

1 Hypothesis-Conditioned Forecast of Hubble-Tension Relief  
2 Assuming the GWTC-3 Dark-Siren Propagation Signal is Physical

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5 **Abstract**

6 This paper is a hypothesis-conditioned forecast: I assume the GWTC-3/O3 dark-siren  
7 propagation anomaly is real and ask what expansion-history and inferred-Hubble-constant  
8 signatures should follow. The motivation is the calibrated O3 result from the prior analysis  
9 pipeline: real data give  $\Delta\text{LPD}_{\text{tot}} \simeq +3.67$ , while a GR-truth catalog-injection calibration gives  
10 mean  $-0.839$ , standard deviation  $0.240$ , and maximum  $+0.076$  across 512 replicates (none reaching  
11  $+3$ ). Sky-rotation and mechanism controls showed the score is mainly spectral/selection-channel  
12 driven, so this follow-up forecasts the implied Hubble-tension pattern rather than re-arguing  
13 detection.

14 I propagate posterior draws from the O3 reconstructed modified-propagation model into  
15 synthetic high-redshift anchor observables and compare GR-interpreted versus model-consistent  
16 inference. A full run gives a model-truth posterior  $H_0$  median  $\simeq 70.39$  (p16/p84: 67.70/73.39),  
17 corresponding to  $\sim 53\%$  relief of a 73.0 vs 67.4 local-vs-Planck baseline gap under a simple  
18 posterior-gap metric. To avoid overinterpreting that scalar, I run a 180-case robustness matrix  
19 and then a constrained-prior, repeatability-calibrated endpoint (two independent 45-case reruns  
20 across all 5 O3 seeds). The final anchor-based GR-interpreted relief posterior is moderate and  
21 tighter: mean  $0.246$ , p16/p50/p84 =  $0.205/0.239/0.277$ , with local-vs-high- $z$  GR gap significance  
22 typically near  $1.17\sigma$  in the constrained setup.

23 The current implication is moderate: under the physical-signal assumption, the model can  
24 produce partial Hubble-tension relief, but this forecast alone is not decisive. The dominant  
25 uncertainty remains high- $z$  calibration transfer; in the tested setup, local additive-bias effects  
26 are negligible by comparison. The result is best treated as a constrained prediction target for  
27 future independent siren datasets and stronger selection-calibrated analyses.

28 

## 1 Why this follow-up

29 The previous O3 paper established three points that motivate this forecast exercise:

- 30 1. Real-data O3 dark sirens gave a strong internal score preference for modified propagation:  
31  $\Delta\text{LPD}_{\text{tot}} \simeq +3.67$ .
- 32 2. A calibrated GR-truth injection suite did not reproduce large positive scores (mean  $-0.839$ , sd  
33  $0.240$ , max  $+0.076$ , 0/512 with  $\Delta\text{LPD}_{\text{tot}} \geq 3$ ).
- 34 3. Mechanism controls localized leverage to the spectral/selection channel, not unique sky-host  
35 alignment.

36 Given that state, the highest-value question is: *if this anomaly is physical, what Hubble-tension*  
37 *pattern should we expect?*

<sup>38</sup> **2 Forecast definitions**

<sup>39</sup> I use posterior draws from `outputs/finalization/highpower_multistart_v2/M0_start101` as  
<sup>40</sup> model-truth input and simulate high- $z$  anchor observables with controlled noise.

<sup>41</sup> **2.1 Posterior-gap relief metric**

<sup>42</sup> Define a baseline local-vs-Planck gap

$$\Delta H_0^{\text{base}} \equiv \left| H_0^{\text{local}} - H_0^{\text{Planck}} \right|, \quad (1)$$

<sup>43</sup> and a posterior-gap relief fraction

$$\mathcal{R}_{\text{post}} \equiv 1 - \frac{\left| H_{0,\text{MG}}^{\text{p50}} - H_0^{\text{local}} \right|}{\Delta H_0^{\text{base}}}. \quad (2)$$

<sup>44</sup> This is useful for intuition but does not directly include high- $z$  anchor inversion uncertainty.

