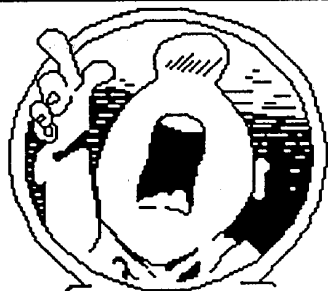


# SJAA EPHEMERIS

VOLUME 6 NUMBER 9 OFFICIAL PUBLICATION OF THE SAN JOSE ASTRONOMICAL ASSOCIATION September 1995



**The Eyepiece**  
by Bob Madden

All is quite on the Western Front. Reports from Rich Neuschaefer about Fremont Peak were sketchy. Some nice telescopes mentioned. The fog did not come in so the viewing was less than excellent. Paul Barton was there and mentioned the "no see-ums" eating him up. That is a hazard at the Peak in the summer; the deer flies. A solution is the use of some "bug-off" (Skin so Soft, DEET, Cutters, etc), but allows the opportunity of some of it getting onto your optics. Here is another thing to try (if you can stand it) reported from Alaska on the internet; pin a sheet of Bounce (the smelly kind) to your shirt collar. I haven't tried it yet, but intend to on my fishing trip to Canada. If anyone tries this suggestion please let me know the results. Skin so Soft is reported to be excellent also. It is distributed through Avon.

We are working diligently this month to publish the Ephemeris early because of vacation plans. Everything is falling into place and it should be finished before vacation time.

The US Postal Service is implementing new rules governing non-profit organizations beginning in October. Actually they were postponed from September of last year. We have submitted a copy of our newsletter for review and beginning in October I have been advised we will not be allowed to have an

**Sept 1:** Hough Park Public Star Party, Sset 7:37 pm, 48% moon, Mset 11:53 pm.

**Sept 2:** No activity, 1st Qtr moon

**Sept 9:** General Meeting 8:00 pm, preceded by Board Mtg at 6:00 pssm. This is our annual Slide/Equipment night. Bring your favorite project and show it to the group.

**Sept 16:** Observational Astronomy Class at Hough Park, 8:00 pm.

**Sept 23:** 3 Star parties: H. Coe, Fremont Peak. Sset 7:02 pm, No moon, also, Halls Valley's public star party at Grant Ranch County Park.

**Sept 29:** Star Party, Hough Park, second this month, Sset 6:54 pm, 34% moon, Mset 10:44 pm.

**Oct 7:** No activity, full moon.

**Oct 14:** General Meeting 8:00 pm, preceded by Board Mtg at 6:15 pm.

**Oct 21:** Observational astronomy class will meet at Fremont Peak, at the 30 inch telescope.

**Oct 21:** Star parties at Coe and the Peak, also HVAG's at Ranch, Sset 6:22 pm, 5% moon rises 5:42 am.

**Oct 27:** Hough park star party, Sset 6:16 pm, 21% moon sets 9:36 pm.

**Oct 28:** 32% moon, too much for a star party.

**October General Meeting will be a dinner at the King's Table (Hamilton/Bascom) to celebrate the SJAA's 40th year. Call Bob Madden 264-4488.**

**If there isn't an active response, this activity will be cancelled**

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Forty Years Ago this month  
by Jim Van Nuland

The September meeting at San Jose State was called to order by President Walt Krumm. Leading off was Steve Bieda, with a short review of the more interesting papers from the recent convention at Yosemite. He was followed by Bob Cunningham, Miss Margaret Krumm, and Walt.

Dr. Geisler presented the constellations of the month, Pegasus and Andromeda. He pointed out many features, including the mythology, location, nebulae, clusters, etc. His discourse on Andromeda was particularly interesting.

Walt Krumm completed his series of talks on telescope mountings, giving us specific details on the best ways to make a tripod. These talks have been very informative, and are greatly appreciated by all members.

The meeting adjourned at 9:45 pm. Most met in Columbus Park afterward and enjoyed an hour of good seeing through the three scopes present.

Tom Nelson, Secretary

Astro Ad column. As I understand it announcements for the benefit of all the members will be allowed and not sale items (for example, individual sale of equipment benefits the buyer and seller only). I will continue to question this ruling for a re-reading up until its implementation. Every thing else generally passed muster.

There is quite a lot of activity going on within our association these days; telescope making, CCD, public star parties, internet and much more. Come to our General meetings and find out. You may meet someone who can help you in your endeavor.

## MOONWATCH

by Don Watson

The article on Tracking Mir/Atlantis in the August SJAA Ephemeris got me to reminiscing about the early days of Satellite tracking.

