SIMPRO, SPACE DEBRIS SIMULATOR AND PROPAGATION - USER GUIDE

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I. Introduction

SIMPRO is a space debris simulator of break-up events, based on NASA's model EVOLVE 4.0 ([3], see also [1]); it also allows to propagate the orbits using different formalisms (Cartesian or Hamiltonian). SIM-PRO represents a useful tool in the analysis of satellites and space debris dynamics, with the aim to provide information on break-up events, either collisions or explosions, and their subsequent dynamics. The application is written in JAVA and can be executed on a wide set of operating systems.

These guidelines are intended to give the user basic information to run the program; details can be found in the accompanying paper [2].

The executable program is publicly available and it can be downloaded from the following link:

SIMPRO GitHub repository

IMPORTANT NOTE: When using the results of this simulator in articles, books, videos, or any other mean, acknowledge the use making reference to:

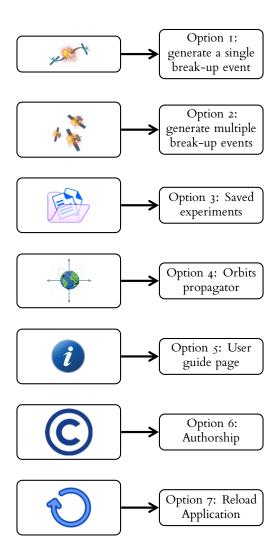
M. Apetrii, A. Celletti, E. Efthymiopoulos, C. Gales, T. Vartolomei, Simulating a breakup event and propagating the orbits of space debris, Celestial Mechanics and Dynamical Astronomy 136, 35 (2024). https://doi.org/10.1007/s10569-024-10205-3

Any question on SIMPRO can be addressed on GitHub or writing to: simpro.project@gmail.com

Refer to these notes as: Version 1.0 - September 2024.

2. QUICK DESCRIPTION OF THE MAIN MENU

This section reproduces the left column that appears when opening SIMPRO and explains the meaning of the symbols, which correspond to the different options that the user can select. The following sections will provide further details on each option.

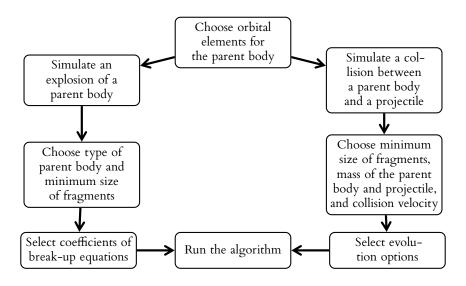


3. Option 1: Generate a single break-up event

The first option of SIMPRO reproduces the break-up model Evolve 4.0 provided by NASA for a single break-up event.

It allows one to determine the cross-sections, masses, and imparted velocities of the fragments after an explosion or a collision ([1], [3]). Section 3.1 describes the work-flow, Section 3.2 illustrates how to choose the initial orbital elements, Sections 3.3 and 3.4 explain, respectively, the procedure to simulate a collision or an explosion, Section 3.5 shows how to get the evolution of each fragment on a given time interval, Section 3.6 explains how to change some settings with respect to [3], Section 3.7 describes the different outputs generated by SIMPRO, Section 3.8 shows how to save the data of an experiment.

3.1. **Work-flow.** The work-flow of SIMPRO is illustrated below.



- 3.2. Choosing Orbital Elements. Choose Mean elements (TLE) and provide the orbital elements of the parent body (where the break-up event will happen).
- Semi-major axis in km: a > 6400 km and a < 45000 km.
- Eccentricity $e \in [0, 0.75]$.
- Inclination in degrees: $i \in [0, 180^o]$.
- Mean anomaly in degrees: $M \in [0, 360^{\circ}]$.
- Argument of perigee in degrees: $\omega \in [0, 360^{\circ}]$.
- Argument of the ascending node in degrees: $\Omega \in [0, 360^{\circ}]$.

An example is given in Figure 1.

Parent Body Position					
Mean Elements (TLE)	Two Line Elements				
Semi-major axis [km]	21600				
Eccentricity	0.1				
Inclination [°]	63.5				
Mean anomaly [°]	80				
Argument of perigee [°]	70				
Longitude of the ascending node [°]	60				

FIGURE 1. Orbital Elements of Parent Body.