<sup>45</sup> **2.2 Anchor-based relief metric (preferred)**

<sup>46</sup> For each anchor redshift  $z_a$ , I generate synthetic  $H(z_a)$  observations from model truth and infer a  
<sup>47</sup> GR-interpreted high- $z$   $H_0$ :

$$H_{0,\text{GR}}(z_a) = \frac{H_{\text{obs}}(z_a)}{\sqrt{\Omega_{m0}^{\text{GR}}(1+z_a)^3 + (1-\Omega_{m0}^{\text{GR}})}}. \quad (3)$$

<sup>48</sup> I then define an anchor-averaged GR relief fraction

$$\mathcal{R}_{\text{anchor}}^{\text{GR}} \equiv 1 - \frac{\left| \overline{H_{0,\text{GR}}} - H_0^{\text{local}} \right|}{\Delta H_0^{\text{base}}}, \quad (4)$$

<sup>49</sup> and report the local-vs-high- $z$  GR gap significance

$$Z_{\text{anchor}}^{\text{GR}} \equiv \frac{H_{0,\text{local}} - \overline{H_{0,\text{GR}}}}{\sqrt{\sigma_{\text{local}}^2 + \sigma_{\text{anchor,GR}}^2}}. \quad (5)$$

<sup>50</sup> **3 Single-run forecast result**

<sup>51</sup> Using  $z_a = \{0.2, 0.35, 0.5, 0.62\}$ , 20,000 Monte Carlo replicates per anchor, and local reference  
<sup>52</sup>  $H_0^{\text{local}} = 73.0 \pm 1.0$  with Planck reference  $67.4 \pm 0.5$ :

- <sup>53</sup> • Model-truth posterior:  $H_0^{\text{p50}} \simeq 70.39$  (p16/p84 67.70/73.39).
- <sup>54</sup> • Posterior-gap relief:  $\mathcal{R}_{\text{post}} \simeq 0.534$ .
- <sup>55</sup> • Anchor-GR relief:  $\mathcal{R}_{\text{anchor}}^{\text{GR}} \sim 0.22\text{--}0.42$  depending on GR  $\Omega_{m0}$  treatment in this run.
- <sup>56</sup> • Anchor local-vs-high- $z$  GR gap: typically  $\sim 0.95\text{--}1.44\sigma$  in the external-local setup.