Back before 1960 I was living in Cincinnati, Ohio and was member of the Cincinnati Astronomical Society. I was working in the "Investigations" department of GE at the time for astronomer Dr. Herb Grosch. This department was the computer programming department for GE's Jet Engine division. They had started out with an IBM 701, and had progressed through the 704 and 709 to the 7090/94.

When we had all been caught flat-footed by the launch of the first Russian satellite, the amateur astronomers were asked to form teams to monitor the skies and keep track of the satellites that were up there. These teams were called "Moonwatch" teams presumably because we were tracking the artificial moons that were being put up. The basic equipment consisted of a wide field telescope roughly the equivalent of half of a 7x50 binocular which was mounted on a small table. Several of these tables were lined up in a row so they would stand a good chance of catching the satellites as they went overhead. (Hopefully, you would have WWV on a radio near by to give accurate times for the transits.) The observations were then turned into a central collection point for processing.

Fortunately for our Moonwatch team there was in one of the local high schools a very bright student, Tom van Flandern, that was really "into" celestial mechanics. (He later got his PhD in Astronomy.) He wrote a Fortran program to track satellites and we managed to get him on board at GE as a summer student. I helped him with his Fortran programming and got him a few hours a week to run off prediction for our observing site. (Tracksat runs off the equivalent predictions in a minute or so on my PC.)

The Cincinnati Astronomical Society had a site out in the country with a 16 or 18 in. Newtonian, a Clark refractor (I think), and some others. With the

aid of the program, we could set up the Newtonian and catch the smaller satellites as they went over.

There was considerable competition with a group in Southern California that had a similar program running on computers at one of the aerospace companies. They often beat us out for the highest number of observations turned in, mainly, I think because they had better weather.

On one of my more interesting evenings at the Newtonian I managed to get two passages of a small satellite that had been classified as lost. However when I got back with my prize catch I had to admit that I had somehow managed to miss a good passage of the brightest satellite up at the time.

My memory is a bit fuzzy about the timing, but I think it was when this operation was wrapped up that we received a very detailed small silver pin showing a row of observers setting at their tables.

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### Henry Coe State Park

by Jim Van Nuland

The July 22 star party at Henry Coe State Park was a success. The grader and back-hoe have been removed, but the water tanks remain. The flip side is that they block some of the stray light.

We had a good turnout, about 10 people, I think, with my 8 inch and Bill O'Shaughnessy's 8 inch the largest apertures. Ken Miura brought the longest scope, his 6 inch f/15 refractor, giving good views of Jupiter. There were a few campers that joined us at dusk, having come by earlier in the day. Jay Freeman came up for a little while. It has been some time since he's been at Coe.

By dark, the temp was down to 56 degrees, but a nasty wind made it seem much colder! Quite a change from 95 when I arrived. I switched to my heavy coat, the one that I bought for Decembers at Coe. With the wind shaking the scopes, it was hard to use high power. But the air was clear, and most of the show objects, M4, M22, M24, M11 were excellent in the moments of stability. Jupiter and Saturn showed only moderately well, the air having gone

soft, even when the wind let up for a moment.

Early on, I noted that the south-east quadrant of Jupiter seemed rather dark, but this was not pursued, unfortunately. It may have been residue of Comet Shoemaker-Levy 9 from a year ago. Perhaps others observed it too?

The highlight of the night was the search and finding of Comet 6P/d'Arrest, predicted at magnitude 9 but seemingly much fainter. After some running back and forth between the two 8 inch scopes, we decided that we were looking at the same stars, and that there was a faint, faint fuzzy there, adjacent to a small "keystone" of faint stars. Watching for motion can fool the eye, so I went off and looked at other things for a while. Bill had already left when I returned to the keystone 1.5 hours later — and the fuzzy was now inside the keystone! The next day, calculations verified the motion at 2.3 arc-minutes per hour, to the south-east, just as seen in the scope.

The wind eventually stopped, and I sought meteors for a little while. There were a few nice ones, but my eyes were tired, and I sacked out about 2 a.m. All in all, a good outing! I hope the Peak was not so windy.

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### Astronomical Humour

by Bob Brauer

Here's a couple of lightly humorous quotes from Isaac Asimov.

What with one thing and another, I've gotten used to explaining various subtle puzzles that arise in connection with the scientific view of the universe. For instance, I have disposed of the matter in which electrons and photons can be waves part of the time and particles the rest of the time in a dozen different ways and by use of a dozen different analogies.

I've gotten so good at it, in fact, that at dinner parties the word nervously goes about, "For heaven's sake, don't ask Asimov anything about wave-particle duality."