Alternatively, choose Two Line Elements and provide the data in LINE 1 and LINE 2 using the NORAD format from CelesTrak.

3.3. Collisions options.

- Choose the size of the smallest fragment generated, with the constraint
 ≥ 1 cm.
- Set the masses of parent body between 300 and 15 000 kg.
- Set the masses of parent body between 0.001 and 5 kg.
- Choose the impact velocity between 1 and 10 000 m/s.
- Select the type of parent body: Drop down menu Upper Stage or Drop down menu Spacecraft

An example is given in Figure 2.

The "Output" frame gives the number of generated fragments and the number of fragments with elliptic orbits, taking into account that some fragments can fall on the Earth. Besides, it gives "Catastrophic: false" or "Catastrophic: true", the latter case occurring whenever the energy of break-up is greater than 40 kJ, which means that the parent body is completely disintegrated. This could happen only in case of a collision with a heavy projectile at a high relative velocity.

3.4. Explosions options. After having chosen the mean elements of the parent body as in Section 3.3, proceed to fix the parameters of break-up.

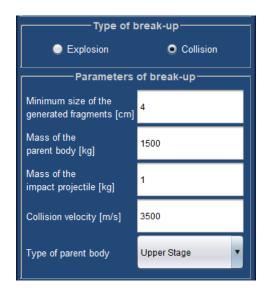


Figure 2. Collision options.

- Choose the size L_c of the smallest fragment generated, with the constraint ≥ 1 cm.
- Select the type of parent body:



More details on the different types are described on [2]. The "Output" frame gives the number of generated fragments and the number of fragments with elliptic orbits, taking into account that some fragments can fall on the Earth.

3.5. Fragments' propagation. To activate the propagation of the fragments, select Toggle activation with a LEFT CLICK and set the options with a RIGHT CLICK. One can choose "short-term" and "long-term" propagation, providing the period and step, the initial time and selecting the forces one wants to include in the propagation, namely the J_2 and J_3 terms of the geopotential, the effects of Moon and Sun, the Solar radiation pressure, and the drag effect due to the Earth's atmosphere, see Figure 3.

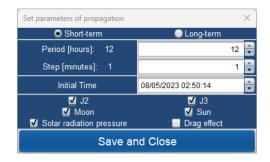


FIGURE 3. Options for the propagation of each fragment for a period of time.

- 3.6. **Setting equations' parameters.** You can change the original coefficients and parameters of the NASA Evolve 4.0 equations described in [1], [3], see Figure 5.
- Open the pop-up window by pressing Set Parameters More Info button.



FIGURE 4. Set NASA Evolve 4.0 parameters.

- Find the equation you want to modify and set the value of the parameter(s) within the editable boxes. The boxes provide the default values; a pop-up window that opens passing on the editable boxes provides the allowed intervals for each parameter.
- You can change one or more parameters.
- Save your choice parameters using Save and close button on top of the window.

Notice that a slightly change in the NASA Evolve 4.0 parameters could give major differences in the results. Please refer to [1] and [2] for more details.

- 3.7. Output. The output panel has two different frameworks, according to whether the orbits are propagated or not.
- 3.7.1. Orbital Elements Plots. The orbital elements a, e, i are plotted in the 2-dimensional planes of a i, a e, and e i; Figure 6 shows the output without (left panel) and with propagation of the orbits (right panel).

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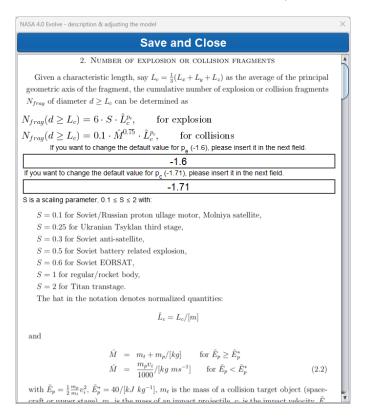


Figure 5. Set parameters window

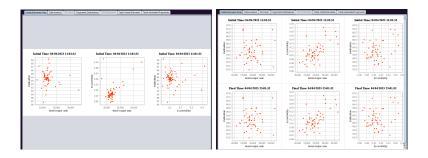


FIGURE 6. Orbital elements plots.