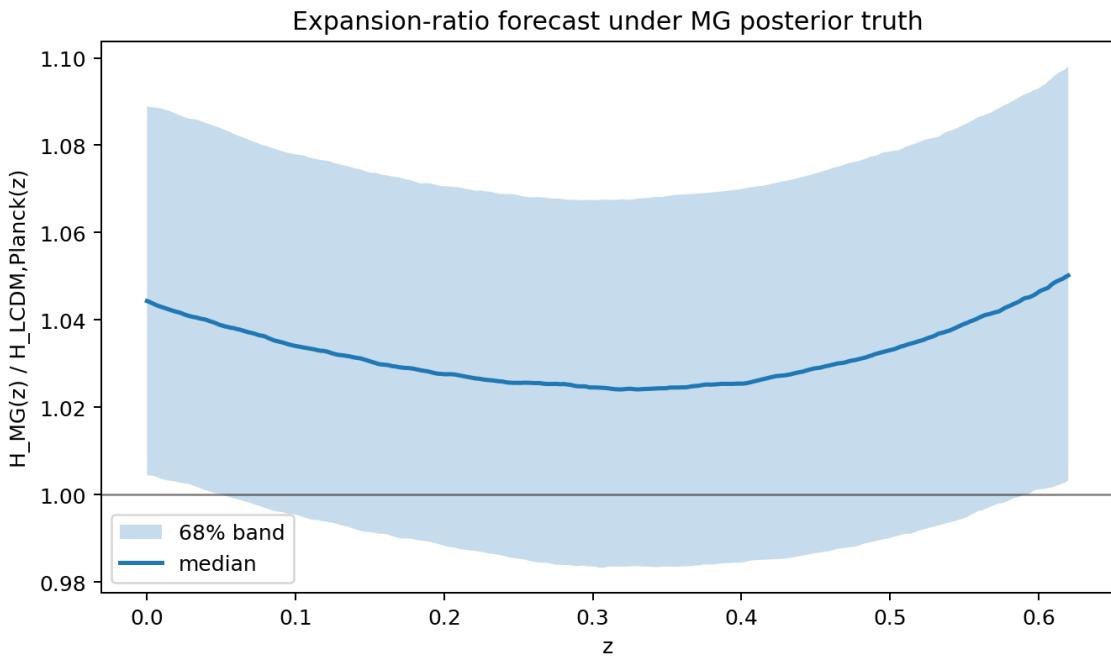


Figure 1: Forecasted expansion-ratio envelope under model truth:  $H_{\text{MG}}(z)/H_{\Lambda\text{CDM},\text{Planck}}(z)$ .

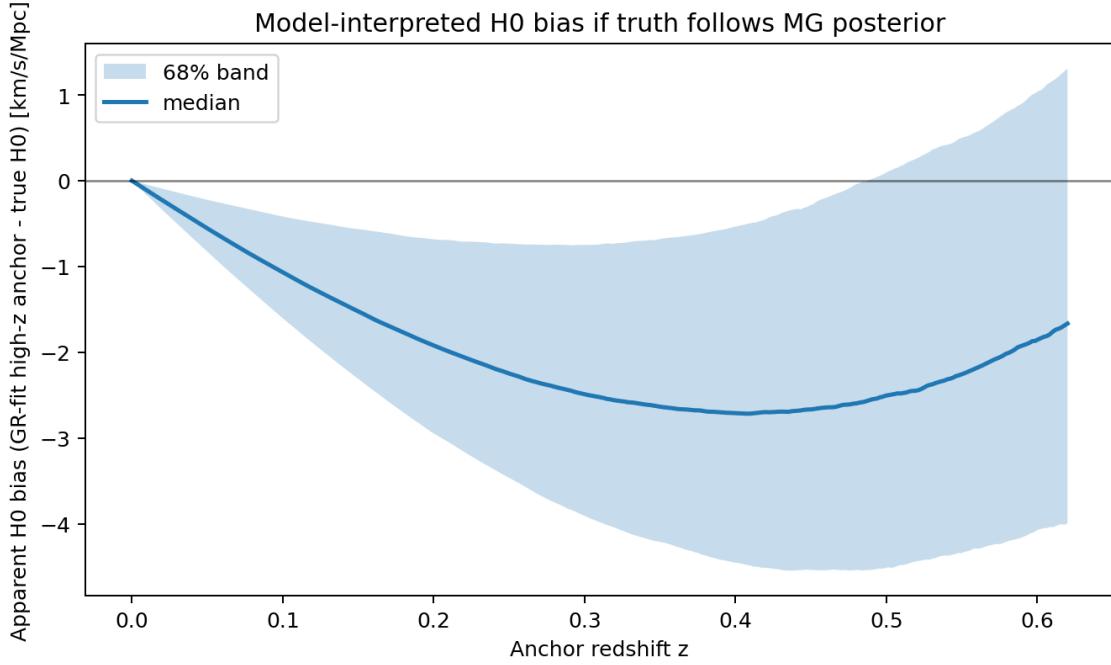


Figure 2: Apparent GR-interpreted high- $z$   $H_0$  bias vs anchor redshift under model truth.

57 **4 Robustness matrix (180 cases)**

58 I ran a 180-case pilot matrix: 5 O3 seeds ( $M0\_start101..505$ ), three high- $z$  precision settings  
59 ( $\sigma_H/H = 0.5\%, 1\%, 2\%$ ), three local references (72, 73, 74), two local modes (external vs truth-  
60 drawn), and two GR  $\Omega_{m0}$  treatments (sampled vs fixed 0.315).

