DENSITY MODEL    TROU002
COMMON /CONST/ PI,PIT2,PID2,DUM(5)                   TROU003
COMMON /XXX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX   TROU004
COMMON R(6) /WW/ ID(10),W0,W(400)                  TROU005
EQUIVALENCE (A,W(151)),(B,W(152)),(ALAT,W(153)),(FACTOR,W(154)) TROU006
DATA (MODX(2)=6HTROUGH)                            TROU007
ENTRY ELECT1                                         TROU008
IF (X.EQ.0...AND.PXPR.EQ.0...AND.PXPTH.EQ.0...AND.PXPPH.EQ.0...) RETURN TROU009
IF (A.EQ.0.) RETURN                                 TROU010
ANGLE=R(2)+ALAT-PID2                               TROU011
WIDTH=B                                           TROU012
IF (ANGLE.GT.0.) WIDTH=FACTOR*B                  TROU013
ANGLE=ANGLE/WIDTH                                    TROU014
DELTA=A*EXP(-ANGLE**2)                           TROU015
DELI=DELTA+1.                                     TROU016
PXPR=PXPR*DELI                                    TROU017
PXPTH=PXPTH*DELI-2.*X*ANGLE*DELTA/WIDTH        TROU018
PXPPH=PXPPH*DELI                                    TROU019
X=X*DELI                                         TROU020
RETURN                                            TROU021
END                                              TROU022-

```

## INPUT PARAMETER FORM FOR SUBROUTINE SHOCK

A perturbation to an ionospheric electron density model consisting of an increase in electron density produced by a shock wave

$$N(R, \theta, \varphi) = N_0(R, \theta, \varphi) [1 + P \exp(-9 \left( \frac{\rho_c - \rho}{w} \right)^2)]$$

$$\rho_c = s(h - h_0) - w$$

$$\rho = R |\cos^{-1}[\cos(\varphi - \varphi_0) \cos(\lambda - \lambda_0)]|$$

$N_0(R, \theta, \varphi)$  is the ambient electron density specified by any electron density model.

$R, \theta, \varphi$  give the position in spherical polar coordinates.

$h = R - a$  is the height above the surface of the earth.

$a$  is the radius of the earth.

$\lambda = \frac{\pi}{2} - \theta$  is the latitude.

Specify:

Relative increase in electron density,  $P = \underline{\hspace{2cm}}$  (W151).

Width of the disturbance,  $w = \underline{\hspace{2cm}}$  km (W152).

Latitude of the center of the disturbance,  $\lambda_0 = \underline{\hspace{2cm}}$  radians or degrees (W153).

Longitude of the center of the disturbance,  $\varphi_0 = \underline{\hspace{2cm}}$  radians or degrees (W154).

Slope measured from vertical - rate of increase of  $\rho_c$  with height,  
 $s = \underline{\hspace{2cm}}$  (W155).

Height to the bottom of the disturbance,  $h_0 = \underline{\hspace{2cm}}$  km (W156).

(W150: = 1. to use perturbation, = 0. to ignore perturbation)

```

C   SUBROUTINE SHOCK                               SHOC001
A   A PERTURBATION TO AN ELECTRON DENSITY MODEL SIMULATING A SHOCK WAVE SHOC002
COMMON /CONST/ PI,PIT2,PID2,DUM(5)                SHOC003
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX SHOC004
COMMON R(6) /WW/ ID(10),W0,W(400)                 SHOC005
EQUIVALENCE (EARTH,R(2)),(P,W(151)),(WW,W(152)),(ALAT,W(153)), SHOC006
1 (ALON,W(154)),(S,W(155)),(HO,W(156))          SHOC007
REAL LAT,LON                                      SHOC008
DATA (MODX(2)=6H SHOCK)                          SHOC009
ENTRY ELECT1                                     SHOC010
IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0..) RETURN SHOC011
IF (P.EQ.0..OR.WW.EQ.0..) RETURN                  SHOC012
H=R(1)-EARTH
RHOC=S*(H-HO)-WW                                 SHOC013
LON=R(3)-ALON                                     SHOC014
LAT=PID2-R(2)-ALAT                                SHOC015
COSLON=COS(LON)                                   SHOC016
COSLAT=COS(LAT)                                   SHOC017
U=COSLON*COSLAT                                  SHOC018
RHO=R(1)*ACOS(U)                                 SHOC019
DIF=RHOC-RHO                                     SHOC020
CON=-9./WW**2                                    SHOC021
CONS=P*EXP(CON*DIF**2)                           SHOC022
CONST=1.+CONS                                     SHOC023
CON=2.*CON*CONS*DIF                            SHOC024
PXPR=PXPR*CONST+X*CON*(S-RHO/R(1))             SHOC025
CONS=R(1)*(1./SQRT(1.-U**2))                   SHOC026
PXPTH=PXPTH*CONST+X*CON*CONS*COSLON*SIN(LAT)  SHOC027
PXPPH=PXPPH*CONST-X*CON*CONS*COSLAT*SIN(LON)  SHOC028
X=X*CONST                                         SHOC029
RETURN                                           SHOC030
END                                              SHOC031
                                                SHOC032-

```

INPUT PARAMETER FORM FOR SUBROUTINE WAVE

A perturbation to an ionospheric electron density model consisting of a "gravity-wave" irregularity traveling from north pole to south pole

$$N = N_0(1 + \Delta)$$

$$\Delta = \delta \exp \left\{ - [ (R - R_0 - z_0) / H ]^2 \right\}.$$

$$\cos \left\{ 2\pi \left[ t' + (\pi/2 - \theta) \frac{R_0}{\lambda_x} + (R - R_0) / \lambda_z \right] \right\}$$

$$\frac{\partial N}{\partial t} = \frac{-2\pi}{\lambda_x} V_x N_0 \delta \exp - [ (R - R_0 - z_0) / H ]^2 .$$

$$\sin 2\pi \left[ t' + (\pi/2 - \theta) \frac{R_0}{\lambda_x} + (R - R_0) / \lambda_z \right]$$

$R_0$  is the radius of the earth.

$R$ ,  $\theta$ ,  $\varphi$  are the spherical (earth-centered) polar coordinates  
 $(\Delta$  is independent of  $\varphi$ ).

$N_0$  ( $R$ ,  $\theta$ ,  $\varphi$ ) is any electron density model.

Specify:

the height of maximum wave amplitude,  $z_0 = \underline{\hspace{2cm}}$  km (W151)

wave-amplitude "scale height,"  $H = \underline{\hspace{2cm}}$  km (W152)

wave perturbation amplitude,  $\delta = \underline{\hspace{2cm}}$  [0. to 1.] (W153)

horizontal trace velocity,  $V_x = \underline{\hspace{2cm}}$  km/sec (W154)  
 (needed only if Doppler shift is calculated)

horizontal wavelength,  $\lambda_x = \underline{\hspace{2cm}}$  km (W155)

vertical wavelength,  $\lambda_z = \underline{\hspace{2cm}}$  km (W156)

time in wave periods,  $t' = \underline{\hspace{2cm}}$  [0. to 1.] (W157)

(W150: = 1. to use perturbation, = 0. to ignore perturbation)

```

C SUBROUTINE WAVE          WAVE001
      PERTURBATION TO AN ALPHA-CHAPMAN ELECTRON DENSITY MODEL   WAVE002
COMMON /CONST/ PI,PIT2,PID2,DUM(5)                           WAVE003
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX           WAVE004
COMMON R(6) /WW/ ID(10),W0,W(400)                         WAVE005
EQUIVALENCE (EARTH,R(2)),(Z0,W(151)),(SH,W(152)),(DELTA,W(153)), WAVE006
1 (VSUBX,W(154)),(LAMBDAX,W(155)),(LAMDAZ,W(156)),(TP,W(157)) WAVE007
REAL LAMBDAX,LAMDAZ                                         WAVE008
DATA (MODX(2)=6H WAVE )                                     WAVE009
ENTRY ELECT1                                                 WAVE010
IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0..) RETURN WAVE011
IF (DELTA.EQ.0..OR.SH.EQ.0..) RETURN                         WAVE012
H=R(1)-EARTH
EXPQ=EXP(-((H-Z0)/SH)**2)                                    WAVE013
TMP=PIT2*(TP+(PID2-R(2)*EARTH/LAMBDAX+H/LAMDAZ)           WAVE014
SINW=SIN(TMP)                                              WAVE015
COSW=SIN(PID2-TMP)                                         WAVE016
CONS=1.0+DELTA*EXPQ*COSW                                    WAVE017
IF (H.NE.0..) PXPR=PXPR*CONS-X*DELTA*EXPQ*(2.0/SH**2*(H-Z0)*COSW WAVE018
1 +PIT2/LAMDAZ*SINW)                                       WAVE019
PXPTH=PXPTH*CONS+X*DELTA*PIT2*EARTH/LAMBDAX*SINW*EXPQ        WAVE020
PXPPH=PXPPH*CONS                                         WAVE021
PXPT=0..                                                 WAVE022
IF (VSUBX.NE.0..) PXPT=-PIT2*VSUBX/LAMBDAX*X*DELTA*EXPQ*SINW WAVE023
X=X*CONS                                              WAVE024
RETURN                                                 WAVE025
END                                                   WAVE026
                                         WAVE027-

```

INPUT PARAMETER FORM FOR SUBROUTINE WAVE 2  
PERTURBATION TO AN IONOSPHERIC ELECTRON DENSITY MODEL

A "gravity-wave" irregularity traveling from north pole to south pole - same as WAVE 1, but with Gaussian amplitude variations in latitude and longitude, and provision for a horizontal "group velocity"

$$N = N_0 (1 + AC)$$

$$A = \delta \exp - \left( \frac{r - r_0 - z_0}{H} \right)^2 \cdot \exp - \left( \frac{\theta - \theta_0(t)}{\Theta} \right)^2 \cdot \exp - \left( \frac{\phi}{\Phi} \right)^2$$

$$C = \cos 2\pi \left[ t' + (\pi/2 - \theta) \frac{r_0}{\lambda_x} + (r - r_0)/\lambda_z \right]$$

$$\theta_0 = \theta_\infty + V_g t / r_0$$

$r_0$  is the radius of the earth.

$r$ ,  $\theta$ ,  $\phi$  are spherical (earth-centered) polar coordinates.

$N_0(r, \theta, \phi)$  is any electron density model.

Specify:

use perturbation \_\_\_\_\_ (W150 = 1.)

ignore perturbation \_\_\_\_\_ (W150 = 0.)

the height of maximum wave amplitude,  $z_0$  = \_\_\_\_\_ km (W151)

wave-amplitude "scale height,"  $H$  = \_\_\_\_\_ km (W152)

wave perturbation amplitude,  $\delta$  = \_\_\_\_\_ (0 to 1) (W153)

horizontal trace velocity,  $V_x$  = \_\_\_\_\_ km/sec. (W154)  
(needed only if Doppler shift is calculated)

horizontal wavelength,  $\lambda_x$  = \_\_\_\_\_ km (W155)

vertical wavelength,  $\lambda_z$  = \_\_\_\_\_ km (W156)

time in wave periods,  $t'$  = \_\_\_\_\_ (W157)

amplitude "scale distance" in latitude,  $\Theta$  = \_\_\_\_\_ degrees (W159)

amplitude "scale distance" in longitude,  $\Phi$  = \_\_\_\_\_ degrees (W160)

latitude of maximum amplitude at  $t = 0$ ,  $\theta_\infty$  = \_\_\_\_\_ degrees (W158)

southward group velocity,  $V_g$  = \_\_\_\_\_ km/sec (W161)  
(needed even if Doppler shift is not calculated)

```

C      SUBROUTINE WAVE2          WAV2001
      PERTURBATION TO AN ANY ELECTRON DENSITY MODEL   WAV2002
      COMMON /CONST/ PI,PIT2,PID2,DUM(5)           WAV2003
      COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT   WAV2004
      COMMON R(6) /WW/ ID(10),W0,W(400)           WAV2005
      EQUIVALENCE (EARTH,R(2)),(Z0,W(151)),(SH,W(152)),(DELT,A,W(153)),   WAV2006
1  (VSUBX,W(154)),(LAMBDAX,W(155)),(LAMBDAZ,W(156)),(TP,W(157)),   WAV2007
2  (TH00,W(158)),(THC,W(159)),(PHIC,W(160)),(VGX,W(161))           WAV2008
      REAL LAMBDAX,LAMBDAZ           WAV2009
      DATA (MODX(2)=6H WAVE2)        WAV2010
      ENTRY ELECT1                  WAV2011
      IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0..) RETURN  WAV2012
      IF (DELT,A.EQ.0..OR.SH.EQ.0..) RETURN           WAV2013
      H=R(1)-EARTH
      TH0=TH00+LAMBDAX*TP*VGX/VSUBX/EARTH
      EXPR=EXP(-((H-Z0)/SH)**2)           WAV2014
      EXPPTH=EXP(-((R(2)-TH0)/THC)**2)     WAV2015
      EXPPhi=EXP(-((R(3)/PHIC)**2)         WAV2016
      WW=PIT2*(TP+(PID2-R(2))*EARTH/LAMBDAX+H/LAMBDAZ)   WAV2017
      SINW=SIN(WW)                      WAV2018
      COSW=COS(WW)                     WAV2019
      E=DELT,A*EXPR*EXPPTH*EXPPhi       WAV2020
      CONS=1.0+E*COSW                 WAV2021
      PXPR=PXPR*CONS-X*E*2.*((COSW*(H-Z0)/SH)**2+PI/LAMBDAZ*SINW)   WAV2022
      PXPTH=PXPTH*CONS+2.*E*(X*PI*EARTH*R(SINW/LAMBDAX-(R(2)-TH0)/   WAV2023
1  THC)**2*COSW)                   WAV2024
      PXPPH=PXPPH*CONS-X*2.*E*R(3)/PHIC**2*COSW   WAV2025
      PXPT=-PIT2*VSUBX*E/LAMBDAX*SINW+2.0*E*VGX/EARTH*R(COSW*(R(2)-TH0)   WAV2026
1  -LAMBDAX*TP/EARTH)/THC           WAV2027
      X=X*CONS                         WAV2028
      RETURN                           WAV2029
      END                             WAV2030
                                         WAV2031
                                         WAV2032-

```

INPUT PARAMETER FORM FOR SUBROUTINE DOPPLER  
HEIGHT PROFILE OF  $dN/dt$

(A perturbation to an ionospheric electron density model which calculates the time derivative of electron density for calculating Doppler shifts)

First card tells how many profile points in I4 format. The cards following the first card give the height and dN/dt of the profile points one point per card in F8.2, E12.4 format. The heights must be in increasing order. Set W151 = 1.0 to read in a new profile. After the cards are read, DOPPLER will reset W151=0. This subroutine makes an exponential extrapolation down using the bottom 2 points in the profile.