3.7.2. *Data Analysis*. It provides a statistical analysis of the data after a break-up event by computing mean, median, min, max, range, variance, standard deviation, mean standard error, coefficient of variation of the orbital elements, size, area-to mass ratio and velocity distributions. Figure 7 shows the output after a propagation.

The output in Figure 7 is described as follows:

• First row: statistics at break-up time;

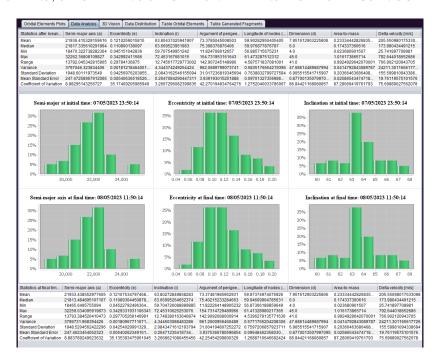


Figure 7. Data analysis window.

- Second row: histograms of *a*, *e*, *i* at break-up;
- Third row: histograms of a, e, i after evolution period;
- Forth row: statistics after evolution period.

3.7.3. 3D Orbit Vision. This page allows to visualize the position of the generated fragments in a 3-dimensional space, whenever the propagation has been computed. An example is given in Figure 8 with the following description, where the blue dots make the contour plot of the Earth surface:

- Top-left: parent body position at break-up time.
- Top-right: fragments' orbits after one fourth of period.
- Bottom-left: fragments' position after three fourths of period.
- Bottom-right: fragments' orbits at final time.

Each panel has additional options to zoom and move the graphs, rotate the plot, reset the zoom, modify the scale settings, save the picture, have a list of Cartesian coordinates (X,Y,Z) of each fragment at a given time, the coordinates of the parent body that gives the orbit, the coordinates of the blue dots representing the Earth. An option allows one to decide whether visualize the fragments, parent body, and the Earth.

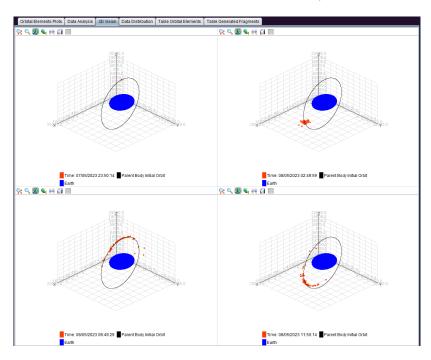


Figure 8. 3D plots.

- 3.7.4. Data Distribution. This page shows the ξ -distribution described in the NASA model ([3]), where $\chi = \log_{10} \frac{A}{m}$, see Figure 9. This page is activated only for catastrophic events, which are obtained for a large mass of the impact projectile and a large collision velocity. The page shows:
- First row: Histograms of size, area-to-mass ratio and Δv ;
- Second row (only for catastrophic events): Histograms of χ for the generated data and for different values of L_c versus the theoretical Probability density function (PDF) of χ .
- 3.7.5. Table of Orbital Elements. This page gives the data about the orbital position $(a, e, i, M, \omega, \Omega)$, the area-to-mass ratio, the dimension, and the ID sequence number of each generated fragment after the break-up, see Figure 10.
- 3.7.6. Table of Generated Fragments. This page gives the state vector and the characteristics of each generated fragment after the break-up; the ID sequence number, the size, the area-to-mass ratio, the area, the mass, the delta velocity, the components of the velocity (V_x, V_y, V_z) , the observation dynamical type, which could be elliptic or hyperbolic orbits, and escape or re-entry orbits, see Figure 11.



FIGURE 9. Data distribution window.