And no one ever does. I sit there all primed and aching to explain, and no

**Continued on Page five—see Laughter**

Heartbreak of Optics  
by Bill Marriott

Well, you'd think that after 15 yrs, I'd know what I was doing, but.....

I'm constantly reminding my employees at the optic shop, "When working with Bk-7, be careful using heat, it is thermal sensitive...."

So, I decided to heat up my 14 lb slab of Bk-7 to mount on a aluminum tooling to slice down the middle, so I could end up with two plates, as opposed to milling down a 1.5" thick slab to make a 1/4 or 1/2 inch corrector or window.

Heated both slab and tooling on the same stove, to the same temperature, picked up the tooling to apply mounting wax, must of cooled just enough, so when I attached it to the big slab, It instantly CRACKED in half....like a knife plunged deep into my soul!!

Suddenly a picture comes to mind: Elihu Thompson had poured the big 60" fused silica blank, working with George Hale on trying to figure out how to make a 200" mirror. Someone opened the annealing oven to soon, and when they examined the disc, it was cracked!!

Supposedly, Elihu turned to A.L. Ellis and all he said was "Better Lay down another 60" disc: We have proved we could do it"

Not quite the same analogy, but I pulled myself out of the deep pit of despair, got on the phone to Glass Fab in N.Y., got our salesman on the line and asked if he had anything close to a 9" dia plate of BK-7. (I rather use fused silica, but at 6x the price, can't rationalize it.)

Low and behold, turns out they had a 11" dia x 1/2" thick plate, and a 15" dia x 1/4" thick plate, both for not much more than I paid Ohara for the big slab. So, the cracked slab goes back on the inventory shelf, where it'll become thousands of dollars worth of small optical parts, and the two plates are on there way from Glass Fab....

So, what to do now...

- 1) Have a 8" f5 primary ready for polishing
- 2) Have a 11"x1/2" blank, perfect for a optical window for a 10" mirror (I do have a 10" mirror

- 3) blank sitting around"  
Have a 15"x1/4" blank Bk-7

Obviously, the 11" could be trimmed to 9", but I think that's a waste, and I should use that with the 10" mirror.

The 15" plate has some defects on the edge, but could be trimmed to 13" for use with a 12.5, but by the time its ground and polished, it'll probably be more like .200", which maybe to thin for a schmidt corrector. Besides, I promised the wife to not build anything that size till we're settled down permanently.

Would .200" be too thin for an 9" optical window (plano/plano), I know it doesn't meet the 1/20 ratio, but what would the effects be??

I'll come to some decision at some point, I'm still educating myself in design mathematics, polishing the 8" primary to perfection (hopefully), building a much more accurate knife edge tester, etc, so no hurry...

Comments welcome as always...

Bill Marriott  
Forest Knolls, Ca. USA  
btk@ix.netcom.com

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**METEOR SCATTER - the FAQs**  
Rev. 1, Jul 25, 1995.

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0. About these FAQs

These FAQs are the result of your various questions and remarks to the monthly RMOB (Radio Meteor Observation Bulletin). Constantly amateur astronomers, radio amateurs and schools looking for new projects, get involved in observing meteors by radio methods, and need the same basic information.

0.1. Revision / posting frequency

I am sure there will be many questions and remarks to this first issue. The section with references has to be expanded considerably. And undoubtly, we will hear from existing radio setups for the first time. Rev. 2 can be expected at the end of August.

0.2. Distribution

These FAQs are posted on:  
sci.astro.amateur  
rec.radio.amateur.space  
Compuserve's Astronomy Forum,  
Section 11, Radio Astronomy.

Please let me know if you feel that other newsgroup are interesting too or more appropriate.

0.3. Acknowledgement

Many thanks to Tom Ashcraft (72632.1427@compuserve.com) and other participants to Compuserve's Astronomy Forum.

1. What is meteor scatter (MS) ?

Meteor scatter is a form of electromagnetic wave propagation. The ionized trail of meteors (typically appearing between 100 km and 80 km high) acts temporarily as a reflector for radio waves. A meteor does NOT generate detectable radio emission itself!

In the case of forward scatter, the transmitter and receiver are at different locations. Backscatter can be seen as a special case, whereby the sender and receiver coincide (typically the case of radar).

Meteor scatter can be used day and night, allowing daylight meteor streams to be studied. Meteor scatter can also be used for secure communication.

Historically, meteor streams were conclusively detected by means of radar at the end of World War II.