```

SUBROUTINE DOPPLER
C      COMPUTES DN/DT           FROM PROFILES HAVING THE SAME FORM DOPP001
C      AS THOSE USED BY SUBROUTINE TABLE X          DOPP002
C MAKES AN EXPONENTIAL EXTRAPOLATION DOWN USING THE BOTTOM TWO POINTS DOPP003
C      NEEDS SUBROUTINE GAUSEL                      DOPP004
C
C      DIMENSION HPC(250),FN2C(250),ALPHA(250),BETA(250),GAMMA(250), DOPP005
1     DELTA(250),SLOPE(250),MAT(4,5)                DOPP006
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2)        DOPP007
COMMON /XX/ MODX(2),XDUM,PXPR,PXPT4,PXPPH,X,HMAX    DOPP008
COMMON R(6)/HW/ ID(10),W0,W(40)                    DOPP009
EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(READFN,W(151)) DOPP010
REAL MAT,K                                         DOPP011
DATA (MODX(2)=7D0)                                DOPP012
ENTRY ELECT1                                       DOPP013
IF (READFN.EQ.0.) GO TO 10                         DOPP014
READFN=0.                                         DOPP015
READ 1000, NOC,(HPC(I),FN2C(I),I=1,NOC)          DOPP016
1000 FORMAT (I4/(F8.2,E12.4))                     DOPP017
PRINT 1200, (HPC(I),FN2C(I),I=1,NOC)             DOPP018
1200 FORMAT(1H1,14K,6HHEIGHT,4K,16H DN/DT          /(1X,F20.10,E20.1)) DOPP019
A=0.                                                 DOPP020
IF(FN2C(1).NE.0.) A= ALOG(FN2C(2)/FN2C(1))/(HPC(2)-HPC(1)) DOPP021
FN2C(1)=K*FN2C(1)                                DOPP022
FN2C(2)=K*FN2C(2)                                DOPP023
SLOPE(1)=A*FN2C(1)                                DOPP024
SLOPE(NOC)=0.                                     DOPP025
NMAX=1                                             DOPP026
DO 6 I=2,NOC                                     DOPP027
IF (FN2C(I).GT.FN2C(NMAX)) NMAX=I               DOPP028
IF (I.EQ.NOC) GO TO 4                           DOPP029
FN2C(I+1)=K*FN2C(I+1)                          DOPP030
DO 3 J=1,3                                     DOPP031
J=I+J-2                                         DOPP032
MAT(J,1)=1.                                      DOPP033
MAT(J,2)=HPC(M)                                 DOPP034
MAT(J,3)=HPC(M)**2                            DOPP035
3   MAT(J,4)=FN2C(M)                            DOPP036
CALL GAUSEL (MAT,4,3,4,NRANK)                   DOPP037
IF (NRANK.LT.3) GO TO 60                         DOPP038
SLOPE(I)=MAT(2,4)+2.*MAT(3,4)*HPC(I)           DOPP039
4   DO 5 J=1,2                                     DOPP040
M=I+J-2                                         DOPP041
MAT(J,1)=1.                                      DOPP042
MAT(J,2)=HPC(M)                                 DOPP043
MAT(J,3)=HPC(M)**2                            DOPP044
MAT(J,4)=HPC(M)**3                            DOPP045
MAT(J,5)=FN2C(M)                                DOPP046
L=J+2                                           DOPP047
MAT(L,1)=0.                                      DOPP048
MAT(L,2)=1.                                      DOPP049
MAT(L,3)=2.*HPC(M)                            DOPP050
MAT(L,4)=3.*HPC(M)**2                          DOPP051
5   MAT(L,5)=SLOPE(M)                            DOPP052
CALL GAUSEL (MAT,4,4,5,NRANK)                   DOPP053
IF (NRANK.LT.4) GO TO 60                         DOPP054
ALPHA(I)=MAT(1,5)                               DOPP055
BETA(I)=MAT(2,5)                                DOPP056
GAMMA(I)=MAT(3,5)                               DOPP057
6   DELTA(I)=MAT(4,5)                            DOPP058
HMAX=AMAX1(HMAX,HPC(NMAX))                     DOPP059
NH=2                                              DOPP060
10  H=R(1)-EARTH          DOPP061
    F2=F**F                                     DOPP062
    IF (H.GE.HPC(1)) GO TO 12                  DOPP063
11  NH=2                                         DOPP064

```

```

      IF(FN2C(1).EQ.0.) GO TO 50          DOPP066
      X=FN2C(1)*EXP(A*(H-HPC(1)))/F2    DOPP067
      GO TO 50                            DOPP068
12   IF (H.GE.HPC(NOC)) GO TO 18        DOPP069
      NSTEP=1                            DOPP070
      IF (H.LT.HPC(NH-1)) NSTEP=-1       DOPP071
15   IF (HPC(NH-1).LE.H.AND.H.LT.HPC(NH)) GO TO 16  DOPP072
      NH=NH+NSTEP                         DOPP073
      GO TO 15                            DOPP074
16   X=(ALPHA(NH)+H*(BETA(NH)+H*(GAMMA(NH)+H*DELTA(NH))))/F2 DOPP075
      GO TO 50                            DOPP077
18   X=FN2C(NOC)/F2                     DOPP078
50   CONTINUE                           DOPP079
      RETURN                             DOPP080
60   PRINT 6000, I,HPC(I)                DOPP081
6000 FORMAT(4H THE,I4,55HTH POINT IN THE DN/DT PROFILE HAS T DOPP082
1HE HEIGHT,F8.2,40H KM, WHICH IS THE SAME AS ANOTHER POINT.) DOPP083
      CALL EXIT                          DOPP084
      END                                DOPP085

```

## APPENDIX 5. MODELS OF THE EARTH'S MAGNETIC FIELD WITH INPUT PARAMETER FORMS

The following models of the earth's magnetic field are available. The input parameter forms, which describe the model, and the subroutine listings are given on the pages shown.

a.	Constant dip and gyrofrequency (CONSTY)	142
b.	Earth-centered dipole (DIPOLY)	143
c.	Constant dip. Gyrofrequency varies as the inverse cube of the distance from the center of the earth (CUBEY)	144
d.	Spherical harmonic expansion (HARMONY)	145

To add other models of the earth's magnetic field the user must write a subroutine that will calculate the normalized strength and direction of the earth's magnetic field ( $Y$ ,  $Y_r$ ,  $Y_\theta$ ,  $Y_\phi$ ) and their gradients ( $\partial Y / \partial r$ ,  $\partial Y / \partial \theta$ ,  $\partial Y / \partial \varphi$ ,  $\partial Y_r / \partial r$ ,  $\partial Y_r / \partial \theta$ ,  $\partial Y_r / \partial \varphi$ ,  $\partial Y_\theta / \partial r$ ,  $\partial Y_\theta / \partial \theta$ ,  $\partial Y_\theta / \partial \varphi$ ,  $\partial Y_\phi / \partial r$ ,  $\partial Y_\phi / \partial \theta$ ,  $\partial Y_\phi / \partial \varphi$ ) as a function of position in spherical polar coordinates ( $r$ ,  $\theta$ ,  $\varphi$ ). ( $Y = f_H / f$ , where  $f_H$  is the electron gyrofrequency and  $f$  is the wave frequency.)

The restrictions on electron density models also apply to models of the earth's magnetic field. The coordinates  $r$ ,  $\theta$ ,  $\varphi$  refer to the computational coordinate system, which is not necessarily the same as geographic coordinates. W24 and W25 give the geographic latitude and longitude of the north pole of the computational coordinate system.

The input to the subroutine ( $r$ ,  $\theta$ ,  $\varphi$ ) is through blank common. (See Table 3.) The output is through common block /YY/. (See Table 9.) It is useful if the name of the subroutine suggests the model to which it corresponds. It should have an entry point MAGY so that other subroutines in the program can call it. Any parameters needed by the subroutine should be input into W201 through W249 of the W array. (See Table 2.) If the subroutine needs massive amounts of data, these should be read in by the subroutine following the example of subroutine HARMONY.

## INPUT PARAMETER FORM FOR SUBROUTINE CONSTY

An ionospheric model of the earth's magnetic field consisting of constant dip and gyrofrequency

Specify:

gyrofrequency,  $f_H$  = \_\_\_\_\_ MHz (W201)

dip, I = \_\_\_\_\_ degrees (W202)  
radians

The magnetic meridian is defined by the geographic coordinates  
of the north magnetic pole:

latitude = \_\_\_\_\_ radians  
degrees north (W24)

longitude = \_\_\_\_\_ radians  
degrees east (W25)

```
C      SUBROUTINE CONSTY                               CONY001
      CONSTANT DIP AND GYROFREQUENCY                 CONY002
      COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPFH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPRCONY003
      1,PYTPT,PYTTPP,YPH,PYPPR,PYPPT,PYPPP          CONY004
      COMMON /WW/ ID(10),W0,W(400)                   CONY005
      EQUIVALENCE (F,W(6)),(FH,W(201)),(DIP,W(202)) CONY006
      DATA (MODY=6HCONSTY)                           CONY007
      ENTRY MAGY                                     CONY008
      Y=FH/F                                         CONY009
      YR=Y*SIN(DIP)                                CONY010
      YTH=Y*COS(DIP)                                CONY011
      RETURN                                         CONY012
      END                                            CONY013-
```

## INPUT PARAMETER FORM FOR SUBROUTINE DIPOLY

An ionospheric model of the earth's magnetic field consisting of an earth centered dipole

The gyrofrequency is given by:

$$f_H = f_{H_0} \left( \frac{R_0 + h}{R_0} \right)^3 \left( 1 + 3 \cos^2 \lambda \right)^{\frac{1}{2}}$$

The magnetic dip angle, I, is given by

$$\tan I = 2 \cot \lambda$$

h is the height above the ground

$R_0$  is the radius of the earth

$\lambda$  is the geomagnetic colatitude

Specify:

the gyrofrequency at the equator on the ground,  $f_{H_0} = \underline{\hspace{2cm}}$  MHz (W201)

the geographic coordinates of the north magnetic pole  
radians

latitude =  $\underline{\hspace{2cm}}$  degrees north (W24)

radians

longitude =  $\underline{\hspace{2cm}}$  degrees east (W25)

```

SUBROUTINE DIPOLY
COMMON /CONST/ PI,PIT2,PID2,DUM(5)                               DIPO001
COMMON /YY/ MODY,Y,PYPR,PYPHT,PYPFH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPRDIP0003
1,PYTPT,PYTPP,YPH,PYPPR,PYPPT,PYPPP                                DIPO002
COMMON R(6) /WW/ ID(10),W0,W(400)                                 DIPO003
EQUIVALENCE (EARTH,R(2)),(F,W(6)),(FH,W(201))                   DIPO004
DATA (MODY=6HDIPOLY)                                              DIPO005
ENTRY MAGY                                                       DIPO006
SINTH=SIN(R(2))                                                 DIPO007
COSTH=SIN(PID2-R(2))                                             DIPO008
TERM9=SQRT(1.+3.*COSTH**2)                                         DIPO009
T1=FH*(EARTH/R(1))**3/F                                         DIPO010
Y=T1*TERM9                                                       DIPO011
YR= 2.*T1*COSTH                                                 DIPO012
YTH= T1*SINTH                                                   DIPO013
PYRPR=-3.*YR/R(1)                                                DIPO014
PYRPT=-2.*YTH                                                   DIPO015
PYTPR=-3.*YTH/R(1)                                               DIPO016
PYTPT=.5*YR                                                    DIPO017
PYPR=-3.*Y/R(1)                                                 DIPO018
PYPTH=-3.*Y*SINTH*COSTH/TERM9**2                                DIPO019
RETURN                                                       DIPO020
END                                                               DIPO021
                                                               DIPO022
                                                               DIPO023-

```

## INPUT PARAMETER FORM FOR SUBROUTINE CUBEY

A model of the earth's magnetic field consisting of a constant dip and a gyrofrequency which varies as the inverse cube of the distance from the center of the earth

This model has the same height variation as a dipole magnetic field.

The gyrofrequency is given by:

$$f_H = f_{Ho} \left( \frac{a}{r} \right)^3$$

a is the radius of the earth.

r is the distance from the center of the earth.

Specify:

gyrofrequency at the ground,  $f_{Ho} = \underline{\hspace{2cm}}$  MHz (W201)

dip, I =  $\underline{\hspace{2cm}}$  radians  
degrees (W202)

The magnetic meridian is defined by the geographic coordinates of the north magnetic pole:

latitude =  $\underline{\hspace{2cm}}$  radians  
degrees north (W24)  
km

longitude =  $\underline{\hspace{2cm}}$  radians  
degrees east (W25)  
km

```

C          SUBROUTINE CUBEY                      CUBE001
C          CONSTANT DIP.                      CUBE002
C          GYROFREQ DECREASES AS CUBE OF DISTANCE FROM CENTER OF EARTH.   CUBE003
C          THIS MODEL HAS SAME HEIGHT VARIATION AS A DIPOLE FIELD.      CUBE004
COMMON /YY/ MODY,Y,PYPR,PYPYR,PYPTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPRCUBE005
1,PYPTP,PYTPP,YPH,PYPPR,PYPPT,PYPPP                                         CUBE006
COMMON /WW/ ID(10),W0,W(400)                                              CUBE007
EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(FH,W(201)),(DIP,W(202))           CUBE008
DATA(MODY=5HCUBEY)                                                       CUBE009
ENTRY MAGY
Y=(EARTH/R)**3 *FH/F
YR= Y*SIN(DIP)
YTH= Y*COS(DIP)
PYPR=-3.*Y/R
PYRPR=-3.*YR/R
PYTPR=-3.*YTH/R
RETURN
END

```

INPUT PARAMETER FORM FOR SUBROUTINE HARMONY

A model of the earth's magnetic field based on a spherical harmonic expansion

The upward, southerly, and easterly components of the earth's magnetic field are given by:

$$H_r = - \sum_{n=0}^{6} (n+1) \left(\frac{a}{r}\right)^{n+2} \sum_{m=0}^n H_n^m(\theta) \left( g_n^m \cos m\varphi + h_n^m \sin m\varphi \right)$$

$$H_\theta = - \frac{1}{\sin \theta} \sum_{n=0}^{6} \left(\frac{a}{r}\right)^{n+2} \sum_{m=0}^n G_n^m(\theta) \left( g_n^m \cos m\varphi + h_n^m \sin m\varphi \right)$$

$$H_\varphi = \frac{1}{\sin \theta} \sum_{n=0}^{6} \left(\frac{a}{r}\right)^{n+2} \sum_{m=0}^n H_n^m(\theta) \left( h_n^m \cos m\varphi - g_n^m \sin m\varphi \right)$$

where

$a$  is the radius of the earth.

$r, \theta, \varphi$  are spherical (earth-centered) polar coordinates.

$$H_0^0(\theta) = 1$$

$$H_1^0(\theta) = \cos \theta$$

$$H_1^1(\theta) = \sin \theta$$

$$H_{m+1}^m(\theta) = H_m^m(\theta) \cos \theta$$

$$H_{m+1}^{m+1}(\theta) = H_m^m(\theta) \sin \theta$$

$$H_{n+2}^m(\theta) = H_{n+1}^m(\theta) \cos \theta - \frac{(n+m+1)(n-m+1)}{(2n+3)(2n+1)} H_n^m(\theta)$$

$$G_n^m(\theta) = - \frac{d}{d\theta} H_n^m(\theta) \sin \theta$$

$$G_m^m(\theta) = -m H_m^m(\theta) \cos\theta$$

$$G_{n+1}^m(\theta) = -(n+1) H_{n+1}^m(\theta) \cos\theta + \frac{(n+m+1)(n-m+1)}{2n+1} H_n^m(\theta)$$

The recursion formulas for calculating  $H_n^m(\theta)$  and  $G_n^m(\theta)$  are from Eckhouse (1964).

This subroutine uses coefficients  $g_n^m$  and  $h_n^m$  for Gauss normalization. Some coefficients are now being published for Schmidt normalization (e. g. Cain and Sweeney, 1970). The factors  $S_{n,m}$  used for converting the "Schmidt normalized" coefficients to the "Gauss normalized" coefficients are as follows (Cain, et. al., 1968, Chapman and Bartels, 1940):

$$S_{0,0} = -1$$

$$S_{n,0} = S_{n-1,0} \left[ \frac{2n-1}{n} \right]$$

$$S_{n,1} = S_{n,0} \sqrt{\frac{2n}{n+1}}$$

$$S_{n,m} = S_{n,m-1} \sqrt{\frac{(n-m+1)}{n+m}} \quad \text{for } m > 1$$

By convention, the "Gauss normalized" coefficient  $g_1^0$  is positive, whereas the "Schmidt normalized" coefficient  $g_1^0$  is negative. Coefficients based on more recent data on the earth's magnetic field including more satellite data are in the POGO 8/69 model.

Specify below the Gauss coefficients  $g_n^m$  and  $h_n^m$  in gauss.

columns						
2 → 10	11 → 20	21 → 30	31 → 40	41 → 50	51 → 60	61 → 70

1st card     $g_0^0 = \underline{\hspace{2cm}}$

2nd card     $g_1^0 = \underline{\hspace{2cm}}$      $g_1^1 = \underline{\hspace{2cm}}$

3rd card     $g_2^0 = \underline{\hspace{2cm}}$      $g_2^1 = \underline{\hspace{2cm}}$      $g_2^2 = \underline{\hspace{2cm}}$

4th card     $g_3^0 = \underline{\hspace{2cm}}$      $g_3^1 = \underline{\hspace{2cm}}$      $g_3^2 = \underline{\hspace{2cm}}$      $g_3^3 = \underline{\hspace{2cm}}$

5th card     $g_4^0 = \underline{\hspace{2cm}}$      $g_4^1 = \underline{\hspace{2cm}}$      $g_4^2 = \underline{\hspace{2cm}}$      $g_4^3 = \underline{\hspace{2cm}}$      $g_4^4 = \underline{\hspace{2cm}}$

6th card     $g_5^0 = \underline{\hspace{2cm}}$      $g_5^1 = \underline{\hspace{2cm}}$      $g_5^2 = \underline{\hspace{2cm}}$      $g_5^3 = \underline{\hspace{2cm}}$      $g_5^4 = \underline{\hspace{2cm}}$      $g_5^5 = \underline{\hspace{2cm}}$

7th card     $g_6^0 = \underline{\hspace{2cm}}$      $g_6^1 = \underline{\hspace{2cm}}$      $g_6^2 = \underline{\hspace{2cm}}$      $g_6^3 = \underline{\hspace{2cm}}$      $g_6^4 = \underline{\hspace{2cm}}$      $g_6^5 = \underline{\hspace{2cm}}$      $g_6^6 = \underline{\hspace{2cm}}$

8th card  $h_0^0 = \underline{\hspace{2cm}}$   
 9th card  $h_1^0 = \underline{\hspace{2cm}} h_1^1 = \underline{\hspace{2cm}}$   
 10th card  $h_2^0 = \underline{\hspace{2cm}} h_2^1 = \underline{\hspace{2cm}} h_2^2 = \underline{\hspace{2cm}}$   
 11th card  $h_3^0 = \underline{\hspace{2cm}} h_3^1 = \underline{\hspace{2cm}} h_3^2 = \underline{\hspace{2cm}} h_3^3 = \underline{\hspace{2cm}}$   
 12th card  $h_4^0 = \underline{\hspace{2cm}} h_4^1 = \underline{\hspace{2cm}} h_4^2 = \underline{\hspace{2cm}} h_4^3 = \underline{\hspace{2cm}} h_4^4 = \underline{\hspace{2cm}}$   
 13th card  $h_5^0 = \underline{\hspace{2cm}} h_5^1 = \underline{\hspace{2cm}} h_5^2 = \underline{\hspace{2cm}} h_5^3 = \underline{\hspace{2cm}} h_5^4 = \underline{\hspace{2cm}} h_5^5 = \underline{\hspace{2cm}}$   
 14th card  $h_6^0 = \underline{\hspace{2cm}} h_6^1 = \underline{\hspace{2cm}} h_6^2 = \underline{\hspace{2cm}} h_6^3 = \underline{\hspace{2cm}} h_6^4 = \underline{\hspace{2cm}} h_6^5 = \underline{\hspace{2cm}} h_6^6 = \underline{\hspace{2cm}}$

Set W200 = 1. to read in a set of coefficients.

This subroutine represents:

$H_n^m(\theta)$  by  $H(m+1, n+1)$

$G_n^m(\theta)$  by  $G(m+1, n+1)$

$g_n^m$  by  $GG(m+1, n+1)$

$h_n^m$  by  $HH(m+1, n+1)$