Orbital Elements Plots	Data Analysis	3D Vision	Fragmer	nts Distributions Ci	i Distribution T	Table O	rbital Elements	Table Generated Fr	agments		
	ccentricity (e)	Inclination	_	Mean anomaly (M)	Argument of pe		Longitude of nod			ID	
20870.511918737 (0.104179933512	61.71090	0305356	99.355785630240	50.38895093	3240	60.6608269224	87 0.164226760	367 45.00	1	- 1
22569 714663864 (0.171224104430	65 93865	4984447	63 585004447253	79 25996130	9397	59.0896085407	61 0.393407463	753 30.00	2	
	0.100925170727	62.263969	9865814	90.495770866025	59.33218433	5535	60.4482089113		366 23.00	3	
	0.108700654924	69.43635		25.593845525215	130.65008546		57.8707355708			4	
	0.154544369825	62.749483		53.659334313395	91.83943115		60.2634349558			5	
	0.039313711323	62.94622		-31.168678530572	195.02388229		60.1890467435			6	
	0.085511536283	63.31454		91.711474908302	60.05670392		60.0505205786			7	
	0.158367867604	61,76634		50.476360127126	94.93971377		60.6394057689			8	
20680.167351988 (0.130320208968	63.39525		99.156671196886	48.11818353		60.0202904519			9	
	0.100404218160	63.67667		75.589109140210	74.53024752		59.9152398342			10	
	0.096969122422	63.22863		74.802500257432	75.66144796		60.0827448679			11	
	0.106450418110	63.57530		86.468222196869	62.87693484		59.9530173857			12	
	0.112201445836	63.84449		78.322681110993	70.40032000		59.8528516161			13	
	0.098431776605	63.22991		73.351870983977	76.98844649		60.0822648566			14	
	0.085610705173	63.32567		67.337101977032	84.75067272		60.0463491913			15	
	0.110398824267	63.68273		72.562812172360	76.49434802		59.9129849974			16	
	0.093220731885	63.78308		94.379150295783	56.68532966		59.8756599773			17	
	0.092602685700	62,74812		61.913410765069	89.62353288		60.2639490930			18	
	0.073096669011	62.80106		79.003970654112	74.05456534		60.2439062747			19	
	0.102925222242	63.18102		79.260279615141	70.41377198		60.1006258957			20	
	0.096146099685	62.81951		89.565304216880	60.87499088		60.2369225410			21	
	0.102865267300	64.67860		73.263969955538	76.80164614		59.5455302006			22	
	0.111809815609	63.09979		123.086897601181			60.1311710665			23	
	0.119458654092	62.50842		139.115823628540			60.3549618999			24	
	0.100687220713	63.72539		70.136425346165	80.15475620		59.8971108408			25	
	0.098188748614	63.942174		69.221678175515	81.43625268		59.8166259413			26	
	0.101310007028	63.28394		62.733298906545	87.85948125		60.0619937437			27	
	0.111127121969	64.45344		67.975227096490	81.34044885		59.6280430339			28	
	0.109713495124	61.85057		74.413672486533	74.34207288		60.6069120528			29	
	0.144480599099	62.236679		61.580007907360	83.99523414		60.4586464442			30	
	0.066134603553	64.78973		43.758200466934	112.40084306		59.5049251208			31	
	0.111181851170	64.09013		87.460761596208	61.44996956		59.7618733840			32	
	0.096705164052	63.03086		66.301326866834	84.53237226		60.1571290221			33	
	0.118967361571	64.13187		65.531722486816	82.98324208		59.7464548004			34	
	0.124292883717	63.09826		57.586860305516	90.86293415		60.1317471499			35	
	0.128621032532	62,73028		60.521232224454	87.07038955		60.2707099588			36	
	0.128621032532	63.13788		115.113540588779			60.2707099588			36	-
										38	_
	0.150877642810	66.38297		44.944381356091 62.279370345762	90.52733392		59.8502335985			38	
	0.086101406792						58.9311049905			39 40	
	0.177955687081	66.259492		36.276014035304	111.15339772		58.9750414908				
	0.196345915288	61.26096- 57.44543I		53.051564062048	87.46168505 41.23536190		60.8355267079			41 42	
				102.760928321898			62.3861520658				
	0.114802439505	63.57188		58.055684899204	91.48388203		59.9542931239			43 44	_
	0.116133092574	63.31386		55.999457144681	93.56894209		60.0507767855				
	0.073738508109	62.96793		88.052192984141	64.93288050		60.1808548520			45	_
	0.075661459624	64.08031		89.368030059624	63.59728290		59.7655046519			46	
	0.085465582591	63.50716		55.527423231549	97.42976326		59.9784516490			47	
23033 224165992	0.106628507941	63 35651	4558347	44 837615992011	107 18748540	14614	60.0347955539	12 0.158164930	403 400	48	

FIGURE 10. Table of orbital elements.