2. How to observe MS ?

Tune to an "empty" frequency in the VHF band, and wait until a suitable oriented meteor trail reflects a distant broadcast station. You hear a more or less short (a fraction of a second to tens of seconds) fragment of the remote broadcast. Occasionally, more than one station (together or one after the other) can be heard. Routinely, stations of 300 km to 1000 km away can be observed. The problem however in densely populated  
**Continued on page four - see Meteors**

**Meteors** - continued from page three  
areas can be to find a non-locally used frequency.

#### 2.1. Which frequencies to use ?

The lower the frequency, the better... but at too low frequencies, the radio-waves bend and follow the curvature of the Earth. A lower practical limit is 40 MHz. More commonly, the VHF is suitable. VHF Band I ranges from 41 to 68 MHz, and is used for TV. The FM band (88 MHz - 108 MHz) is also used frequently. Try as low as possible in the band, and avoid free radio's in the upper band (104-108 MHz). Not to forget are beacons at e.g. 75 MHz. Radio hams listen also to MS in the 2 m (144 MHz) band. The reflections are shorter and weaker than at the lower frequencies. There is little use in trying MS at still higher frequencies.

#### 2.2. Which equipment to use ?

A regular FM receiver is sufficient. An all-band receiver is of course better in terms of sensitivity and features such as bandwidth selection. Connect a horizontally polarized Yagi antenna with 4 to 6 elements (not too directive) to your receiver. The antenna can be pointed horizontally, or at a certain angle (the latter avoids direct reception from nearby stations). Keep the antenna feed cable short to minimize losses, or add an antenna amplifier. With a good setup, it should be easy to obtain 10 to 20 meteor reflections per hour when no stream is active (the sporadic background). During streams, this number can get as high as several hundreds. Due to the daily variation in meteor activity, more reflections are received during the morning hours (5h to 6h local time) than in the evening (18h). This can serve as a proof that you are actually observing meteors!

#### 2.3. Always possible to observe ?

Unfortunately, other forms of propagation interfere with MS. The worst one is sporadic-E or Es, consisting of conduction clouds in the high atmosphere that make permanent reception of remote stations possible during minutes to tens

of minutes. The sporadic-E season is from May to July in large parts of the northern hemisphere. In some regions however, sporadic-E is unknown!

There are also tropospheric influences. A temperature inversion can also cause reception over wider ranges than normal.

Thunderstorms cause very sharp peaks. On FM, most of these (amplitude modulated) spikes are suppressed. However, DO NOT connect your antenna to your receiver during thunderstorms or when you are absent for longer time. Lightning strikes have ruined several radio shacks.

#### 2.4. One step further: recording meteor reflections

Meteor reflections can easily be recognized by listening. But a better measure of the shape of the reflection can be obtained in accessing the signal strength and recording it, either on a pen recorder, either feeding it into a computer via an analog/digital (A/D) convertor. Special care has to be taken to shield well the computer, power supplies etc., or the computer signal will be more or less noticed in the receiver.

Meteor reflections can broadly be divided in two classes - underdense and overdense - with different reflection profiles. These profiles can be related to the physical characteristics of the meteors.

#### 3. Who else is observing ?

During recent years, amateurs in following countries have been recording meteor reflections: Austria, Belgium, Denmark, Germany, Finland, the Netherlands, the UK, the USA.

Some of them observe only during the periods of the great streams, others run continuous automated stations.

Only a few radars are still used for monitoring streams. One is run by O. Belkovitch (oleg@astro.kazan.su), at Kazan University.

#### 4. How to relate radio- to visual observations ?

Hearing and seeing a meteor is a unique experience. There is a fairly simple relation between the visual magnitude of a meteor and the duration of the reflection. At 70 MHz, a zero magnitude meteor gives typically rise to a 1 second reflection.

#### 5. How to correct MS counts to standard conditions ?

The number of meteor reflections that will be observed from a certain stream is not a simple function of radiant height, as is the case with visual counts (cfr. the zenithal hourly rate). Conversely, deriving the true number of meteors from the observed number of meteors is not a trivial matter, certainly not when more than one transmitter is involved. Certainly, a stream does not necessarily peak when the observed number of meteors is highest!

The various parameters, such as antenna characteristics, position of transmitter and receiver, radiant position and other stream characteristics, power of the transmitter etc. can be fed into a simulation program, such as FORWARD (by the undersigned). Although the general activity pattern can be reconstructed, one has to be careful.

For visual meteors, the observed number is highest when the radiant is in the zenith (all other remaining the same). This is not so for the number of meteor reflections: the maximum number is observed for a radiant elevation of 45 to 50 degr. When the radiant is too high (say more than 80 degr), very few reflections are received!