```

SUBROUTINE HARMONY
C MODEL OF THE EARTH'S MAGNETIC FIELD BASED ON A HARMONIC ANALYSIS      HARM001
      DIMENSION PHPTH(7,7),PGPTH(7,7),A1(7,7),B1(7,7)                      HARM002
      DIMENSION H(7,7),G(7,7),GG(7,7),HH(7,7),SINP(7),COSP(7)            HARM003
      COMMON /YY/ MODY,Y,PYPR,PYPHT,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTSHARM004
      1,PYTPT,PYTPP,YPH,PYPPR,PYPPT,PYPPP
      COMMON R(6)/WW/ ID(10),W0,W(400)                                     HARM005
      COMMON /CONST/ PI,PIT2,PID2,DUM(5)                                    HARM006
      EQUIVALENCE (THETA,R(2)),(PHI,R(3))                                HARM007
      EQUIVALENCE (EARTH,R(2)),(F,W(6)),(READFH,W(200))                  HARM008
C RATIO OF CHARGE TO MASS FOR ELECTRON                                 HARM009
      DATA(EOM=1.7589E7)                                                 HARM010
      DATA (SET=0.),(H=1.,48(0.)),(G=49(0.)),(PHPTH=49(0.))           HARM011
      1 ,(PGPTH=49(0.)),(MODY=7HHARMONY)                                HARM012
      ENTRY MAGY
      IF(SET) GO TO 2
      DO 1 M=1,7
      DO 1 N=1,7
      B1(M,N)=(N+M-1)*(N-M+1)/(2*N-1.)
      1 A1(M,N)=B1(M,N)/(2*N+1)
      SET=1.

```

```

2 IF(READFH.EQ.0.) GO TO 3 HARM022
READ 2000,GG,HH HARM023
2000 FORMAT(1X,F9.4,6F10.4) HARM024
PRINT 2100,GG HARM025
2100 FORMAT(1H1,10X,1H0,14X,1H1,14X,1H2,14X,1H3,14X,1H4,14X,1H5,14X,1H6) HARM026
1 /9X,7(1HG,14X)/10X,7(1HN,14X)//(1X,7F15.6)) HARM027
PRINT 2200,HH HARM028
2200 FORMAT(// 11X,1H0,14X,1H1,14X,1H2,14X,1H3,14X,1H4,14X,1H5,14X,1H6) HARM029
1 /9X,7(1HH,14X)/10X,7(1HN,14X)//(1X,7F15.6)) HARM030
READFH=0. HARM031
3 COSTHE=COS(THETA) HARM032
SINTHE=SIN(THETA) HARM033
AOR=EARTH/R(1) HARM034
PAORPR=-AOR/R(1) HARM035
CNST2=AOR HARM036
PCNSPR=PAORPR HARM037
FIN1=PFIN1R=PFIN1T=PFIN1P=0. HARM038
FIN2=PFIN2R=PFIN2T=PFIN2P=0. HARM039
FIN3=PFIN3R=PFIN3T=PFIN3P=0. HARM040
DO 4 M=1,7 HARM041
SINP(M)=SIN((M-1)*PHI) HARM042
4 COSP(M)=COS((M-1)*PHI) HARM043
H(1,2)=COSTHE HARM044
H(2,2)=SINTHE HARM045
DO 5 M=1,5 HARM046
H(M+1,M+2)=COSTHE*H(M+1,M+1) HARM047
H(M+2,M+2)=SINTHE*H(M+1,M+1) HARM048
DO 5 N=M,5 HARM049
5 H(M,N+2)=COSTHE*H(M,N+1)-A1(M,N)*H(M,N) HARM050
DO 6 M=1,6 HARM051
G(M+1,M+1)=-M*COSTHE*H(M+1,M+1) HARM052
PHPTH(M+1,M+1)=-G(M+1,M+1)/SINTHE HARM053
PGPTH(M+1,M+1)=M*SINTHE*H(M+1,M+1)-M*COSTHE*PHPTH(M+1,M+1) HARM054
DO 6 N=M,6 HARM055
G(M,N+1)=-N*COSTHE*H(M,N+1)+B1(M,N)*H(M,N) HARM056
PHPTH(M,N+1)=-G(M,N+1)/SINTHE HARM057
6 PGPTH(M,N+1)=N*SINTHE*H(M,N+1)-N*COSTHE*PHPTH(M,N+1)+B1(M,N)*PHPTH(M,N+1) HARM058
1 (M,N)
DO 8 N=1,7 HARM059
CR=PCRPTH=PCRPPH=0. HARM060
CTH=PCTHPT=PCTHPP=0. HARM061
CPH=PCPHPT=PCPHPP=0. HARM062
DO 7 M=1,N HARM063
TEMP1=GG(M,N)*COSP(M)+HH(M,N)*SINP(M) HARM064
TEMP2=(M-1)*(HH(M,N)*COSP(M)-GG(M,N)*SINP(M)) HARM065
CR =CR +H(M,N)*TEMP1 HARM066
PCRPTH=PCRPTH+PHPTH(M,N)*TEMP1 HARM067
PCRPPH=PCRPPH+H(M,N)*TEMP2 HARM068
CTH =CTH +G(M,N)*TEMP1 HARM069
PCTHPT=PCTHPT+PGPTH(M,N)*TEMP1 HARM070
PCTHPP=PCTHPP+G(M,N)*TEMP2 HARM071
CPH =CPH +H(M,N)*TEMP2 HARM072
PCPHPT=PCPHPT+PHPTH(M,N)*TEMP2 HARM073
7 PCPHPP=PCPHPP-H(M,N)*(M-1)**2*TEMP1 HARM074
CNST2=CNST2*AOR HARM075
CNST2=CNST2*PAORPR+AOR*PCNSPR HARM076
FIN1=FIN1+N*CNST2*CR HARM077
PFIN1R=PFIN1R+N*PCNSPR*CR HARM078
PFIN1T=PFIN1T+N*CNST2*PCRPTH HARM079
PFIN1P=PFIN1P+N*CNST2*PCRPPH HARM080
FIN2=FIN2+CNST2*CTH HARM081
PFIN2R=PFIN2R+PCNSPR*CTH HARM082
PFIN2T=PFIN2T+CNST2*PCTHPT HARM083
PFIN2P=PFIN2P+CNST2*PCTHPP HARM084
FIN3=FIN3+CNST2*CPH HARM085

```

```

PFIN3R=PFIN3R+PCNSPR*CPH          HARM087
PFIN3T=PFIN3T+CNST2*PCPHPT       HARM088
8 PFIN3P=PFIN3P+CNST2*PCPHP>     HARM089
HTHETA=-FIN2/SINTHE               HARM090
HPHI=FIN3/SINTHE                 HARM091
C***** CONVERT FROM MAG FIELD IN GAUSS TO GYROFREQ IN MHZ   HARM092
CONST=-EOM/PIT2*1.E-6/F           HARM093
YR=-CONST*FIN1                   HARM094
YTH=CONST*HTHETA                HARM095
YPH=CONST*HPHI                  HARM096
Y=SQRT(YR**2+YTH**2+YPH**2)      HARM097
PYRPR=-CONST*PFIN1R              HARM098
PYTPR=-CONST*PFIN2R/SINTHE       HARM099
PYPPR=CONST*PFIN3R/SINTHE        HARM100
PYPR=(YR*PYRPR+YTH*PYTPR+YPH*PYPPR)/Y    HARM101
PYRPT=-CONST*PFIN1T              HARM102
PYTPT=-CONST*(PFIN2T/SINTHE+HTHETA*COSTHE/SINTHE)  HARM103
PYPPT=CONST*(PFIN3T/SINTHE-HP4I*COSTHE/SINTHE)      HARM104
PYPTH=(YR*PYRPT+YTH*PYTPT+YPH*PYPPT)/Y            HARM105
PYRPP=-CONST*PFIN1P              HARM106
PYTPP=-CONST*PFIN2P/SINTHE       HARM107
PYPPP=CONST*PFIN3P/SINTHE        HARM108
PYPPH=(YR*PYRPP+YTH*PYTPP+YPH*PYPPP)/Y            HARM109
RETURN                           HARM110
C COEFFICIENTS IN GAUSSIAN UNITS FROM JONES AND MELOTTE (1953). HARM111
C THE FOLLOWING 14 CARDS CAN BE USED AS DATA CARDS FOR THIS SUBROUTINE HARM112
C 0.                                HARM113
C .3039   .0218                   HARM114
C .0176   -.0509     -.0135       HARM115
C -.0255   .0515     -.0236     -.0074       HARM116
C -.0393   -.0397     -.0238     .0087     -.0018       HARM117
C .0293   -.0329     -.0130     .0031     .0034     .0005       HARM118
C -.0211   -.0073     -.0007     .0210     .0017     -.0004     .0006       HARM119
C 0.                                HARM120
C     -.0555                   HARM121
C     .0260     -.0044       HARM122
C     .0190     -.0033     -.0001       HARM123
C     -.0139     .0076     .0019     .0010       HARM124
C     .0057     -.0018     .0009     .0032     -.0004       HARM125
C     -.0026     -.0204     .0018     .0009     .0004     .0002       HARM126
C THE FOLLOWING SET OF GAUSS NORMALIZED COEFFICIENTS WERE CONVERTED HARM127
C FROM THE SCHMIDT NORMALIZED COEFFICIENTS CALCULATED BY LINEARLY HARM128
C EXTRAPOLATING TO EPOCH 1974 THE COEFFICIENTS PUBLISHED FOR EPOCH HARM129
C 1960 BY CAIN AND SWEENEY (1970). (USES EARTH RADIUS = 6371.2) HARM130
C .000000                           HARM131
C+.300953 +.020298               HARM132
C+.028106 -.05214    -.014435     HARM133
C-.0308   +.06560     -.025252     -.006952     HARM134
C-.041243 -.043956     -.016897     +.008021     -.002525     HARM135
C+.014742 -.037078     -.018906     +.002819     +.003656     +.000036     HARM136
C-.006713 -.012234     -.004364     +.02137     +.001593     -.000072     +.000068     HARM137
C .000000                           HARM138
C .000000     -.057886       HARM139
C .000000     +.035942     +.001129       HARM140
C .000000     +.011084     -.004421     +.001180       HARM141
C .000000     -.010299     +.008794     -.000086     +.002256       HARM142
C .000000     -.003849     -.012615     +.007845     +.002207     -.000328       HARM143
C .000000     +.003157     -.012670     -.009281     +.002286     -.000135     +.000243       HARM144
END                               HARM145-

```



## APPENDIX 6. COLLISION FREQUENCY MODELS WITH INPUT PARAMETER FORMS

The following collision frequency models are available. The input parameter forms, which describe the model, and the subroutine listings are given on the pages shown.

a.	Tabular profiles (TABLEZ)	152
b.	Constant collision frequency (CONSTZ)	155
c.	Exponential profile (EXPZ)	156
d.	Combination of two exponential profiles (EXPZ2)	157

To add other collision frequency models the user must write a subroutine that will calculate the normalized collision frequency ( $Z$ ) and its gradient ( $\partial Z / \partial r$ ,  $\partial Z / \partial \theta$ ,  $\partial Z / \partial \phi$ ) as a function of position in spherical polar coordinates ( $r, \theta, \phi$ ). ( $Z = v / 2\pi f$ , where  $v$  is the collision frequency between electrons and neutral air molecules and  $f$  is the wave frequency. If the Sen-Wyller formula for refractive index is used, then  $Z = v_m / 2\pi f$ , where  $v_m$  is the mean collision frequency.)

The restrictions on electron density models also apply to collision frequency models. The coordinates  $r, \theta, \phi$  refer to the computational coordinates system, which may not be the same as geographic coordinates. In particular, they are geomagnetic coordinates when the earth-centered dipole model of the earth's magnetic field is used.

The input to the subroutine ( $r, \theta, \phi$ ) is through blank common. (See Table 3.) The output is through common block /ZZ/. (See Table 10.) It is useful if the name of the subroutine suggests the model to which it corresponds. It should have an entry point COLFRZ so that other subroutines in the program can call it. Any parameter needed by the subroutine should be input into W251 through W299 of the W array. (See Table 2.) If the model needs massive amounts of data, these should be read in by the subroutine following the example of subroutine TABLEZ.

INPUT PARAMETER FORM FOR SUBROUTINE TABLEZ

## IONOSPHERIC COLLISION FREQUENCY PROFILE

The first card tells how many profile points in I4 format. The cards following the first card give the height and collision frequency of the profile points one point per card in F8.2, E12.4 format. The heights must be in increasing order. Set W250 = 1.0 to read in a new profile. After the cards are read, TABLEZ will reset W250 = 0.0. This subroutine makes an exponential extrapolation down using the bottom 2 points in the profile.