3.8. Save Experiment. To save the current experiment (including the evolution), press the button below the output text, see Figure 12.

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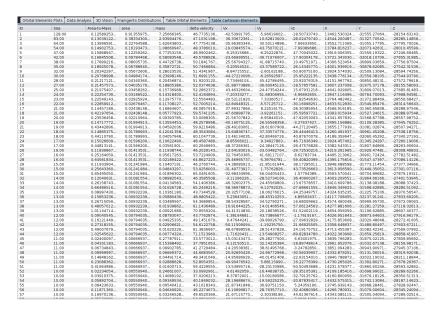


FIGURE 11. Table of generated fragments.



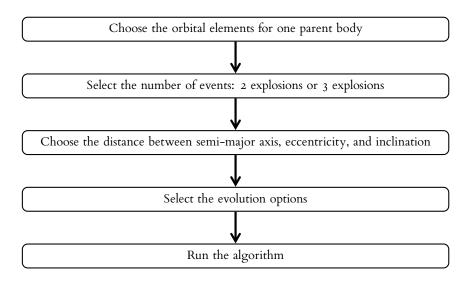
FIGURE 12. Save experiment button.

4. Option 2: generate multiple break-up events

The second option of SIMPRO reproduces the break-up model Evolve 4.0 provided by NASA in [3] to generate two or three explosions and to analyze their evolution.

The work-flow is described in Section 4.1, the choice of the orbital elements is shown in Section 4.2, the selection of the number of breakup events is found in Section 4.3, while Section 4.4 shows how to get the evolution of the generated fragment on a given time interval, and Section 6.7 describes the different outputs generated by SIMPRO for multiple break-up events.

4.1. Work-flow. The work-flow below reproduces the left column that appears when opening SIMPRO for the second option of multiple break-up events.



- 4.2. Choosing Orbital Elements. Choose the orbital elements of the parent body (where the break-up event will happen), see Figure 1.
- Semi-major axis in km: a > 6400 km and a < 45000 km.
- Eccentricity $e \in [0, 0.75]$.
- Inclination in degrees: $i \in [0, 180^{\circ}]$.
- Mean anomaly in degrees: $M \in [0, 360^{\circ}]$.
- Argument of perigee in degrees: $\omega \in [0, 360^{\circ}]$.
- Argument of the ascending node in degrees: $\Omega \in [0, 360^{\circ}]$.
- 4.3. Number and Distance between Events. Select the number of break-up events, either 2 or 3 bodies (see Figure 14).

Parent Body Position				
Mean Elements (TLE)	Two Line Elements			
Semi-major axis [km]	21600			
Eccentricity	0.			
Inclination [°]	8888			
Mean anomaly [°]	80			
Argument of perigee [°]	70			
Longitude of the ascending node [°]	60			

FIGURE 13. Choice of the orbital elements.



FIGURE 14. Selection of the number of break-up events.

Choose the distance in terms of a, e, i from the orbital elements of the first parent body (see Figure 15).



FIGURE 15. Distances of the orbital elements.

4.4. **Fragments' propagation.** To propagate the generated fragments, select Toggle activation with a LEFT CLICK and set the options with a RIGHT CLICK. One can choose "short-term" and "long-term" propagation, providing the period and step, the initial time and selecting the forces one wants to include in the propagation, namely the J_2 and J_3 terms of the geopotential, the effects of Moon and Sun, the Solar radiation pressure, and the drag effect due to the Earth's atmosphere, see Figure 16.



Figure 16. Propagation options.

- 4.5. Output. The output panel has two different frameworks, according to whether the orbits are propagated or not.
- 4.5.1. Orbital Elements Plots. The orbital elements a, e, i are plotted in the 2-dimensional planes of a-e, a-i, and e-i; Figure 17 shows the output without (left panel) and with propagation of the orbits (right panel).

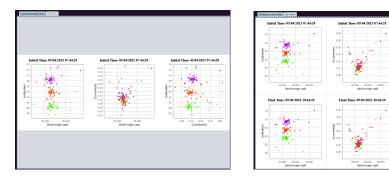


Figure 17. Plots of the orbital elements without and with propagation.

