

1           Hypothesis-Conditioned Forecast of Hubble-Tension Relief  
2           Assuming the GWTC-3 Dark-Siren Propagation Signal is Physical

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4           February 10, 2026

5           **Abstract**

6     This work presents hypothesis-conditioned forecasts for the Hubble tension under the assumption  
7     that the O3 dark-siren modified-propagation anomaly is physical. The upstream anomaly  
8     analysis, archived on Zenodo (DOI: 10.5281/zenodo.18584705), reports  $\Delta\text{LPD}_{\text{tot}} \simeq +3.67$ ,  
9     whereas a GR-truth injection calibration gives mean  $-0.839$ , standard deviation  $0.240$ , and  
10    maximum  $+0.076$  over 512 realizations. Posterior draws from the reconstructed modified-gravity  
11    model are propagated into late-time anchor observables, CMB-lensing forecasts, and compressed  
12    early-universe inversions.

13    In the constrained, repeatability-calibrated endpoint, the preferred anchor-based relief posterior  
14    is moderate:  $\mathcal{R}_{\text{anchor}}^{\text{GR}}$  has mean  $0.246$  with  $p_{16}/p_{50}/p_{84} = 0.205/0.239/0.277$ , and the  
15    local-versus-high- $z$  GR gap is typically  $\sim 1.17\sigma$ . A joint SN+BAO+CC transfer-bias fit with O3  
16    metadata yields  $\log \text{BF}_{\text{transfer/no-transfer}} = -0.533$ , indicating no preference for explicit transfer  
17    terms in this setup.

18    For CMB signatures, CAMB-based propagation to Planck 2018 lensing bandpowers predicts  
19    suppressed lensing power, with median shifts of about  $-14.9\%$  near  $L \simeq 100$  and  $-8.4\%$  near  
20     $L \simeq 300$  in a direct 16-draw pilot. Compressed  $\theta_*$  inversion under GR assumptions raises inferred  
21     $H_0$  relative to model truth: median inferred  $H_0 \simeq 72.55$  (fixed  $\Omega_m$ ) or  $75.34$  (lensing-proxy  $\Omega_m$ ).  
22    Under the physical-signal hypothesis, partial tension relief is plausible, but decisive resolution  
23    requires full Boltzmann-level modified-gravity refits and independent siren data.

24    **1 Motivation and Scope**

25    The present analysis is conditional: it does not re-establish the O3 anomaly detection claim, but asks  
26    what follows if that signal is physical. The O3 anomaly replication repository is public on Zenodo  
27    (DOI: 10.5281/zenodo.18584705) and provides the calibrated baseline used here. The follow-up  
28    objective is to convert that hypothesis into quantitative predictions for:

- 29       1. late-time inferred- $H_0$  behavior,
- 30       2. transfer-bias robustness across SN+BAO+CC,
- 31       3. CMB-lensing and compressed early-universe inference shifts under GR interpretation.

32    **2 Forecast Definitions**

33    Posterior draws are taken from `outputs/finalization/highpower_multistart_v2/M0_start101`  
34    and propagated through synthetic anchor and CMB inference pipelines.

<sup>35</sup> **2.1 Late-time relief metrics**

<sup>36</sup> Define the baseline local-versus-Planck gap

$$\Delta H_0^{\text{base}} \equiv |H_0^{\text{local}} - H_0^{\text{Planck}}|, \quad (1)$$

<sup>37</sup> and the posterior-gap relief fraction

$$\mathcal{R}_{\text{post}} \equiv 1 - \frac{|H_{0,\text{MG}}^{\text{p50}} - H_0^{\text{local}}|}{\Delta H_0^{\text{base}}}. \quad (2)$$

<sup>38</sup> The preferred estimator is anchor-based. For each anchor redshift  $z_a$ , a synthetic  $H(z_a)$  is  
<sup>39</sup> generated under model truth and inverted with GR assumptions:

$$H_{0,\text{GR}}(z_a) = \frac{H_{\text{obs}}(z_a)}{\sqrt{\Omega_{m0}^{\text{GR}}(1+z_a)^3 + (1-\Omega_{m0}^{\text{GR}})}}. \quad (3)$$