#### 6. Where to find meteor stream data ?

The International Meteor Organization (IMO), specializing in visual meteor observations, issues a yearly meteor stream calendar, which can be found on several places on the net. Contact otherwise Juergen Rendtel (jrendtel@aip.de)

**Continued on page five - see Meteors**

**Meteors** - Continued from page four  
or Peter Brown  
(peter@canlon.physics.uwo.ca).

Gary Kronk's home page [http://wums.wustl.edu/~kronk/index.html] gives also a good overview of the various streams.

#### 7. Where to find observational data and reduction software ?

Ilkka Yrjola, OH51Y  
(ILKKA.YRJOLA  
@915000.KY.KYMMENE.MEMONET.mailnet.fi)  
makes his data available on ftp server  
FTP.FUNET.FI in directory pub/ham/vhf-  
work/mssoft42.zip. The software for  
viewing is available too, as well as stream  
data.

The author edits the monthly RMOB,  
which gives summaries of recent obser-  
vations from all over the world. Included  
are the setup characteristics. Just con-  
tact me for putting you on the distribution  
list.

#### 8. References

##### 8.1. Books

THE reference is still "Meteor Science  
and Engineering" by D.W.R. McKinley,  
McGrawHill 1961. A fantastic book, giv-  
ing the theory and practical results of the  
heydays of the radio meteor science.

A recent, less astronomical and more  
communication oriented work is: Meteor  
Burst Communications (Theory and  
Practice) Ed. Donald L. Schilling, Wiley  
Series in Communications, 1993, ISBN  
0-471-52212-0.

##### 8.2. Periodicals

The regular astronomy magazines, such  
as Sky & Telescope, Astronomy, or As-  
tronomy Now! run occasionally contri-  
butions about forward scatter  
experiments (due to lack of time: com-  
plete list in Rev. 2 of these FAQs).

WGN, the Journal of IMO, contain some-  
times contributions and letters about  
radio observations of meteors.

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B 2400 Mol, Belgium 51 d N, 5 d E

**Laughter** - continued from page two  
one ever asks. It kills the party for me.  
—The Magazine of Fantasy and Sci-  
ence Fiction, May 1966

I suppose many of you are fa-  
miliar with the comic strip "Peanuts." My  
daughter Robyn (now in the fourth grade)  
is very fond of it, as I am myself.

She came to me one day, de-  
lighted with a particular sequence in  
which one of the little characters in "Pea-  
nuts" asks his bad-tempered older sis-  
ter, "Why is the sky blue?" and she  
snaps back, "Because it isn't green!"

When Robyn was all through  
laughing, I thought I would seize the  
occasion to maneuver the conversation  
in the direction of a deep and subtle  
scientific discussion (entirely for Robyn's  
own good, you understand), So I said,  
"Well, tell me Robyn, why is the night sky  
black?"

And she answered at once (I  
suppose I ought to have foreseen it),  
"Because it isn't purple!"

—The Magazine of Fantasy and Sci-  
ence Fiction, November 1964

Submitted for your amusement and in  
memory of Dr. Asimov.  
Bob Brauer

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other participating astronomy clubs, pub-  
lishers and selected individuals.

**My new 1.16" f/34 Schiefspiegler**  
by BOB ROGERS  
(rogers@comp.tamu.edu)

Hi there y'all,

Frustration can sometimes lead  
to desperate measures. The frustration  
in this case was trying to wade through  
humid Texas' skies and get a decent  
look at Jupiter. With my Meade 2045D  
SC (4", f/10), I could see one dark zone  
on Jupiter except on the best of nights  
when there was an inkling of the South-  
ern zone present.

After reading about all the ben-  
efits attributable to Schiefspiegler de-  
signs, I determined that I might be able  
to improve the image contrast with a  
simple disk with a hole in it. I measured  
the appropriate diameters on the correc-  
tor plate and secondary mirror mount  
and then cut a suitable disc of card stock  
with a 1.16" hole located on one side.  
The hole just grazed the outside edge of  
the corrector plate and the near edge of  
the secondary mount. To my way of  
thinking, this is the equivalent of a 1.16"  
f/34 Schiefspiegler.

Anyway, with the collimator in  
place I enjoyed seeing both the Northern  
and Southern dark zones. Necessarily  
the image wasn't bright but with a 12.5  
mm SP (80X) it looked very reminiscent  
of the photograph that appeared in the  
June issue of Astronomy on the page  
where they gave the transit times for the  
red spot. With just a little better sky and  
some luck I may get to see that red spot.

It was simple to make, it seemed  
to work, and it must set some sort of  
record for a small aperture  
Schiefspiegler.

To those who may quibble that  
this isn't a Schiefspiegler at all but simply  
a stopped down aperture, I shall quibble  
only after receiving a box of homemade  
chocolate chip cookies from you.