```

C      SUBROUTINE TABLEZ                               TABZ001
C      CALCULATES COLLISION FREQUENCY AND ITS GRADIENT FROM PROFILES   TABZ002
C      HAVING THE SAME FORM AS THOSE USED BY CROFTS RAY TRACING PROGRAM   TABZ003
C MAKES AN EXPONENTIAL EXTRAPOLATION DOWN USING THE BOTTOM TWO POINTS   TABZ004
C      NEEDS SJROUTINE GAUSEL                                         TABZ005
C
C      DIMENSION HPC(100),FN2C(100),ALPHA(100),BETA(100),                TABZ006
1     GAMMA(100),DELTA(100),MAT(4,5),SLOPE(100)                         TABZ007
COMMON /CONST/ PI,PIT2,PID2,DJM(5)                                       TABZ008
COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH                                     TABZ009
COMMON R(6) /WWS/ ID(10),W0,W(400)                                      TABZ010
EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(READNU,W(250))                   TABZ011
REAL MAT
DATA (MODZ=6HTABLEZ)
ENTRY COLFRZ
IF (.NOT.READNU) GO TO 10
READNU=0.
READ 2, NOC,(HPC(I),FN2C(I),I=1,NOC)                                    TABZ012
2 FORMAT(I4/(F8.2,E12.4))
PRINT 1200, (HPC(I),FN2C(I), I=1,NOC)                                     TABZ013
1200 FORMAT(1H1,14X,6HHEIGHT,4X,20HCOLLISION FREQUENCY                  TABZ014
1(1X,F20.10,E20.10))
A=0.
IF(FN2C(1).NE.0.) A= ALOG(FN2C(2)/FN2C(1))/(HPC(2)-HPC(1))          TABZ015
FN2C(1)=FN2C(1)/PIT2*1.E-6                                              TABZ016
FN2C(2)=FN2C(2)/PIT2*1.E-6                                              TABZ017
SLOPE(1)=A*FN2C(1)                                                       TABZ018
SLOPE(NOC)=0.                                                               TABZ019
DO 5 I=2,NOC
IF(I.EQ.NOC) GO TO 6
FN2C(I+1)= FN2C(I+1)/PIT2*1.E-6                                         TABZ020
DO 3 J=1,3
M=I+J-2
MAT(J,1)=1.                                                               TABZ021
MAT(J,2)=HPC(M)                                                       TABZ022
MAT(J,3)=HPC(M)**2                                                       TABZ023
3 MAT(J,4)=FN2C(M)                                                       TABZ024
CALL GAUSEL (MAT,4,3,4,NRANK)                                              TABZ025
IF (NRANK.LT.3) GO TO 20
SLOPE(I)=MAT(2,4)+2.*MAT(3,4)*HPC(I)                                     TABZ026
5 CONTINUE
DO 4 J=1,2
M=I+J-2
MAT(J,1)=1.                                                               TABZ027
MAT(J,2)=HPC(M)                                                       TABZ028
MAT(J,3)=HPC(M)**2                                                       TABZ029
MAT(J,4)=HPC(M)**3                                                       TABZ030
MAT(J,5)=FN2C(M)                                                       TABZ031
L=J+2
MAT(L,1)=0.
MAT(L,2)=1.                                                               TABZ032
MAT(L,3)=2.*HPC(M)                                                       TABZ033
MAT(L,4)=3.*HPC(M)**2                                                     TABZ034
4 MAT(L,5)=SLOPE(M)                                                       TABZ035
CALL GAUSEL (MAT,4,4,5,NRANK)                                              TABZ036
IF (NRANK.LT.4) GO TO 20
ALPHA(I)=MAT(1,5)                                                       TABZ037
BETA(I)=MAT(2,5)                                                       TABZ038
GAMMA(I)=MAT(3,5)                                                       TABZ039
5 DELTA(I)=MAT(4,5)                                                       TABZ040
JUP=2
10 H=R(1)-EARTH
IF (H.GE.HPC(1)) GO TO 12
11 JUP=2
Z=FN2C(1)*EXP(A*(H-HPC(1)))/F
PZPR=A*Z

```

```

      RETURN                               TABZ066
12 IF (H.GE.HPC(NOC)) GO TO 18          TABZ067
      NSTEP=1                             TABZ068
      IF (H.LT.HPC(JUP-1)) NSTEP=-1       TABZ069
15 IF (HPC(JUP-1).GT.H.OR.H.GE.HPC(JUP)) GO TO 16   TABZ070
      Z=(ALPHA(JUP)+H*(BETA(JUP)+H*(GAMMA(JUP)+H*DELTA(JUP))))/F TABZ071
      PZPR=(BETA(JUP)+H*(2.*GAMMA(JUP)+H*3.*DELTA(JUP)))/F TABZ072
      RETURN                               TABZ073
16 JUP=JUP+NSTEP                         TABZ074
      IF (JUP.LT.2) GO TO 11              TABZ075
      IF (JUP.LT.NOC) GO TO 15           TABZ076
18 JUP=NOC                               TABZ077
      Z=FN2C(NOC)/F                     TABZ078
      PZPR=0.                            TABZ079
      RETURN                               TABZ080
20 PRINT 21, I,HPC(I)                   TABZ081
21 FORMAT(4H THE,I4+58HTH POINT IN THE COLLISION FREQUENCY PROFILE HATABZ082
      IS THE HEIGHT,F8.2,40H KM, WHICH IS THE SAME AS ANOTHER POINT.) TABZ083
      CALL EXIT                           TABZ084
      END                                TABZ085

```

## INPUT PARAMETER FORM FOR SUBROUTINE CONSTZ

An ionospheric collision frequency model consisting of a constant collision frequency

$v = 0$  for  $h \leq h_{\min}$

$v = v_0$  for  $h > h_{\min}$

Specify:

$v_0$  = \_\_\_\_\_ collisions per second (W251)

$h_{\min}$  = \_\_\_\_\_ km (W252)

```
C      SUBROUTINE CONSTZ                               CONZ001
      CONSTANT COLLISION FREQUENCY                  CONZ002
      COMMON /CONST/ PI,PIT2,PID2,DUM(5)           CONZ003
      COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH        CONZ004
      COMMON R(6) /WW/ ID(10),W0,W(400)          CONZ005
      EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(NU,W(251)),(HMIN,W(252)) CONZ006
      REAL NU                                         CONZ007
      DATA (MODZ=6HCONSTZ)                         CONZ008
      ENTRY COLFRZ                                CONZ009
      H=R(1)-EARTH
      Z=0.
      IF (H.GT.HMIN) Z=NU/(PIT2*F)*1.E-6
      RETURN
      END                                            CONZ014-
```

## INPUT PARAMETER FORM FOR SUBROUTINE EXPZ

An ionospheric collision frequency model consisting of an exponential profile

$$\nu = \nu_0 e^{-a(h-h_0)}$$

$h$  is the height above the ground

Specify:

The collision frequency at the height  $h_0$ ,  $\nu_0$  = \_\_\_\_\_ collisions per second (W251)

The reference height,  $h_0$  = \_\_\_\_\_ km (W252)

The exponential decrease of  $\nu$  with height,  $a$  = \_\_\_\_\_  $\text{km}^{-1}$  (W253)

```

C      SUBROUTINE EXPZ          EXPZ001
      EXPONENTIAL COLLISION FREQUENCY MODEL    EXPZ002
      COMMON /CONST/ PI,PIT2,PID2,DUM(5)        EXPZ003
      COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH      EXPZ004
      COMMON R(6) /WW/ ID(10),W0,W(400)        EXPZ005
      REAL NU,NU0                                EXPZ006
      EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(NU0,W(251)),(H0,W(252)),
1 (A,W(253))
      DATA (MODZ=6H EXPZ )
      ENTRY COLFRZ
      H=R(1)-EARTH
      NU=NU0/EXP (A*(H-H0))
      Z=NU/(PIT2*F*1.E6)
      PZPR  ==A*Z
      RETURN
      END

```

## INPUT PARAMETER FORM FOR SUBROUTINE EXPZ2

An ionospheric collision frequency model consisting of a combination of two exponential profiles

$$\nu = \nu_1 e^{-a_1(h-h_1)} + \nu_2 e^{-a_2(h-h_2)}$$

where  $h$  is the height above the ground.

Specify for the first exponential:

Collision frequency at height  $h_1$ ,  $\nu_1 = \underline{\hspace{2cm}}$  collisions per second (W251)

Reference height,  $h_1 = \underline{\hspace{2cm}}$  km (W252)

Exponential decrease of  $\nu$  with height,  $a_1 = \underline{\hspace{2cm}}$   $\text{km}^{-1}$  (W253)

Specify for the second exponential:

Collision frequency at height  $h_2$ ,  $\nu_2 = \underline{\hspace{2cm}}$  collisions per second (W254)

Reference height,  $h_2 = \underline{\hspace{2cm}}$  km (W255)

Exponential decrease of  $\nu$  with height,  $a_2 = \underline{\hspace{2cm}}$   $\text{km}^{-1}$  (W256)

```

C      SUBROUTINE EXPZ2                               XPZ2001
      COLLISION FREQUENCY PROFILE FROM TWO EXPONENTIALS   XPZ2002
      COMMON /CONST/ PI,PIT2,PID2,DUM(5)                 XPZ2003
      COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH               XPZ2004
      COMMON R(6) /WW/ ID(10),W0,W(400)                XPZ2005
      EQUIVALENCE (EARTH,R(2)),(F,W(6)),(NU1,W(251)),(H1,W(252)),
1 (A1,W(253)),(NU2,W(254)),(H2,W(255)),(A2,W(256))  XPZ2006
      RREAL NU1,NU2                                     XPZ2007
      DATA (MODZ=6H EXPZ2)                            XPZ2008
      FENTRY COLFRZ                                 XPZ2009
      H=R(1)-EARTH
      EXP1= NU1* EXP(-A1*(H-H1))                   XPZ2010
      EXP2= NU2* EXP(-A2*(H-H2))                   XPZ2011
      Z=(EXP1+EXP2)/(PIT2*F*1.E6)                  XPZ2012
      PZPR=(-A1*F*XP1-A2*EXP2)/(PIT2*F*1.F6)    XPZ2013
      RETURN
      END                                              XPZ2014
                                                       XPZ2015
                                                       XPZ2016
                                                       XPZ2017-
```



## APPENDIX 7. CDC 250 PLOT PACKAGE

This appendix describes the plotting routines used by the Three-Dimensional Ray Tracing Program. The information was taken from "User's Guide to Cathode Ray Plotter Subroutines," ESSA Technical Memorandum ERLTM-ORSS 5, by L. David Lewis, January, 1970, and is printed with the permission of the author.

If you have access to a plotter, you may obtain plots by converting the following plotting commands to comparable commands on your system.

The CDC-250 Microfilm Recorder, under control of the NOAA Boulder CDC-3800 computer, plots data on the face of a high resolution cathode ray tube, which is photographed onto standard sized perforated, 35 mm film.

The plotting area, called a frame, is a square. Plotting positions are described in rectangular coordinates. Coordinate values are integers in the range 0 - 1023; (0, 0) is the "lower left hand corner".

Plotting specifications are transmitted to the plot routines via the following COMMON.

COMMON /DD/ IN, IOR, IT, IS, IC, ICC, IX, IY

The usage of each of the eight variables is listed below, followed by an explanation of the subroutine calls.

IN	Intensity.
	IN=0 specifies normal intensity.
	IN=1 specifies high intensity.
IOR	Orientation.
	IOR=0 specifies upright orientation.
	IOR=1 specifies rotated orientation (90° counter-clockwise).
IT	Italics (Font).
	IT=0 specifies non-Italic (Roman) symbols.
	IT=1 specifies Italic symbols.

IS	Symbol size. IS=0 specifies miniature size. IS=1 specifies small size. IS=2 specifies medium size. IS=3 specifies large size.
IC	Symbol case. IC=0 specifies upper case. IC=1 specifies lower case.
ICC	Character code, 0-63 (R1 format). ICC and IC together specify the symbol plotted.
IX	X-coordinate, 0-1023.
IY	Y-coordinate, 0-1023.

CALL DDINIT (N, ID) is required to initialize the plotting process.

CALL DDBP defines a vector origin at position IX, IY.

CALL DDVC plots a vector (straight line), with intensity IN, from the vector origin defined by the previous DDBP or DDVC call, to the vector end position at IX, IY. A single call to DDBP followed by successive calls to DDVC (with changing IX and IY) plots connected vectors.

CALL DDTAB initializes tabular plotting.

CALL DDTEXT (N, NT) plots a given array in a tabular mode, after initiating tabular plotting via DDTAB, as described above. NT is an array of length N, containing "text" for tabular plotting. Text consists of character codes, packed 8 per word (A8 Format). Text characters are plotted as tabular symbols until the command character # (octal code 14, card code 4, 8, or the alphabetic shift counterpart of the = on the keypunch) occurs. The command character is not plotted. DDTEXT interprets the next character as a command; and after the command is processed, tabular plotting resumes until # is again encountered.  
# . means end of text: DDTEXT returns to the calling routine.

CALL DDFR causes a frame advance operation. Plotting on the current frame is completed, and the film advances to the next frame.

## APPENDIX 8. SAMPLE CASE

A sample case is included with the description of the program for three reasons. First, it demonstrates the use of the program. Second, it illustrates the three types of output available (printout, punched cards, and ray path plots). Finally, it serves as a test case to verify that the user's copy of the program is running correctly. This last point is especially important if the user has had to make many modifications in converting the program to run on a computer other than a CDC 3800.

Although the ionospheric models in the sample case demonstrate the use of the program, they don't give realistic absorption for the radio waves. The absorption in the sample case is too low for two reasons. First, although the Chapman layer has a realistic electron density for the F region, it has much too low an electron density for the D region, where most of the absorption occurs. Second, the collision frequency profile in the sample case is designed for use with the Sen-Wyller formula for refractive index rather than the Appleton-Hartree formula used in the sample case. Multiplying the collision frequency profile in the sample case by 2.5 gives an effective collision frequency profile for use with the Appleton-Hartree formula that will give nearly the correct absorption for HF radio waves (Davies, 1965, p. 89).

### Appendix 8a. Input Parameter Forms for the Sample Case

Filled-out input parameter forms are included to describe the sample case (i. e., show what ray paths are requested for which ionospheric models and what type of output is wanted). Furthermore, comparing them with Appendix 8b illustrates the relationship between the forms and the input data cards.

## INPUT PARAMETER FORM FOR THREE-DIMENSIONAL RAY PATHS

Name \_\_\_\_\_ Project No. \_\_\_\_\_ Date \_\_\_\_\_

Ionospheric ID (3 characters) X01Title (75 characters) Test Case

Models:	Electron density	<u>CHAPX</u>
Perturbation		<u>WAVE</u>
Magnetic field		<u>DIPOLY</u>
Ordinary		(W1 = + 1.)
Extraordinary	<input checked="" type="checkbox"/>	(W1 = - 1.)
Collision frequency		<u>EXPZ2</u>
Transmitter:	Height	<u>0</u> km, nautical miles, feet (W3)
	Latitude	<u>40</u> rad, <u>deg</u> km (W4)
	Longitude	<u>-105</u> rad, <u>deg</u> km (W5)
	Frequency, initial	<u>1</u> MHz (W7)
	final	(W8)
	step	(W9)
	Azimuth angle, initial	<u>45</u> rad, <u>deg</u> clockwise of north (W11)
	final	(W12)
	step	(W13)
	Elevation angle, initial	<u>0</u> rad, <u>deg</u> (W15)
	final	<u>90</u> (W16)
	step	<u>15</u> (W17)
Receiver:	Height	<u>200</u> km, nautical miles, feet (W20)
Penetrating rays:	Wanted	<input checked="" type="checkbox"/> (W21 = 0.)
	Not wanted	<input type="checkbox"/> (W21 = 1.)
Maximum number of hops		<u>3</u> (W22)
Maximum number of steps per hop		<u>1000</u> (W23)
Maximum allowable error per step		<u>10<sup>-4</sup></u> (W42)
Additional calculations:		= 1. to integrate = 2. to integrate and print
Phase path	<u>2</u>	(W57)
Absorption	<u>2</u>	(W58)
Doppler shift		<u>1</u> (W59)
Path length		<u>1</u> (W60)
Other		<u>1</u>
Printout:	Every <u>5</u> steps of the ray trace (W71)	
Punched cards (raysets):	<input checked="" type="checkbox"/>	(W72 = 1.)

INPUT PARAMETER FORM FOR PLOTTING THE PROJECTION  
OF THE RAY PATH ON A VERTICAL PLANE

Coordinates of the left edge of the graph:

Latitude = 40. rad  
(deg) north (W83)  
km

Longitude = -105. rad  
(deg) east (W84)  
km

Coordinates of the right edge of the graph:

Latitude = 52.12 rad  
(deg) north (W85)  
km

Longitude = -81.8 rad  
(deg) east (W86)  
km

Height above the ground of the bottom of the graph = 0. km (W88)

Distance between tic marks = 100. rad  
deg (W87)  
km

(W81 = 1.)

INPUT PARAMETER FORM FOR PLOTTING THE PROJECTION  
OF THE RAY PATH ON THE GROUND

Coordinates of the left edge of the graph:

Latitude = 40. rad  
 deg   north (W83)  
km

Longitude = -105. rad  
 deg   east (W84)  
km

Coordinates of the right edge of the graph:

Latitude = 52.12 rad  
 deg   north (W85)  
km

Longitude = -81.8 rad  
 deg   east (W86)  
km

Factor to expand lateral deviation scale by = 200. (W82)

Distance between tic marks on range scale = 100. rad  
deg   (W87)  
 km

(W81 = 2.)

## INPUT PARAMETER FORM FOR SUBROUTINE CHAPX

An ionospheric electron density model consisting of a Chapman layer with tilts, ripples, and gradients

$$\begin{aligned}
 f_N^2 &= f_c^2 \exp\left(\alpha(1-z-e^{-z})\right) \\
 z &= \frac{h - h_{max}}{H} \\
 f_c^2 &= f_{co}^2 \left(1 + A \sin\left(2\pi\left(\theta - \frac{\pi}{2}\right)/B\right) + C\left(\theta - \frac{\pi}{2}\right)\right) \\
 h_{max} &= h_{max_0} + E\left(\theta - \frac{\pi}{2}\right) R_0
 \end{aligned}$$

$f_N$  is the plasma frequency

$h$  is the height above the ground

$R_0$  is the radius of the earth in km

and  $\theta$  is the colatitude in radians.

Specify:

Critical frequency at the equator,  $f_{co} = \underline{6.5}$  MHz (W101)

Height of the maximum electron density at the equator,  $h_{max_0} = \underline{300}$  km (W102)

Scale height,  $H = \underline{62.}$  km (W103)

$\alpha = \underline{0.5}$  (W104, 0.5 for an  $\alpha$  Chapman layer, 1.0 for a  
8 Chapman layer)

Amplitude of periodic variation of  $f_c^2$  with latitude,  $A = \underline{0.}$  (W105)

Period of variation of  $f_c^2$  with latitude,  $B = \underline{0.}$  rad  
deg (W106)  
km

Coefficient of linear variation of  $f_c^2$  with latitude,  $C = \underline{0.}$  rad $^{-1}$  (W107)

Tilt of the layer,  $E = \underline{0.}$  rad (W108)  
deg

INPUT PARAMETER FORM FOR SUBROUTINE WAVE

A perturbation to an ionospheric electron density model consisting of a "gravity-wave" irregularity traveling from north pole to south pole

$$N = N_0(1 + \Delta)$$

$$\Delta = \delta \exp - [(R - R_0 - z_0)/H]^2 .$$

$$\cos 2\pi \left[ t' + (\pi/2 - \theta) \frac{R_0}{\lambda_x} + (R - R_0)/\lambda_z \right]$$

$$\frac{\partial N}{\partial t} = \frac{-2\pi}{\lambda_x} V_x N_0 \delta \exp - [(R - R_0 - z_0)/H]^2 .$$

$$\sin 2\pi \left[ t' + (\pi/2 - \theta) \frac{R_0}{\lambda_x} + (R - R_0)/\lambda_z \right]$$

$R_0$  is the radius of the earth.

$R$ ,  $\theta$ ,  $\varphi$  are the spherical (earth-centered) polar coordinates  
 $(\Delta$  is independent of  $\varphi$ ).

$N_0(R, \theta, \varphi)$  is any electron density model.

Specify:

the height of maximum wave amplitude,  $z_0 = \underline{250}$  km (W151)

wave-amplitude "scale height,"  $H = \underline{100}$  km (W152)

wave perturbation amplitude,  $\delta = \underline{0.1}$  [0. to 1.] (W153)

horizontal trace velocity,  $V_x = \underline{-}$  km/sec (W154)  
 (needed only if Doppler shift is calculated)

horizontal wavelength,  $\lambda_x = \underline{100}$  km (W155)

vertical wavelength,  $\lambda_z = \underline{100}$  km (W156)

time in wave periods,  $t' = \underline{0.}$  [0. to 1.] (W157)

## INPUT PARAMETER FORM FOR SUBROUTINE DIPOLY

An ionospheric model of the earth's magnetic field consisting of an earth centered dipole

The gyrofrequency is given by:

$$f_H = f_{H_0} \left( \frac{R_0 + h}{R_0} \right)^3 \left( 1 + 3 \cos^2 \lambda \right)^{\frac{1}{2}}$$

The magnetic dip angle, I, is given by

$$\tan I = 2 \cot \lambda$$

h is the height above the ground

$R_0$  is the radius of the earth

$\lambda$  is the geomagnetic colatitude

Specify:

the gyrofrequency at the equator on the ground,  $f_{H_0} = \underline{0.8}$  MHz (W201)

the geographic coordinates of the north magnetic pole

latitude = 78.5 radians degrees north (W24)

longitude = 291. radians degrees east (W25)

## INPUT PARAMETER FORM FOR SUBROUTINE EXPZ 2

An ionospheric collision frequency model consisting of a double exponential profile

$$\nu = \nu_1 e^{-a_1(h-h_1)} + \nu_2 e^{-a_2(h-h_2)}$$

where  $h$  is the height above the ground.

Specify for the first exponential:

Collision frequency at height  $h_1$ ,  $\nu_1 = \underline{3.65 \times 10^4}$  collisions per second (W251)

Reference height,  $h_1 = \underline{100}$  km (W252)

Exponential decrease of  $\nu$  with height,  $a_1 = \underline{0.148}$   $\text{km}^{-1}$  (W253)

Specify for the second exponential:

Collision frequency at height  $h_2$ ,  $\nu_2 = \underline{30}$  collisions per second (W254)

Reference height,  $h_2 = \underline{140}$  km (W255)

Exponential decrease of  $\nu$  with height,  $a_2 = \underline{0.0183}$   $\text{km}^{-1}$  (W256)

## Appendix 8b. Listing of Input Cards for the Sample Case