<sup>40</sup> The anchor-averaged relief statistic is

$$\mathcal{R}_{\text{anchor}}^{\text{GR}} \equiv 1 - \frac{|\overline{H}_{0,\text{GR}} - H_0^{\text{local}}|}{\Delta H_0^{\text{base}}}. \quad (4)$$

<sup>41</sup> **2.2 CMB-focused tests**

<sup>42</sup> Two CMB-oriented tests are used:

<sup>43</sup> 1. draw-level propagation of  $(H_0, \Omega_{m0}, \Omega_{k0}, \sigma_8)$  to Planck 2018 lensing bandpowers (template-proxy  
<sup>44</sup> and direct CAMB modes),

<sup>45</sup> 2. compressed early-universe inversion using  $\theta_\star = r_d/D_M(z_\star)$  under GR assumptions, with  
<sup>46</sup> alternative assumptions for inferred  $\Omega_m$ .

<sup>47</sup> These tests target inference shifts and predicted signatures; they are not full TT/TE/EE Boltzmann  
<sup>48</sup> likelihood refits.

<sup>49</sup> **3 Results**

<sup>50</sup> **3.1 Late-time forecast and robustness**

<sup>51</sup> Using  $z_a = \{0.2, 0.35, 0.5, 0.62\}$ , 20,000 Monte Carlo replicates per anchor, and reference values  
<sup>52</sup>  $H_0^{\text{local}} = 73.0 \pm 1.0$  and  $H_0^{\text{Planck}} = 67.4 \pm 0.5$ :

<sup>53</sup> • model-truth posterior gives  $H_0^{\text{p50}} \simeq 70.39$  (p16/p84 = 67.70/73.39),

<sup>54</sup> • single-run posterior-gap relief is  $\mathcal{R}_{\text{post}} \simeq 0.534$ ,

<sup>55</sup> • constrained endpoint gives  $\mathcal{R}_{\text{anchor}}^{\text{GR}}$  mean 0.246 with p16/p50/p84 = 0.205/0.239/0.277,

<sup>56</sup> • typical local-versus-high- $z$  GR gap significance is  $\sim 1.17\sigma$ .

<sup>57</sup> The joint transfer-bias fit over SN+BAO+CC (with O3 support as metadata) yields

$$\log \text{BF}_{\text{transfer/no-transfer}} = -0.533, \quad (5)$$

<sup>58</sup> so explicit transfer terms are not preferred by these data in this configuration.

