

Sheet 1

Lena Rathmann 370357
Max Stramb 355175

$$1) 1 \text{ dyn} = 1 \frac{\text{g cm}}{\text{s}^2} = 10^{-3} \text{ kg} \cdot 10^{-2} \text{ m} \cdot \frac{1}{\text{s}^2} = 10^{-5} \text{ N} \quad \checkmark$$

$$1 \text{ erg} = 1 \frac{\text{cm}^2 \text{ g}}{\text{s}^2} = 10^{-3} \text{ kg} (10^{-2} \text{ m})^2 \frac{1}{\text{s}^2} = 10^{-7} \text{ J} \quad \checkmark$$

$$1 \frac{\text{erg}}{\text{s}} = 10^{-7} \frac{\text{J}}{\text{s}} = 10^{-7} \text{ W} \quad \checkmark$$

2

3.) $\Omega_0 = 0,3036$

a.) $H_0 = 68,14$ 2.1

z	Mpc
0,01	44,58
0,1	476,48
1	6775,56

\checkmark

-0,5

2.2

Up to about $z=0,1$ the linear Hubble law still produces 2.3 reliable results.

home.fnal.gov/~gnedin/cc/

b.) $1 \text{ kpc} = 3262 \text{ ly}$

$1 \text{ Mpc} = 3262 \cdot 10^3 \text{ ly} ?$

(from Wolfram alpha)

\checkmark

$$c.) \frac{c}{H_0} \cdot z = 2359 \text{ Mpc} \quad \boxed{H_0 = 68.14}$$

Luminosity distance: 3168 Mpc ✓

(with H_0 and values from previous exercise)

Proper distance:

$$d_e = \frac{d_L}{(1+z)} = 2062 \text{ Mpc} \quad \checkmark$$

4.5

4)

a) $S_e = \epsilon \cdot \frac{d\Phi(\epsilon)}{d\epsilon}$ ✓

b) $\frac{d\Phi(\epsilon)}{d\epsilon} = \Phi_0 \epsilon^{-\gamma}$ ✓

$\Rightarrow \Phi(\epsilon) = \int_{\epsilon}^{\infty} \Phi_0 \epsilon'^{-\gamma} d\epsilon' = \left[-\frac{\Phi_0}{\gamma-1} \epsilon'^{-\gamma+1} \right]_{\epsilon}^{\infty}$
 $= \frac{\Phi_0}{(\gamma-1)} \epsilon^{-(\gamma-1)}$ ✓

$\Rightarrow \alpha = \gamma - 1, \quad \Phi_0 = \frac{\Phi_0}{(\gamma-1)}$ ✓

c) $S = \int_{\epsilon}^{\infty} \epsilon \cdot \frac{d\Phi(\epsilon)}{d\epsilon} d\epsilon = \int_{\epsilon}^{\infty} (\gamma-1) \Phi(\epsilon) d\epsilon$ ✓ 4.1 -0.3

d) $\psi(\epsilon) = \epsilon \frac{d\Phi(\epsilon)}{d\epsilon} = (\gamma-1) \Phi(\epsilon) = \frac{(\gamma-2)}{\epsilon} S(\epsilon)$ ✓ = S

e) $\left[\frac{d\Phi(\epsilon)}{d\epsilon} \right] = \text{s}^{-1} \text{cm}^{-2} \text{sr}^{-1} \text{GeV}^{-1}$ ✓ describes the flux of particles for a specific energy interval $d\epsilon$. ✓ 4.2

$[\Phi(\epsilon)] = \text{s}^{-1} \text{cm}^{-2} \text{sr}^{-1}$ ✓ describes the flux of particles starting at an energy of ϵ to infinity. 4.3

$[S_e] = \text{s}^{-1} \text{cm}^{-2} \text{sr}^{-1}$ ✓ describes the flux of energy for a specific energy interval $d\epsilon$.