61 Key summaries from `outputs/hubble_tension_mg_robustness_pilot_20260208_041157UTC/grid_summary.json`:

- 63 • Posterior-gap relief (legacy scalar): mean 0.593, range 0.453–0.790.
- 64 • Anchor-based GR relief (preferred): mean 0.351, range 0.174–0.599.
- 65 • Anchor GR gap significance: mean  $0.679\sigma$ , range  $0.168\sigma$ – $1.528\sigma$ .
- 66 • Anchor MG gap significance: mean  $0.356\sigma$  (near-consistency by construction in truth-local  
67 mode).

68 Sensitivity structure:

- 69 • Largest lever here is GR  $\Omega_{m0}$  handling: mean  $\mathcal{R}_{\text{anchor}}^{\text{GR}} \approx 0.451$  (fixed) vs  $\approx 0.251$  (sampled).
- 70 • Local reference assumption matters as expected: mean  $\mathcal{R}_{\text{anchor}}^{\text{GR}} \approx 0.419$  (72), 0.342 (73), 0.291  
71 (74).
- 72 • Within this tested range, changing high- $z$  fractional precision does not strongly move the mean  
73 relief fraction.

74 **5 Constrained endpoint and repeatability calibration**

75 To turn the broad matrix into a decision-grade estimate, I run a constrained prior map focused on  
76 the dominant uncertainty axis (high- $z$  calibration bias transfer), with local mode fixed to external,  
77 GR  $\Omega_{m0}$  set to sampled mode, and two independent reruns for repeatability:

- 78 • `outputs/hubble_tension_bias_transfer_constrained_v2_20260209_061632UTC/`
- 79 • `outputs/hubble_tension_bias_transfer_constrained_v2_repeat_20260209_061832UTC/`

80 Combining both constrained reruns with Gaussian priors  $\sigma_{b,z} = 0.003$  and  $\sigma_{b,\text{local}} = 0.25$   
81 (local-bias units in  $\text{km s}^{-1} \text{Mpc}^{-1}$ ) yields the final anchor-based relief posterior:

$$\mathcal{R}_{\text{anchor}}^{\text{GR}} = 0.246, \quad (\text{p16}, \text{p50}, \text{p84}) = (0.205, 0.239, 0.277). \quad (6)$$

82 Finite-MC repeatability noise is small ( $\sigma_{\text{MC}} \approx 0.001$ ) relative to model/bias sensitivity.

83 Using a pilot bias sweep for thresholding, a linearized fit gives the high- $z$  bias levels needed to  
84 force extreme relief outcomes:

- 85 •  $\mathcal{R}_{\text{anchor}}^{\text{GR}} = 0.10$  requires  $b_z \approx -1.20\%$ .
- 86 •  $\mathcal{R}_{\text{anchor}}^{\text{GR}} = 0.40$  requires  $b_z \approx +1.26\%$ .