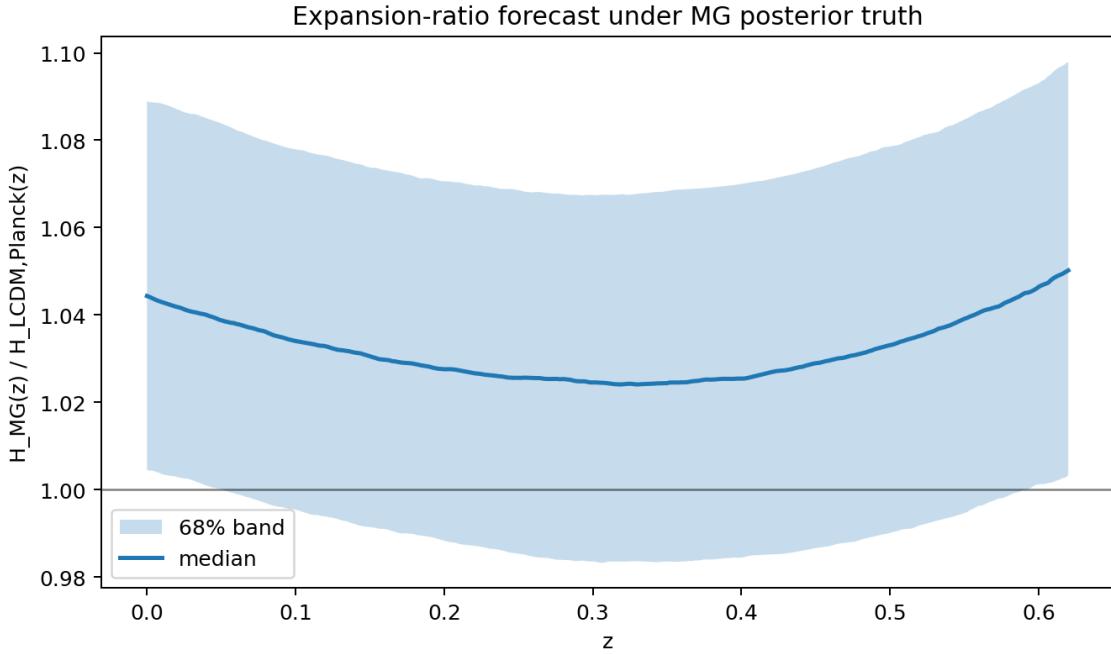


Figure 1: Forecasted expansion-ratio envelope under model truth:  $H_{\text{MG}}(z)/H_{\Lambda\text{CDM},\text{Planck}}(z)$ .

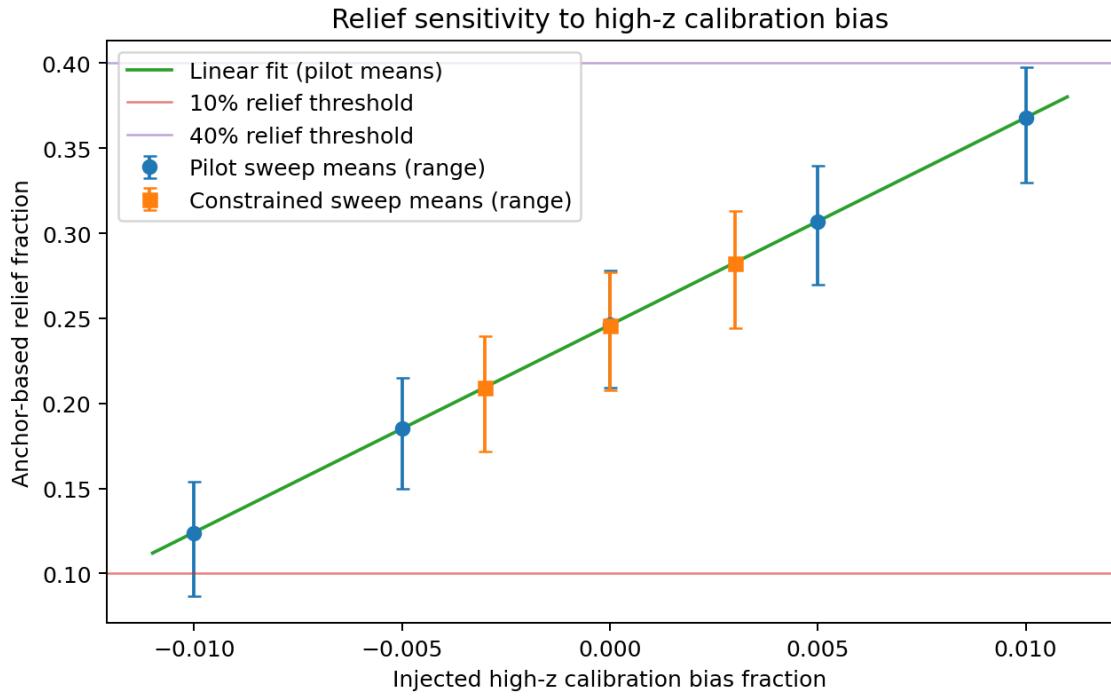


Figure 2: Anchor-based relief sensitivity to injected high- $z$  calibration bias (pilot and constrained sweeps).

<sup>59</sup> **3.2 CMB lensing signature forecast**

<sup>60</sup> The direct CAMB pilot run (16 posterior draws, Planck 2018 lensing bandpowers) finds median  
<sup>61</sup> suppression of the lensing spectrum relative to the Planck-reference model:

$$\left. \frac{C_L^{\phi\phi}(\text{MG})}{C_L^{\phi\phi}(\text{Planck ref})} \right|_{L \approx 100} \simeq 0.851, \quad \left. \frac{C_L^{\phi\phi}(\text{MG})}{C_L^{\phi\phi}(\text{Planck ref})} \right|_{L \approx 300} \simeq 0.916. \quad (6)$$

<sup>62</sup> In the same pilot, only 12.5% of draws outperform the Planck-reference model in lensing-bandpower  
<sup>63</sup>  $\chi^2$ .