$[S] = \text{s}^{-1} \text{cm}^{-2} \text{sr}^{-1} \text{GeV}$ ✓ describes the integrated flux of energy starting from ϵ to infinity.

$[\psi(\epsilon)] = \text{s}^{-1} \text{cm}^{-2} \text{sr}^{-1}$ ✓ describes a flux of particles. 4.4

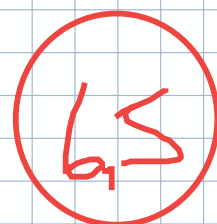
f) 1st case: $[\Phi_0] = \text{s}^{-1} \text{cm}^{-1} \text{sr}^{-1} \text{GeV}^{\gamma-1}$ ✓ -0.5

$[\Phi_0] = \text{s}^{-1} \text{cm}^{-1} \text{sr}^{-1} \text{GeV}^{\gamma-1} = [S_0] = [\psi_0]$ ✓

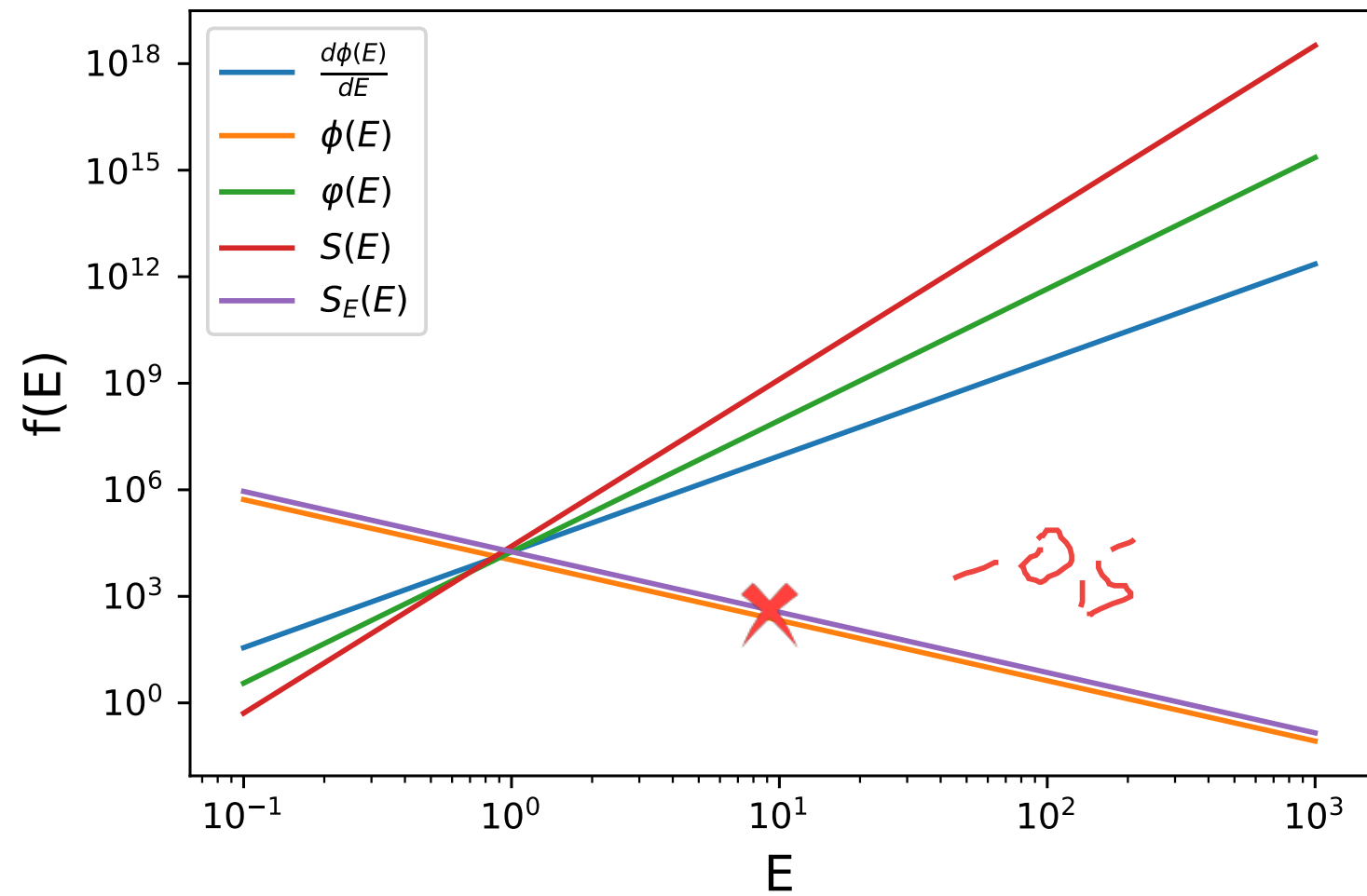
2nd case: $[\Phi_0] = \text{s}^{-1} \text{cm}^{-1} \text{sr}^{-1} \text{GeV}^{-1}$ ✓