- 4.5.2. 3D Orbit Vision. This page allows to visualize the position of the generated fragments in a 3-dimensional space, whenever the propagation has been computed. An example is given in Figure 18 with the following description, where the blue dots make the contour plot of the Earth surface:
- Top-left: parent body position at break-up time.
- Top-right: fragments' orbits after 3 hours.
- Bottom-left: fragments' position after 6 more hours.
- Bottom-right: fragments' orbits after 3 more hours.

Each panel has additional options to zoom and move the graphs, rotate the plot, reset the zoom, modify the scale settings, save the picture, have a list of Cartesian coordinates (X, Y, Z) of each fragment at a given time, the coordinates of the parent body that gives the orbit, the coordinates of the blue dots representing the Earth. An option allows to decide whether to visualize the fragments, parent body, and the Earth.

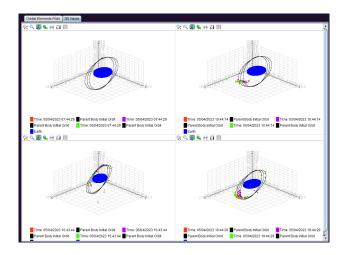


FIGURE 18. 3D plots.

5. OPTION 3: SAVED EXPERIMENTS

If an experiment is saved using Option 1, it can be retrieved using the third option, which allows one to see the elements and plots of the saved experiment, see Figure 19.

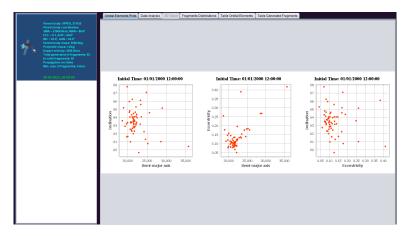


FIGURE 19. Saved experiments.

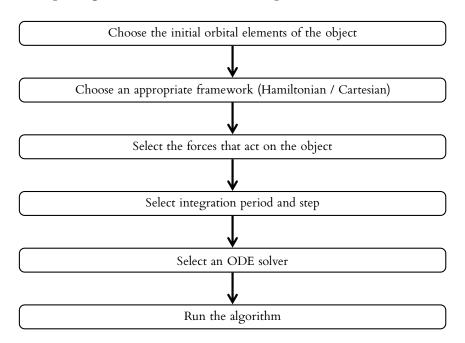
Pressing double click on one element on the left list will open the page from Option 1 using the input data of the selected experiment.

6. Option 4: Orbits' propagator

This option allows one to propagate the orbit of an object around the Earth, taking into account different forces.

The work-flow is described in Section 6.1, the choice of the orbital elements is given in Section 6.2, the selection of the framework, either Cartesian or Hamiltonian, is given in Section 6.3, the selection of the forces is shown in Section 6.4, the choice of the parameters is described in Section 6.5, the integration method can be chose as described in Section 6.6, Section 6.7 describes the different outputs generated by SIMPRO, while Section 6.8 gives the propagation history.

6.1. **Workflow.** This section reproduces the left column that appears when opening SIMPRO for the fourth option.



A summary of the equations and parameters used to propagate the orbits is shown when clicking on Formulas - More Info, see Figure 20.



FIGURE 20. Equations and parameters for the propagation.

- 6.2. Choosing Orbital Elements. Choose Mean elements (TLE), which allows to select the orbital elements to be propagated.
- Semi-major axis in km: a > 6400 km and a < 45000 km.
- Eccentricity $e \in [0, 0.75]$.
- Inclination in degrees: $i \in [0, 180^{\circ}]$.
- Mean anomaly in degrees: $M \in [0, 360^{\circ}]$.
- Argument of perigee in degrees: $\omega \in [0, 360^{\circ}]$.
- Argument of the ascending node in degrees: $\Omega \in [0, 360^{\circ}]$.

An example is given in Figure 21.

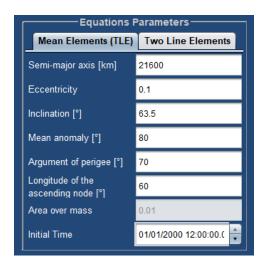


FIGURE 21. Orbital elements of the parent body.

Alternatively, choose Two Line Elements and provide the data in LINE 1 and LINE 2 using the NORAD format from CelesTrak.