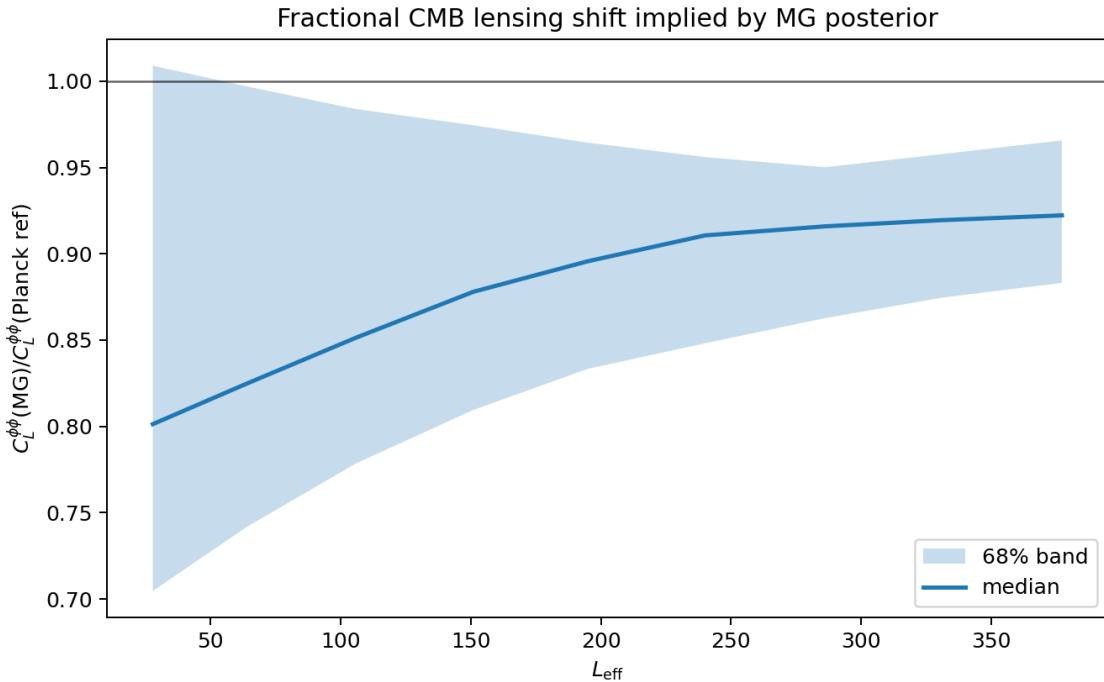


Figure 3: Predicted CMB lensing ratio  $C_L^{\phi\phi}(\text{MG})/C_L^{\phi\phi}(\text{Planck ref})$  from the direct CAMB pilot.

<sup>64</sup> **3.3 Early-universe GR mis-inference test**

<sup>65</sup> Compressed  $\theta_*$  inversion under GR assumptions gives systematic upward shifts in inferred  $H_0$   
<sup>66</sup> relative to model-truth draws:

- <sup>67</sup> • fixed-Planck  $\Omega_m$  assumption: inferred  $H_0$  mean 71.72, p50 72.55, mean  $\Delta H_0 \equiv H_0^{\text{inf}} - H_0^{\text{true}} \simeq$   
<sup>68</sup> +1.14 km s<sup>-1</sup> Mpc<sup>-1</sup>,
- <sup>69</sup> • lensing-proxy  $\Omega_m$  assumption: inferred  $H_0$  mean 75.15, p50 75.34, mean  $\Delta H_0 \simeq +4.57$  km s<sup>-1</sup>  
<sup>70</sup> Mpc<sup>-1</sup>.

<sup>71</sup> The implied sound-horizon shift required to force exact local- $H_0$  matching is modest in central  
<sup>72</sup> tendency but broad in distribution: mean  $\Delta r_d/r_d \approx -1.74\%$  (fixed- $\Omega_m$  mode) or +2.92% (lensing-  
<sup>73</sup> proxy mode).