Best regards and clear skies.

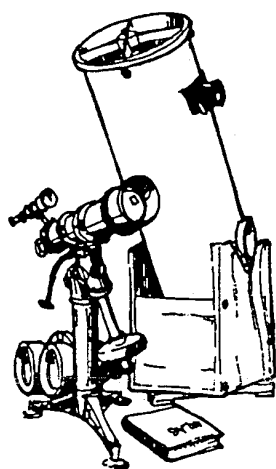
Bob

# 1995 SJAA Calendar

| General Meeting | Houge Park Star Party | Observational Astronomy Class |
|-----------------|-----------------------|-------------------------------|
| Sept 9          | 1 and 29              | 16                            |
| Oct 14          | 27                    | 7 (Last -one Fremont Peak)    |
| Nov -           | -                     | None                          |
| Dec -           | -                     | None                          |

Begin 1/1996

Please read your *Ephemeris* each month for changes



## Telescope Loaner Status by Paul Barton

| NO. | Name                | User               | Due Date   |
|-----|---------------------|--------------------|------------|
| 1   | 4-1/2" Newt/P Mount | ----->             | available  |
| 2   | 6" Dobson           | John Paul Dasilvia | 10/3/95    |
| 3   | 4" Quantum          | Albert Chen        | 9/18/95    |
| 6   | C-8 Celestron       | Bob Maillot        | 9/22/95    |
| 7   | 12-1/2" Dobson      | Tom Rice           | indefinite |
| 8   | 14" Dobson          | Lee Courtney       | 9/8/95     |
| 15  | 8" Dobson           | Bob Elsberry       | 9/8/95     |
| 18  | 8" Newt/P Mount     | Jerry Lovelace     | 10/10/95   |
| 19  | 6" Newt/P Mount     | ----->             | Availiable |
| 21  | 10" Dobson          | Richard Lee        | 9/5/95     |
| 23  | 6" Newt/P mount     | Jim Marquis        | 9/8/95     |

Solar telescope. Available only to experienced members for special occasions such as day time public star parties, etc. Call.

Wait List

Steve Wincor C-8

If you want to borrow a telescope call Paul Barton (number is on the credit Marque) and get your name on a general list (any telescope) or on a specific telescope list.

## ASTRO ADS

ASTRO ADS are free to all noncommercial advertisers wishing to sell astronomically related products or services. Please send your ad directly to the Editor:

Bob Madden

1616 Inglis Lane

San Jose, Ca. 95118-2825

**Lost: Battery ....** Left at Fremont Peak, Coulter Camp, Saturday, June 17 or 24 Deep Cycle, Trojan Marine, Red Top, White Case, Call Paul Krukar at:

(h) 286-5728

(w) 277-4638

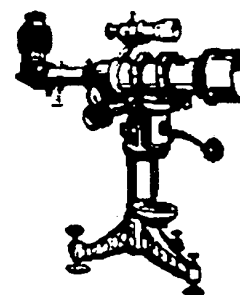
or Paul Barton at:

377-0148

or any board member. The battery is SJAA property, part of the C-11.

**Star Instruments 12-1/2 inch, f/9, Ritchey Chretien, GEM - stepper driven, 2-inch star diag. and 3-2-inch EP. Complete w/ battery and drive. Chip on primary; doesn't hurt image. Newly recoted primary. All overhauled and working. Exceptionally easy to collimate. Very compact for a large scope (like a Cassegrain). Designed for photography. \$2,000 or \$1500 w/out EPs. Call Paul Barton (408) 377-0148** 8/95

**8" f8 Neutonian** on a Mead mount with a clock drive. It has a cave mirror. \$500 Call Paul Mancuso (408) 946-0738 8/95



## Notice

During the month of September John Dobson will be having his 80th birthday. come to the September 1 Star Party at Houge Park to dedicate this activity in honor of John. John has given so much to amateur astronomy and astronomy education.

# Celestial Calendar - Sept 1995

by Richard Stanton

| Lunar Phase | Date  | Rise | Trans | Set         |
|-------------|-------|------|-------|-------------|
| FQ          | 02:04 | 02   | 14:27 | 19:38 00:53 |
| FM          | 20:36 | 08   | 19:05 | 00:18 06:15 |
| LQ          | 14:26 | 16   | 23:46 | 06:45 14:02 |
| NM          | 09:54 | 24   | 06:56 | 13:02 19:02 |