```

X01 TEST CASE
1 0.          OF DUPLICATE W CARDS, THE LAST ONE DOMINATES
1 -1.          EXTRAORDINARY RAY
3 0.          TRANSMITTER HEIGHT, KM
4 40.          1  TRANSMITTER LATITUDE, DEG NORTH
5 -105.        1  TRANSMITTER LONGITUDE, DEG EAST
7 6.0          INITIAL FREQUENCY, MC/S
9 0.          DONT STEP FREQUENCY
11 45.0        1  INITIAL AZIMUTH ANGLE, DEGS CLOCKWISE FROM NORTH POLE
13 0.          DONT STEP AZIMUTH ANGLE
15 0.          1  INITIAL ELEVATION ANGLE, DEG
16 90.0        1  FINAL ELEVATION ANGLE, DEG
17 15.0        1  STEP IN ELEVATION ANGLE, DEG
20 200.        RECEIVER HEIGHT ABOVE THE EARTH, KM
22 3.          NUMBER OF HOPS
57 2.          INTEGRATE AND PRINT PHASE PATH
58 2.          INTEGRATE AND PRINT ABSORPTION
71 5.0          NUMBER OF STEPS FOR EACH PRINTING
72 1.          PUNCH RAYSETS
81 1.          PLOT PROJECTION OF RAY PATH ON A VERTICAL PLANE
83 40.0        1  LEFT LATITUDE OF PLOT, DEG
84 -105.        1  LEFT LONGITUDE OF PLOT, DEG
85 52.12        1  RIGHT LATITUDE OF PLOT, DEG
86 -81.8        1  RIGHT LONGITUDE OF PLOT, DEG
87 100.0        1  DISTANCE BETWEEN TIC MARKS, KM
101 6.5          CRITICAL FREQUENCY, MC/S
102 300.0        HMAX, KM
103 62.          SCALE HEIGHT, KM
104 0.5          ALPHA CHAPMAN LAYER
150 1.          CALL PERTURBATION SUBROUTINE
151 250.          Z0, KM
152 100.          SH, SCALE HEIGHT, KM
153 0.1          DELTA
155 100.          LAMBDAx, HORIZONTAL WAVELENGTH, KM
156 100.          LAMBDAz, VERTICAL WAVELENGTH, KM
201 0.8          GYROFREQUENCY ON THE GROUND AT THE EQUATOR, MHZ
24 78.5          1  ACCEPTED STANDARD LAT. OF NORTH MAGNETIC POLE, DEG NORTH
25 291.          1  ACCEPTED STANDARD LONG. OF NORTH MAGNETIC POLE, DEG EAST
251 3.65          E4  COLLISION FREQUENCY AT H1, /SEC
252 100.0        H1, REFERENCE HEIGHT, KM
253 .148         A1, EXPONENTIAL DECREASE OF NU WITH HEIGHT, /KM
254 30.           COLLISION FREQUENCY AT H2, /SEC
255 140.          H2, REFERENCE HEIGHT, KM
256 .0183        A2, EXPONENTIAL DECREASE OF NU WITH HEIGHT, /KM
                  (A BLANK IN COL. 1-3 ENDS THE CURRENT W ARRAY)

X01 TEST CASE
71 0.          NO PERIODIC PRINTOUT
72 0.          DO NOT PUNCH RAYSETS
81 2.          PLOT PROJECTION OF RAY PATH ON THE GROUND
82 10.0         LATERAL DEVIATION EXPANSION FACTOR
                  (A BLANK IN COL. 1-3 ENDS THE CURRENT W ARRAY)

```

- Col. 1-3 Identification number  
 Col. 4-17 Data in E14.6 format  
 Col. 18 A 1 indicates an angle in degrees  
 Col. 19 A 1 indicates a central earth angle in kilometers  
 Col. 20 A 1 indicates a distance in nautical miles  
 Col. 21 A 1 indicates a distance in feet  
 Col. 22-24 Left for other conversions  
 Col. 25-80 Description of the data

Appendix 8c. Sample Printout

X01 TEST CASE  
CHAPX HAVE DIPOLY EXPZZ 11/05/74  
APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

INITIAL VALUES FOR THE W ARRAY -- ALL ANGLES IN RADIANS, ONLY NONZERO VALUES PRINTED