- 6.3. Selecting the framework. Select the framework to be used for the propagation of the orbits (see Figure 22).
- (1) For a long-period integration, select Hamilton's equations.
- (2) For a short-period integration, select Cartesian equations.
- (3) For a simplified perturbation model, select SGP4.
- (4) For a comparison between Hamiltonian and Cartesian propagation, select "Hamiltonian vs Cartesian".
- (5) For a comparison between Cartesian and SGP4 propagation, select "Cartesian vs SGP4".



Figure 22. Selection of the integration framework.

- 6.4. **Selecting the forces.** Choose the forces that act on the object, among the following options (see Figure 23).
- (1) Geopotential: J2 and J3 effects.
- (2) Moon and Sun attraction.
- (3) Solar radiation pressure (for a given A/m).
- (4) Drag effect.



Figure 23. Selection of the forces.

- 6.5. Choosing the integration time and step. In the box shown in Figure 24, select the integration period in days and the integration step in minutes.
- (1) In case of Cartesian equations, it is convenient to select a time period between 1 day and 366 days with a time step of 1-5 minutes.
- (2) In case of Hamilton's equations, it is convenient to select a time period between 1 year and 500 years with a time step of 1-10 days.



Figure 24. Select the integration options.

6.6. Choosing the integration method. In the box shown in Figure 25, select the integration method: Runge-Kutta or Adams-Bashforth-Moulton.



FIGURE 25. Integration method.

6.7. **Output.** The output of the integration consists in plotting the time evolution of the orbital elements $a, e, i, M, \omega, \Omega$, see Figure 27. For each run, one experiment page is created.

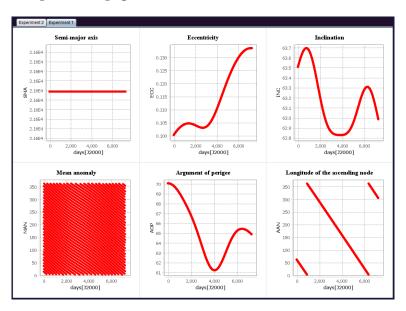


FIGURE 26. Output of the propagation.

6.8. **Propagation history.** The propagator stores up to 10 experiments on a single session (see Figure 27). Once the application is closed, all the experiments are erased.



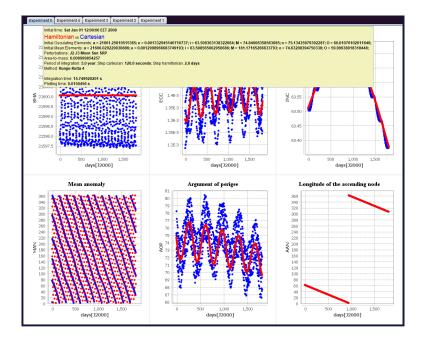


FIGURE 27. Different saved experiments.

7. OPTION 5: USER GUIDE PAGE

The option n. 5 allows the user to get access to an information page containing slides to illustrate graphically the overall program and each option, see Figure 29.

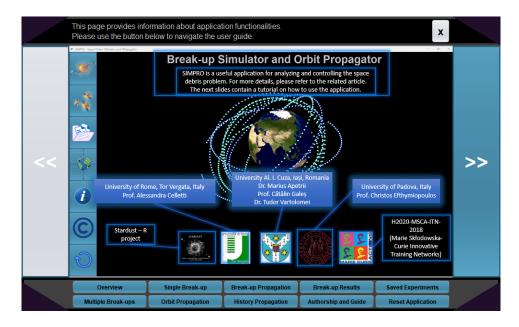


FIGURE 28. User guide page.

8. Option 5: Copyright page

This page displays the detail of the related article and authorship of SIMPRO, see Figure 29.



FIGURE 29. Copyright and contact page.

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9. OPTION 6: RELOAD APPLICATION BUTTON

The last option reloads the program, removing any existing experiment from Option 1 and Option 2 and erasing all experiments from Option 4, see Figure 30.



FIGURE 30. The reload button.

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- [1] AAVV, NASA Standard Breakup Model 1998 Revision, prepared by Lockheed Martin Space Mission Systems & Services for NASA, July 1998.
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- [3] N. L. Johnson, P. H. Krisko, J.-C. Lieu, and P. D. Am-Meador, NASA's new breakup model of EVOLVE 4.0, Adv. Space Res., 28, n. 9, 1377-1384 (2001)

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