## Nearer Planets

|           |    |       |       |       |
|-----------|----|-------|-------|-------|
| Mercury   | 07 | 08:58 | 14:41 | 20:23 |
| 0.82 A.U. | 17 | 08:58 | 14:28 | 19:57 |
| Mag. -1.2 | 27 | 08:17 | 13:45 | 19:14 |

|           |    |       |       |       |
|-----------|----|-------|-------|-------|
| Venus     | 07 | 07:04 | 13:26 | 19:47 |
| 1.70 A.U. | 17 | 07:26 | 13:32 | 19:37 |
| Mag. -4.0 | 27 | 07:48 | 13:38 | 19:27 |

|           |    |       |       |       |
|-----------|----|-------|-------|-------|
| Mars      | 07 | 10:28 | 15:55 | 21:21 |
| 2.09 A.U. | 17 | 10:22 | 15:41 | 20:59 |
| Mag. +11  | 27 | 10:17 | 15:28 | 20:39 |

|           |    |       |       |       |
|-----------|----|-------|-------|-------|
| Jupiter   | 07 | 13:30 | 18:24 | 23:18 |
| 5.48 A.U. | 17 | 12:57 | 17:50 | 22:43 |
| Mag. -2.0 | 27 | 12:24 | 17:16 | 22:08 |

|           |    |       |       |       |
|-----------|----|-------|-------|-------|
| Saturn    | 07 | 19:47 | 01:38 | 07:24 |
| 8.60 A.U. | 17 | 19:06 | 00:56 | 06:41 |
| Mag. +0.7 | 27 | 18:25 | 00:14 | 05:58 |

## SOL Star Type G2V

| RA           | Dec |       |       |       |
|--------------|-----|-------|-------|-------|
| 11:02 +06:12 | 07  | 06:41 | 13:06 | 19:30 |
| 11:37 +02:23 | 17  | 06:50 | 13:02 | 19:14 |
| 12:13 -01:30 | 27  | 06:58 | 12:59 | 18:58 |

## Astronomical Twilight

|              |    | Begin | End   |
|--------------|----|-------|-------|
| JD 2,449,967 | 07 | 05:12 | 20:59 |
| 977          | 17 | 05:22 | 20:42 |
| 987          | 27 | 05:32 | 20:25 |

## Sidereal Time

|                |    |       |   |       |
|----------------|----|-------|---|-------|
| Transit Right  | 07 | 00:00 | = | 21:55 |
| Ascension at   | 17 | 00:00 | = | 22:35 |
| Local Midnight | 27 | 00:00 | = | 23:14 |

## Darkest Saturday Night: 23-Sep-1995

|              |       |
|--------------|-------|
| Sunset       | 19:05 |
| Twilight End | 20:31 |
| Moon Set     | 18:27 |
| Dawn Begin   | 05:29 |



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## COMET COMMENTS

by Don Machholz

Periodic Comet d'Arrest dims while Periodic Comet Jackson-Neujmin brightens slightly. Meanwhile, a new comet has been visually found. Comet 1995 01 (Hale-Bopp): Alan Hale of Cloudcroft, New Mexico and Thomas Bopp of Glendale, Arizona discovered this comet while observing M 70 on July 23. Hale is a well-known comet observer who has done some comet hunting but was not actually searching for comets when he found this. At that same hour Bopp was observing M 70 through his friend's (Jim Stevens) 17" telescope when he noticed the comet nearby. He promptly drove home (90 miles) to report it. Bopp doesn't own anything bigger than a spotting scope, but has been involved in astronomy for some 25 years, mostly in Ohio. Two days later Gerry Rattley of Gilbert, Arizona also discovered it. The orbit of this new object is difficult to determine, but the positions below should help you follow it through September. Early indications are that the comet is distant and will be around for a long time.