```
1 -1.000000000000+000
2 6.37000050000+003
4 6.98131700803-001
5 -1.83259571460+000
7 6.00000000000+000
11 7.85390153390-001
16 1.57079032679+000
17 2.61799367791-001
20 2.00000000000+002
22 3.00000000000+000
23 1.00000000000+003
24 1.37008346260+000
25 5.07890812531+000
41 3.00003500000+000
42 1.00000000000-004
43 5.00000000000+001
44 1.00000000000+000
45 1.00000000000+002
46 1.00000000000-008
47 5.00000000000-001
57 2.00000000000+000
58 2.00000000000+000
71 5.00000000000+000
72 1.00000000000+000
81 1.00000000000+000
83 6.98131700803-001
84 -1.83259571460+000
85 9.09065606124-001
86 -1.42767932613+000
87 1.55985871273-002
101 5.50000010000+000
102 3.00000000000+002
103 6.139999999995+001
104 5.00000000000-001
150 1.00000000000+000
151 2.50000000000+002
152 1.00000000000+002
153 1.00000000000-001
155 1.00000000000+002
156 1.00000000000+002
201 8.00000000000-001
251 3.650000000002+004
252 1.00000000000+002
253 1.460000000001-001
254 3.00000000000+001
255 1.339999999999+002
256 1.833000000002-002
```



11/05/74

X01 TEST CASE	JIPOLY	$\Sigma$ XPZ	APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS
CH4PX			
WAVE			
	FREQUENCY =	6.00000 MHz, AZIMUTH ANGLE OF TRANSMISSION =	45.00000 DEG
		ELEVATION ANGLE OF TRANSMISSION =	15.00000 DEG

AZIMUTH DEVIATION				ELEVATION				POLARIZATION				GROUP PATH		PHASE PATH		ABSORPTIVE	
HEIGHT	RANGE	XMTR	LOCAL	XMTR	LOCAL	DEG	DEG	REAL	IMAG	KM	KM	DB	DB	C	G		
0+000	0+000	XMTR	0.0000	0.0000	0.0000	-0.0003	-0.0000	15.000	15.000	0.001	-1.000	0.0000	0.0000	194.1401	194.1401		
0+000	0+000	ENTR ION	5.2+852	1.9+0.0044	0.0000	-0.0003	-0.0000	15.000	16.000	3.0113	+0.3568	1.94.1401	0.0000	0.0000	0.0000		
0+000	0+000		6.7+815	2.1+1.6243	-0.203	-0.000	-0.000	15.000	16.903	2.323	-7.474	221.1401	0.0000	0.0000	0.0000		
0+000	0+000		6.6+6275	2.3+1.5627	-0.203	-0.000	-0.000	15.000	17.074	1.297	-9.793	241.1401	0.0000	0.0000	0.0000		
0+000	0+000		7.9+0.632	2.7+1.1042	-0.200	-0.000	-0.000	15.000	17.447	0.368	-18.990	285.1401	0.0000	0.0000	0.0000		
0+000	0+000		9.1+7.890	3.0+9.7220	-0.200	-0.000	-0.000	15.000	17.785	0.3511.019	1.019	325.1401	0.0000	0.0000	0.0000		
1-000	1-000		10.4+1.175	3.6+7.1973	-0.100	-0.000	-0.000	15.000	18.116	-0.015	29.112	365.1401	0.0000	0.0000	0.0000		
1-000	1-000		1.26+7.922	4.1+2.8687	-0.100	-0.001	-0.001	14.930	16.525	-0.000	10.310	437.1401	0.0000	0.0000	0.0000		
1-000	1-000		1.51+7.725	4.3+1.4663	-0.005	-0.001	-0.002	14.931	16.907	-0.001	4.591192	517.1401	0.0000	0.0000	0.0000		
1-000	1-000		1.61+4.111	5.2+4.9617	-0.001	-0.001	-0.002	14.559	13.612	-0.000	-4.495	557.1401	0.0057	0.0000	0.0000		
1-000	1-000		1.69+1.102	5.6+1.6061	0.005	-0.078	-0.078	14.015	8.649	-0.000	-2.160	597.1401	0.0068	0.0000	0.0000		
1-000	1-000		1.75+7.075	5.9+9.2553	0.015	-0.155	-0.155	13.155	1.378	-0.000	-5.337	637.1401	0.0077	0.0000	0.0000		
1-000	1-000		1.77+14.18	6.0+10.34	0.015	-0.134	-0.134	12.975	0.000	-0.000	-1.481	643.1528	0.0084	0.0000	0.0000		
1-000	1-000		1.77+3.011	6.1+3.1034	0.015	-0.134	-0.134	12.975	0.000	-0.000	-1.481	643.1528	0.0084	0.0000	0.0000		
1-000	1-000		1.78+3.011	6.1+3.1034	0.015	-0.134	-0.134	12.975	0.000	-0.000	-1.481	643.1528	0.0084	0.0000	0.0000		
2+010	2+010		HAVE REV	7.0+7.2225	7.0+7.2225	0.15	0.198	12.894	-0.320	-0.000	-1.452	647.1528	0.0084	0.0000	0.0000		
2+010	2+010		1.53+2.276	7.1+9.5590	0.125	0.078	0.078	12.894	-15.755	-0.003	-1.173	756.1528	0.0114	0.0000	0.0000		
2+010	2+010		1.30+0.051	7.50+9.017	0.033	-0.036	-0.036	5.835	-18.205	-0.000	-1.154	836.1528	0.0135	0.0000	0.0000		
2+010	2+010		1.05+1.515	8.03+5.089	0.037	-0.032	-0.032	3.033	1.167	-0.000	-1.162	916.1528	0.0160	0.0000	0.0000		
2+010	2+010		81.5+547	1.30+6.6930	0.239	-0.030	-0.030	0.752	-17.194	-0.003	-1.174	913.1602	0.0177	0.0000	0.0000		
3+035	3+035		83.7+3.333	1.03+6.673	0.042	-0.027	-0.027	-1.983	-16.239	-0.076	-0.637	1095.1602	0.0174	0.0000	0.0000		
3+035	3+035		GRND REF	0.0+0.000	1.21+2.9251	0.045	-0.023	-0.023	-16.656	-0.002	1.000	1292.1934	0.0177	0.0000	0.0000		
3+035	3+035		ENTR ION	5.2+9.655	1.0+0.040	0.145	-0.020	-0.020	-4.165	-1.366	-1.306	1.885	1490.3081	0.0177	0.0000	0.0000	
3+035	3+035			5.4+11.58	1.40+8.664	0.449	-0.020	-0.020	-4.142	16.400	-1.305	2.105	1471.2614	0.0193	0.0000	0.0000	
3+035	3+035			11.0+4.072	1.56+5.1016	0.052	-0.016	-0.016	-3.213	17.984	-0.000	2.342	1683.0181	0.0193	0.0000	0.0000	
3+035	3+035			1.34+3.372	1.55+5.5945	0.052	-0.018	-0.018	-2.890	18.138	-0.000	2.221	1763.3081	0.0214	0.0000	0.0000	
3+035	3+035			1.56+1.139	1.26+2.0642	0.052	-0.021	-0.021	-1.983	15.432	-0.000	3.024	1835.3081	0.0223	0.0000	0.0000	
3+035	3+035			1.95+2.114	1.76+2.7025	0.051	-0.021	-0.021	-2.675	11.173	-0.000	14.789	1875.1081	0.0244	0.0000	0.0000	
3+035	3+035			1.70+7.131	1.799+2.958	1.154	-0.203	-0.203	-2.779	5.562	-0.003	-3.471	1915.3081	0.0244	0.0000	0.0000	
3+035	3+035			1.77+0.005	1.76+6.6204	1.059	-0.065	-0.065	-2.959	-0.000	-0.003	-1.952	1947+4.017	0.025	0.0000	0.0000	
4+000	MIN J1ST		171.9566													THIS RAY CAL-CALULATION FOUK	11-2+7 SEC

X01 TEST CASE  
CHAPX HAVE DIPOLY EXPZ2  
FREQUENCY = 6.00010 MHz, AZIMUTH ANGLE OF TRANSMISSION = 45.00000 DEG  
ELEVATION ANGLz OF TRANSMISSION = 30.00000 DEG

	HEIGHT	DANGE	DEVIATION	ELEVATION	POLARIZATION	GROUP PATH	PHASE PATH	ABSORPTION
	KM	XMTR DEG	LOCAL DEG	XMTR DEG	REAL IMAG	KM	KM	DB
3-011 YMTR	0.00000	0.00000	-0.000	30.000	-0.000	1.000	0.0000	0.0000
0+0.03 ENTR ION	52.6739	89.9010	-0.100	30.000	-0.409	1.259	104.6686	0.0000
-3-011	60.6634	102.6627	-0.100	30.000	-0.218	1.562	119.6686	0.0000
0+0.00	65.8141	111.4525	-0.000	30.000	-0.110	1.614	129.6686	0.0000
-2-010	78.2077	131.4761	-0.002	30.000	-0.017	1.609	153.6686	0.0000
-2-009	88.3857	148.3519	-0.100	16.000	-0.004	1.592	173.6686	0.0002
-1-007	107.4766	178.5914	-0.000	30.000	-0.000	1.562	205.6671	0.0008
-1-006	128.3954	211.9316	-0.000	29.993	-0.000	1.539	249.6686	0.0020
-5-005	149.2460	245.1205	-0.102	0.009	-0.003	1.574	289.6686	0.0029
-2-005	168.5912	277.9003	-0.111	-0.068	29.599	-0.000	329.6686	0.0038
2-005	184.0395	310.3608	-0.016	0.029	28.909	19.956	-1.895	326.6686
1-005	191.3569	142.2014	-0.027	0.444	27.306	5.952	7.331	360.9884
0+000 MN JIST	191.5641	754.9608	-0.119	26.398	0.000	-1.911	409.6686	0.0052
0+000 MN UIST	191.5641	354.9468	-0.123	0.219	26.398	0.000	-1.530	390.0049
6-000 HAVE REV	131.4066	358.41065	-0.020	0.137	26.154	-1.395	0.000	0.0082
4-016	172.2654	426.8680	0.135	1.783	-23.513	0.000	-1.113	429.7924
2-005	135.5960	493.714d	0.23c	-0.521	12.986	-29.961	0.000	0.0125
2-005	95.7861	561.7775	0.29f	-0.455	7.074	-29.718	0.000	0.147
2-005	56.5019	630.5340	0.346	-0.405	2.257	-29.100	0.039	674.7924
2-005 EXIT ION	4.6.1356	644.4661	0.357	-0.397	1.406	-28.375	0.035	704.5457
-3-011 GND REF	J.0000	733.6060	0.405	-0.349	-3.239	28.173	0.000	0.0177
0+000 ENTR IUN	52.9781	830.3382	0.446	-0.308	+0.104	29.043	-0.277	822.6383
-4-017	92.6836	910.1376	1.470	-0.284	1.779	29.671	-3.002	933.3256
-2-005	132.5904	364.2234	0.490	-0.266	3.349	30.046	-0.000	1014.0786
-1-005	163.3876	162.1970	0.501	-0.223	4.356	26.978	-0.000	1144.3256
-2-005	179.0552	165.0139	0.507	-0.452	4.735	20.718	-1.003	1093.9864
-9-006	186.8742	187.7704	0.51t	0.102	4.754	9.433	-5.108	1208.8256
0+000 MN JIST	169.8217	1107.5272	0.51i	0.454	4.531	0.000	-1.771	1156.1125
								0.0228
								1248.3256
								1191.4144
								1226.3256
								1240.1070
								0.0259
								1312.8736
								0.642 SEC

X01 TEST CASE  
CHAPX WAVE

DIPOLY EXPZ2

11/05/74

APFLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

FREQUENCY = 6.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELEVATION ANGLE OF TRANSMISSION = 45.000000 DEG

	HEIGHT	RANGE	AZIMUTH		ELEVATION		POLARIZATION	GROUP	PATH	PHASE	PATH	ABSORPTION
			XMT	LOCAL	XMT	LOCAL						
0+000 XMTR	0.0000	0.0000			45.000	45.471	-0.000	1.000	0.0000	0.0000	0.0000	0.0000
0+000 ENTR ION	52.9759	52.3246	-0.103	-0.000	45.000	45.552	-0.059	1.207	87.6128	87.6128	97.6128	0.0000
3-011	52.2499	61.3530	-1.000	-0.000	45.000	45.564	-0.021	1.219	97.6128	97.6128	97.6128	0.0000
0+000	69.3926	63.2802	0.000	0.000	45.000	45.614	-0.002	1.217	119.6128	119.6128	139.6128	0.0001
-3-109	65.1331	83.4659	-0.000	-0.000	45.000	45.751	-0.000	1.213	139.6128	139.6125	159.6091	0.0004
-4-008	39.4743	97.2070	-0.100	-0.100	45.000	45.873	-0.000	1.209	159.6128	159.6128	159.6091	0.0009
-1-007	113.8433	110.8868	-0.700	-0.000	45.000	45.386	-0.000	1.260	239.6128	237.3102	237.3102	0.0030
5-005	142.5566	138.0456	-0.002	0.003	44.980	45.657	-0.000	1.209	199.6128	199.4233	199.4233	0.0021
-2-005	170.1037	164.7826	-0.114	-0.106	44.789	43.268	-0.000	1.260	279.6128	268.9418	268.9418	0.0048
6-005	193.2703	190.6257	-1.071	0.369	44.104	34.673	-0.000	1.601	295.1649	278.4398	304.3962	0.0058
-3-011 RCVR	200.0000	200.2014	-0.132	0.607	43.622	28.480	-0.000	2.293	348.1649	348.1649	348.1649	0.0095
2-006	209.5843	233.2599	-0.160	-1.590	40.436	4.780	0.000	-1.976	352.1649	306.3131	306.3131	0.0098
2-006 APOGEE	203.6634	235.8629	-0.040	-1.746	40.109	3.110	0.000	-1.831	360.1649	310.1892	310.1892	0.0103
3-006 WAVE REV	209.4360	241.0969	0.021	-1.982	39.431	-0.266	0.000	-1.623	388.1649	324.3166	324.3166	0.0122
3-007	205.9618	259.5199	1.244	-2.001	36.819	-12.892	0.000	-1.268	410.6543	336.8372	336.8372	0.0137
3-011 RCVR	260.0000	274.3768	0.385	-1.415	34.427	-23.456	0.000	-1.142	467.6543	377.6491	454.5212	0.0169
1-005	172.6329	311.8736	0.547	-0.183	27.233	-41.318	0.003	-1.041	547.6543	494.5131	587.6543	0.0193
4-006	117.8520	366.4860	0.563	-0.033	16.021	-45.154	0.000	-1.031	659.6543	566.5130	659.6543	0.0206
6-005	89.5472	394.2849	1.570	-0.030	10.932	-44.927	0.001	-1.632	791.3392	778.9960	812.3392	0.0214
6-005	38.3026	44.9524	0.574	-0.027	2.978	-44.472	0.004	-1.000	872.3392	818.8774	857.2445	0.0226
6-005 EXIT ION	38.3026	44.9524	0.574	-0.027	2.978	-44.472	0.004	-1.000	912.3392	857.2445	890.3367	0.0236
0+000 GRND REF	0.0000	484.7060	3.576	-0.025	-2.180	44.114	-0.003	1.000	952.3392	905.3752	905.3752	0.0252
0+000 ENTR ION	52.9777	538.6621	0.578	-0.322	3.168	44.599	-0.105	1.094	952.3392	905.3752	905.3752	0.0266
-6-007	110.1068	535.3574	0.581	-0.020	7.705	45.103	-0.000	1.196	952.3392	905.3752	905.3752	0.0266
-6-007	130.4297	622.9739	0.581	-0.016	9.586	45.095	-0.000	1.194	952.3392	905.3752	905.3752	0.0266
-4-006	165.6750	650.1889	0.579	-0.068	11.199	42.989	-0.000	1.230	952.3392	905.3752	905.3752	0.0266
7-005	183.5935	676.6632	0.572	0.057	12.376	35.811	-0.003	1.434	992.3392	905.3752	905.3752	0.0266
-3-011 RCVR	200.0000	491.5163	0.548	0.796	12.769	26.353	-0.000	2.159	1015.8617	1015.8617	1015.8617	0.0266

THIS RAY CALCULATION TOOK 10.310 SEC



X01 TEST CASE  
CHAPX WAVE

DIPOLY EXPZ2

11/05/74

APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

FREQUENCY = 6.000300 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELLEVATION ANGLE OF TRANSMISSION = 75.000000 DEG

		AZIMUTH	DEVIATION	ELEVATION		POLARIZATION	GROUP	PATH	PHASE PATH	ABSORPTION
	HEIGHT	RANGE	XMTR LOCAL	XMTR LOCAL		REAL	IMAG	KM	KM	DB
	KM	KM	DEG DEG	DEG DEG						
0+000 XMTR	0.0000	0.0000			75.000	-0.000	1.000	0.0000	0.0000	0.0000
0+000 ENTR ION	52.9646	14.0759	-0.000 -0.000	75.000 75.127		-0.024	1.023	54.8374	54.8374	0.0000
-3-011	60.716t	16.1100	-0.000 -0.000	75.000 75.145		-0.014	1.044	62.8374	62.8374	0.0000
-3-011	65.5439	17.3768	-0.000 -0.000	75.000 75.156		-0.003	1.047	67.8374	67.8374	0.0000
-2-006	87.7850	23.1903	-0.000 -0.000	75.000 75.209		-0.003	1.048	90.8374	90.8374	0.0001
-1-007	127.1237	26.2124	-0.000 -0.000	75.000 75.253		-0.000	1.047	110.8374	110.8365	0.0005
-9-007	126.4556	33.2025	-0.001 0.000	74.999 75.279		-0.000	1.046	130.8374	130.8161	0.0013
4-007	145.6921	36.1498	-0.007 -0.012	74.934 75.188		-0.000	1.047	150.8374	150.6143	0.0019
3-005	178.7737	46.8081	-0.062 -0.321	74.920 73.849		-0.000	1.057	186.8374	183.5011	0.0030
0+000 RCVR	200.1030	52.6875	-0.302 1.334	74.783 71.357		-0.000	1.087	213.9181	202.1284	0.0046
6-005	224.9440	60.6695	-1.577 1.370	74.379 52.074		-0.000	1.568	266.9181	217.2922	0.0091
1-004	230.7936	66.1623	-1.711 -7.396	73.432 15.569		0.000	-9.825	306.9181	220.0213	0.0127
1-004 APOGEE	230.9183	67.5504	-1.511 -8.853	73.111 8.237		0.000	-3.876	314.9181	220.4673	0.0134
9-003 HAVE REV	230.2593	70.7641	-0.922 -11.181	72.307 -4.474		0.003	-1.950	330.9181	221.5762	0.0147
5-005	222.0226	74.2620	-1.43 -14.206	68.508 -29.337		0.000	-1.143	378.9181	228.5798	0.0187
0+000 RCVR	200.0000	105.6297	5.511 -13.543	61.084 -46.475		0.000	-1.024	434.9673	249.2747	0.0232