$[\Phi_0] = \text{s}^{-1} \text{cm}^{-1} \text{sr}^{-1} = [\psi_0]$ ✓

$[S_0] = \text{s}^{-1} \text{cm}^{-1} \text{sr}^{-1} \text{GeV}$ ✓



g)



Sheet 1, Task 2

Rathmann, Straub

In [37]:

```
import astropy.units as u
```

Define Custom Units:

In []:

```
GU = u.def_unit("GU", 8.2e3 * u.parsec)
SU = u.def_unit("SU", 4.244 * u.lyr)
```

GU is the "Galactic Unit"[1], SU is the "Stellar Unit"[2].

In [34]:

```
dia_milky = 710e3*u.parsec # diameter of milky way [3]
canis_major = 25e3*u.lyr # distance to nearest galaxy, canis major [4]
markarian = 140e6*u.parsec # distance to blazar markarian 501 [5]
quasar = 13.1e9*u.lyr # distance to most distant known quasar ULAS J1342+0928 [6]
```

In [35]:

```
for dist in [dia_milky, canis_major, markarian, quasar]:
    print(dist.to(u.m))
    print(dist.to(GU))
    print(dist.to(SU))
    print("_"*10)
```

```
2.190831082841706e+22 m
86.58536585365853 GU
545643.3274671854 SU
```

```
2.3651826181452e+20 m
0.9347603469144901 GU
5890.66918001885 SU
```

```
4.3199486140540683e+24 m
17073.170731707316 GU
107591642.03578304 SU
```

```
1.2393556919080848e+26 m
489814.4217831928 GU
3086710650.3298774 SU
```



References

1: doi:10.1017/S1743921312021060

2: doi:10.1007/978-3-642-22839-1_10

3: <https://www.sciencenews.org/article/astronomers-have-found-edge-milky-way-size>
(<https://www.sciencenews.org/article/astronomers-have-found-edge-milky-way-size>).

4: https://imagine.gsfc.nasa.gov/features/cosmic/nearest_galaxy_info.html
(https://imagine.gsfc.nasa.gov/features/cosmic/nearest_galaxy_info.html)

5: https://en.wikipedia.org/wiki/Markarian_501 (https://en.wikipedia.org/wiki/Markarian_501) (couldn't find a better source, sorry)

6: <https://www.space.com/39000-oldest-farthest-monster-black-hole-yet.html> (<https://www.space.com/39000-oldest-farthest-monster-black-hole-yet.html>)

In []:

Index der Kommentare

- 2.1 unit
- 2.2 values too large. where are the calculated values?
- 2.3 up to $z=0.4-0.5$

- 4.1 incomplete
- 4.2 differential flux
- 4.3 integrated particle flux
- 4.4 energy weighted particle flux, differential

- 6.1 visible diameter is 30kpc
- 6.2 it is a dwarf galaxy, not a galaxy. the nearest is Andromeda
- 6.3 this is light traveled distance, not proper distance