## EPHEMERIS

| 6P/d'ARREST   |          |         |      |     |      | 58P/JACKSON-NEUJMIN |          |         |      |     |      |
|---------------|----------|---------|------|-----|------|---------------------|----------|---------|------|-----|------|
| DATE          | R.A.     | DEC     | EL   | SKY | MAG  | DATE                | R.A.     | DEC     | EL   | SKY | MAG  |
| (00UT) (2000) |          |         |      |     |      | (00UT) (2000)       |          |         |      |     |      |
| 08-06         | 23h46.6m | -08d23m | 139d | M   | 9.1  | 08-06               | 21h28.7m | +0kd09m | 161d | M   | 12.7 |
| 08-11         | 23h58.8m | -11d58m | 142d | M   | 9.1  | 08-11               | 21h30.7m | -00d05m | 165d | M   | 12.4 |
| 08-16         | 00h09.8m | -15d35m | 144d | M   | 9.2  | 08-16               | 21h32.8m | -0kd38m | 168d | M   | 12.2 |
| 08-21         | 00h19.5m | -19d06m | 146d | M   | 9.3  | 08-21               | 21h35.2m | -03d28m | 169d | M   | 12.0 |
| 08-26         | 00h27.7m | -22d24m | 147d | M   | 9.5  | 08-26               | 21h38.2m | -05d34m | 169d | E   | 11.9 |
| 08-31         | 00h34.4m | -25d24m | 147d | M   | 9.7  | 08-31               | 21h41.8m | -07d53m | 167d | E   | 11.7 |
| 09-05         | 00h39.6m | -28d01m | 147d | M   | 9.9  | 09-05               | 21h46.2m | -10d21m | 163d | E   | 11.5 |
| 09-10         | 00h43.5m | -30d13m | 147d | M   | 10.2 | 09-10               | 21h51.7m | -12d52m | 159d | E   | 11.4 |
| 09-15         | 00h46.1m | -32d00m | 146d | E   | 10.5 | 09-15               | 21h58.2m | -15d20m | 154d | E   | 11.4 |
| 09-20         | 00h47.8m | -33d21m | 145d | E   | 10.7 | 09-20               | 22h05.8m | -17d40m | 150d | E   | 11.3 |
| 09-25         | 00h48.6m | -34d18m | 144d | E   | 11.0 | 09-25               | 22h14.6m | -19d48m | 146d | E   | 11.3 |
| 09-30         | 00h49.0m | -34d52m | 142d | E   | 11.3 | 09-30               | 22h24.4m | -21d37m | 142d | E   | 11.4 |
| 10-05         | 00h49.0m | -35d06m | 140d | E   | 11.6 | 10-05               | 22h35.1m | -23d06m | 139d | E   | 11.4 |
| 10-10         | 00h49.0m | -35d00m | 139d | E   | 11.9 | 10-10               | 22h46.6m | -24d13m | 135d | E   | 11.5 |

## 1994 01 (HALE-BOPP)

| DATE          | R.A.     | DEC     | EL   | SKY | MAG  | DATE          | R.A.     | DEC     | EL   | SKY | MAG  |
|---------------|----------|---------|------|-----|------|---------------|----------|---------|------|-----|------|
| (00UT) (2000) |          |         |      |     |      | (00UT) (2000) |          |         |      |     |      |
| 07-27         | 18h40.6m | -32d02m | 153d | E   | 10.6 | 09-05         | 18h20.3m | -30d26m | 111d | E   | 10.4 |
| 08-01         | 18h37.2m | -31d53m | 147d | E   | 10.5 | 09-10         | 18h19.1m | -30d11m | 106d | E   | 10.3 |
| 08-06         | 18h34.0m | -31d43m | 142d | E   | 10.5 | 09-15         | 18h18.2m | -29d57m | 101d | E   | 10.3 |
| 08-11         | 18h31.0m | -31d31m | 137d | E   | 10.5 | 09-20         | 18h17.6m | -29d42m | 96d  | E   | 10.3 |
| 08-16         | 18h28.3m | -31d19m | 132d | E   | 10.4 | 09-25         | 18h17.4m | -29d28m | 91d  | E   | 10.3 |
| 08-21         | 18h25.8m | -31d07m | 127d | E   | 10.4 | 09-30         | 18h17.5m | -29d14m | 86d  | E   | 10.3 |
| 08-26         | 18h23.7m | -30d53m | 121d | E   | 10.4 | 10-05         | 18h17.9m | -28d59m | 81d  | E   | 10.3 |
| 08-31         | 18h21.8m | -30d40m | 116d | E   | 10.4 | 10-10         | 18h18.6m | -28d45m | 77d  | E   | 10.3 |

| Object              | P/d'ARREST       | P/Jackson-Neujmin | 199501 HALE-BOPP   |
|---------------------|------------------|-------------------|--|
| Peri. Date          | 1995 07 27.36197 | 1995 10 06.61876  |  |
| Peri. Dist.(AU)     | 1.34587 AU       | 1.381125 AU       | The orbital elements for this comet are not known well enough for publication or long-term prediction. The positions and magnitudes given above are approximate. |
| Arg of Peri.(2000)  | 178.0504 deg     | 200.3470 deg.     |  |
| Ascend. Node (2000) | 138.9874 deg     | 160.7177 deg.     |  |
| Inclination (2000)  | 019.5232 deg     | 013.4779 deg.     |  |
| Eccentricity        | 0.6140404        | 0.6614285         |  |
| Orbital Period      | 6.51 yrs.        | 8.24 yrs.         |  |
| Source              | MPC 20122        | MPC 20123         |  |

Don Machholz (916) 346-8963

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