2-005	161.0693	134.6131	3.542 -11.278	49.256 -55.293		0.000	-1.007	491.9673	291.7804	0.0263
-2-005	129.0135	159.6964	10.092 -9.754	38.661 -56.407		0.000	-1.006	531.9673	330.5977	0.0275
-3-006	95.7312	177.1855	11.285 -8.562	27.405 -56.282		0.000	-1.006	571.9673	370.5601	0.0289
-1-006	62.4974	196.9508	12.22t -7.620	16.463 -56.084		0.001	-1.006	611.9573	410.5600	0.0292
-1-006 EXIT ION	29.3420	220.9908	12.392 -6.456	6.551 -55.885		0.000	-1.000	651.9573	450.5600	0.0292
0+000 GRNG REF	0.0000	240.7569	13.557 -6.291	-1.083 55.706		-0.000	1.000	687.4459	486.0386	0.0292
0+000 ENTR ION	52.9836	276.3443	14.37t -5.479	9.565 56.127		-1.069	1.064	751.4563	550.0469	0.0292
-8-008	57.0257	296.8465	14.78t -5.055	14.785 56.231		-0.001	1.134	792.4563	591.0469	0.0293
-6-006	120.3697	320.5864	15.12t -4.721	18.948 56.406		-0.000	1.150	832.4563	631.0396	0.0303
2-005	153.3778	342.0624	15.42t -4.425	22.353 55.904		-0.000	1.135	872.4563	670.5260	0.0314
4-005	183.3935	362.9372	15.57t -4.214	24.315 51.894		-0.000	1.195	912.4563	705.8795	0.0327
-3-011 RCVR	200.0000	375.3507	15.794 -3.644	25.937 45.376		-0.000	1.374	937.9577	722.8911	0.0343

THIS RAY CALCULATION TOOK 9.624 SEC

#01 TEST CASE  
 CHA0X JIPJLY C\*0.22  
 FREQUENCY = E LOCATION MHz, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG  
 ELEVATION ANGLE OF TRANSMISSION = 90.000000 DEG

HEIGHT KM	ORIGIN KM	AZIMUTH		ELEVATION		POLARIZATION LOCAL DEG	IMAG KM	GROUP PATH KM	PHASE PATH KM	ABSORPTION 0.0
		XTR DEG.	YTR DEG.	XTR DEG.	YTR DEG.					
0+0.00 X*TR	J+0.010	0.0000	0.0000	90.0000	-0.0003	1.000	0.0000	52.9008	52.9008	0.0000
-3+0.11 ENTR ION	52.1938	0.0003	0.0000	90.0000	-0.0003	1.*019	0.0000	54.9008	54.9008	0.0000
-3+0.11 MAX LAT	54.1938	0.0000	0.0000	90.0000	-0.0003	1.*012	0.0000	56.9008	56.9008	0.0000
-3+0.11	50.1938	0.0000	0.0000	90.0000	-0.0003	1.*014	0.0000	62.4008	62.4008	0.0000
-6+0.11	62.4938	0.0000	0.0000	90.0000	-0.0004	1.*015	0.0000	64.9008	64.9008	0.0000
-9+0.11	64.1938	0.0003	0.0000	90.0000	-0.0003	1.016	0.0000	67.4008	67.4008	0.0000
-6+0.11	67.4938	0.0000	0.0000	90.0000	-0.0002	1.*019	0.0000	69.4008	69.4008	0.0000
-6+0.11 WAVE REV	69.2438	0.0000	0.0000	90.0000	-0.0002	1.*013	0.0000	69.4008	69.4008	0.0000
-6+0.11	69.2438	0.0000	0.0000	90.0000	-0.0001	1.*015	0.0000	71.7008	71.7008	0.0000
-6+0.11	71.7438	0.0000	0.0000	90.0000	-0.0001	1.019	0.0000	74.2408	74.2408	0.0000
-3+0.11	74.2438	0.0000	0.0000	90.0000	-0.0001	1.019	0.0000	76.7408	76.7408	0.0000
-3+0.11	76.7438	0.0000	0.0000	90.0000	-0.0001	1.*019	0.0000	79.4008	79.4008	0.0000
-8+0.11	79.4438	0.0000	0.0000	90.0000	-0.0001	84.2408	0.0000	84.2408	84.2408	0.0000
-2+0.10	84.2438	0.0003	0.0000	90.0000	-0.0003	1.*019	0.0000	89.2408	89.2408	0.0000
-3+0.09	89.2438	0.0000	0.0000	90.0000	-0.0000	94.2408	0.0000	94.2408	94.2408	0.0000
-3+0.09	94.2438	0.0000	0.0000	90.0000	-0.0000	1.016	0.0000	105.2402	105.2402	0.0005
-1+0.09	105.2438	0.0000	0.0001	94.186178.9202	0.0000	1.016	0.0000	110.2402	110.2402	0.0005
-2+0.09	110.2438	0.0000	0.0001	94.186178.9202	0.0000	1.016	0.0000	115.2402	115.2402	0.0005
-3+0.09	115.2438	0.0000	0.0001	94.186178.9202	0.0000	1.016	0.0000	119.7331	119.7331	0.0000
-3+0.09 WAVE REV	119.7338	0.0003	0.0003	94.186178.9202	0.0004	1.016	0.0000	120.7331	120.7331	0.0000
-3+0.09	120.7338	0.0003	0.0003	94.186178.9202	0.0003	1.016	0.0000	125.7226	125.7226	0.0013
-4+0.09	125.7338	0.0007	0.0007	94.5269	0.0003	1.016	0.0000	138.6413	138.6413	0.0017
-4+0.09	138.6438	0.0039	0.0039	94.5269	0.0001	1.016	0.0000	148.4476	148.4476	0.0019
9+0.08	148.4479	0.0021	0.0021	94.5269	0.0000	1.016	0.0000	157.9965	157.9965	0.0022
3+0.07	158.7327	0.0035	0.0035	94.5269	0.0000	1.016	0.0000	158.7408	158.7408	0.0025
3+0.07	169.7823	0.0070	0.0070	145.2761	179.9357	30.0000	-0.0003	1.016	1.016	0.0029
2+0.07 HAVE REV	175.1532	0.1157	0.145260	140.0003	0.0004	30.0000	-0.0000	1.016	1.016	0.0031
2+0.07	176.1406	0.1263	0.145261	140.0000	0.0003	30.0000	-0.0000	1.016	1.016	0.0040
-1+0.16	192.1777	0.2353	0.145261	140.0000	0.0000	39.9322	-0.0000	1.016	1.016	0.0047
0+0.06 RCVR	200.0000	0.3026	0.145266	140.0000	0.0000	39.9311	89.7779	-0.0000	1.016	0.0063
4+0.05	210.5537	0.3315	0.145266	140.0000	0.0000	39.9388	89.9787	-0.0000	1.016	0.0069
3+0.06	227.9864	0.7769	0.145266	140.0000	0.0000	39.7399	89.3193	-0.0000	1.016	0.0076
4+0.06 HAVE REV	230.2092	0.9818	0.145266	140.0000	0.0000	93.9747	89.9795	-0.0000	1.016	0.0081
-4+0.04	237.3654	2.6623	0.145266	0.0003	0.0000	49.333	89.9761	-0.0000	1.016	0.0115
-3+0.03 WAVE REV	236.2305	3.6109	0.145266	0.0003	0.0000	39.957	89.9797	-0.0000	1.016	0.0127
-2+0.03 APOE	233.0207	4.6333	0.145266	0.0000	0.0000	68.699	89.978	-0.0000	1.016	0.0134
-4+0.04	236.0571	5.1869	0.145266	0.0000	0.0000	68.699	51.733	-0.0000	1.016	0.0146
2+0.05	221.3660	9.3361	0.145266	0.0000	0.0000	67.375	-66.363	-0.0000	1.016	0.0177
-6+0.11 RCVR	151.1549	15.3973	-145.269	0.0000	0.0000	45.160	-75.213	-0.0000	1.016	0.0205
2+0.05	112.0674	21.7823	-145.266	0.0003	0.0000	30.469	-77.011	-0.0000	1.016	0.0232
2+0.05 EXIT ION	25.6745	39.6085	-145.266	0.0000	0.0000	91.133	-79.110	-0.0000	1.016	0.0245
-3+0.11 GRND REF	53.0300	48.9337	-145.266	0.0000	0.0000	27.535	-78.963	-0.0000	1.016	0.0252
-3+0.11 ENTR ION	52.1956	63.9344	-145.266	0.0000	0.0000	39.224	79.011	-0.0003	1.016	0.0252
-1+0.07	85.3595	70.2016	-145.266	0.0000	0.0000	50.073	79.357	-0.0003	1.016	0.0253
-2+0.05	121+0.623	77.6+2.3	-145.256	0.0000	0.0000	57.484	79.121	-0.0000	1.016	0.0255
3+0.05	103.2840	81.+4.935	-145.266	0.0000	0.0000	51.819	78.758	-0.0003	1.016	0.0277
-3+0.06	137.0532	92.0447	-145.256	0.0003	0.0000	54.210	77.445	-0.0003	1.016	0.0297
0+0.06 RCVR	200.0300	32.7116	-145.256	0.0003	0.0000	54.368	77.233	-0.0000	1.016	0.0300

THIS DAY CALCULATION Took

13.937 SEC

X81 TEST CASE  
CHAPX WAVE DIFOLY EXPZ2

APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS  
11/05/74

INITIAL VALUES FOR THE WANGLES == ALL ANGLES IN RADIANS; ONLY NONZERO VALUES PRINTED

1	-1.00000000000+000
2	5.37000000000+003
4	6.38131700000-001
5	-1.83259514660+000
7	5.00000000000+000
11	7.85596133390-001
16	1.5707362679+000
17	2.6179337791+001
20	2.00001000000+002
22	3.00000000000+000
23	1.00000000000+000
24	1.3700639286+000
25	5.37593812331+063
41	3.00000000000+000
42	1.00000000000-004
43	5.00000000000+001
44	1.00000000000+000
45	1.00000000000-002
46	1.00000000000J-002
47	5.00000000000-001
57	2.00000000000+000
58	2.00000000000+000
71	1.00000000000+003
81	2.30000000000+000
82	1.00000000000+000
83	6.38131700000-001
84	-1.83259514660+000
85	9.3865666124+001
86	-1.4276793281+000
87	1.56345514273+002
101	6.50000000000+000
102	3.00000000000+002
103	6.139399999995+001
104	5.00000000000+001
150	1.00000000000+000
151	2.50000000000+002
152	1.00000000000+002
153	1.00000000000+001
155	1.00000000000+002
156	1.00000000000+002
201	8.00000000000+001
251	3.65000000002+004
252	1.00000000000+002
253	1.40000000001+001
254	3.00000000000+001
255	1.339939999999+002
256	1.93000000002+002

X01 TEST CASE  
CHAPTER WAVE DIFOLY EXP22  
FREQUENCY = b.00000C MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.00000 DEG

APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS  
ELEVATION ANGLE OF TRANSMISSION = 0.00000 DEG

HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION REAL DEG	GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
		XMR DEG	LOCAL DEG	XMR DEG	LOCAL DEG				
0+000 XMR	0.0000	3.18.72.0	3.099	0.000	-0.000	0.000	-1.000	-1.000	0.0000
-3-011 ENTR ION	52.0737	1.63.14.65	1.63.15.61	0.002	-0.029	0.7364	0.597	-1.427	823.2533 0.0000
-3-011 MIN DIST	158.1469	1.49.15.61	1.49.15.61	0.002	-0.029	-0.753	0.000	-1.720	151.8.9022 1512.7766 0.0106
-3-011 MIN J151	158.1469	1.49.15.61	1.49.15.61	0.002	-0.029	-0.753	0.000	-1.720	151.8.9022 1512.7766 0.0106
-6-011 WAVE REV	158.1177	1.49.09.5559	0.003	-0.039	-0.736	-0.941	0.000	-1.495	1522.9022 1516.5869 0.0107
6-006 EXIT ION	49.2239	21.85.5251	0.003	0.013	-6.572	-7.38	0.205	-1.093	2231.9022 2227.6590 0.0220
0+000 GRND REF	0.0000	2.90.00.0482	0.000	-1.3.042	0.738	0.001	-1.000	-1.000	2935.4929 2943.2497 0.0220
0+000 ENTR ION	52.9836	3.54.8054	-0.003	0.008	-15.566	7.0.01	-0.333	1.636	3700.7907 3688.5475 0.0220
-1-005 MAX LAT	107.3072	3.785.1958	-0.002	0.008	-16.435	10.475	-0.000	1.435	4033.7907 4044.5426 0.0247
-1-005 WAVE REV	107.3072	3.785.1958	-0.002	0.008	-16.435	10.475	-0.000	1.635	4053.7907 4044.5426 0.0247
0+000 MIN DIST	157.9016	4.305.0842	-0.007	0.075	-17.366	-0.000	-3.000	9.317	4368.0431 4369.8069 0.0330

THIS RAY CALCULATION TOOK 10.723 SEC

HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION REAL DEG	GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
		XMR DEG	LOCAL DEG	XMR DEG	LOCAL DEG				
0+000 XMR	0.0000	186.0004	-0.300	-0.000	15.000	16.073	3.119	-3.368	0.0000
0+000 ENTR ION	52.9852	172.0034	3.015	0.164	12.975	0.000	0.000	-1.481	194.1411 194.1404 0.0000
-3-011 IN DIST	172.0032	172.0032	3.015	0.164	12.975	0.000	0.000	-1.481	663.6528 632.4672 0.0081
-3-011 MIN DIST	172.0032	172.0032	3.015	0.164	12.975	0.000	0.000	-1.481	643.6528 632.4672 0.0081
-2-010 WAVE REV	172.0659	5.07.7222	0.015	0.136	12.664	-0.020	0.000	-1.452	647.6528 635.9500 0.0083
-3-005 EXIT ION	46.8333	1.036.8759	0.042	-0.027	-1.983	-16.439	0.376	-1.037	1108.6528 1085.6502 0.0174
0+000 GRND REF	0.0000	1.212.9251	0.046	-0.023	-5.055	14.656	-0.002	1.000	1292.1574 1269.1466 0.0174
-3-011 ENTR ION	52.9855	140.0048	0.049	-0.020	-4.165	16.366	-1.306	1.885	1400.301 1467.2614 0.0174
0+000 MIN DIST	171.3560	1.52.0024	0.059	-0.065	-22.559	-0.000	0.000	-1.952	1947.4107 1912.3923 0.0258

THIS RAY CALCULATION TOOK 10.756 SEC

X01 TEST CASE  
CHAPX WAVE SIPOLY EXPZ2

11/05/74

APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

FREQUENCY = 6.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELEVATION ANGLE OF TRANSMISSION = 30.000000 DEG

	HEIGHT	RANGE	AZIMUTH DEVIATION		ELEVATION		POLARIZATION	GROUP	PATH	PHASE	PATH	ABSORPTION
			XMT <sub>RM</sub>	LOCAL <sub>DEG</sub>	XMT <sub>RM</sub>	LOCAL <sub>DEG</sub>						
3-011 XMTR	0.0000	0.0000	-0.000	-0.000	30.000	-0.000	1.000	0.0000	0.0000	0.0000	0.0000	0.0000
0+000 ENTR ION	52.9739	89.9010	-0.000	-0.000	30.000	30.609	-0.332	1.259	104.6686	104.6686	0.0000	0.0000
0+000 MIN DIST	191.5641	354.9408	-0.023	0.219	26.398	0.000	0.000	-1.530	425.7924	400.9614	0.0082	0.0082
0+000 MIN DIST	191.5641	354.9408	-0.023	0.219	26.398	0.000	0.000	-1.530	425.7924	400.9614	0.0082	0.0082
6-010 WAVE REV	191.4066	358.1065	-0.020	0.137	26.154	-1.395	0.003	-1.475	429.7924	403.6919	0.0685	0.0685
2-005 EXIT ION	46.7358	644.4661	0.357	-0.397	1.406	-28.975	0.035	-1.018	770.7924	720.5457	0.0177	0.0177
-3-011 GRND REF	0.0000	733.6080	1.405	-0.349	-3.299	28.173	-0.000	1.000	872.6850	822.4383	0.0177	0.0177
0+000 ENTR ION	52.9739	830.3382	0.446	-0.308	-0.104	29.043	-0.277	1.223	983.3256	933.0789	0.0177	0.0177
0+000 MIN DIST	199.8217	1107.5272	3.514	0.454	4.581	0.000	0.000	-1.771	1312.6736	1240.1070	0.0259	0.0259
THIS RAY CALCULATION TOOK 8.113 SEC												

ELEVATION ANGLE OF TRANSMISSION = 45.000000 DEG

	HEIGHT	RANGE	AZIMUTH DEVIATION		ELEVATION		POLARIZATION	GROUP	PATH	PHASE	PATH	ABSORPTION
			XMT <sub>RM</sub>	LOCAL <sub>DEG</sub>	XMT <sub>RM</sub>	LOCAL <sub>DEG</sub>						
0+000 XMTR	0.0000	0.0000	-0.000	-0.000	45.000	-0.000	1.000	0.0000	0.0000	0.0000	0.0000	0.0000
0+000 ENTR ION	52.9759	52.3246	-0.000	-0.000	45.000	45.471	-0.113	1.100	74.6128	74.6128	0.0000	0.0000
-3-011 RCVR	200.0000	200.2014	-0.132	0.607	43.622	26.480	-0.000	2.293	295.1649	278.4398	0.0058	0.0058
2-005 APOGEE	209.6834	235.6629	-0.040	-1.746	40.109	3.110	0.000	-1.831	352.1649	306.3131	0.0098	0.0098
3-006 WAVE REV	239.4360	241.3969	0.021	-1.382	39.431	-0.266	0.000	-1.623	360.1649	310.1892	0.0103	0.0103
3-011 RCVR	200.1000	274.3788	0.389	-1.415	34.427	-23.456	0.000	-1.142	410.8543	336.8372	0.0137	0.0137
6-005 EXIT ION	38.9026	44.9524	0.574	-0.027	2.978	-44.472	0.004	-1.000	659.8543	566.5130	0.0206	0.0206
0+000 GRND REF	3.3010	884.7360	0.576	-0.025	-2.110	44.114	-3.001	1.000	715.5626	622.2214	0.0206	0.0206
0+000 ENTR ION	52.9777	534.6621	0.575	-0.022	3.158	44.599	-0.105	1.094	791.3392	697.9980	0.0206	0.0206
-3-011 RCVR	200.0000	691.5163	0.545	0.796	12.769	26.953	-0.000	2.159	1015.3617	905.3752	0.0266	0.0266
THIS RAY CALCULATION TOOK 9.813 SEC												

X01 TEST CASE  
CHAPX WAVE DIPOLY EXPZ211/05/74  
APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