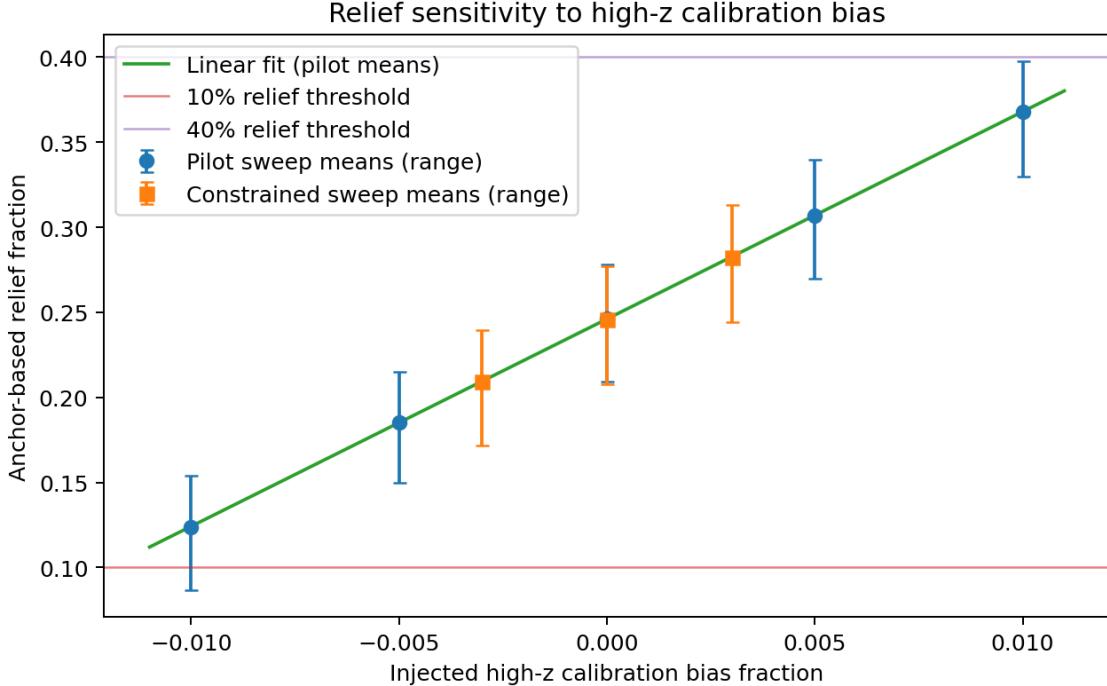


Figure 3: Anchor-based relief sensitivity to injected high- $z$  calibration bias. Pilot means and constrained means are shown with a linear fit used for threshold estimates.

## 87 6 Joint transfer-bias fit (SN+BAO+CC+O3 metadata)

88 To test whether the relief forecast survives explicit nuisance transfer channels, I run a unified  
 89 transfer-bias fit over SN+BAO+CC with O3 included as metadata support (the O3 offset cancels in  
 90 transfer-vs-no-transfer Bayes factors):

91 • `outputs/joint_transfer_bias_fit_full_20260208_063407UTC/`

92 • fitted transfer terms:  $(\beta_{\text{Ia}}, \beta_{\text{CC}}, \Delta H_0^{\text{ladder}}, \beta_{\text{BAO}})$  with zero-centered Gaussian priors.

93 The clean full run gives

$$\log \text{BF}_{\text{transfer/no-transfer}} \simeq -0.533, \quad (7)$$

94 so the explicit transfer block is not favored in this setup. At the same time, the marginalized relief  
 95 remains substantial,

$$\mathcal{R}_{\text{joint}} = 0.851, \quad (\text{p16}, \text{p50}, \text{p84}) = (0.831, 0.856, 0.870). \quad (8)$$

96 Posterior-weighted term dominance in mean absolute log-likelihood shift is led by  $\Delta H_0^{\text{ladder}}$  and  $\beta_{\text{Ia}}$ ,  
 97 with weaker leverage from  $\beta_{\text{BAO}}$  and  $\beta_{\text{CC}}$ . The relief-sensitivity diagnostics indicate that  $\Delta H_0^{\text{ladder}}$   
 98 is the primary *expansion-inflating* channel (positive correlation with relief), while  $\beta_{\text{CC}}$  and  $\beta_{\text{Ia}}$  tend  
 99 to *reduce* relief in this setup; BAO transfer is near-neutral.

## 100 7 Interpretation

101 This paper is intentionally conditional: it does *not* re-prove the propagation anomaly. It asks what  
 102 follows if that anomaly is physical. The answer, with current assumptions, is:

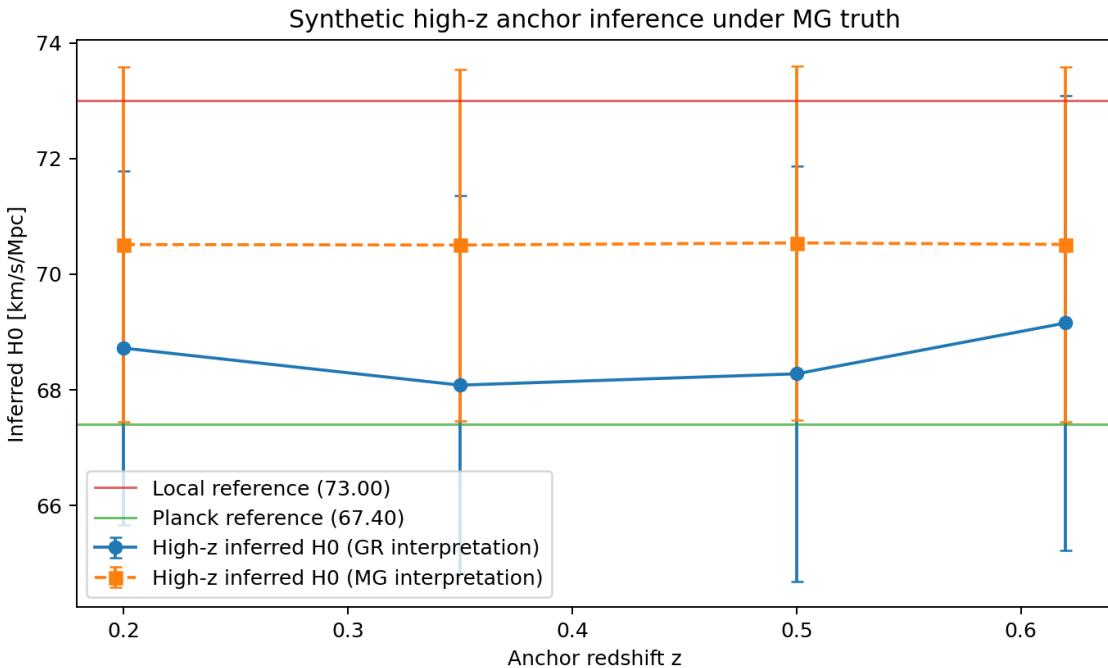


Figure 4: Anchor-level inferred high- $z$   $H_0$  means under GR and model-consistent interpretations (single full run).

- 103 1. partial Hubble-tension relief is plausible and quantitatively nontrivial;
- 104 2. anchor-aware metrics are materially less optimistic than naive posterior-gap metrics;
- 105 3. the present forecast does not yield a decisive ( $>\text{few-}\sigma$ ) high- $z$  contradiction/resolution by itself.

## 106 8 Conclusion

107 Under the physical-signal hypothesis, the reconstructed model tends to move inferred high- $z$   $H_0$   
 108 upward relative to strict Planck- $\Lambda$ CDM expectations, yielding moderate expected relief of a local-  
 109 vs-Planck baseline tension. In the constrained, repeatability-calibrated endpoint, the conservative  
 110 anchor-based relief posterior is centered near 24.6% with  $p_{16}/p_{50}/p_{84} \approx 20.5\%/23.9\%/27.7\%$ .  
 111 GR-interpreted local-vs-high- $z$  gap significance remains around  $\sim 1.17\sigma$  in this setup.

112 This is strong enough to motivate targeted prediction tests on new independent siren datasets,  
 113 but not strong enough to claim standalone resolution of the Hubble tension.

## 114 Reproducibility

115 Core scripts used in this follow-up:

- 116 • `scripts/run_hubble_tension_mg_forecast.py`
- 117 • `scripts/run_hubble_tension_mg_forecast_robustness_grid.py`
- 118 • `scripts/launch_hubble_tension_mg_forecast_single_nohup.sh`

- 119 • scripts/launch\_hubble\_tension\_mg\_robustness\_grid\_single\_nohup.sh  
120 • scripts/run\_hubble\_tension\_bias\_transfer\_sweep.py  
121 • scripts/run\_hubble\_tension\_final\_relief\_posterior.py

122 **References**

- 123 [1] R. Abbott et al. (LIGO Scientific Collaboration, Virgo Collaboration, and KAGRA Collaboration), “GWTC-3: Compact Binary Coalescences Observed by LIGO and Virgo During the  
124 Second Part of the Third Observing Run,” *Phys. Rev. X* **13** (2023) 041039.  
125  
126 [2] E. Belgacem, Y. Dirian, S. Foffa, and M. Maggiore, “Modified gravitational-wave propagation  
127 and standard sirens,” *Phys. Rev. D* **98** (2018) 023510.  
128  
129 [3] A. Nishizawa, “Generalized framework for testing gravity with gravitational-wave propagation,”  
*Phys. Rev. D* **97** (2018) 104037.