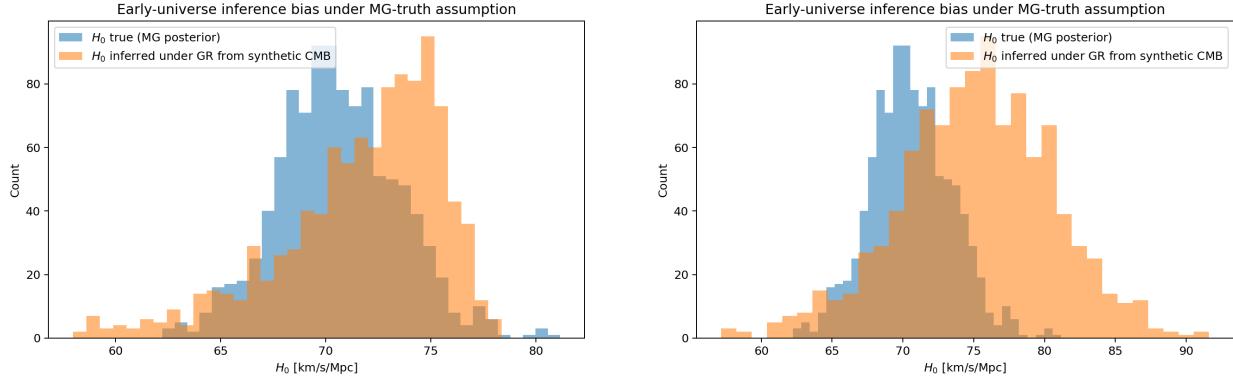


Figure 4: Histogram-level comparison of true and GR-inferred  $H_0$  in compressed early-universe inversion tests. Left: fixed- $\Omega_m$  assumption. Right: lensing-proxy- $\Omega_m$  assumption.

## 74 4 Interpretation

75 Under the physical-signal hypothesis, the model predicts upward pressure on GR-inferred early/high-  
 76  $z$   $H_0$  and non-negligible local-versus-high- $z$  tension relief, but not a standalone full resolution. The  
 77 late-time constrained endpoint remains moderate, and CMB-facing tests indicate detectable but  
 78 model-dependent signatures that require full perturbation-level treatment for definitive statements.

## 79 Reproducibility

80 Core scripts used in this follow-up:

- 81 • `scripts/run_hubble_tension_mg_forecast.py`
- 82 • `scripts/run_hubble_tension_mg_forecast_robustness_grid.py`
- 83 • `scripts/run_hubble_tension_bias_transfer_sweep.py`
- 84 • `scripts/run_hubble_tension_final_relief_posterior.py`
- 85 • `scripts/run_joint_transfer_bias_fit.py`
- 86 • `scripts/run_hubble_tension_cmb_forecast.py`
- 87 • `scripts/run_hubble_tension_early_universe_bias.py`

## 88 Data Availability and DOIs

89 The follow-up uses posterior products from the O3 anomaly pipeline and public cosmology datasets.  
 90 Data provenance and DOIs are:

- 91 • O3 modified-gravity tension replication repository (Zenodo): DOI 10.5281/zenodo.18584705.
- 92 • O3 search-sensitivity injection data used in upstream calibration (Zenodo): DOI 10.5281/zen-  
 93 odo.7890437.
- 94 • GWTC-3 catalog paper: DOI 10.1103/PhysRevX.13.041039.

- Pantheon+ cosmology constraints: DOI 10.3847/1538-4357/ac8e04.
- SH0ES local- $H_0$  reference: DOI 10.3847/2041-8213/ac5c5b.
- SDSS DR12 BOSS consensus BAO (source of `sdss_DR12Consensus_bao.dat`): DOI 10.1093/mnras/stx721.
- eBOSS DR16 cosmological compilation (source class for `sdss_DR16_LRG_BAO_DMDH.dat`): DOI 10.1103/PhysRevD.103.083533.
- DESI 2024 BAO cosmological constraints (source class for `desi_2024_gaussian_bao_ALL_GCcomb_mean.txt`): DOI 10.1088/1475-7516/2025/02/021.
- Cosmic-chronometer compilation components used in `Hz_BC03_all.dat`: DOIs 10.1088/1475-7516/2012/08/006, 10.1103/PhysRevD.71.123001, and 10.1088/1475-7516/2010/02/008.
- Planck 2018 cosmological parameters and lensing references: DOIs 10.1051/0004-6361/201833910 and 10.1051/0004-6361/201833886.

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