FREQUENCY = 6.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELEVATION ANGLE OF TRANSMISSION = 60.000000 DEG

	HEIGHT	RANGE	AZIMUTH		ELEVATION		POLARIZATION	GROUP	PATH	PHASE	PATH	ABSORPTION	
			XMT	LOCAL	XMT	LOCAL							REAL
	KM	KM	DEG	DEG	DEG	DEG							
0+000 XMTR	0.0000	0.0000			60.000	-0.003	1.000	0.0000	0.0000	0.0000	0.0000	0.0000	
0+000 ENTR ION	52.9736	30.2935	-0.000	-0.000	60.272	-0.053	1.049	61.0908	61.0908	0.0000	0.0000	0.0000	
0+000 RCVR	200.0000	114.4152	-0.157	-0.851	59.325	50.578	-0.000	1.250	239.2552	225.6343	0.0050	0.0050	0.0050
-1-004 APOGEE	225.8332	145.6930	-0.354	5.402	55.872	-5.298	0.000	-2.002	324.2552	253.2681	0.0114	0.0114	0.0114
-1-004 WAVE REV	225.8332	145.6930	-0.354	5.402	55.872	-5.298	0.000	-2.002	324.2552	253.2681	0.0114	0.0114	0.0114
0+000 RCVR	200.0000	168.1639	-1.383	16.376	+8.744	-62.499	0.000	-1.025	395.1225	273.3158	0.0173	0.0173	0.0173
-2-005 EXIT ION	26.7751	230.9196	-6.447	13.564	5.561	-59.411	0.000	-1.000	588.1225	453.6124	0.0223	0.0223	0.0223
0+000 GRND REF	0.0000	240.7118	-7.009	13.003	-1.083	59.320	-0.000	1.000	616.7330	482.2229	0.0223	0.0223	0.0223
-3-011 ENTR ION	52.9816	260.0662	-7.992	12.021	10.299	69.499	-0.026	1.025	673.3376	538.8275	0.0223	0.0223	0.0223
0+000 RCVR	200.0000	313.9111	-10.115	8.555	30.682	63.564	-0.000	1.099	837.7013	691.0808	0.0271	0.0271	0.0271
THIS RAY CALCULATION TOOK 7.518 SEC													

ELEVATION ANGLE OF TRANSMISSION = 75.000000 DEG

	HEIGHT	RANGE	AZIMUTH		ELEVATION		POLARIZATION	GROUP	PATH	PHASE	PATH	ABSORPTION	
			XMT	LOCAL	XMT	LOCAL							REAL
	KM	KM	DEG	DEG	DEG	DEG							
0+000 XMTR	0.0000	0.0000			75.000	-0.000	1.000	0.0000	0.0000	0.0000	0.0000	0.0000	
0+000 ENTR ION	52.9846	14.0759	-0.000	-0.000	75.000	75.127	-0.024	1.023	54.8374	54.8374	0.0000	0.0000	0.0000
0+000 RCVR	200.0000	52.6875	-0.302	1.334	74.783	71.357	-0.000	1.087	213.9181	202.1284	0.0046	0.0046	0.0046
1-004 APOGEE	230.9183	57.5564	-1.511	-8.853	73.111	8.237	0.000	-3.876	314.9181	220.4673	0.0134	0.0134	0.0134
9-005 WAVE REV	230.2533	70.7691	-0.322	-11.181	72.307	-6.474	0.000	-1.950	330.9181	221.5782	0.0147	0.0147	0.0147
0+000 RCVR	200.0000	106.6297	5.511	-13.543	61.084	-46.475	0.000	-1.024	434.9673	249.2747	0.0232	0.0232	0.0232
-1-006 EXIT ION	29.3420	220.9908	12.392	-6.856	6.551	-55.885	0.000	-1.000	651.9673	450.5600	0.0292	0.0292	0.0292
0+000 GRND REF	0.0000	240.7589	13.557	-6.291	-1.083	55.706	-0.001	1.000	687.4459	486.0386	0.0292	0.0292	0.0292
0+000 ENTR ION	52.9836	276.3403	14.375	-5.479	9.565	56.027	-0.069	1.064	751.4563	550.0489	0.0292	0.0292	0.0292
-3-011 RCVR	200.0000	375.3507	15.794	-3.644	25.987	45.376	-0.000	1.374	937.9577	722.8911	0.0343	0.0343	0.0343
THIS RAY CALCULATION TOOK 9.049 SEC													

X01 TEST GASE  
 CHAPX WAVE DIPOLY RX2Z2  
 FREQUENCY = 0.0000 MHz, AZIMUTH ANGLE OF TRANSMISSION = 45.00000 DEG  
 ELEVATION ANGLE OF TRANSMISSION = 90.00000 DEG  
 APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS 11/05/74

	AZIMUTH		ELEVATION		POLARIZATION		GROUP PATH		PHASE PATH		ABSORPTION	
	XMT	LOCAL	XMT	LOCAL	REAL	IMAG	KM	KM	OB	OB	OB	
WAVE LENGTH	KM	DEG	KM	DEG								
0+0.00 XMTR ION	3.0000	0.0000	30.0000	-0.0003	1.000	0.000	0.0000	0.0000	0.0000	0.0000	0.0000	
-3-0.11 ENTR ION	5.23993e-01	0.0000	30.0000	-0.0003	1.000	0.000	52.9908	52.9908	52.9908	52.9908	52.9908	
-3-0.11 MAX LAT	5.43993e-01	0.0000	30.0000	-0.0003	1.002	0.002	54.9908	54.9908	54.9908	54.9908	54.9908	
-6-0.11 WAVE REV	6.32405e-01	0.0000	30.0000	-0.0002	1.019	0.002	69.2408	69.2408	69.2408	69.2408	69.2408	
-3-0.09 WAVE REV	1.19735e-01	0.0003	-145.262	0.004	30.0000	-0.000	1.197408	1.197408	1.197408	1.197408	1.197408	
2-0.07 WAVE REV	1.75153e-01	0.1157	-145.262-180.000	0.004	30.0000	-0.000	1.016	1.016	1.016	1.016	1.016	
0+0.00 RCVR	2.03036e-01	0.3028	-145.262-180.000	0.000	89.351	69.993	-0.003	1.021	1.021	1.021	1.021	
4-0.06 WAVE REV	2.31203e-01	0.7913	-145.262-180.000	0.000	59.311	88.779	-0.003	1.030	1.030	1.030	1.030	
-3-0.03 WAVE REV	2.352305e-01	3.6109	-145.262	0.000	59.311	88.779	-0.003	1.052	1.052	1.052	1.052	
-2-0.03 APOGEE	2.384027e-01	6.0933	-145.265	0.000	38.978	-33.291	0.003	4.505	4.505	4.505	4.505	
-6-0.11 RCVR	2.713002e-01	15.3973	-145.265	0.000	35.450	75.113	0.003	-5.605	-5.605	-5.605	-5.605	
2-0.05 EXIT ION	25.677e+00	48.6937	-145.265	0.000	27.355	-78.363	0.001	411.0570	411.0570	411.0570	411.0570	
-3-0.11 CRNU REF	0.0000	53.7019	-145.265	0.000	-0.242	78.918	-0.003	596.0570	596.0570	596.0570	596.0570	
-3-0.11 ENTR ION	6.33994e-01	6.33994	-145.265	0.000	39.224	79.011	-0.003	410.365	410.365	410.365	410.365	
0+0.00 RCVR	2.03036e-01	92.71e-01	-145.265	0.000	04.368	77.233	-1.003	676.2116	676.2116	676.2116	676.2116	
THIS DAY CALCULATION TOOK	13.101 SEC				8.333	279.0		635.5670	635.5670	635.5670	635.5670	

Appendix 8d. Listing of Punched Card Output (ray sets)  
for Sample Case

X01	TEST CASE
CHAPX	6.500+000 3.000+002 6.200+001 5.000-001 0.000+000 0.000+000 0.000+000
WAVE	2.500+002 1.000+002 1.000-001 0.000+000 1.000+002 1.000+002 0.000+000
DIPOLY	8.000-001 0.000+000 0.000+000 0.000+000 0.000+000 0.000+000 0.000+000
EXPZ2	3.650+004 1.000+002 1.480-001 3.000+001 1.400+002 1.830-002 0.000+000
X01X	0 40000255000 2000000 60000 4500000 0 0 -1003T
1581469	14911561 2 -29 0 1514389 4513 -1610 11 0 0 -1721M
1581469	14911561 2 -29 0 1514389 4513 -1610 11 0 0 -1722M
1581469	29000482 -0 10 738 2875068 80425 68182 22 0 0 -1003G
1579016	43050842 -7 76 -0 4278561109482 91246 33 0 -0 9323M
X01X	0 40000255000 2000000 60000 4500000 1500000 0 -1003T
1721392	6041034 15 184 0 635731 8122 -3264 8 0 0 -1481M
1721392	6041034 15 184 0 635731 8122 -3264 8 0 0 -1482M
1721418	12129251 46 -23 14656 1211094 81100 58053 17 0 -0 1003G
1719566	18286204 59 -65 -0 1854769 92641 58123 26 0 0 -1953M
X01X	0 40000255000 2000000 60000 4500000 3000000 -0 1003T
1915641	3549408 -23 219 0 407964 17828 -7003 8 0 0 -1531M
1915641	3549408 -23 219 0 407964 17828 -7003 8 0 0 -1532M
1916346	7336080 405 -349 28173 733203139482 89236 18 0 -0 1003G
1898217	11075272 514 454 0 1138430174444101677 26 0 0 -1773M
X01X	0 40000255000 2000000 60000 4500000 4500000 -0 1003T
0	2002014 -132 607 28480 285194 9971 -6754 6 0 -0 2291R
2096843	2743788 389 -1415-23456 342980 67874 -6143 14 0 0 -1142R
2096843	4847060 576 -25 44114 484589230974137632 21 0 -0 1003G
0	6915163 548 796 26953 729880285982175495 27 0 -0 2163R
X01X	0 40000255000 2000000 60000 4500000 6000000 -0 1003T
0	1144182 -157 -851 50578 231305 7950 -5671 5 0 -0 1251R
2258382	1681639 -1383 16876-62499 262993132130 10323 17 0 0 -1032R
2258382	2407118 -7009 13003 69320 240697376036241525 22 0 -0 1003G
0	3139111-10115 8555 63564 376316461386314765 27 0 -0 1103R
X01X	0 40000255000 2000000 60000 4500000 7500000 -0 1003T
0	526875 -302 1334 71357 207034 6884 -4906 5 0 -0 1091R
2309183	1066297 5511-13543-46475 227435207533 21840 23 0 0 -1022R
2309183	2407589 13557 -6291 55706 240745446701245294 29 0 -0 1003G
0	3753507 15794 -3644 45376 430430507528292462 34 0 -0 1373R
X01X	0 40000255000 2000000 60000 4500000 9000000 -0 1003T
0	3028214734180000 88779 200000 7166 -5034 5 0 -0 1031R
2382305	153973214734 0-75213 200610210447 37559 20 0 0 -1082R
2382305	537019214734 0 78918 53702568516382995 25 0 -0 1003G
0	927180214734 0 77233 221057612222414510 30 0 -0 1013R

The first card is the title card.

The second card contains the name of the electron density model plus parameters W101-W107.

The third card contains the name of the perturbation model plus parameters W151-W157.

The fourth card contains the name of the magnetic field model plus parameters W201-W207.

The fifth card contains the name of the collision frequency model plus parameters W251-W257.

For description of remaining cards, see figures 1 and 2.

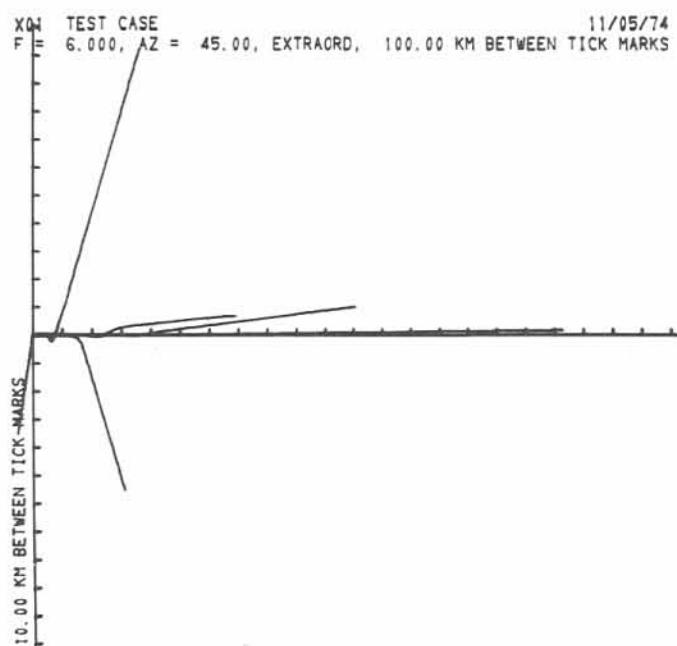
Appendix 8e. Ray Path Plots for Sample Case

Projection of raypath on vertical plane

X01 TEST CASE  
F = 6.000, AZ = 45.00, EXTRAORD, 100.00 KM BETWEEN TICK MARKS 11/05/74



Projection of raypath on ground for sample case





## BIBLIOGRAPHIC DATA SHEET

	1. PUBLICATION OR REPORT NO.  OTR 75-76	2. Gov't Accession No.	3. Recipient's Accession No.
4. TITLE AND SUBTITLE  A versatile three-dimensional ray tracing computer program for radio waves in the ionosphere		5. Publication Date  October 1975	6. Performing Organization Code OT/ITS, Div. 1
7. AUTHOR(S)  R. Michael Jones and Judith J. Stephenson		9. Project/Task/Work Unit No.	
8. PERFORMING ORGANIZATION NAME AND ADDRESS  U. S. Department of Commerce Office of Telecommunications Institute for Telecommunication Sciences Boulder, Colorado 80302		10. Contract/Grant No.	
11. Sponsoring Organization Name and Address  U. S. Department of Commerce Office of Telecommunications Institute for Telecommunication Sciences Boulder, Colorado 80302		12. Type of Report and Period Covered  Technical Report	13.
14. SUPPLEMENTARY NOTES			
<p>15. ABSTRACT (A 200-word or less factual summary of most significant information. If document includes a significant bibliography or literature survey, mention it here.)</p> <p>This report describes an accurate, versatile FORTRAN computer program for tracing rays through an anisotropic medium whose index of refraction varies continuously in three dimensions. Although developed to calculate the propagation of radio waves in the ionosphere, the program can be easily modified to do other types of ray tracing because of its organization into subroutine.</p> <p>The program can represent the refractive index by either the Appleton-Hartree or the Sen-Wyller formula, and has several ionospheric models for electron density perturbations to the electron density (irregularities), the earth's magnetic field and electron collision frequency.</p> <p>For each path, the program can calculate group path length, phase path length, absorption, Doppler shift due to a time-varying ionosphere, and geometrical path length. In addition to printing these parameters and the direction of the wave normal at various points along the ray path, the program can plot the projection of the ray path on any vertical plane or on the ground and punch the main characteristics of each ray path on cards.</p> <p>The documentation includes equations, flow charts, program listings with comments, definitions of program variables, deck set-ups, description of input and output, and a sample case.</p>			
<p>KEY WORDS:</p> <p>Appleton-Hartree formula; computer program; ionosphere; radio waves; ray tracing; Sen-Wyller formula; three-dimensional.</p>			
17. AVAILABILITY STATEMENT  <input checked="" type="checkbox"/> UNLIMITED.  <input type="checkbox"/> FOR OFFICIAL DISTRIBUTION.		18. Security Class (This report)  UNCLASSIFIED	20. Number of pages  197
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ERRATA (23 April 1979)

A VERSATILE THREE-DIMENSIONAL RAY TRACING COMPUTER PROGRAM  
FOR RADIO WAVES IN THE IONOSPHERE

by R. Michael Jones and Judith J. Stephenson

The 8th line from the bottom of page 9 should read:

(1) the dispersion relation cannot be exactly satisfied, or

The first line in the 3rd complete paragraph on page 12 should read:

Similarly, the AHWFNC (Appleton-Hartree, with field, no colli-

Line PRIN125 in SUBROUTINE PRINTR on page 80 should read:

RANGE=EARTH\*ATAN2(RCE,EARTH+EPS+XMTRH)

PRIN125

Line BQNC020 in SUBROUTINE BQWFNC on page 100 should read:

REAL N2,NNP,LPOLAR,LPOLRI,KR,KTH,KPH,K2,KDOTY,K4,KDOTY2,

BQNC020

Line TABX064 in SUBROUTINE TABLEX on page 112 should read:

PXPR=PXPTH=PXPPH=0.

TABX064

Following line CHAP024 in SUBROUTINE CHAPX on page 116, insert the line:

PXPPH=0.

CHAP0245

Line VCHA010 in SUBROUTINE VCHAPX on page 117 should read:

X=PXPR=PXPTH=PXPPH=0.

VCHA010

Line DCHA014 in SUBROUTINE DCHAPT on page 119 should read:

X=PXPR=PXPTH=PXPPH=0.

DCHA014

Line LINE013 in SUBROUTINE LINEAR on page 120 should read:

X=PXPR=PXPTH=PXPPH=0.

LINE013

Line PARA012 in SUBROUTINE QPARAB on page 121 should read:

X=PXPR=PXPTH=PXPPH=0.

PARA012

Following line BULG038 in SUBROUTINE BULGE on page 123, insert the line:

PXPPH=0.

BULG0385

Following line EXPX014 in SUBROUTINE EXPX on page 124, insert the line:

PXPTH=PXPPH=0.

EXPX0145

The equation for the gyrofrequency near the top of page 143 should read:

$$F_H = F_{H_0} (R_0 / (R_0 + h))^3 (1 + 3 \cos^2 \theta)^{1/2}$$

where  $\theta$  is the geomagnetic colatitude.

Line TABZ015 in SUBROUTINE TABLEZ on page 153 should read:

IF (READNU.EQ.0.) GO TO 10

TABZ015